

CLOUD-TRACKED MARTIAN WINDS AS SEEN FROM HST IN OPPOSITION 2003

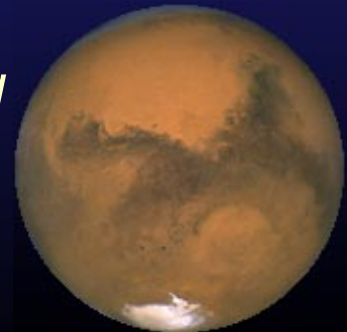
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CO 80303-1058, USA.*

⁴*Cornell Univ., Dept. of Astron. 402 Space Sci. Building
NY 14853-6801, USA.*

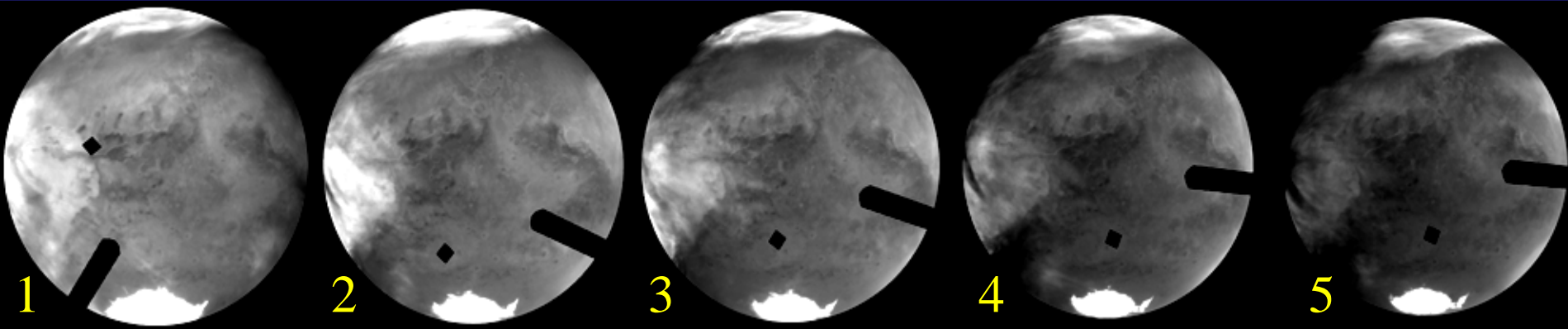


HST observations of Mars Opposition 2003

- Camera: ACS/HRC
- Disk size ~ 1000 pixels
- Highest resolution: (disk center) ~ 7 km/pixel

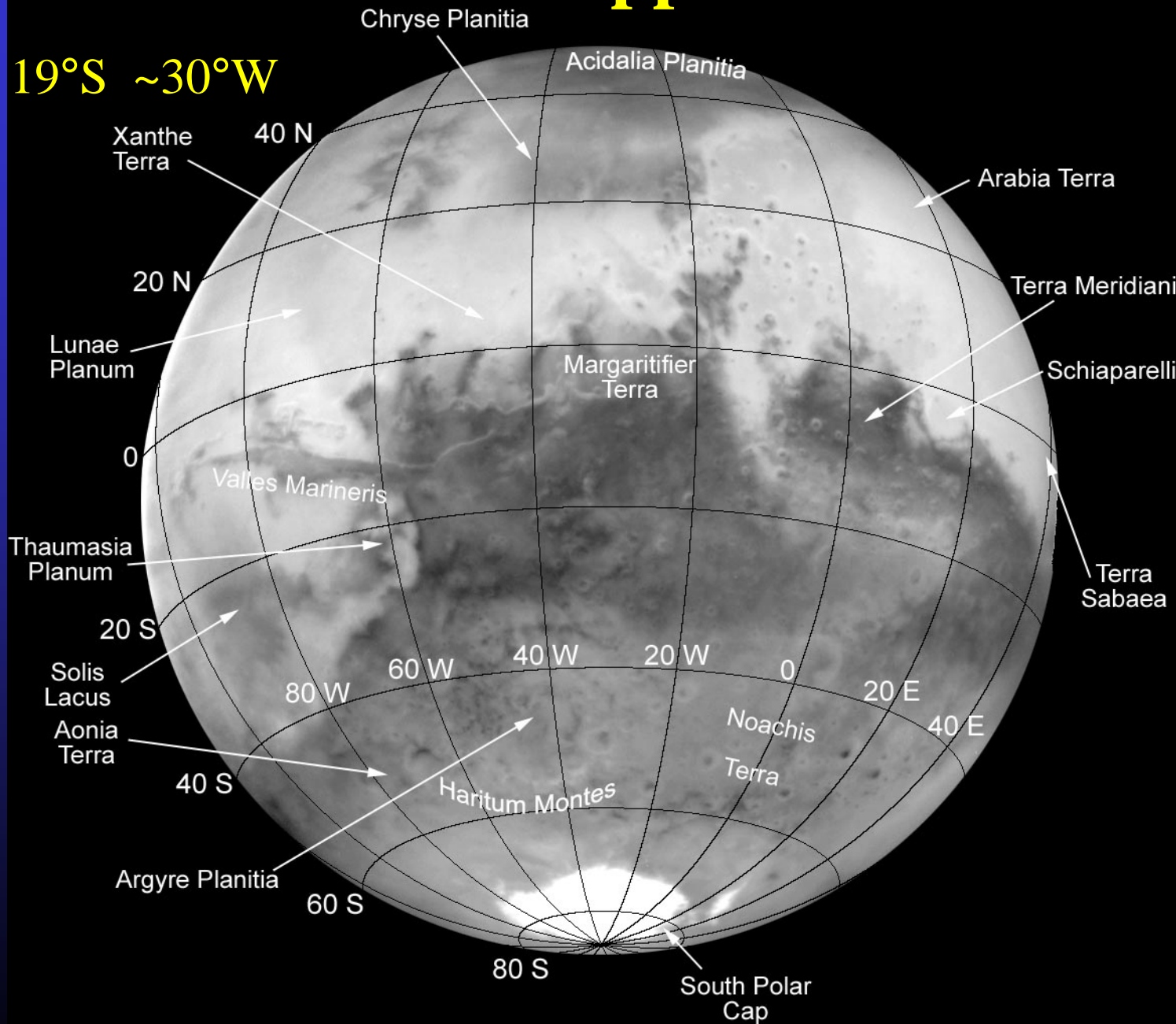
Date of Year 2003		L_s , deg.	Spectral filters	
1	Aug., 24	247	250 nm	UV I (+ 3 polarizations)
2	Sept., 05	254.5	330 nm	UV II (+ 3 polarizations)
3	Sept., 07	255.8	435 nm	Blue (+ 3 polarizations)
			502 nm	Green
4	Sept., 12	259	658 nm	Red
5	Sept., 15	261	892 nm	IR

Clouds: • Northern cloud belt • changing clouds system (west)



HST observations of Mars Opposition 2003

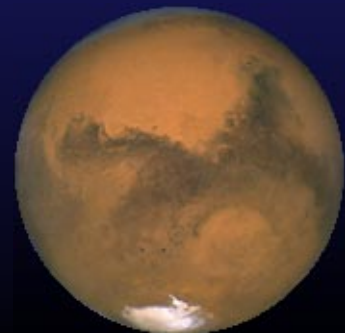
- **Disk center:** $19^{\circ}\text{S} \sim 30^{\circ}\text{W}$
- **Season:** summer in southern hemisphere



Red filter,
Sept. 15, 2003

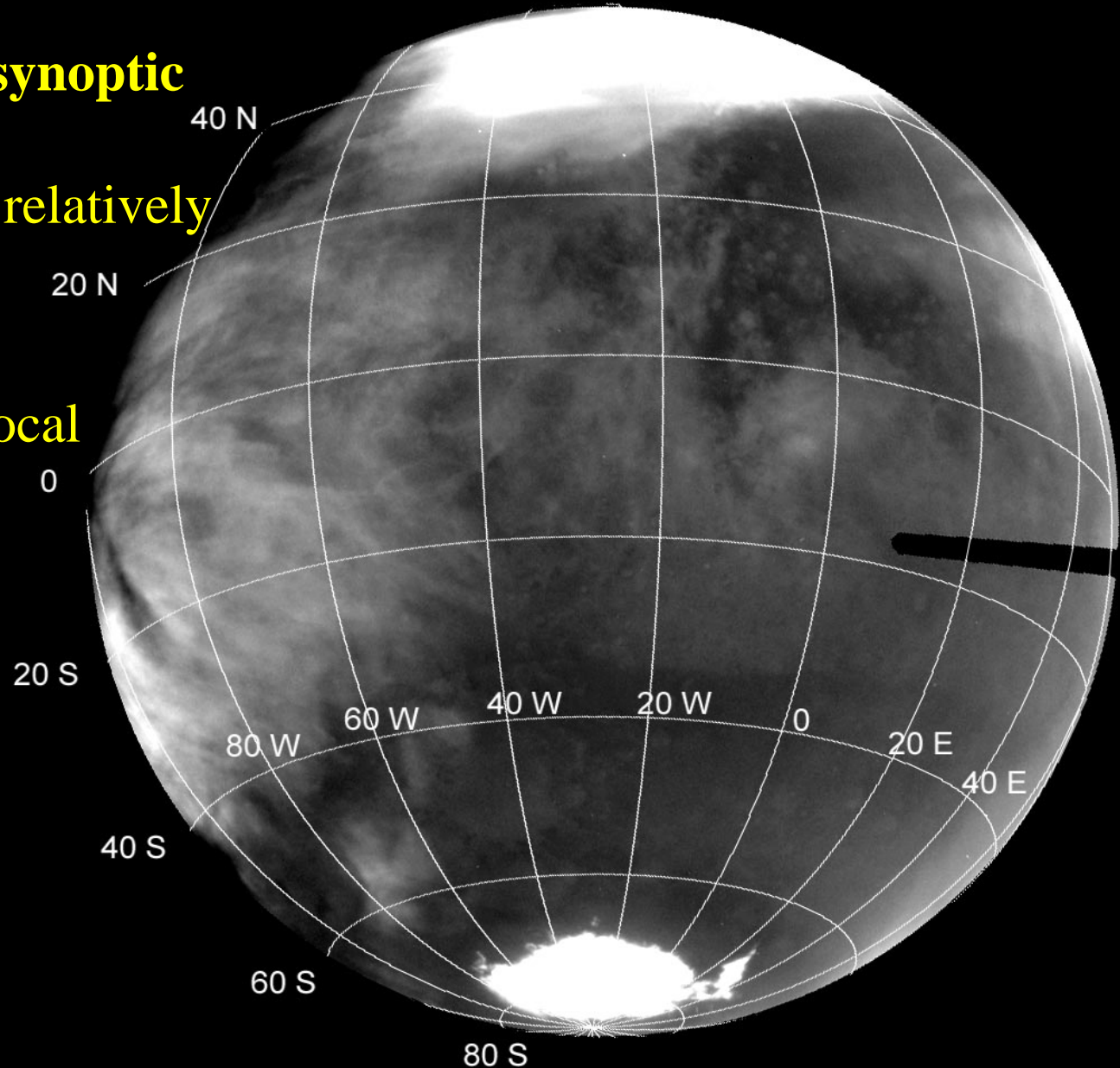
Image data processing

- **Standard pipeline calibration**
 - dark current
 - flat field
 - geometric distortion
- **Cosmic-ray-tracks removal** (abundant in the near-UV)
- **Coregistration**
 - 3 successive images within each near-UV set
 - Time lag ~2 min between images (Mars rotation)
 - Mapping into a similar projection
 - Smoothing



Wind speed measurements: cloud shifts

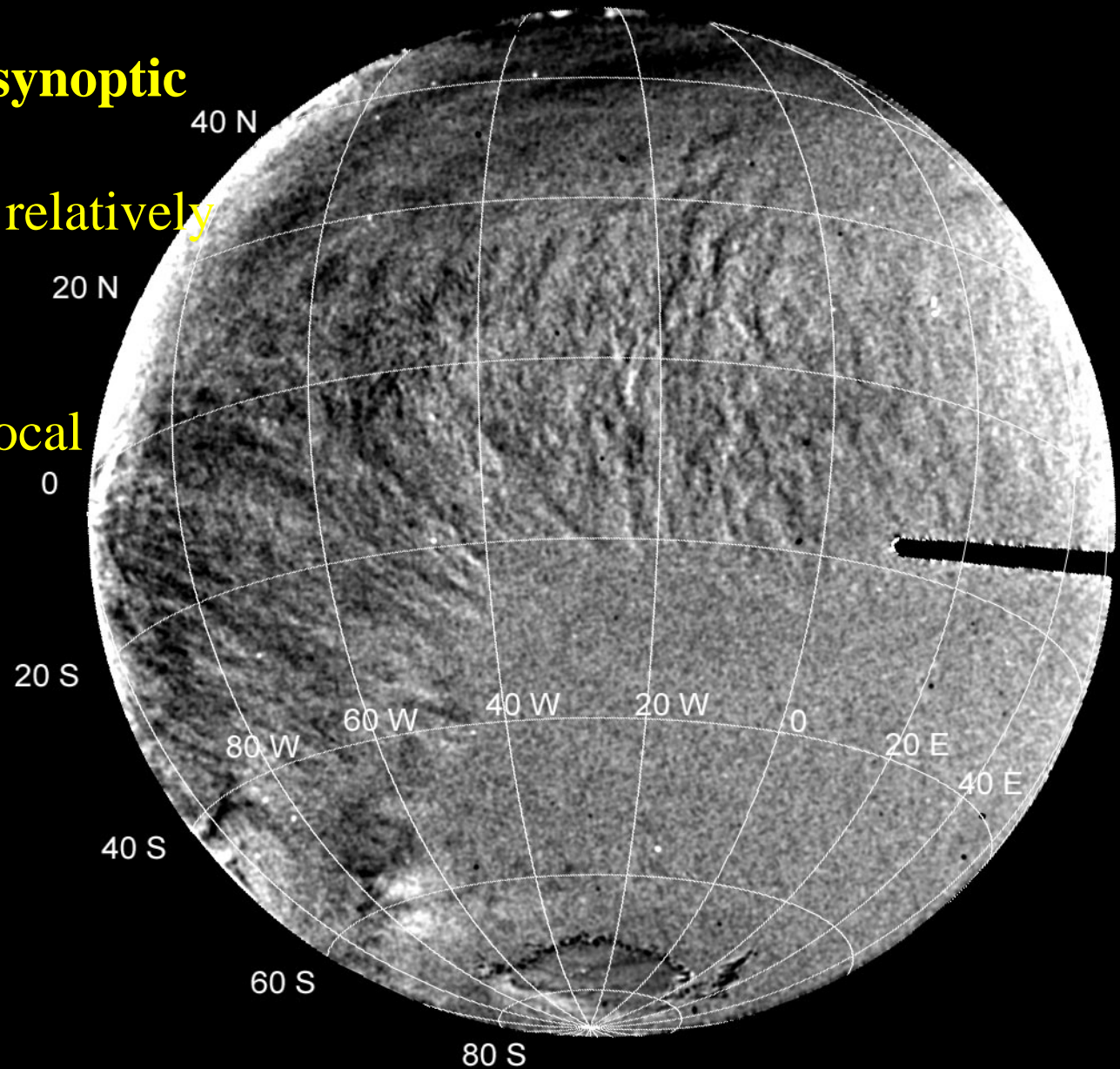
- **Detection of the synoptic motion:**
- clouds movement relatively to the surface
- coregistration – maximization of local covariation of images



UV II,
Sept. 15, 2003

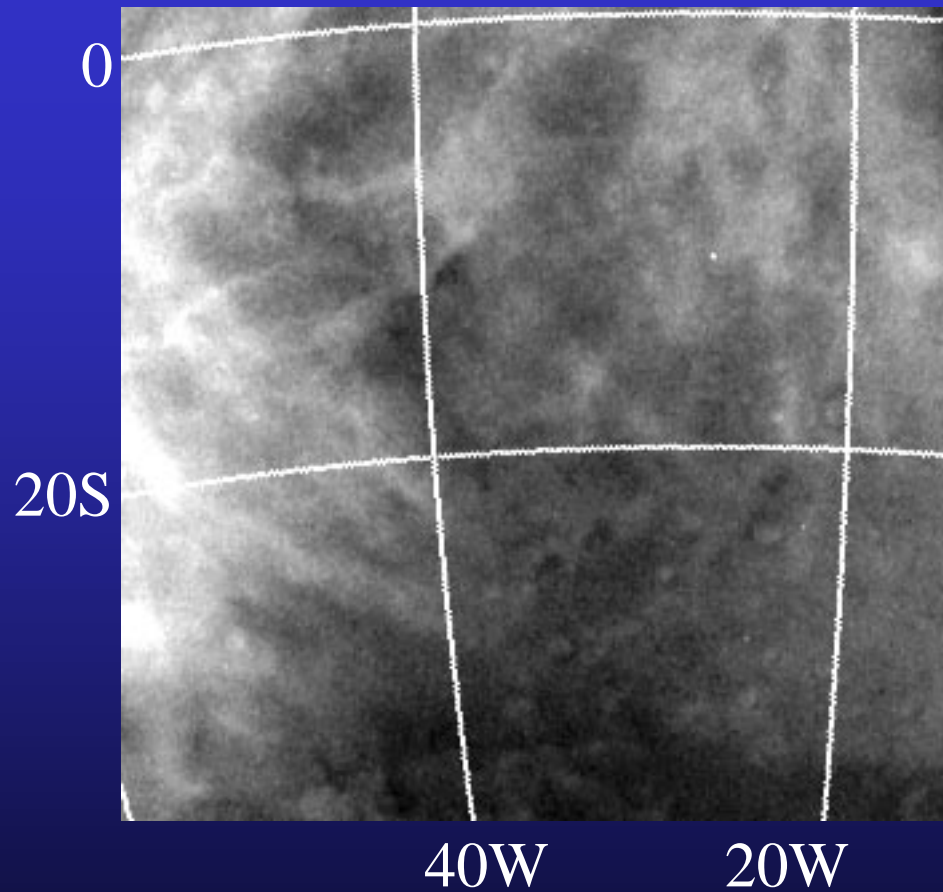
Wind speed measurements: cloud shifts

- **Detection of the synoptic motion:**
- clouds movement relatively to the surface
- coregistration – maximization of local covariation of images

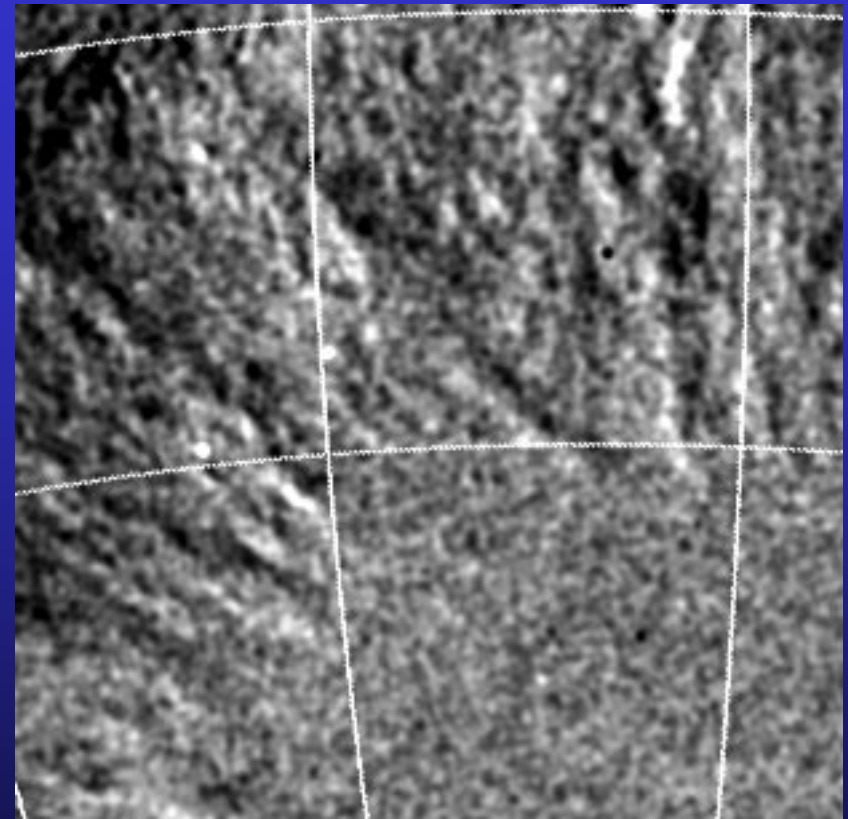


UV II Ratio:
POL120/POL0,
Sept. 15, 2003

Wind speed measurements: cloud shifts



UV II



UV II Ratio: POL120/POL0

Sept. 15, 2003

Results: wind speed measurements (Sept. 7, 2003)

• Determination:

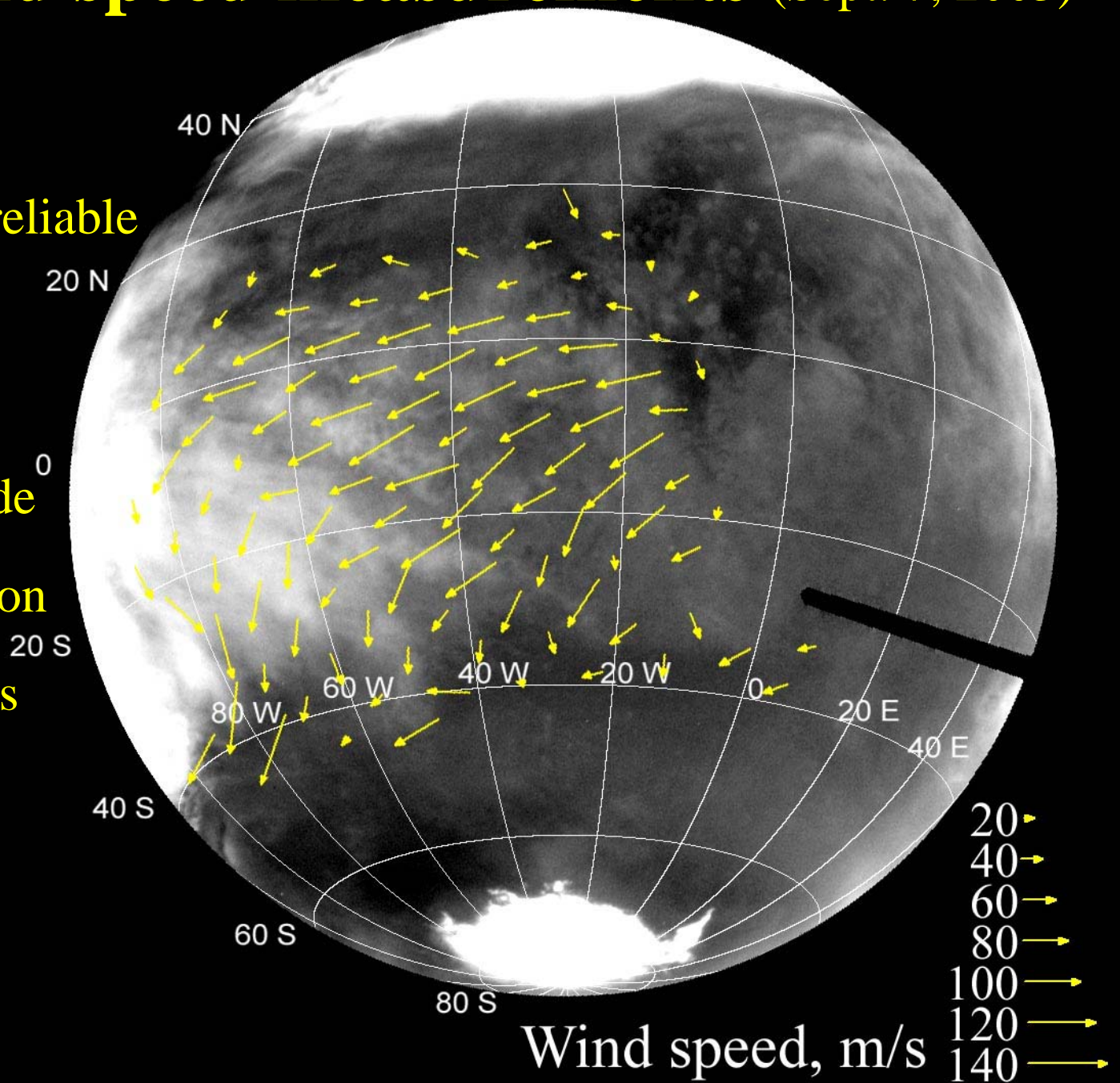
Tharsis rise: most reliable

• Features:

• Blowing retrograde

• Poleward deflection

• Orographic effects
(Tharsis edge)



Results: wind speed measurements (Sept. 15, 2003)

- Determination:

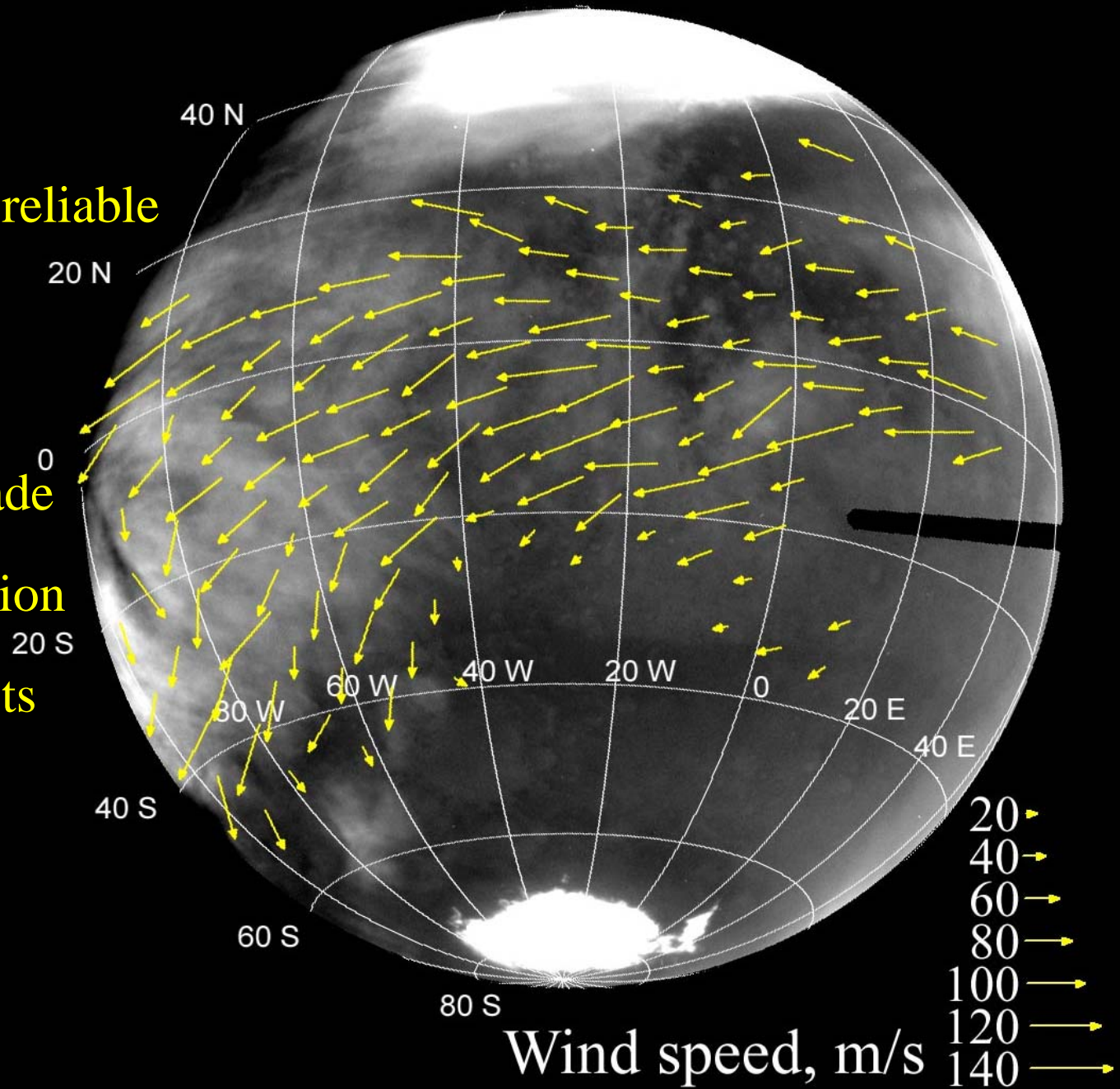
Tharsis rise: most reliable

- Features:

- Blowing retrograde

- Poleward deflection

- Orographic effects
(Tharsis edge)



Results: wind speed measurements (Aug.24, 2003)

- Determination:

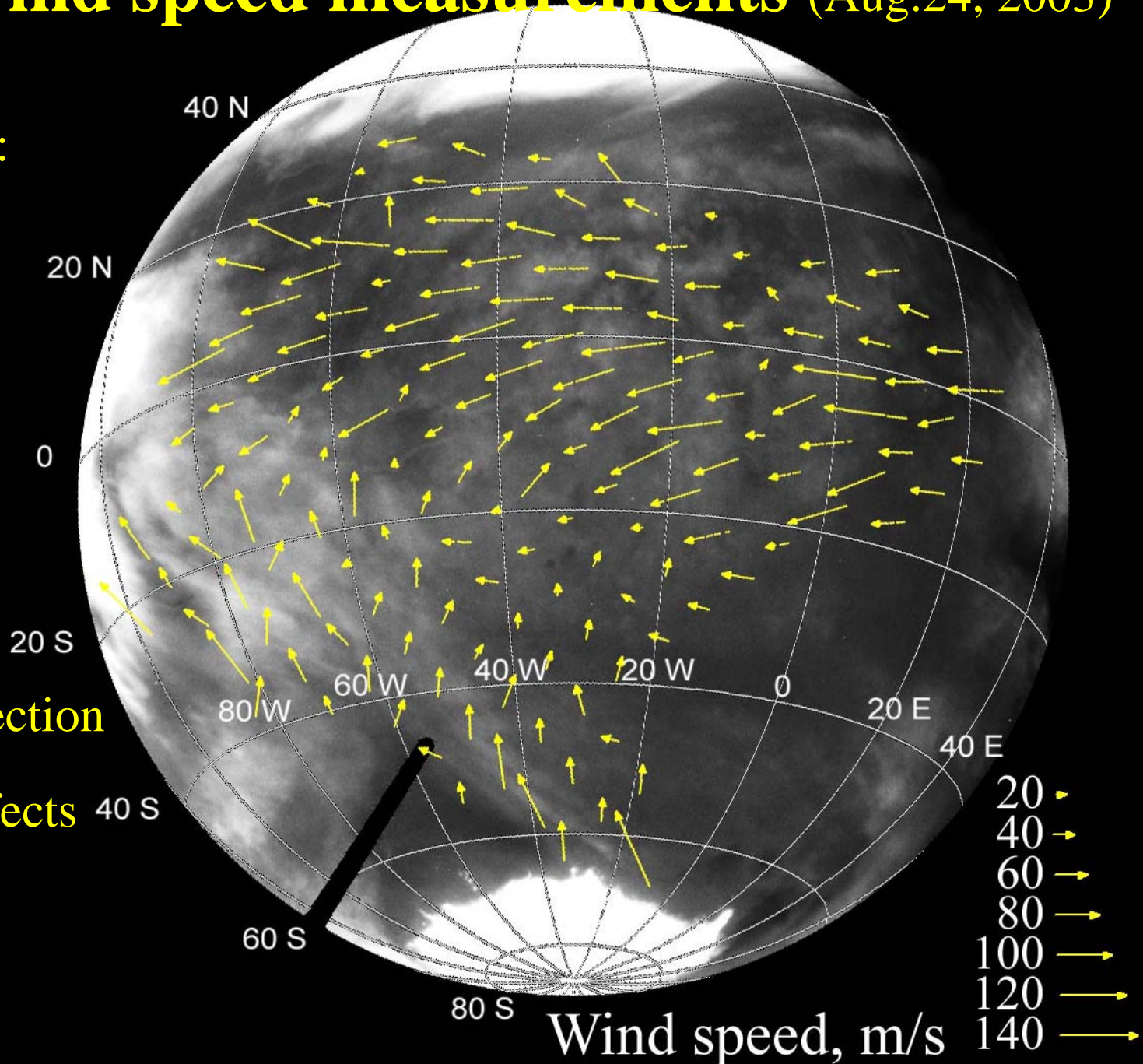
Tharsis rise:
most reliable

- Features:

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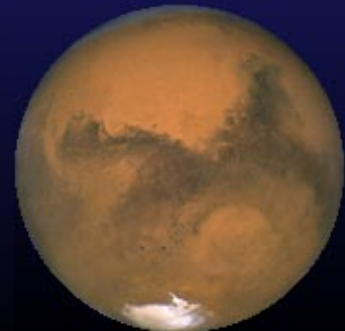
- Poleward deflection

- Orographic effects
(Tharsis edge)



Wind speed measurements: source of errors

- maximization of local covariation of images:
 ± 0.1 px $\rightarrow \sim \pm 3$ m/s & ± 6 deg.
- inaccuracy of scale knowledge (at 10^{-3} level)
- inaccuracy in the location of the Mars disk center
- individual field of view distortion of the polarization filters
(Biretta et al., *ACS 2004-09*, Baltimore: STScI, 2004):
 ± 0.3 px $\rightarrow \sim \pm 10$ m/s & ± 17 deg.
(smoothly spatially varying geometric distortion)

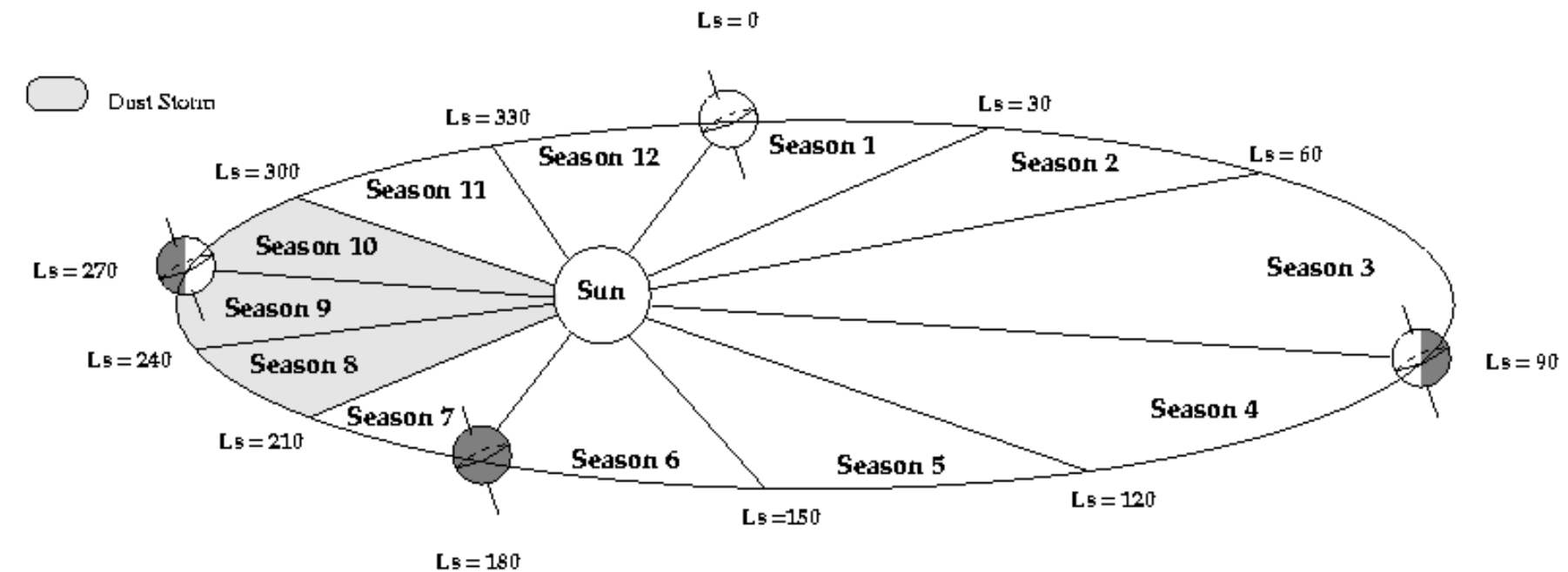


Wind speed: comparison with the Mars climate database

THE EUROPEAN MARS CLIMATE DATABASE

[Lewis, S.R., Collins, M. and Forget, F. (2001a)

Mars Climate Database v3.0] <http://www.lmd.jussieu.fr/mars.html>



$L_s = 240^\circ - 270^\circ$ (Season 9 – perihelion)

Wind speed: comparison with the Mars climate database

Dust scenario: Low / High ??

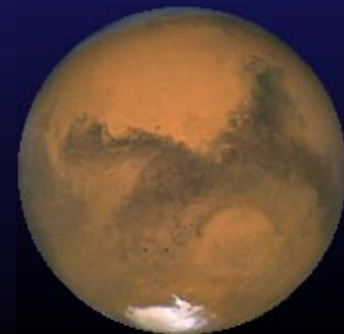
MGS TES atmospheric dust opacity ($9.7 \mu\text{m}$) retrievals (Sept., 2003):

- the atmosphere was relatively dusty in comparison to the aphelion
- the dust is evenly distributed in the atmosphere
- TES data do not show any pronounced dust-lifting events

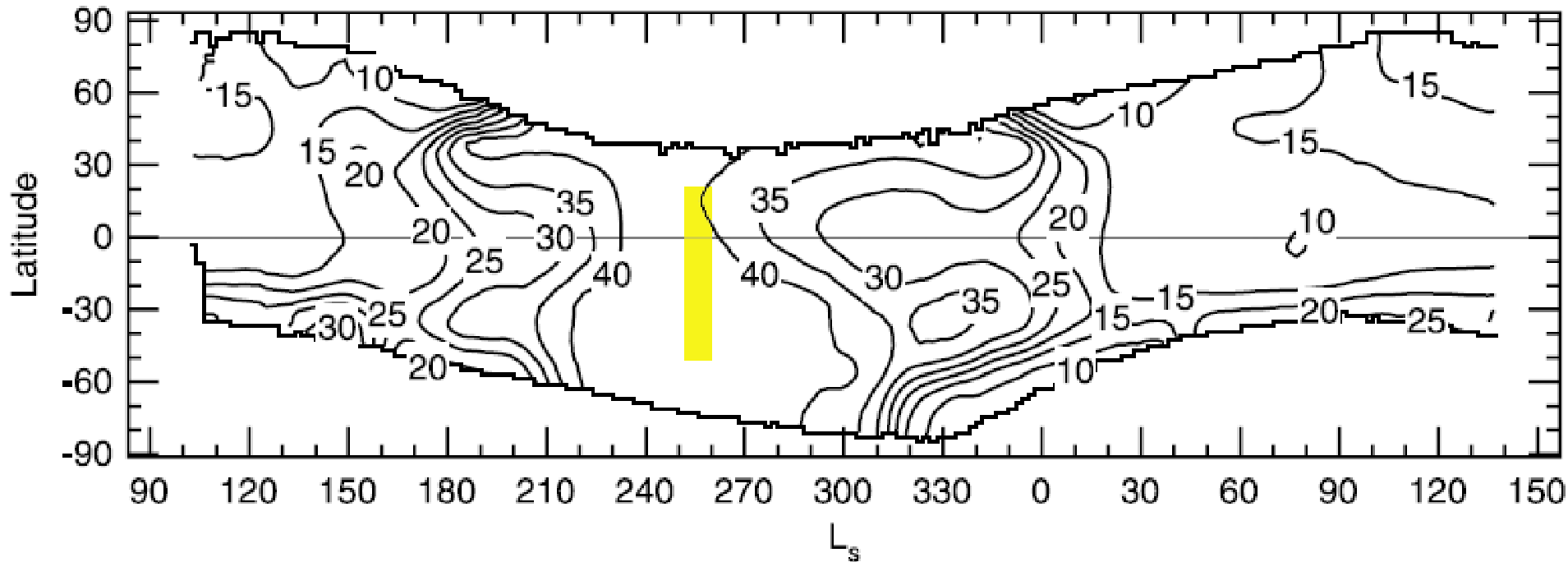
MGS Dust Scenario & “Viking Lander” Dust Scenario

Diurnal cycle:

HST images → morning for western part of disk,
evening for the eastern part



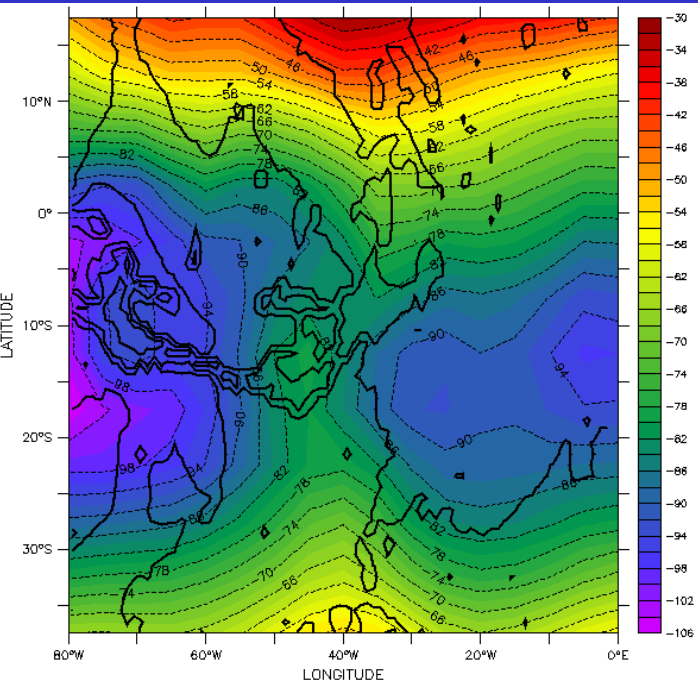
Altitude of upper cloud layer ?



The height in km above the surface of water condensation level
[From: Smith, M.D., JGR, 2002, vol.107, No.E11, 5115]

- Highest condensation levels ≥ 40 km
- Altitude of optically thick clouds for wind tracking ??

MGS dust scenario - Zonal wind

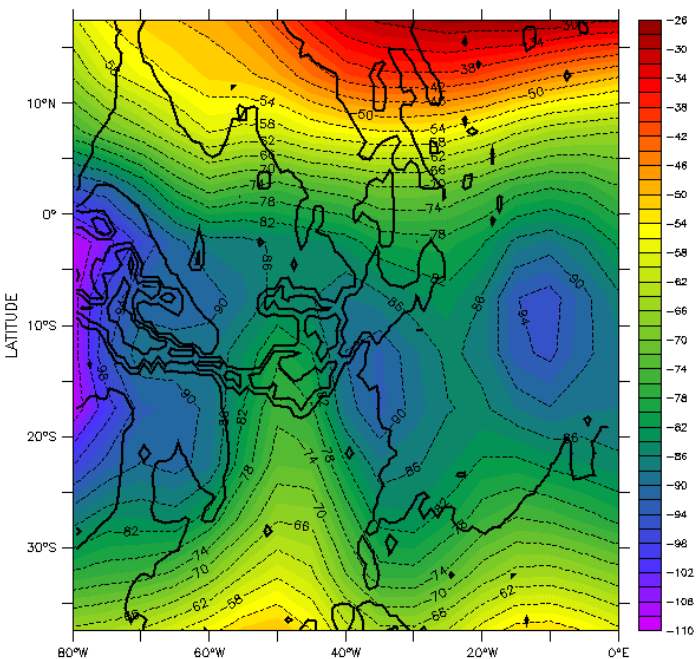


← LT 12:00

U

-100 ... -30 m/s

Layer 17
(43km) –
 $\sigma = 0.011008$

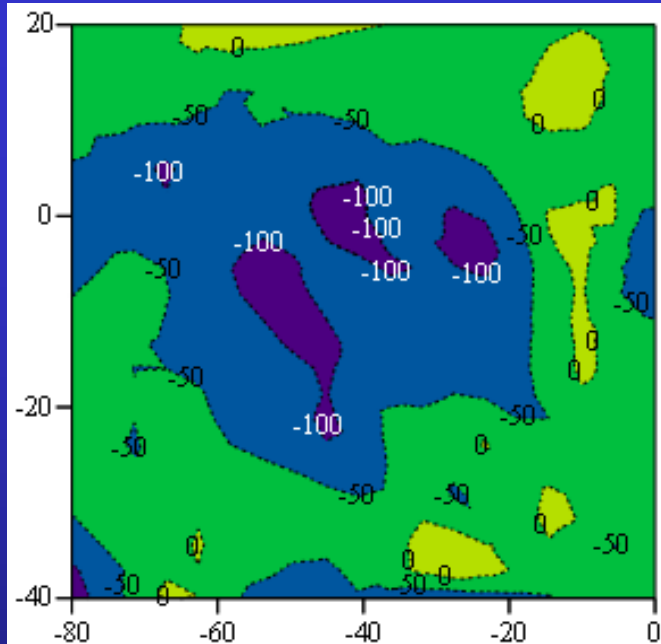


← LT 08:00

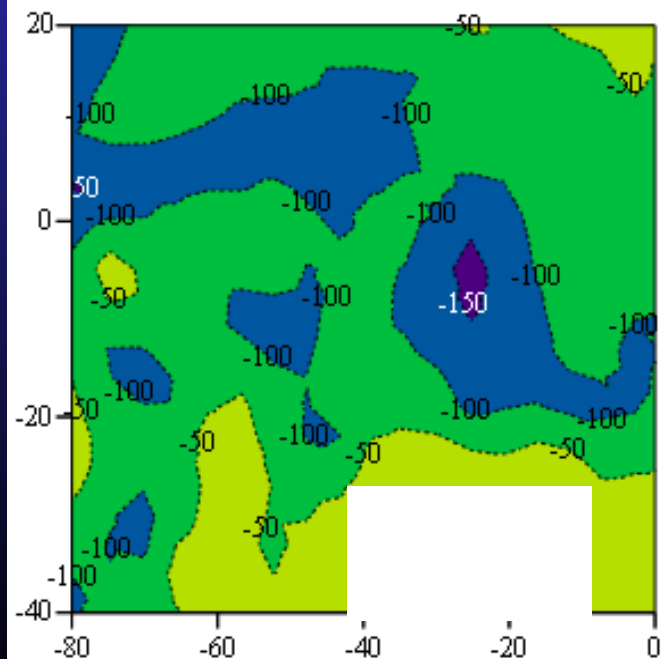
U

-110 ... -30 m/s

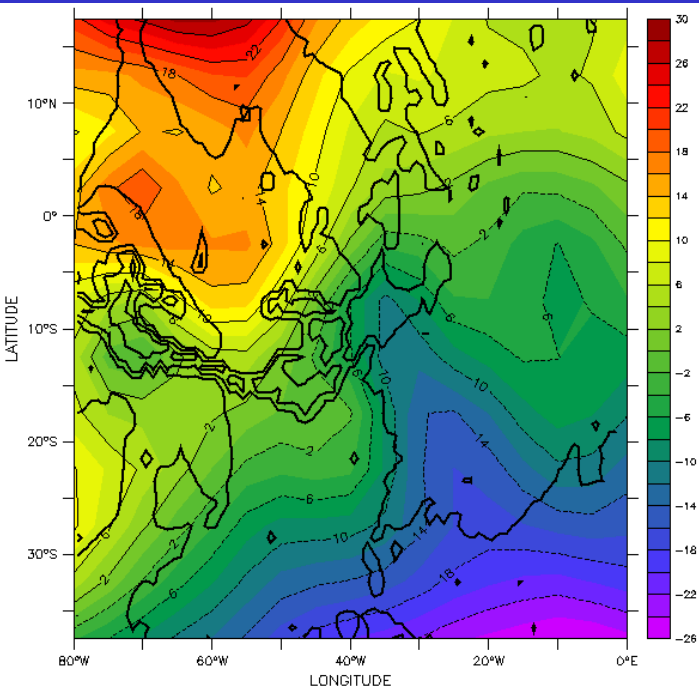
Sept., 7



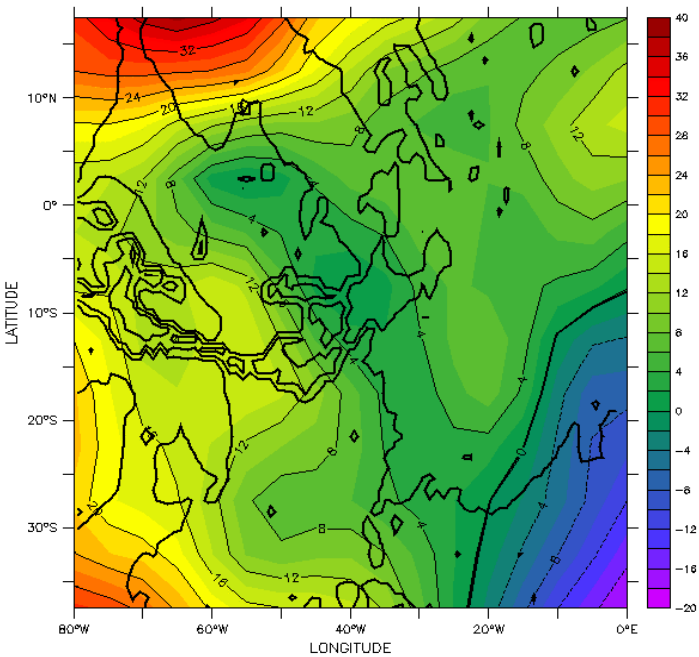
Sept., 13



MGS dust scenario - Meridional wind

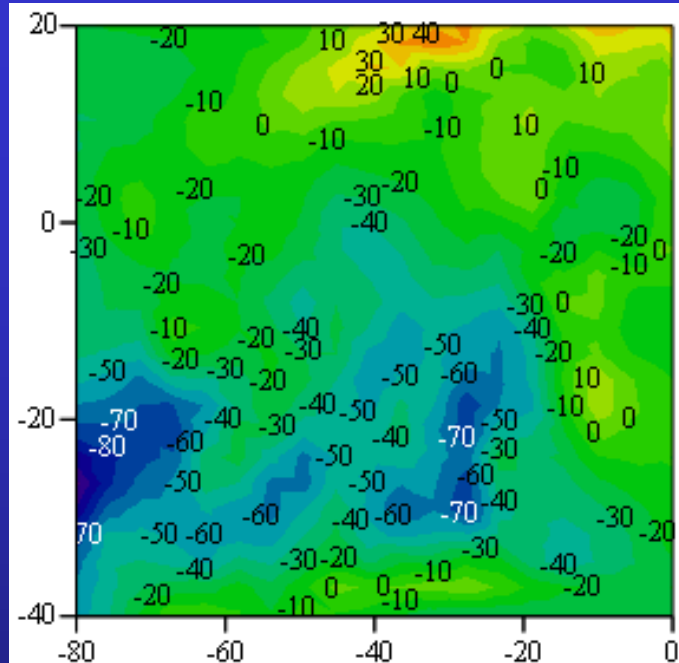


← LT 12:00
V
-30 ... 30 m/s

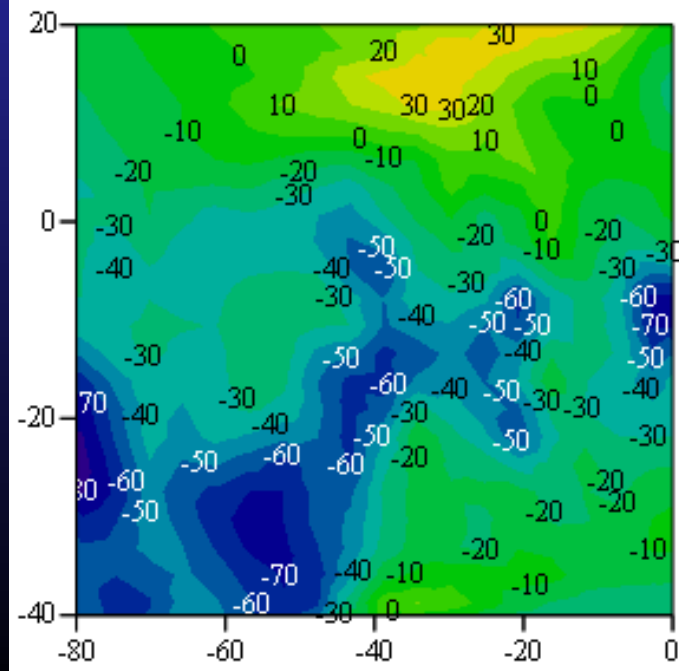


Layer 17
(43km) –
 $\sigma = 0.011008$

← LT 08:00
V
-20 ... 40 m/s

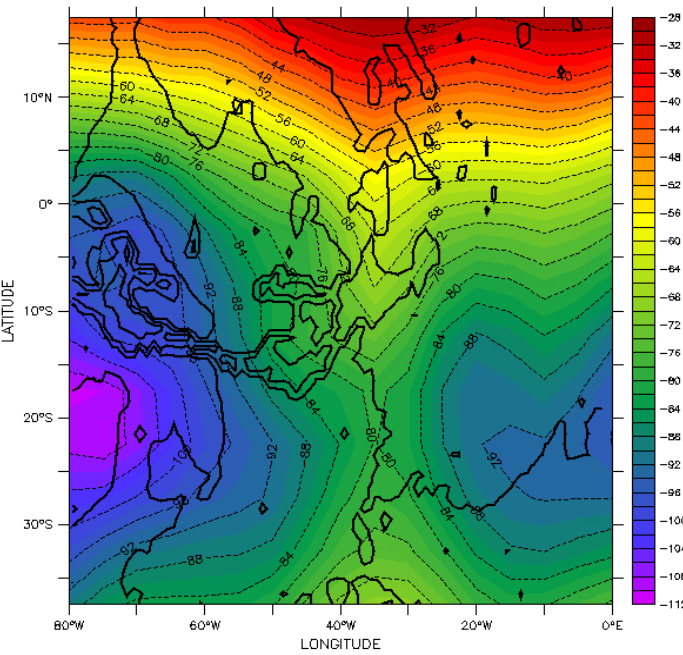


Sept, 7



Sept, 13

Viking dust scenario - Zonal wind

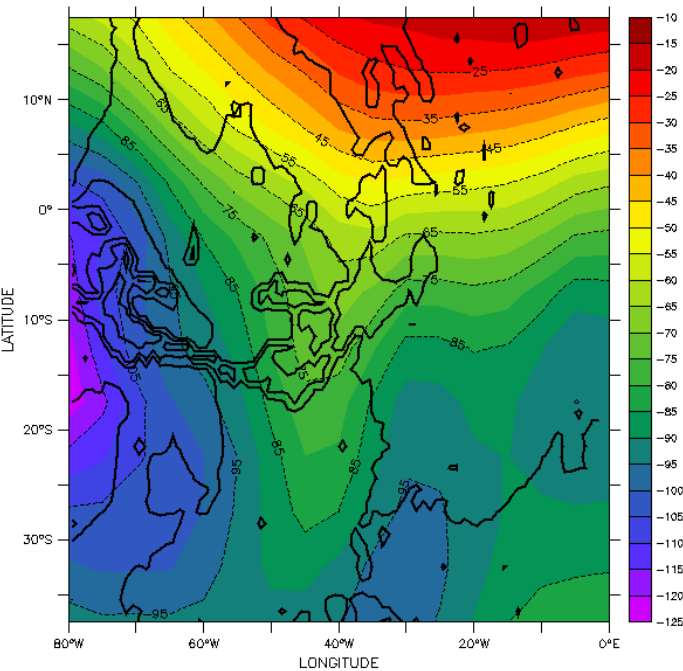


← LT 12:00

U

-110 ... -30 m/s

Layer 17
(43km) –
 $\sigma = 0.011008$

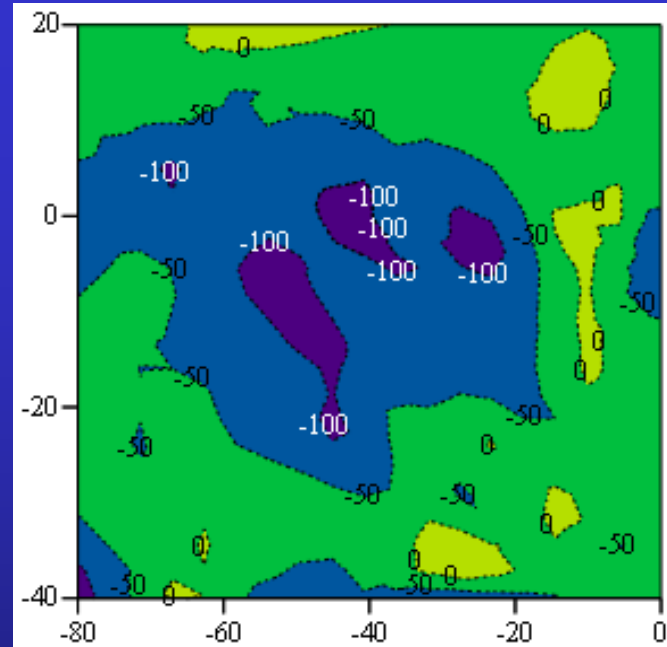


← LT 08:00

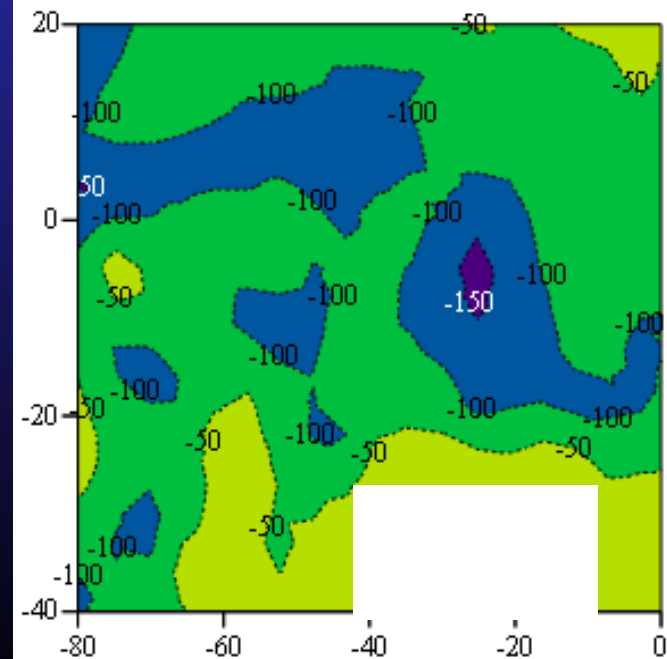
U

-125 ... -10 m/s

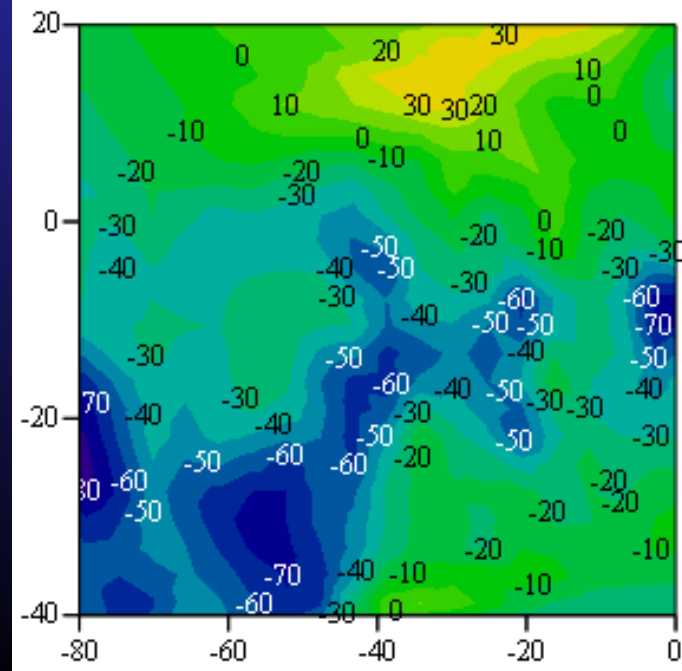
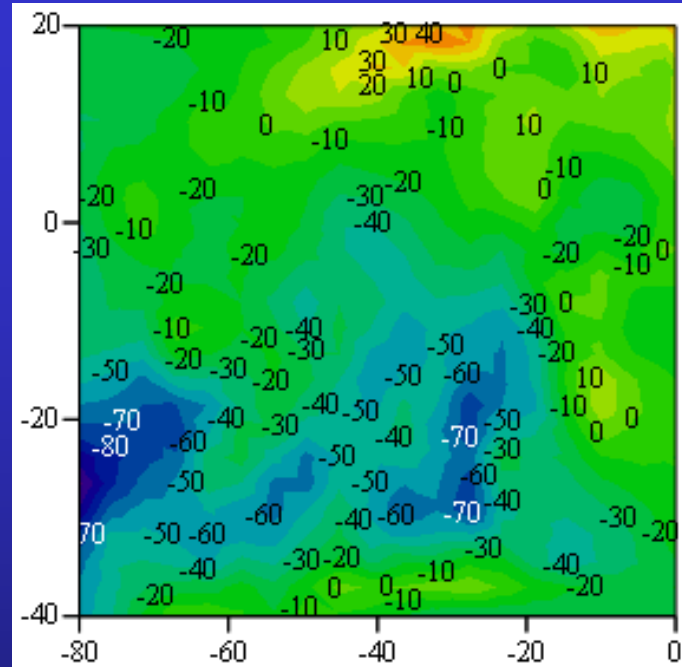
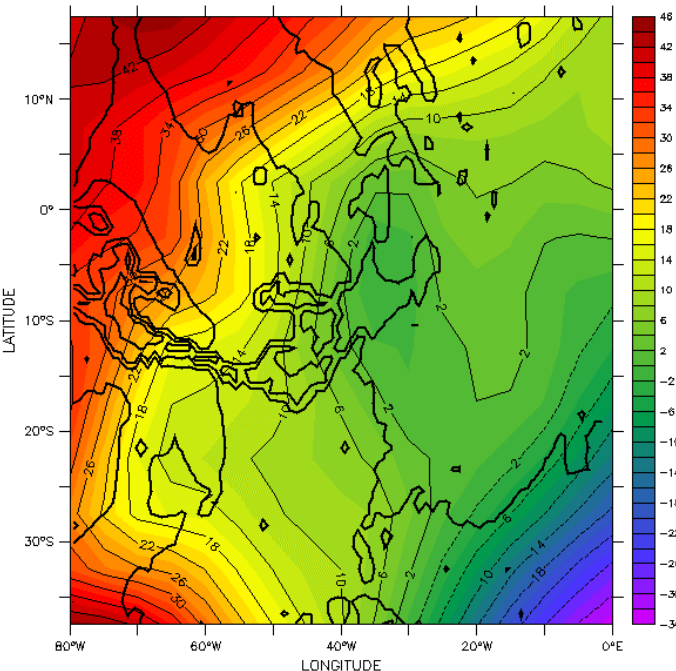
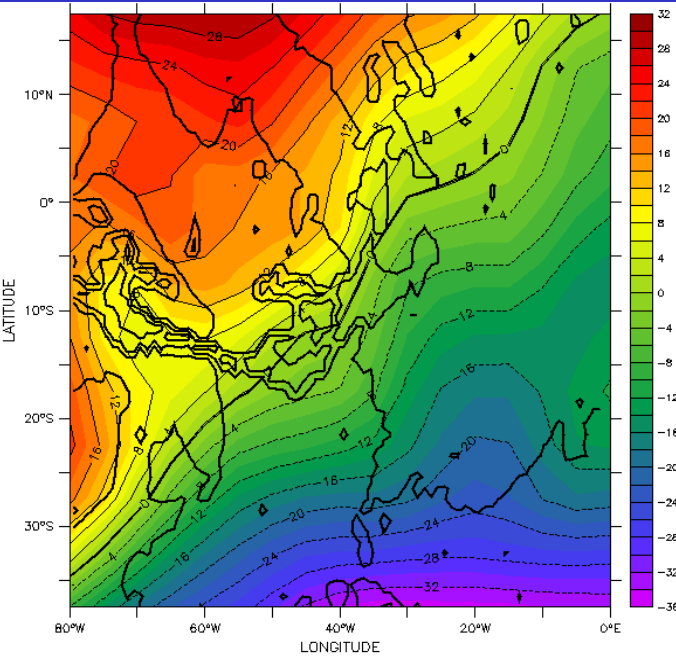
Sept, 7



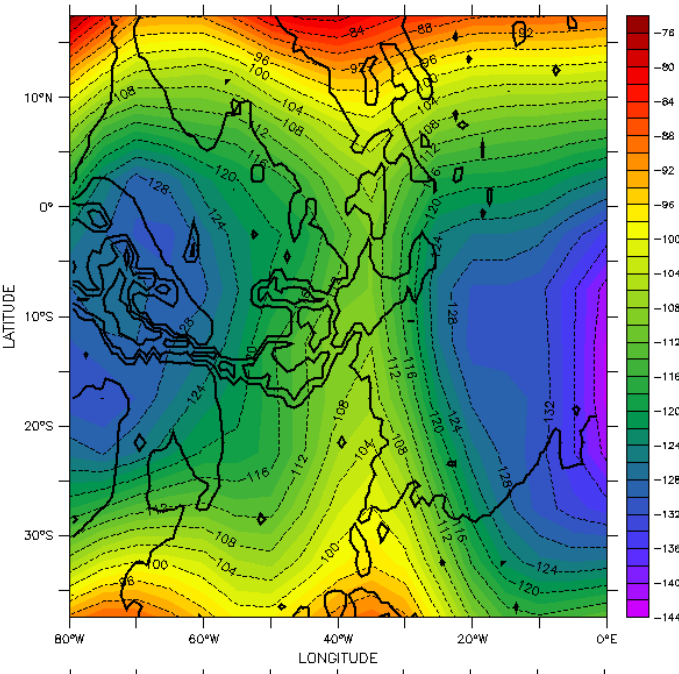
Sept, 13



Viking dust scenario - Meridional wind

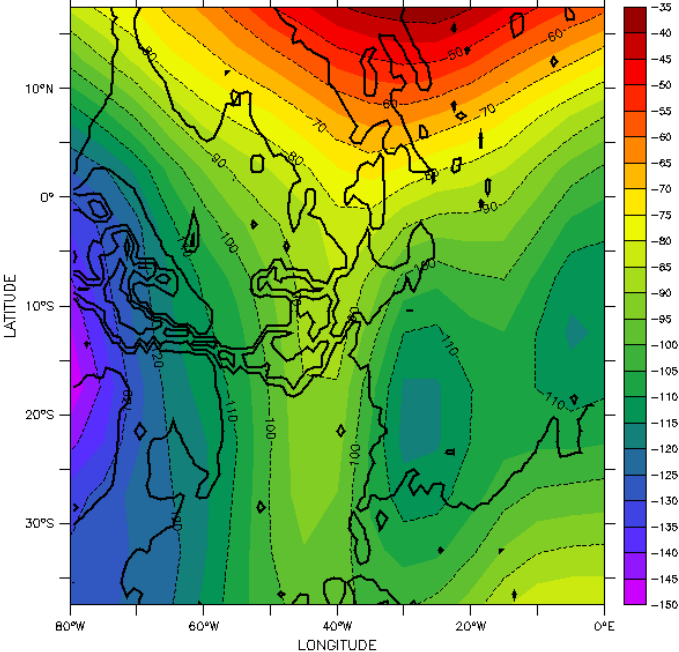


Viking dust scenario – Zonal wind



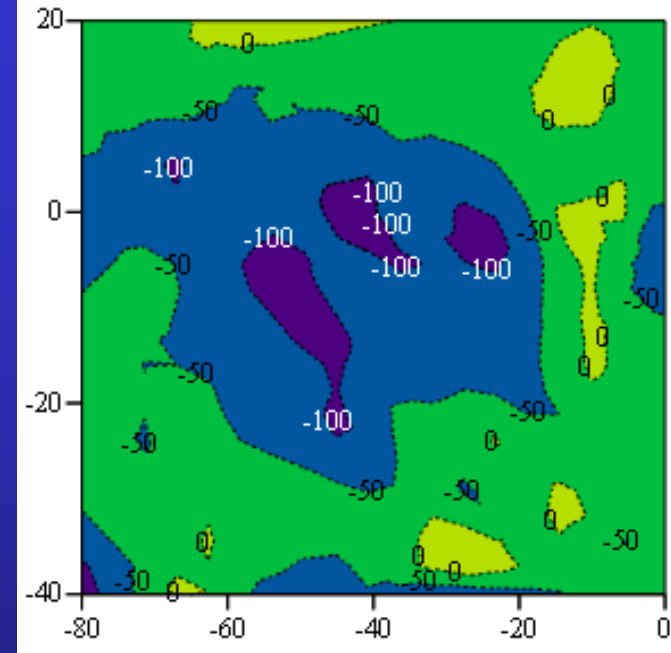
← LT 12:00
U
-145 ... -75 m/s

Layer

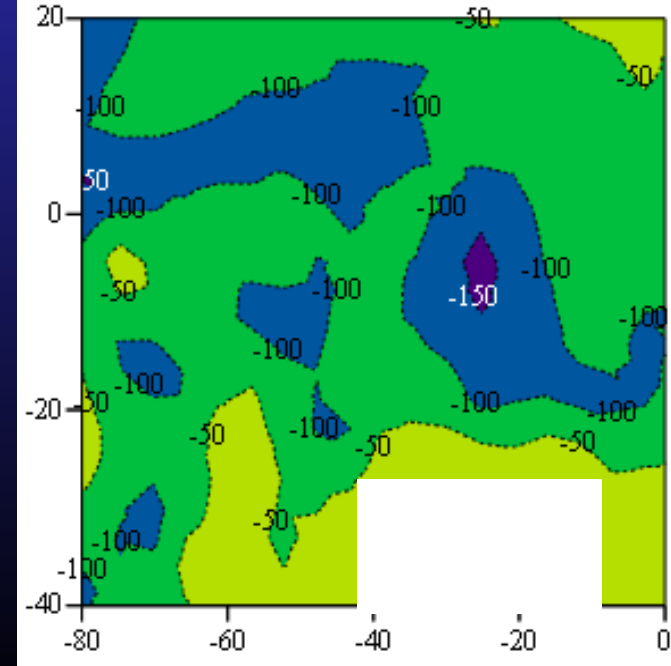


(65km) –
 $\sigma = 0.00277$
← LT 08:00
U
-150 ... -35 m/s

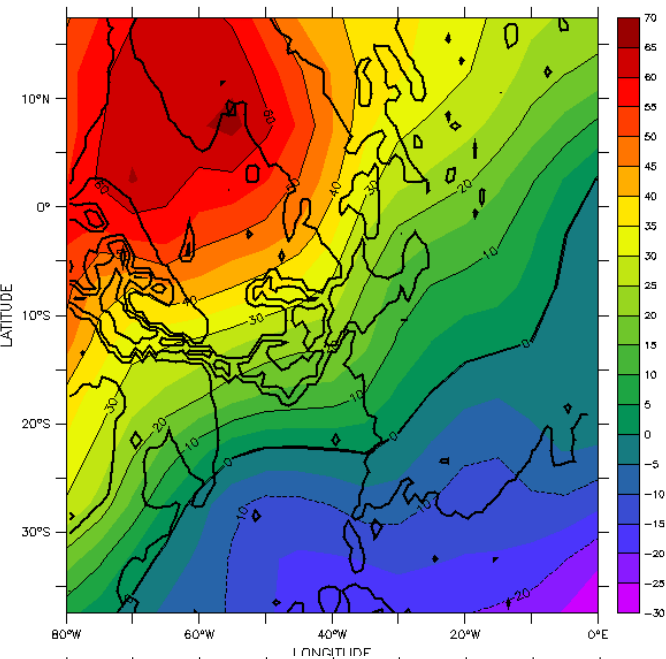
Sept, 7



Sept, 13



Viking dust scenario - Meridional wind

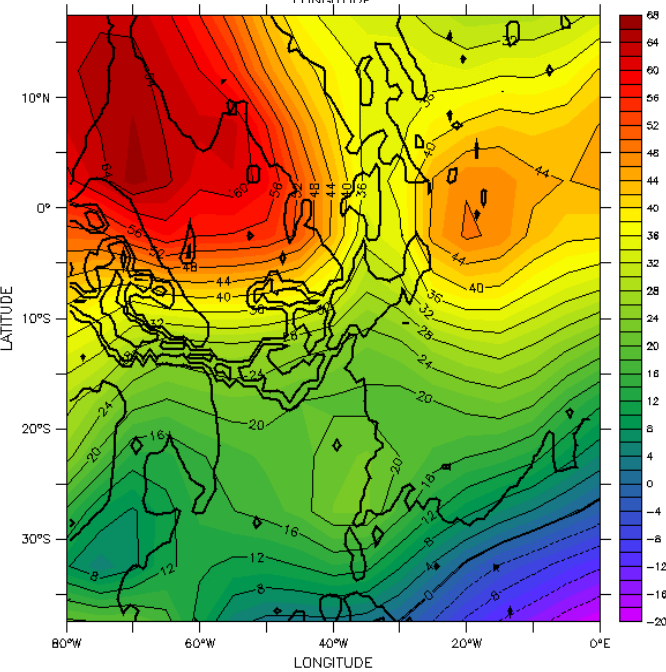


← LT 12:00

V

-30 ... 70 m/s

Layer



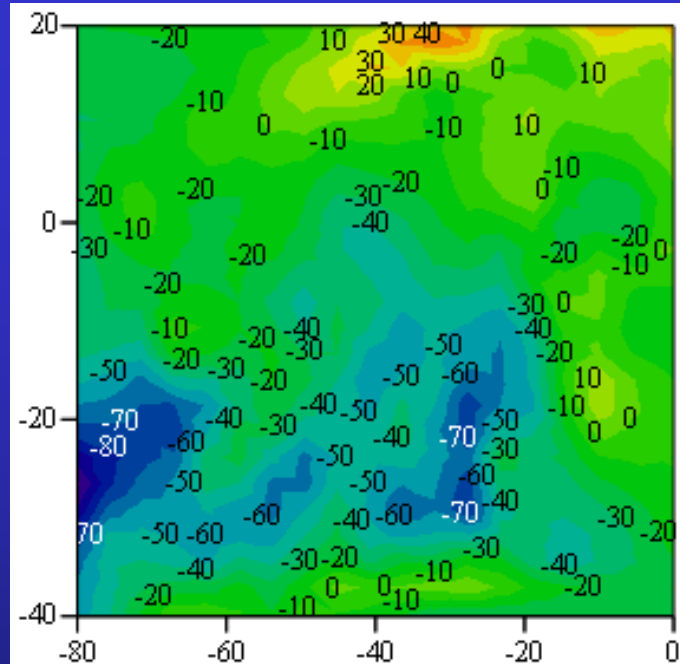
(65km) –
 $\sigma = 0.00277$

← LT 08:00

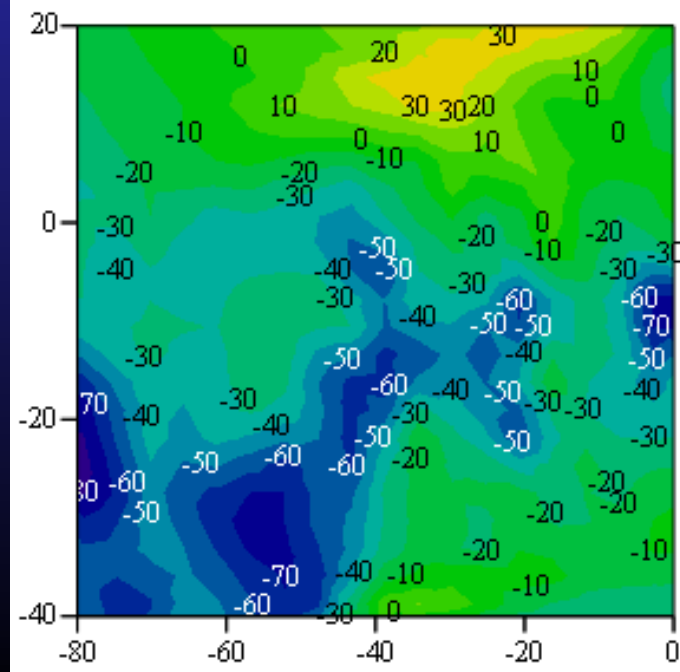
V

-20 ... 65 m/s

Sept, 7



Sept, 13



Conclusions

- **Wind field features:**

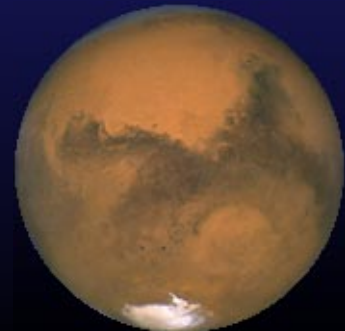
- ✓ Retrograde to the planetary motion
(pole-equator temperature gradient)
- ✓ Poleward deflection (southern latitudes)
- ✓ Deviation of wind direction pattern near Eastern edge of Tharsis rise

- **MARS CLIMATE DATABASE comparison:**

- ✓ General consistency assuming clouds condensation level ≥ 40 km
- ✓ Higher negative U and V observed

- **Synchronous HST and MeX (PSF) observations:**

- ✓ Complementary measurements of Martian winds



Thank you
for attention!

