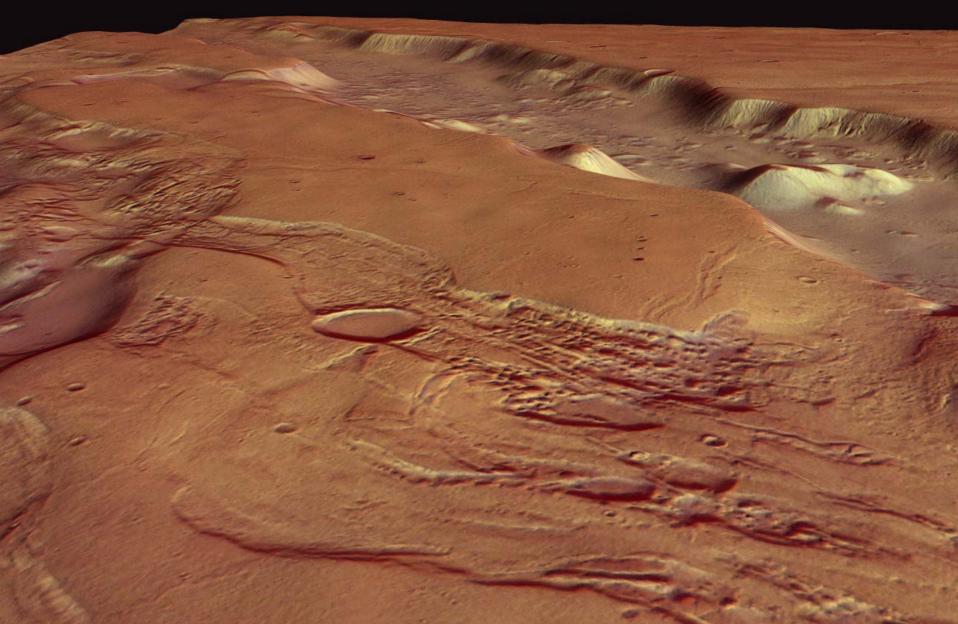
# **PRESS EMBARGO**

The following presentation will be published in Nature on March 17<sup>th</sup>, and is subject to Nature's embargo on press reporting.

# Pack-ice on Mars

#### John Murray, Jan-Peter Muller, Gerhard Neukum & the University University College London HRSC Mars Express team

## Dao & Niger Valles Evidence of catastrophic flooding events





PLANUM

chasma

Perrolla

#### PLANITIA

Viking 1 Landing Site

ER

Chasma .

Aurorae

E

Pathfinder Landing Site

> Masursk Sagan

> > Galilaei

rsinoes

Chaos

# Where does the water end up?

MARINERIS

AUMAS

Lassel

Chié

OP

PLANUM

Tholus

ithonium Chasma

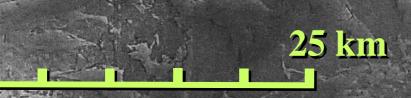
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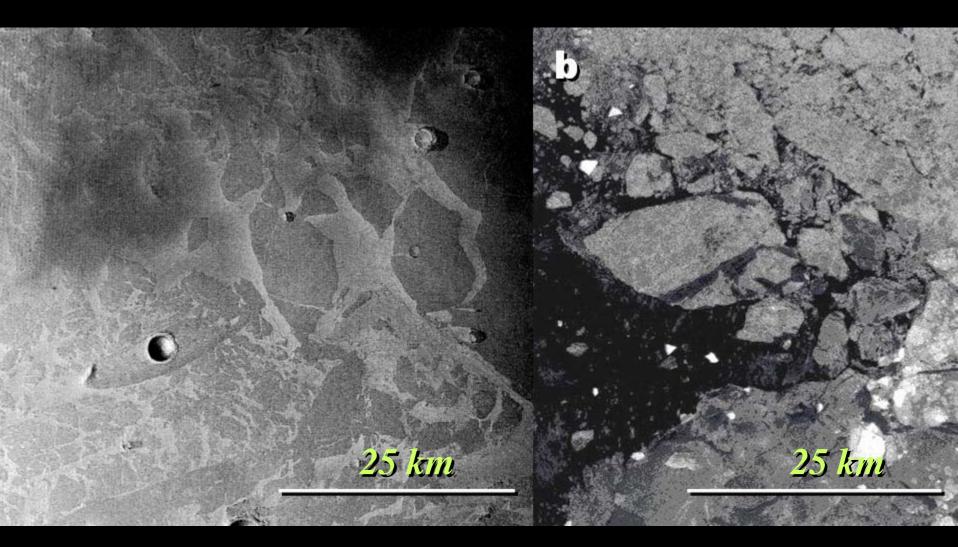
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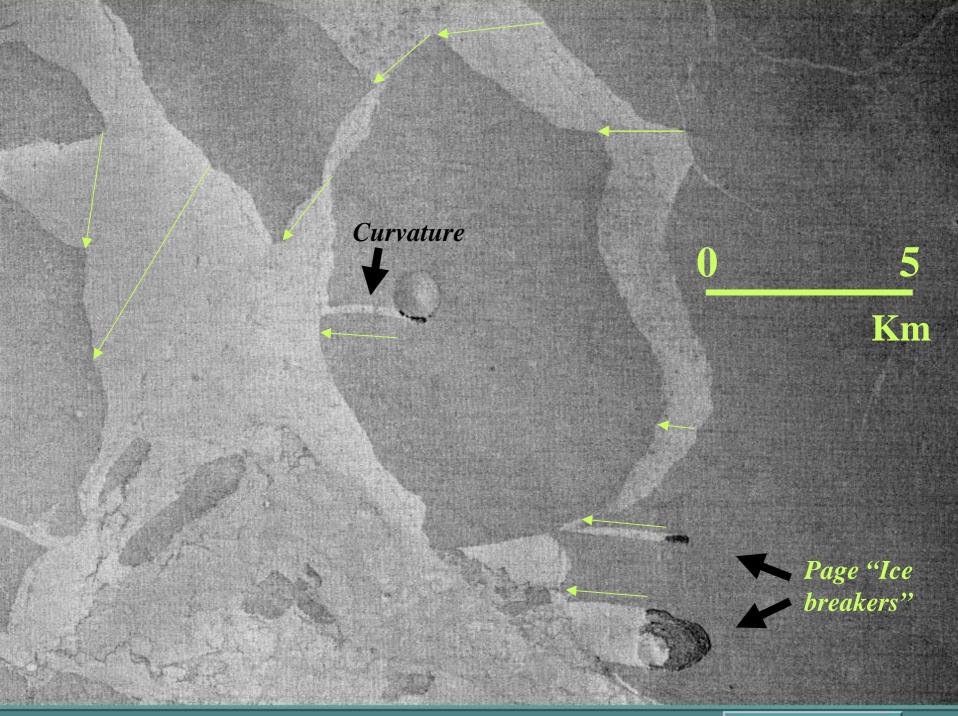
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## MARS



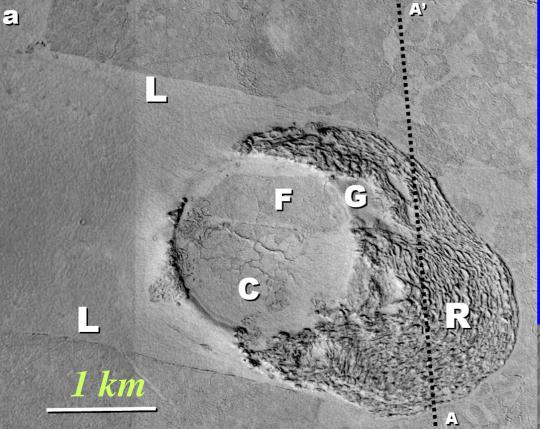


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CALLE OF DEPITCI IC

km

Ice Compression ridges



# MARS Pressure ridges

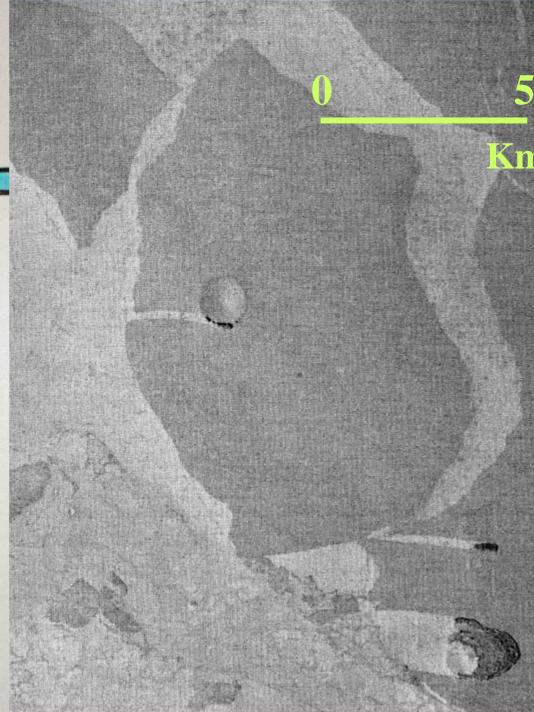
100 m

Arctic Ocean Pressure ridges round an artificial island

# 1100m diameter

Maximum Depth = 45 metres

Approximate Diam. : outer rim height Ratio for fresh impact Craters = 0.04

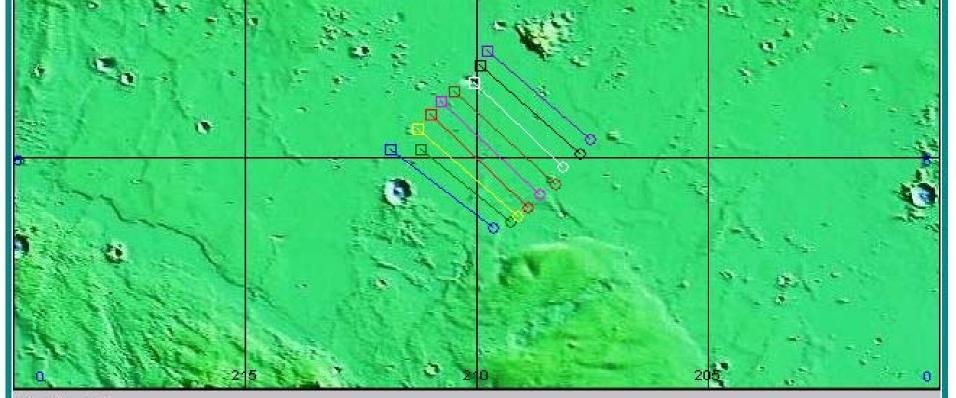


### MOC Image

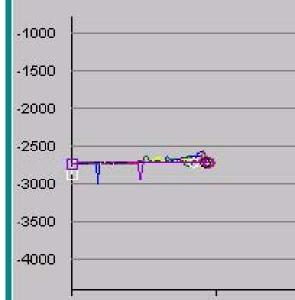
# Age = 5 million years

# (very young)

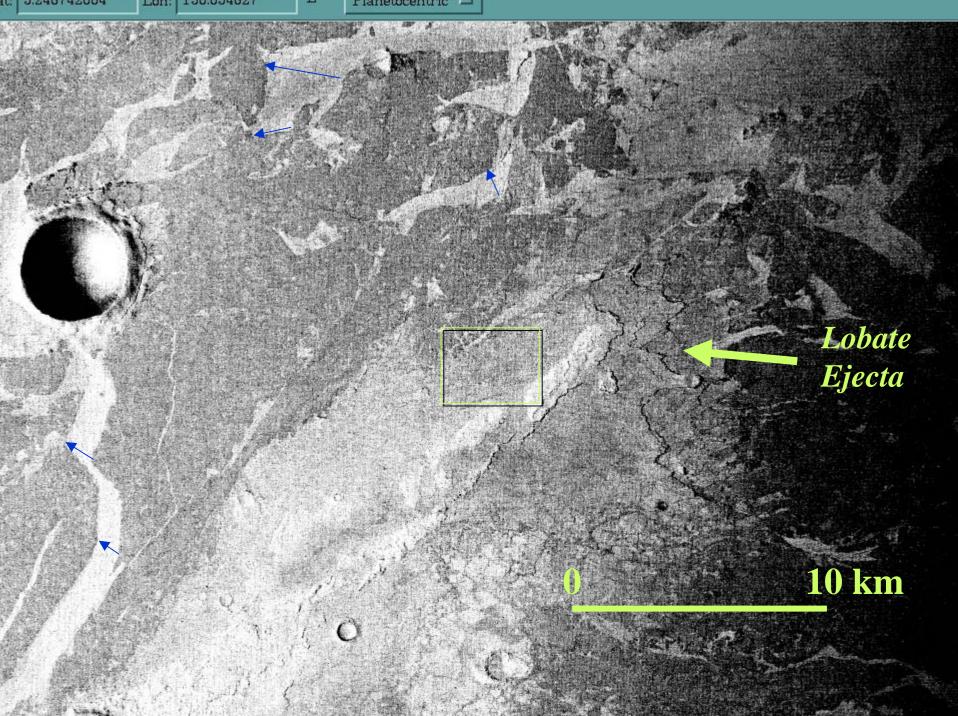
**S** 



#### Elevation (m)



## MOLA altimeter data: 0°.005 slope = EXTREMELY FLAT (Same slope as water surfaces during Tidal events in Bristol Channel)



# Pack-ice

(MOC image)

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Ejecta

1 km

# Alternative ideas: Very fluid lava flows?

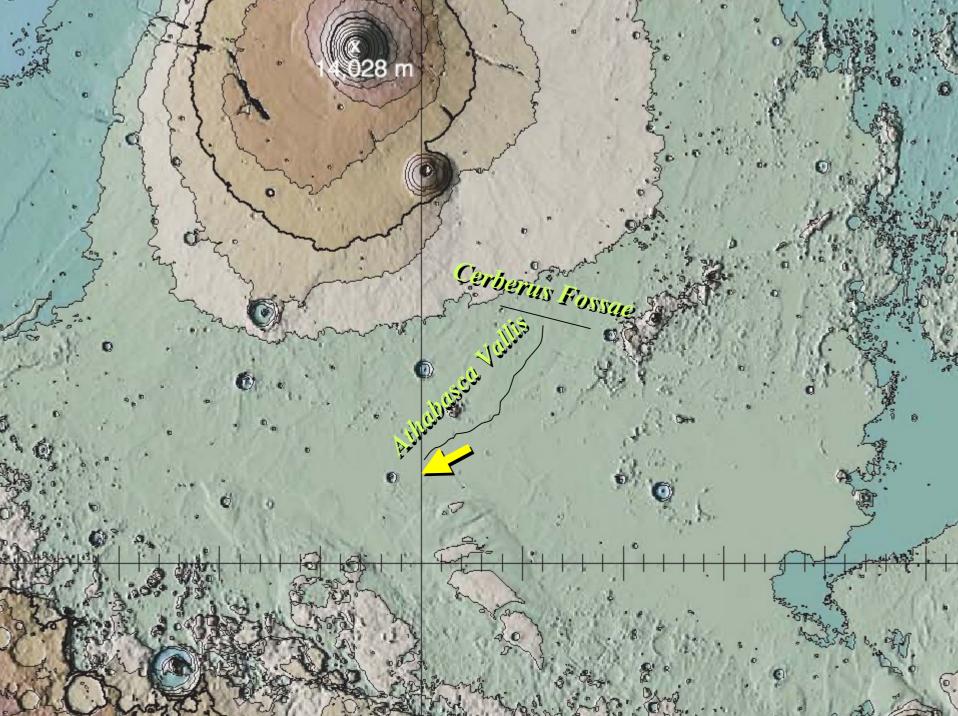


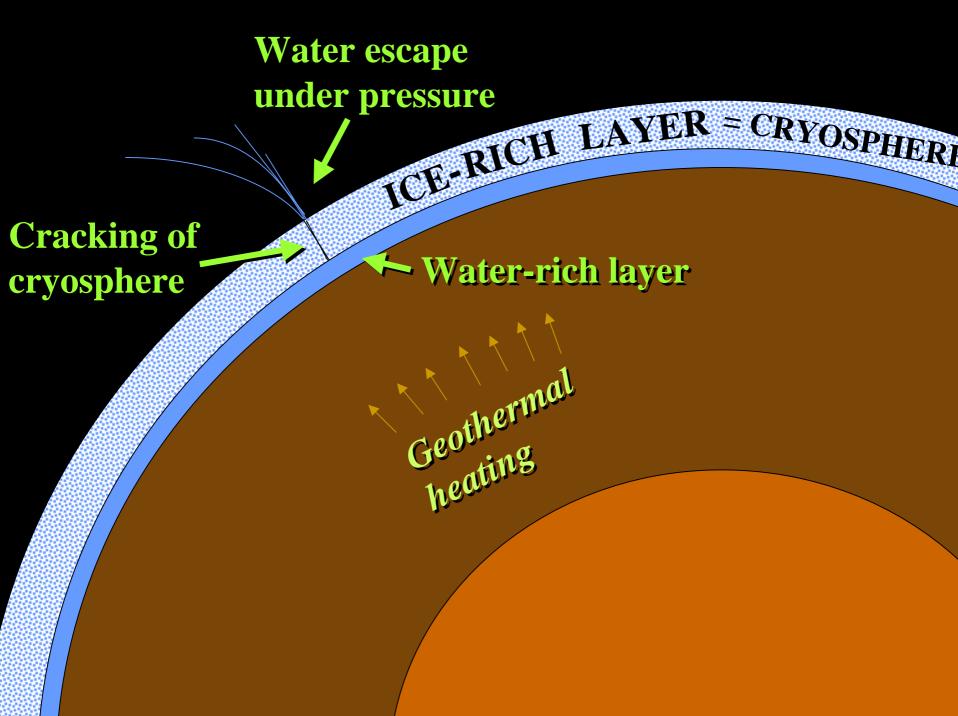
## Ice MUST be dust-covered

0

THE

California 1





- 1. Warm wet subterranean places on Mars have existed throughout Mars' geological history, and life may have developed there.
- 2. Periodically, water eruptions disgorge the contents of these habitats on to the surface, most recently 5 million years ago to form the Elysium frozen sea.
- 3. Elysium is the most likely place to find past or present life on Mars.

