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The Search for Life on Mars

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WHAT IS A HABITABLE PLANET?

- A PLANET WITH LIQUID WATER
- A PLANET WITH POTENTIAL ENERGY SOURCES (SOLAR, GEOCHEMICAL, HYDROTHERMAL, ...)
- A PLANET THAT HAS CARBON
- A PLANET THAT IS GEOLOGICALLY ACTIVE (RENEWAL OF NUTRIENT AND ENERGY SOURCES)

The habitability of a planet is not a static phenomenon

It changes with time as a function of: 1. The position of the planet in the (changing) habitable zone of the Solar System

- 2. The geological evolution of the planet in question
- 3. The co-evolution of life and a geologically active planet

(Westall, 2004)

The Red Planet Mars



MISSIONS TO MARS TOTAL OF 36 MISSIONS - ONLY 11.5 SUCCESSFUL MARS 1960A First Launch - USSR Oct 10, 1960 MARINER 4, 6, 7 - USA Nov. 1964-69- First flybys Dry, Volcanic, Cratered, Dusty, Dead Planet MARS 3 - USSR - First lander May 28, '71 **Only Radio Signal for 20 seconds- Little Data** MARINER 9 - USA First Mars Orbiter Nov. 14, '71 Mars active planet, volcanoes, channels, poles, "alive", 7,000 photographs VIKING 1 and 2 USA, July 20 & Sept. 3, 1976 First **Science Landers & Orbiters PHOBOS-2 USSR 1989 -Minimal data before contact lost** MARS PATHFINDER - USA Rover July 4, 1997 *MARS GLOBAL SURVEYOR - USA Orbiter Sept. 11, 1997 *MARS ODYSSEY - USA Orbiter Oct. 23, 2001 *MARS EXPRESS - ESA Orbiter Dec. 2003 (Beagle 2) ****MARS EXPLORATION ROVERS - USA Jan, Feb '04**

What we have learned...

Schiaparelli@ observations, 1877



Viking discovered water vapor present in the atmosphere; determined polar ice caps are carbon dioxide and water; found landing site surface was highly-oxidized, iron-rich clay

MARS I, II, III*

Mariner 4 Mariner 6 Mariner 7 Mariner 9*

Mariner 9 analyzed the atmosphere and revealed giant volcanoes and remnants of ancient riverbeds



Where we have been









Our Farthest Search for Life Beyond Earth

Trench excavated by Viking 1 lander surface sampler

VIKING LANDER HAD FIVE EXPERIMENTS SEEKING LIFE

- 1. <u>Camera</u> Seeking movements on surface
- 2. <u>Gas Chromatograph-Mass Spectrometer</u>
 - Seeking carbon components 1 ppm detection.
 NO carbon compounds found [Note: May have operated instrument incorrectly for C detection]

THREE DEDICATED LIFE DETECTION EXPERIMENTS

- 3. <u>Gas Exchange</u> seeking oxygen released after metabolism NEGATIVE
- 4. <u>Pyrolytic Release</u> photosynthesis via isotopically labelled ¹⁴CO₂ followed by pyrolysis seeking ¹⁴C NEGATIVE
- 5. <u>Labelled Release</u> "nutrients" added to sample. Possible positive signature obtained BUT because of NO carbon detected via GC-MS - INCONCLUSIVE.

Viking 1 and 2 Label Release Experiment

Evolution of ¹⁴CO₂ behaved as if metabolic activity was going on BUT Viking Team noted signals may be an artifact of "active" Martian soil -**INCONCLUSIVE ? Experiment needs to** be repeated.

QuickTime[™] and a TIFF (Uncompressed) decompressor are needed to see this picture.

AFTER: Levin and Straat, 1986

1983 - 1996 - PRESENT Martian Meteorites



- Martian meteorites strongly suggest that they were "processed" in a wet environment, therefore suggesting that water played a relatively large and important role in the planet in crustal history.
- Meteorite (ALHA84001) is almost as old as the solar system, and contains preserved carbonates.
- Mars (unlike Earth) has preserved samples of its earliest rocks (and hence environments) and therefore may harbor clues to how LIFE gets started on any planet in our Solar System.

ALH84001 Carbonate Globule



Orthopyroxene Groundmass

100 µm

Carbonate/

Magnetite Rim

Four Lines of Evidence Presented in 1996

Carbonates formed at Low-Temperatures, Indigenous Organic Compounds Present, Magnetites (Fe₃O₄) of possible biogenic origin, Unusual Morphologies of "fossilized features"

HYPOTHESIS: NSIGNATURES OF BIOGENIC ACTIVITYÓ

After Eight Years of Extensive Research

All Four Lines of Evidence Still Stand

Magnetite Continues to be our Strongest Evidence

MAGNETITES AS SIGNATURES OF LIFE

MAGNETOTACTIC BACTERIA Magnetite Chains in MV-1

Scale same as below

Magnetite Chains in ALH84001



Friedmann et al. PNAS (2001)

Six Properties of Biogenic Magnetite



Thomas-Keprta et al. 2000, 2001, 2002; Kirschvink, AGU 2003

Morphology of Biogenic and Martian Magnetites



Thomas-Keprta et al. 2000, 2001, 2002; Kirschvink, AGU 2003

MAGNETITES IN MARTIAN CARBONATES

Unique Magnetite morphologies in ALH84001 Carbonates
 Identical to Magnetites from MV-1
 Six Properties which are unique.

Chains of Magnetites in 84001 carbonates (Friedmann et al., PNAS, 2001)

Synthetic Magnetites of Golden et al. are NOT identical to magnetites in 84001 carbonates-Violate several unique properties of Martian magnetites (i.e. compositions, thermal histories, trace elements, etc.)

QUESTION: The properties of magnetite have been used to identify signatures of biogenic activity in terrestrial samples for thirty years. When the same signatures exist in Martian carbonates, why do the identification rules not apply?

To be Biology: Duplicates must exist. Features Found in Two Younger Martian Meteorites of different ages!

Unique Morphologies In two Younger Martian Meteorites Compared to Recognized Terrestrial Fossilized Bacteria in Columbia River Basalt (Basalt age = 10 my)

Formation Age 165 My On Mars

Formation Age 10 My On Earth





Both Martian Meteorites Are Fresh Falls NakhlaĜ Fm. Age 1.3 Gy Shergotty Fm. Age 165 my

Water on Mars?

--you bet!!



(Notes: all meteorites shown here except for ALH84001 were <u>observed falls</u>; all examples here are from interiors of the meteorites) Wentworth et al., Icarus (2005)

CARBON IN MARTIAN METEORITES

Lithosphere Primary Magmatic silicates

. <u>Hydrosphere and Atmosphere</u> Secondary Carbonates

. <u>Biosphere (?)</u> Reduced organic carbon components

MARTIAN SAMPLE



ALH84001

CARBON COMPONENTS

Data from stepped combustion

From Grady et al, 1997

Carbon Isotopic Compositions

Mars	Atmosphere Carbonates – 			Magmatic Organic (?)				
Earth	• •			Atmosphere Marine Carbonates Mantle Typical organic				
Other		HED — Moon						
-4	0 -30 -20	0 -10	0	+10	+20	+30	+4	
δ ¹³ C (‰)								

After Grady and Wright, 2004

REVIEW OF CARBON COMPONENTS

⇒ PRIMARY CARBON

Present as reduced, well crystalline carbon δ^{13} C of -20 <u>+</u> 4‰ (lighter than terrestrial basalts of ~ -5‰)

\Rightarrow EITHER:

- δ^{13} C of Mars is different from Earth
- Martian meteorites contain reduced carbon with δ¹³C established by isotopic fractionation during emplacement (loss of heavy component during degassing)
- Secondary Carbonates have δ^{13} C of ~ +24 to +40‰
- Martian meteorites contain a recycled component that could, in principle, reflect a Martian biological signature (range of ~ -18‰ to -60‰)

[After Grady and Wright, 2004]

MARTIAN METEORITES' REDUCED CARBON COMPONENTS

Indigenous reduced carbon identified in Martian Meteorites: EET79001 by Wright et al. (Nature, 1989)

¹⁴C Signatures in Martian meteorites suggest indigenous reduced carbon component present:
 ALH84001 15 to 20 % is indigenous and not terrestrial Nakhla 75 to 80 % is indigenous (Jull et al.)

Amino Acid's presence in Martian meteorites is inconclusive ALH84001 has amino acids BUT identified to be contaminants from Antarctic environment.

Nakhla contains amino acids -possible mixture of terrestrial and extraterrestrial amino acids.

Biofilms identified in Martian meteorites: ALH84001, Nakhla and Shergotty (Gibson et al., Precambrian Research, 2001: Westall et al., 2001,2002)

- Successfully demonstrated surface mobility and robust entry & landing system
- Stereoscopic imager and elemental analyses of rocks enhanced surface geology

Mars Pathfinder





Pathfinder - Mars 1997



MARS GLOBAL SURVEYOR

1999 To Present

Martian Gullies: *Watersigns*





AMERICAN ASSOCIATION FOR THE ADVANCEMENT OF SCIENCE

MARS GLOBAL SURVEYOR

MASSIVE SEDIMENTS

Deposition by: Wind? Water?





Mars and Earth both formed 4.5 billion years ago. Did Mars once have a warm, wet environment for life?





Evidence suggests wet planet

Mars Express with Beagle 2 Approaching Mars

Mid-December 2003

Beagle 2 on Surface

(Artist Impression)

December 25, 2003 ??

MAJOR SCIENCE GOALS

Orbiter:

Global high-resolution & selected super-resolution photogeology
Global mineralogical mapping
Subsurface structure (few km)
Global atmospheric circulation and chemical composition
Water/ice reservoirs and cycle
Surface-atmosphere interaction
Interaction of upper atmosphere with solar wind

Lander:

Geology & mineralogy of landing site Organic, mineral, isotopic chemistry Exobiology (i.e. life signatures) Meteorology and climatology


Water Beneath Surface ?



Mars Express Planetary Fourier Spectrometer Spectra of Methane in Mars' Atmosphere

Methane in Mars Atmosphere

Methane concentrations increase over three basins suggesting correlation with water/ice.

OuickTimeTM and a TIFF (Uncompressed) decompressor are needed to see this pic

Water/Ice + Methane = Suggestion of Life??

V. Formisano et al. Science 306, 1758 (2004)

SUGGESTIONS OF REGIONAL VARIATIONS IN CH₄ SOURCES

Water in Martian Atmosphere

Methane in the Martian Atmosphere: Evidence for Life?

Krasnopolsky, Maillard and Owen, ICARUS (in press, 2004)

- ³⁄₄ Telescopic measurements of CH₄ in Martian atmosphere using Fourier Transform Spectrometer at Canada-France-Hawaii Telescope.
- ³/₄ Measured 15 strongest Doppler-shifted lines for CH_4 in Martian atmosphere.
- ³/₄ Observed CH₄ mixing ratio is 10 ± 3 ppb with total photochemical loss of CH₄ in Martian atmosphere is equal to 2.2 x 10^5 cm⁻²s⁻¹.
- $\frac{3}{4}$ CH₄ lifetime is 340 years in the Martian atmosphere.
- ³⁄₄ No processes for CH_4 production in atmosphere, so photochemical loss must be balanced by abiotic and/or biogenic sources.
- ³⁄₄ Possible sources of CH_4 in Mars atmosphere:

No recent Volcanism in past 10 MY, NO thermal emissons or hot spots No current hydrothermal systems on Mars - little CH_4 production if any Cometary impacts deliver < 2% CH_4 and Meteoritic and Interplanetary Dust delivery does not exceed 4% *METHANOGENESIS by living subterranean organisms (i.e. bacteria) is a plausible explanation for the levels of CH_4 in Martian atmosphere.*

METHANE IN MARS ATMOSPHERE

- Mumma et al. 2003 (DPS)
 10 to 20 ppbV
- Krasnopolsky et al. (2004) 11 <u>+</u> 4 ppbV
- PFS-MEX Formisano et al. (2004) 10 + 5 ppbV Plus regional differences above selected basins with concentrations up to 30 ppbV

Sources for Martian Methane

o. Abiotic Sources:

Volcanic or Hydrothermal

No thermal anomalies detection by TES in orbit **Cometary Infall** - Only 2% Contribution - Too Low **Meteorite/Dust Infall** - Only 4% Contribution - Too Low

o. Biogenic Sources

Supported by:

- o Methane in Martian atmosphere Methanogenic Bacteria?
- o Abundance of Martian water "life requires water"
- o Indigenous reduced carbon in Martian meteorites
- o Isotopic composition of reduced carbon- -15 to -18 ‰
- o Biogenic-like magnetite in Martian Meteorites
- o Label Release Experiment-Viking ??

1989: observation of formaldehyde by the Russian PHOBOS satellite



SPICAM And PFS this Meeting NEW RESULTS

[After Muller et al.EANA, 2004]

Horowitz et al., PNAS (1971) showed formaldehyde, acetaldehyde, and glycolic adid should be present in Mars atmosphere because of photochemistry. In Lab: $CH_4 + H_2O + uv = HCOH$ (this meeting) Therefore, formaldehyde's presence is not totally unexpected.

Spirit and Opportunity-Robot Geologists

On The Mast •Panorama Camera (PanCam) •Infrared Spectrometer (Mini-TES)

On The Arm (IDD) •Microscope (MI) •Alpha Particle X-Ray Spectrometer (APXS) •Mössbauer Spectrometer (MB) •Rock Abrasion Tool (RAT)

Opportunity : Evidence of Water at Meridiani Planun



Štone MountainÕ

Opportunity Discovers Evidence of Rocks Deposited in a Body of Water!





Last Chance

ENDURANCE CRATER

"Geyser" or "Fluid Vent"

QuickTime[™] and a TIFF (Uncompressed) decompressor are needed to see this picture.

|<--- ~ 10 cm --->|

Opportunity - Endurance Crater "Geyser Site"



Water Leaking from Beneath the surface in Bonneville Crater

Tufas

Evaporite Deposits ?

QuickTime[™] and a TIFF (Uncompressed) decompressor are needed to see this picture.

UNUSUAL MORPHOLOGICAL FEATURE: MICROFOSSIL - ?? ARTIFACT OF "RAT" ABRASION TOOL??



File: 1M130859941EFF0454P2949M1M1; Opportunity, Sol 30



- \Rightarrow Presence of sulfates suggest very acidic fluids (pH 2-6)
- ⇒ Remember: Most living systems prefer slightly alkaline fluids. However, there are exceptions.



LOST OPPORTUNITIES FOR BIO/GEOSCIENCES

- ANOTHER LOCATION ON MARS
- UV FLUX AT SURFACE SURVIVAL OF ORGANICS
- ISOTOPIC MEASUREMENTS OF MARTIAN H, C, N, AND O In Situ for BIOGENIC ELEMENTS VIA STEPPED COMBUSTION
- METHANE'S ISOTOPIC COMPOSITION AND ABUNDANCE IN ATMOSPHERE - DETERMINE IF BIOTIC OR ABIOTIC ORIGIN
- GROUND TRUTH FOR HYDROGEN ABUNDANCES SUPPORTS MARS ODYSSEY MISSION
- SUBSURFACE AND CORE SAMPLES FOR GAS ANALYSIS AND COMPOSITIONAL STUDIES
- COMPOSITIONAL ANALYSIS OF WEATHERING SURFACES
- AGE DATING SURFACE MATERIALS CALIBRATION OF CRATERING RATES

Evaporites and Salts at Meridiani Site Gypsum MoldsÓ?

A Beagle 2 Instrument Suite could provide information on unanswered questions



GAP has capability to distinguish between hydrates, carbonates, sulfates, etc. along with measuring isotopic compositions which would have assisted in identifying origins of evaporites

Hematite Concretions on Earth often have a N



http://marsrovers .jpl.nasa.gov/newsroom/ pressreleases /20040318a.html





MARS IN THE PAST and TODAY



Recently Active Volcanoes? Living Mars





Pit crater chain of vents

KNOWN VENTS: The *"Yellowstones* " of Mars?

Rille and filled vent



Living Vent, Earth



HOW DOES LIFE BEGIN ? PROPER INGREDIENTS REQUIRED



The Biological Envelope



The biological envelope is defined by the volume of multi-dimensional

known organisms

User Temperature (??) heeds investigating

(Don Cowan, 2004)

On Earth, water, energy, C sources are readily available so we find microbial life on modern Earth anywhere and everywhere....

Water is key to sustaining life as we know it.







Early Mars: Water Energy C source

> Life on Mars?

Rocks are Science's "Gold"

- Samples from the Moon were PROFOUND!
- Meteorites have given us the chemistry of our Solar System
- Meteorites from Mars have shown us the promise that samples from Mars will deliver
- Rock (and dust/soil) samples provide...
 - Chronology of key events
 - Chemistry related to LIFE
 - History of climate
 - Volatile reservoirs
 - Definitive Measurements of Provenance
 - Ground-truth to direct future exploration
- The Record of Life will be preserved in rocks (like artifacts) if it is there

Importance of Sample Return









Oldest Microfossils on Earth?



Warrawoona Group, N. Pole Dome/ Marble Bar, WA; 3.5 Ga





Ancient Life on Earth: The 2002 Debate





NTwo scientists bicker over whether a stone contains the oldest fossils on Earth or just junk.
How will they ever decide about life on Mars?Ó

Oliver Morton, Newsweek, International Issue, March 18, 2002 **Evidence for Life on Mars**

Martian Meteorite Evidence - 3 rocks magnetites, indigenous reduced carbon, lack of ¹⁴C, <u>N</u>structures Ó biofilms

Persistent Presence of Water and Evaporites NLife requires presence of water Ó

Unique ÓFeaturesÓwithin Soils and Rocks (filaments, spirals, etc.)

Presence of Methane in Atmosphere-NEW!
 3 measurements ~10-20 ppbV
 Hydrothermal-?? Biogenic-Possibly
 Label-Release Data from Viking ??









Fossil Biosignatures : What we look for ...



After DesMarais, 2005

ANCIENT LIFE ON MARS OR PRESENT LIFE ON MARS ?

OR BOTH 22