Physics of the hot evolving Universe

Science themes for a New-Generation X-ray Telescope

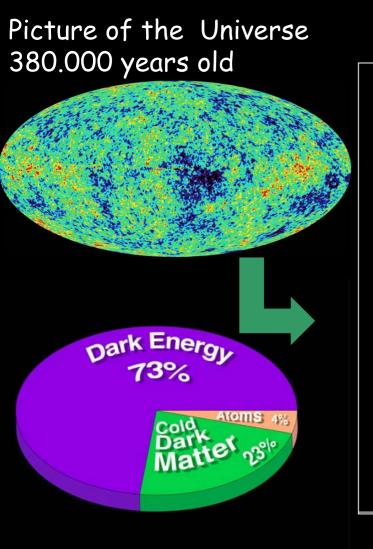
Günther Hasinger Max-Planck-Institut für extraterrestrische Physik Garching

ESA Cosmic Vision 2015-2025 Workshop, Paris, September 15-16

MPE

1. Quest for the first massive Black Holes

Formation of Large-Scale Structure



Klypin, Kravtsov, Gottlöber

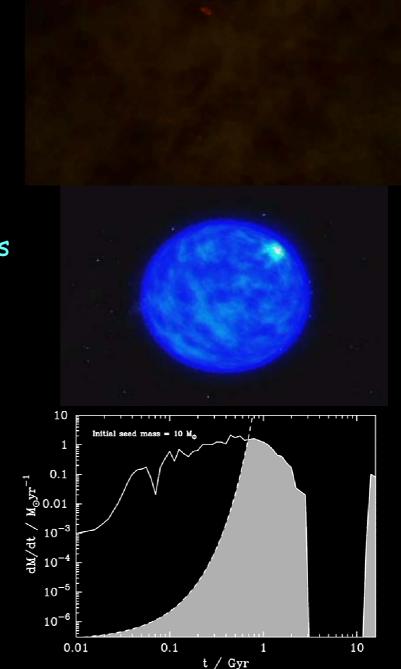
The first Black Hole

Before the first star can form, the universe has to cool down to ~100K to allow molecular hydrogen cooling.

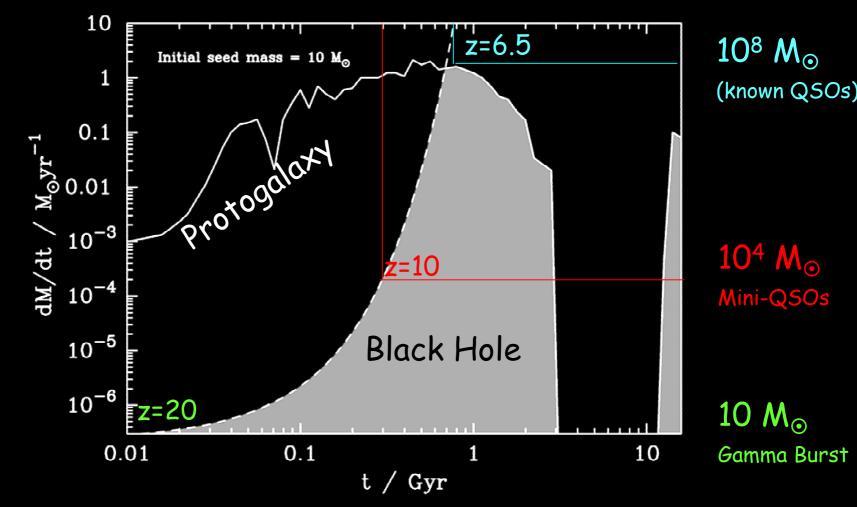
The first star is expected to be massive (~300 M_{\odot}), shines for ~1 Million years, sterilizes its cosmic environment, explodes in a GRB hypernova, pollutes its environment with heavy elements and leaves a seed Black Hole.

While the galaxy forms, the BH continues to grow exponentially, quickly producing a powerful quasar, if enough fuel can be provided.

Sensitive X-ray observations can study the first GRB explosions and can detect mini-QSOs with $10^4~M_{\odot}$



QSO exponential feeding

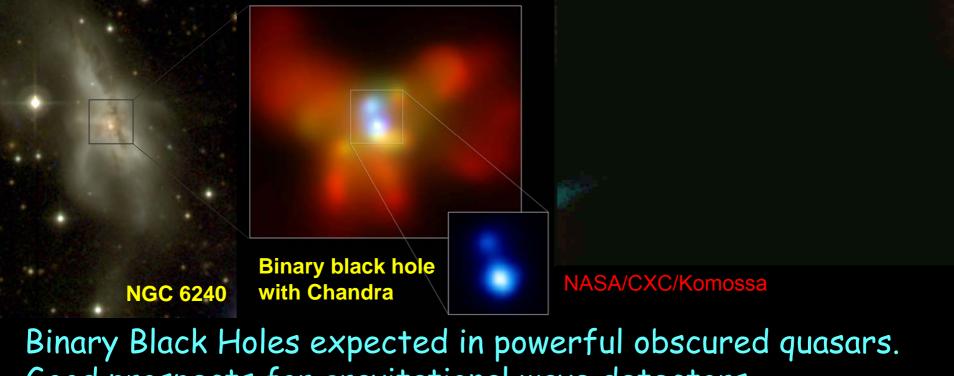


Need New Generation X-ray Telescope to detect and study BH in conjunction with forming galaxy ($S_{min} \sim 10^{-18} \text{ erg cm}^{-2} \text{ s}^{-1}$). 10⁴ M_o @ redshift 10 detectable.

Binary BH mergers - efficient accretion

Biggest problem: how to funnel down large amount of gas over a galactic scale on a very short timescales

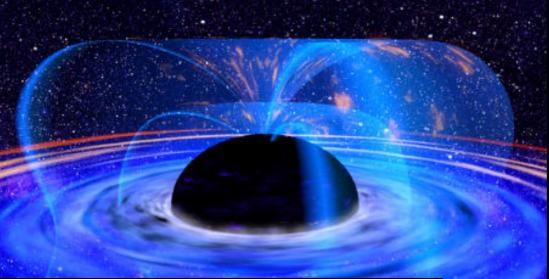
Observations tell us, that gas-rich mergers can do the trick.



Good prospects for gravitational wave detectors.

But NGXT needed for the electromagnetic waves!

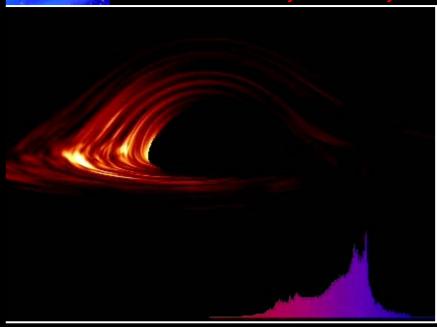
Approaching the Black Hole



Simulation courtesy Chris Reynolds

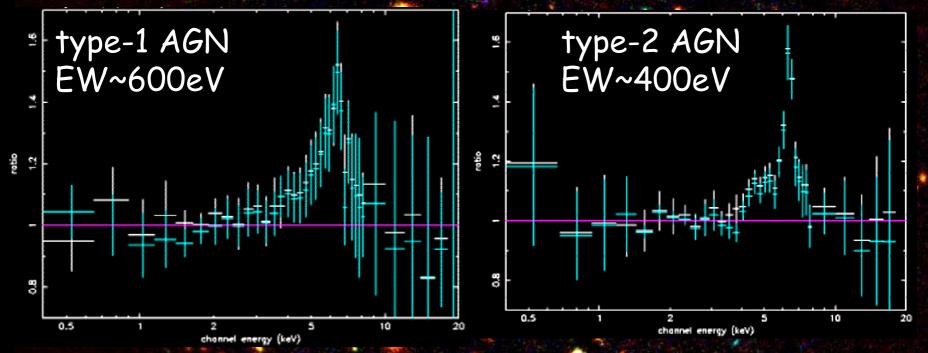
Simulation of precessing binary accretion disks in 3C273 with New Generation X-ray Telescope (Torres et al., 2004)

Ratio



Lockman Hole 800 ks XMM-Newton observation

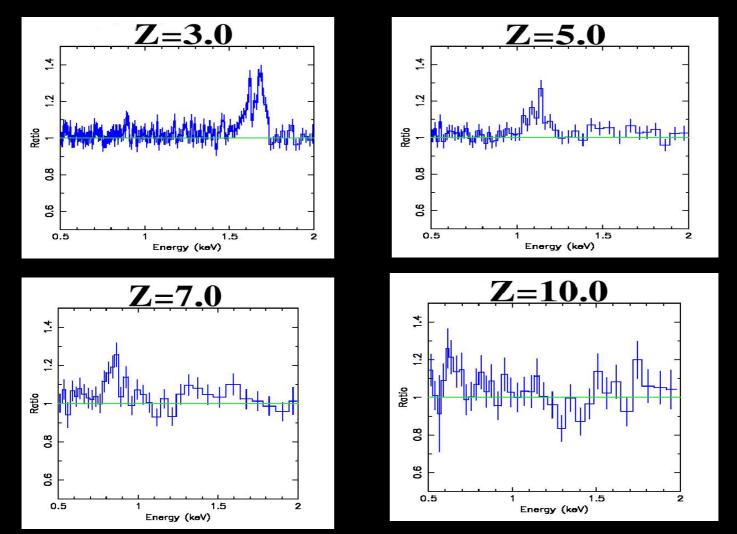
Average rest-frame spectra show relativistic Fe-lines



Streblyanskaya et al., 2004

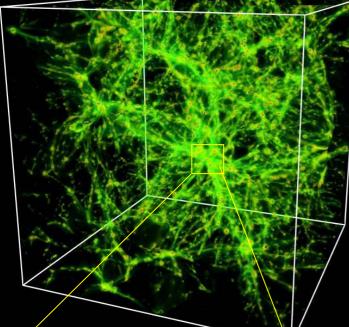
New Generation X-ray Telescope can determine redshifts and study Fe lines in each individual object

NGXT view of relativistic Fe line



We get this or better quality for many objects in each field

2. Evolution of the Cosmic Web and the Life Cycle of Matter

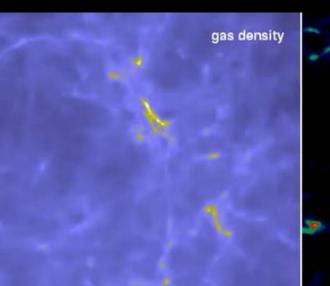


dark matter density

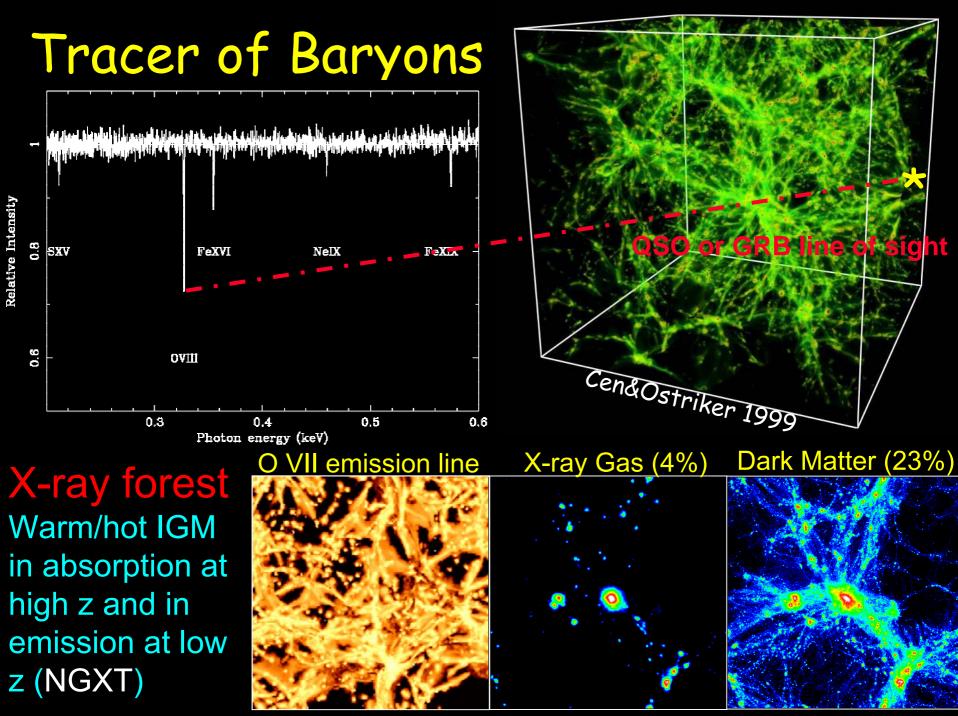
Clusters of Galaxies: Largest Objects in the Universe

courtesy V. Springel MPA

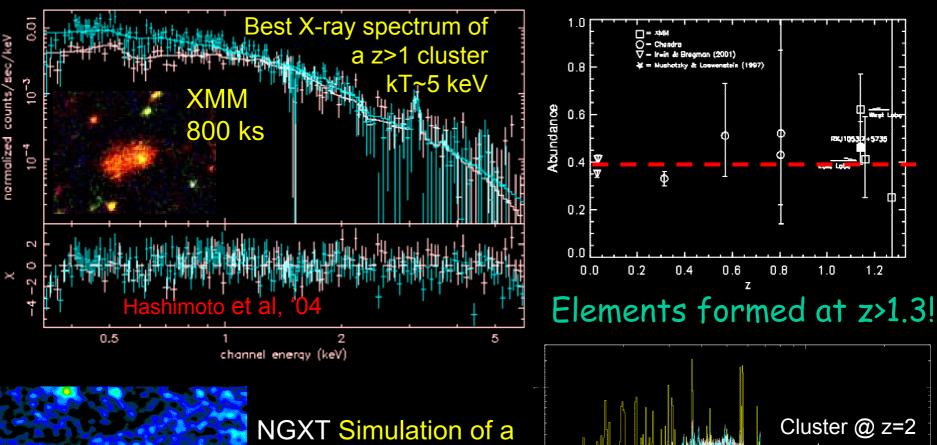
gas temperature



z = 20.0



Early clusters: enrichment of elements



0.1

Our Galaxy

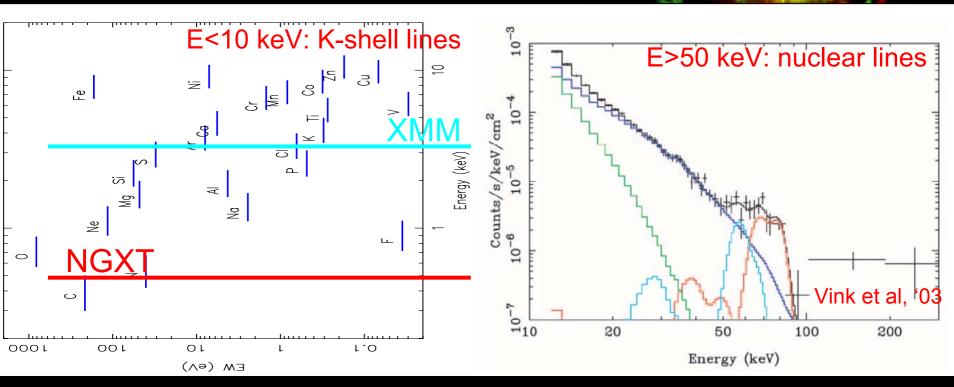
5 keV cluster at z=2

Element abundances of O, Ne, Mg, S, Si, Fe,

Element Abundances from SNRs

Closing the life cycle of matter





Nucleosynthesis measurements in the near future technologically easier in X-rays than in (MeV) Gamma range

Summary

- Scientific Requirements
 - Sensitivity:
 - Energy Band
 - Angular resolution 2-5 arcsec
 - Spectral resolution 1-2 eV

 10^{-18} erg cm⁻² s⁻¹ \rightarrow > 10 m² area @1 keV

- 0.1 to ~100 keV

NGXT science is exciting with strong European leadership! Science has highest priority in national planning excercises (e.g. NAS Decadal Survey, German "Denkschrift") + ~100 scientist signatures Coordinated planning in Europe, Japan and US

- Technology development required:
 - Factor ~ 10 lighter mirrors, high precision micropore optics
 - Formation flying, 1 mm³ accuracy over ~50 m
 - Imaging calorimeter, better than 2 eV
 - Large, fast active pixel detector with µs timing

Proof of concept exists for key technologies, European lead! Need sufficient technology investment NOW to be ready for a mission in the 2015-2020 timeframe

Thank you very much!