A study of small satellites capture in corotation resonance

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46th ESLAB SYMPOSIUM Formation and Evolution Moons

26/06/2012

Pattern Speed:

$$m\Omega_p = mn' + k\kappa' + p\nu'$$

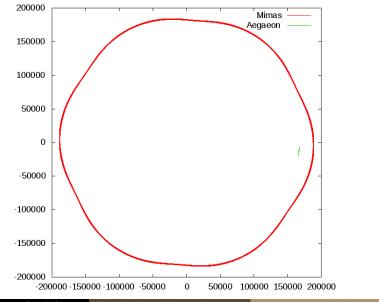
• In Corotational Resonance the frequency of the perturber is the same as the pattern speed:

$$m(n-\Omega_p)=0$$

• Resonant Angle:

$$\varphi_{CR} = m(\lambda' - \lambda) + (m - p)(\lambda - \varpi)$$

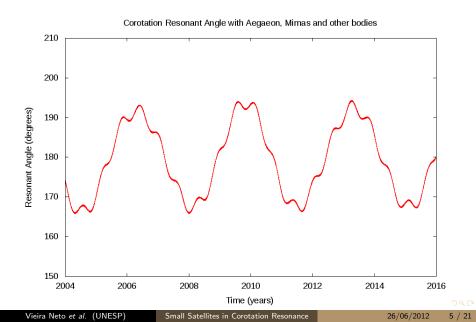
- In the Saturnian System there is three satellites in corotation resonance with Mimas
 - Aegaeon 6:7, 6 $(\lambda' \lambda) (\lambda \varpi)$
 - Anthe 11:10, $11(\lambda' \lambda) + (\lambda \varpi)$
 - Methone 15:14, $15(\lambda' \lambda) + (\lambda \varpi)$

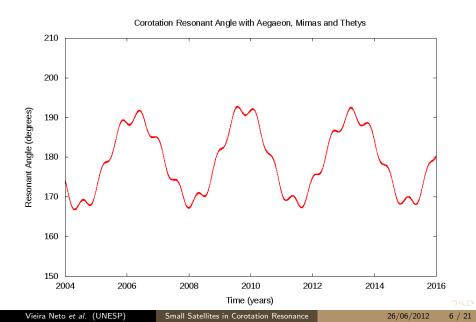


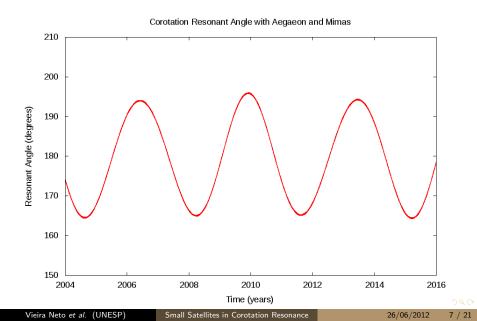
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Small Satellites in Corotation Resonance

4 / 21 26/06/2012





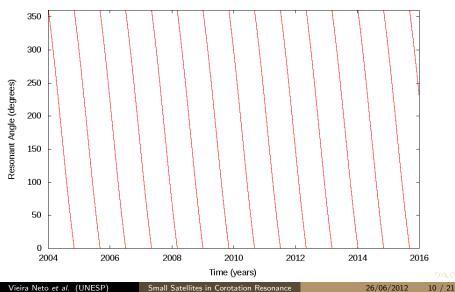


- How Aegaeon enters this resonance?
- Was it a captured satellite?
- Does it born in this corotation resonance?
- Dynamical study of corotation resonance

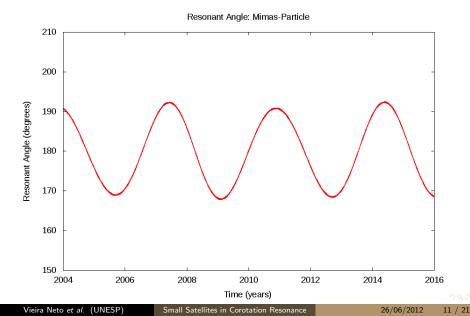
- Resonant angle is measured at each time step
 - if it circulates
 - we say the particle is not in corotation with Mimas
 - otherwise
 - · we say the particle is in corotation resonance

Not in Corotation Resonance

Resonant Angle: Mimas-Particle

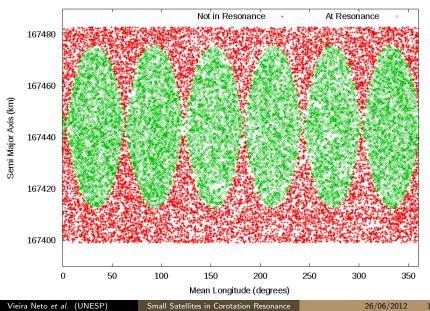


Corotation Resonance



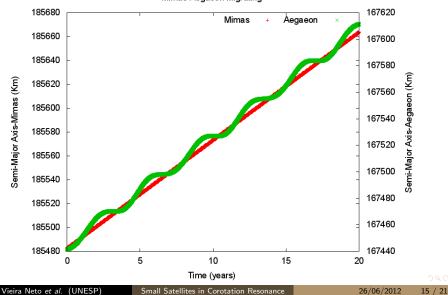
 With this definition we made Monte-Carlo simulations, varying semi-major axis, and mean longitude of the orbital elements of Aegaeon and classify the outcomes.

Classification

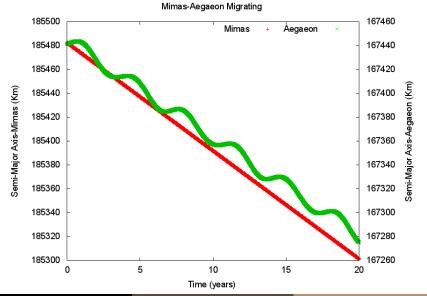


13 / 21

- With the corotation sample we studied the effect of Mimas migration.
 - We put a drag force only in Mimas dynamics
 - First we test with Aegaeon



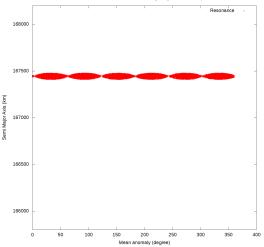
Mimas-Aegaeon Migrating



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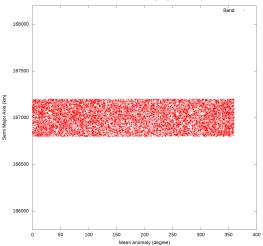
26/06/2012 16 / 21



Particles in Corotation with Mimas Migrating + Band 2004 years

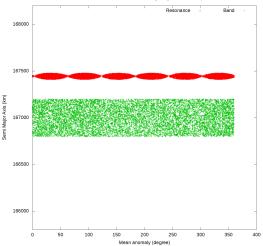
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26/06/2012 17 / 21



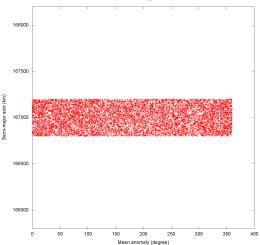
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Particles in Corotation with Mimas Migrating + Band 2004 years

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Band 2184 years

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- Corotation resonance is very robust.
- Although some particles moves in the border of the corotation resonance, none of them really enters the corotation.
- Others possibilities:
 - Slower migration
 - Mimas growth