

Orbiting L₂ Observation Point in Space

Herschel-Planck Mission Analysis Martin Hechler ESOC 19/03/2009

LIBRATION (LANGRANGE) POINTS IN THE SUN-EARTH SYSTEM

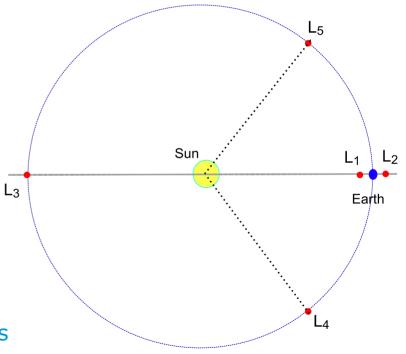


Libration Points:

- 5 Lagrange Points
- L₁ and L₂ of interest for space missions

Satellite at L₂:

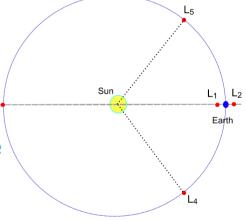
- Centrifugal force (R=1.01 AU) balances
 central force (Sun + Earth)
 - 1 year orbit period at 1.01 AU with
 Sun + Earth attracting
 - \Rightarrow Satellite remains in L₂
- However: Theory only valid if Earth moves on circle and Earth+Moon in one point



WHY DO ASTRONOMY MISSIONS GO TO $\rm L_2$ REGION



- Advantages for Astronomy Missions:
 - Sun and Earth nearly aligned as seen from spacecraft
 - ⇒ stable thermal environment with sun + Earth IR shielding
 - ⇒ only one direction excluded form viewing (moving 360° per year)
 - ⇒ possibly medium gain antenna in sun pointing
 - Low high energy radiation environment
- Drawbacks:
 - > 1.5×10^6 km for communication
 - However development of deep space communications technology (X-band, K-band) ameliorates disadvantage
 - Long transfer duration
 - \Rightarrow Fast transfer in about 30 days with +10 m/s
 - Instable orbits
 - ⇒ Manoeuvres every 30 days (~10 cm/s each)
 - \Rightarrow Escape at end of mission (L₂ region "self-cleaning")



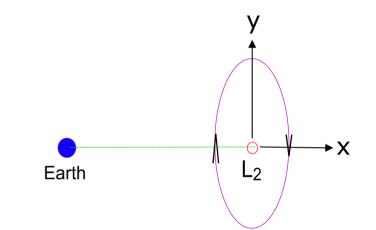
ORBITS AT L₂

Satellite in L₂

- Does not work in exact problem
- Would also be in Earth half-shadow
- And difficult to reach (much propellant)

Satellites in orbits around L₂

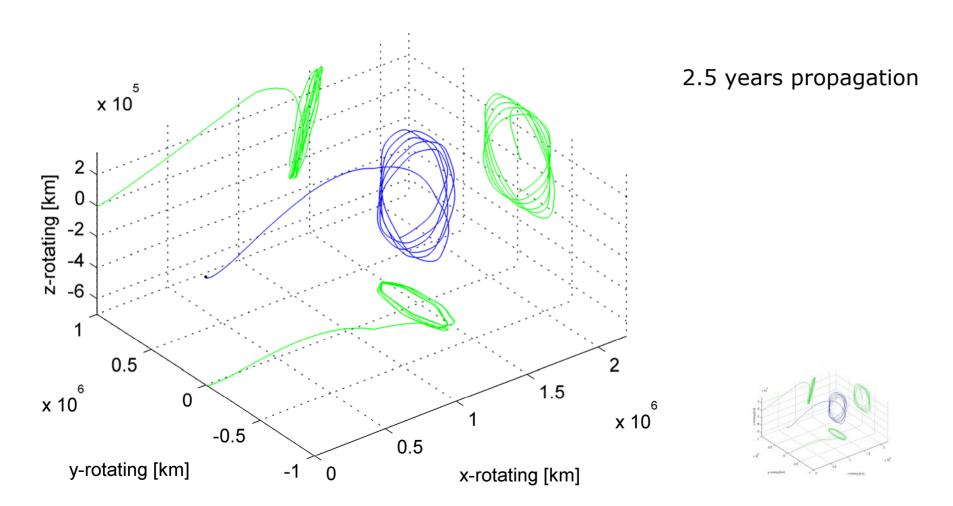
- With certain initial conditions a satellite will remain near L₂ also in exact problem ⇒ called Orbits around L₂
 - Different Types of Orbits classified by their motion in y-z (z=out of ecliptic)
 - ⇒ **Lissajous figure** in y-z for small amplitudes in y and z
 - ⇒ Halo for special combination of amplitudes





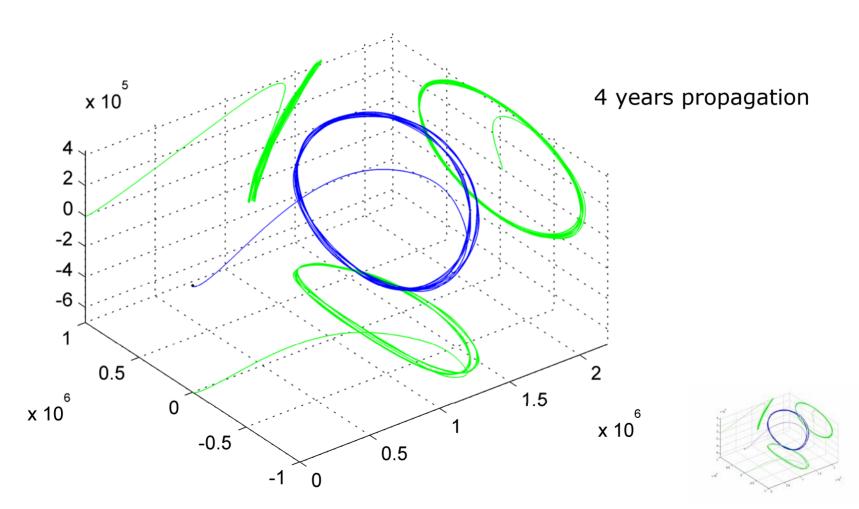
LISSAJOUS ORBIT (PLANCK)





HALO ORBIT (HERSCHEL)



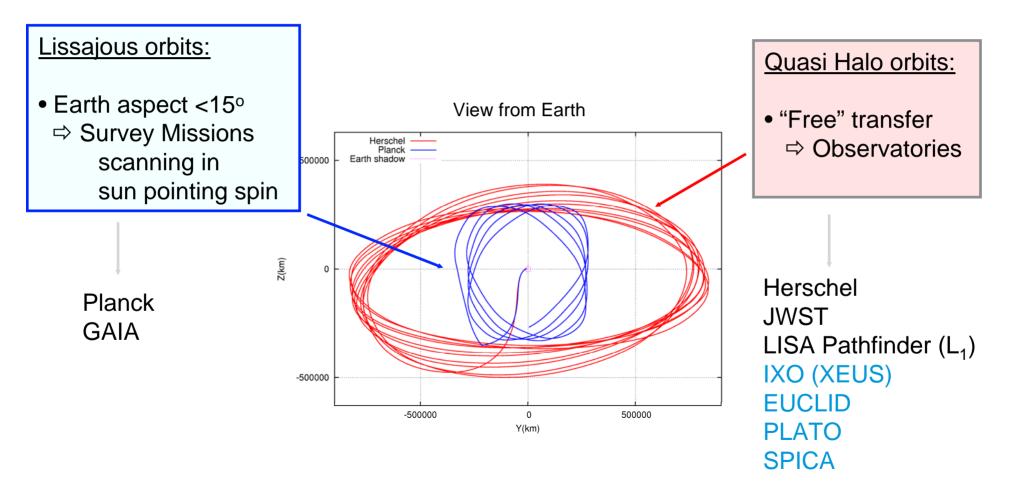


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ORBITS AT L₂ USED FOR SPACE MISSIONS





PAGE FOR MATHEMATICIANS



• Differential equations for relative motion in frame rotating with Earth around sun

$$\ddot{x} - 2 \dot{y} - (1 + 2K) x = 0$$

$$\ddot{y} + 2 \dot{x} - (1 - K) y = 0$$

$$\ddot{z} + K z = 0$$
Capture
• Complete solution of linearised problem (x-y motion and z-motion are uncoupled)

$$x = A_1 e^{\lambda_{xyt}} + A_2 e^{-\lambda_{xyt}} + A_x \cos(\omega_{xy}t + \phi_{xy})$$

$$y = A_1 c_1 e^{\lambda_{xyt}} - A_2 c_1 e^{-\lambda_{xyt}} - A_x c_2 \sin(\omega_{xy}t + \phi_{xy})$$
Periodic

$$z = A_z \cos(\omega_z t + \phi_z)$$

• Choice of initial conditions such that $A_1 = A_2 = 0 \Rightarrow$ Lissajous Orbits

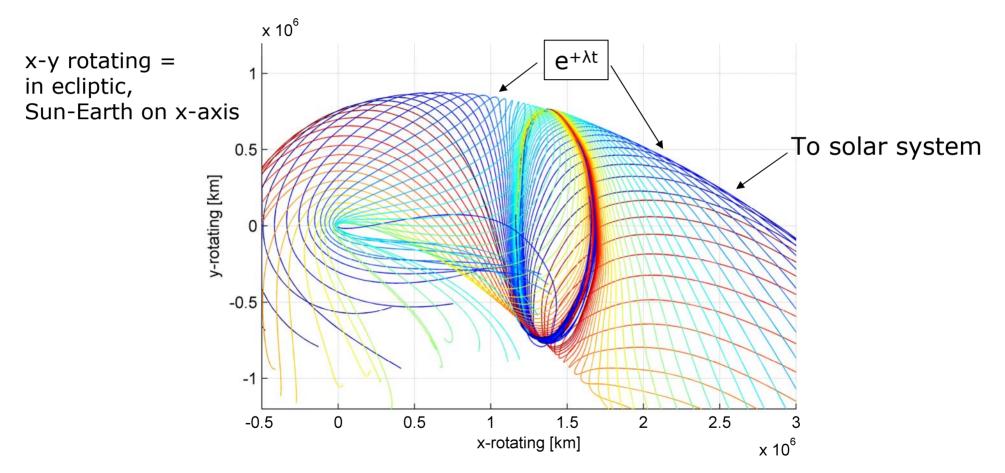
$$\begin{aligned} x &= A_x \cos(\omega_{xy}t) \\ y &= -A_y \sin(\omega_{xy}t), \text{ with } A_y = c_2 A_x \\ z &= A_z \cos(\omega_z t + \phi_z). \end{aligned}$$

$A_2 \neq 0 \Rightarrow$ Used for transfer

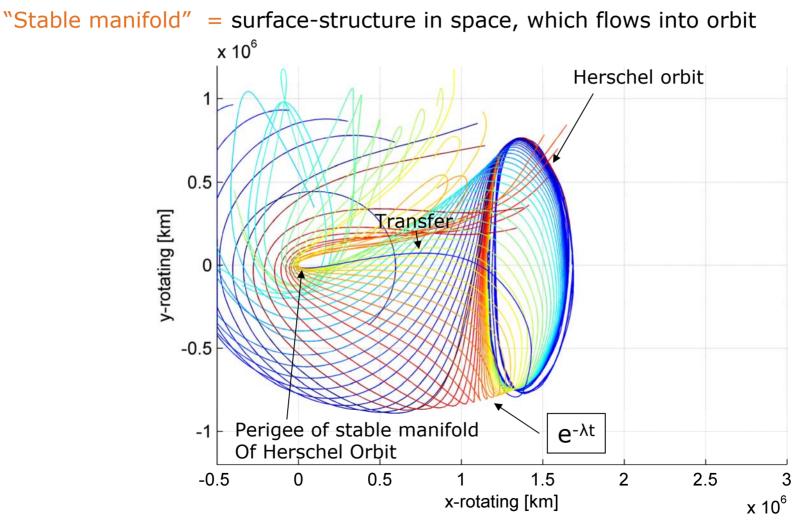
PROPERTIES OF ORBITS AT L₂ : INSTABILITY



Orbits at L₂ are unstable ⇒ escape for small deviation ⇒ unstable manifold



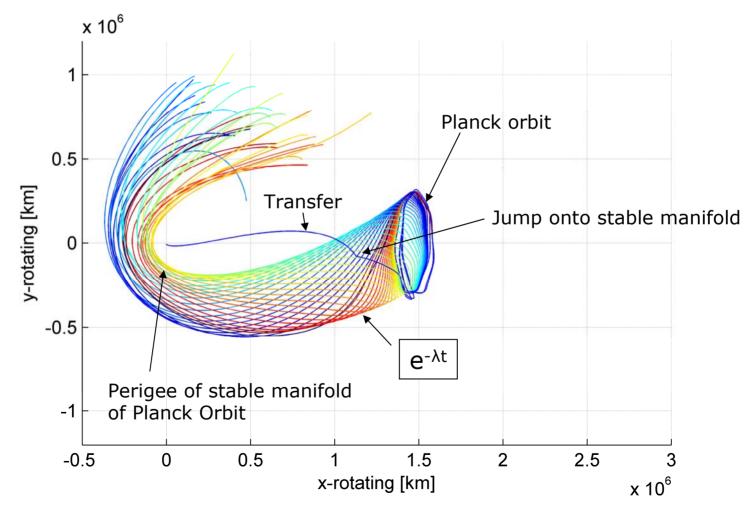




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STABLE MANIFOLD OF OF PLANCK ORBIT





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WHAT IS A LAUNCH WINDOW ?



Definition of Launch Window:

- Dates (seasonal) and hours (daily) for which a launch is possible
- Launcher target conditions (possibly as function of day and hour)

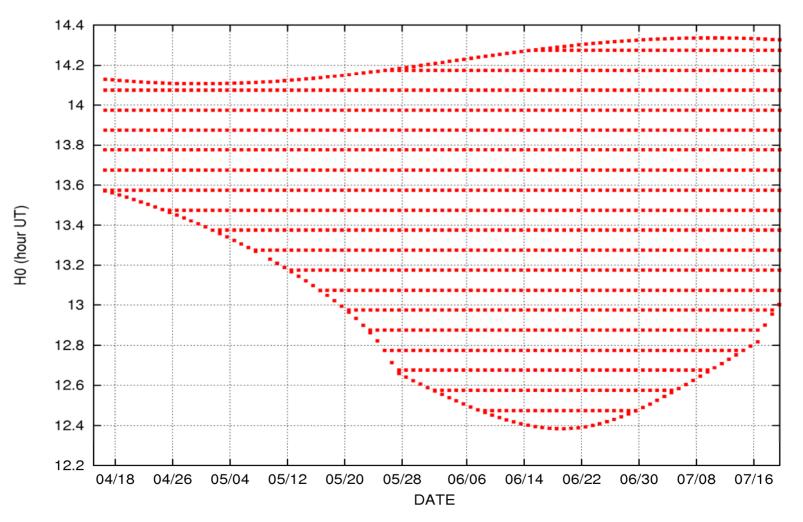
Constraints:

- Propellant on spacecraft required to reach a given orbit (type)
- Geometric conditions:
 - no eclipses during all orbit phases (power)
 - sun shall not shine into telescope during launch (damage)

Typical calculation method:

- Calculate orbits for scan in launch times
- Remove points for which one of the conditions is not satisfied

LAUNCH WINDOW FOR HERSCHEL-PLANCK



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THANK YOU

Martin Hechler Orbiting L₂ martin.hechler@esa.int

European Space Agency