

Composition

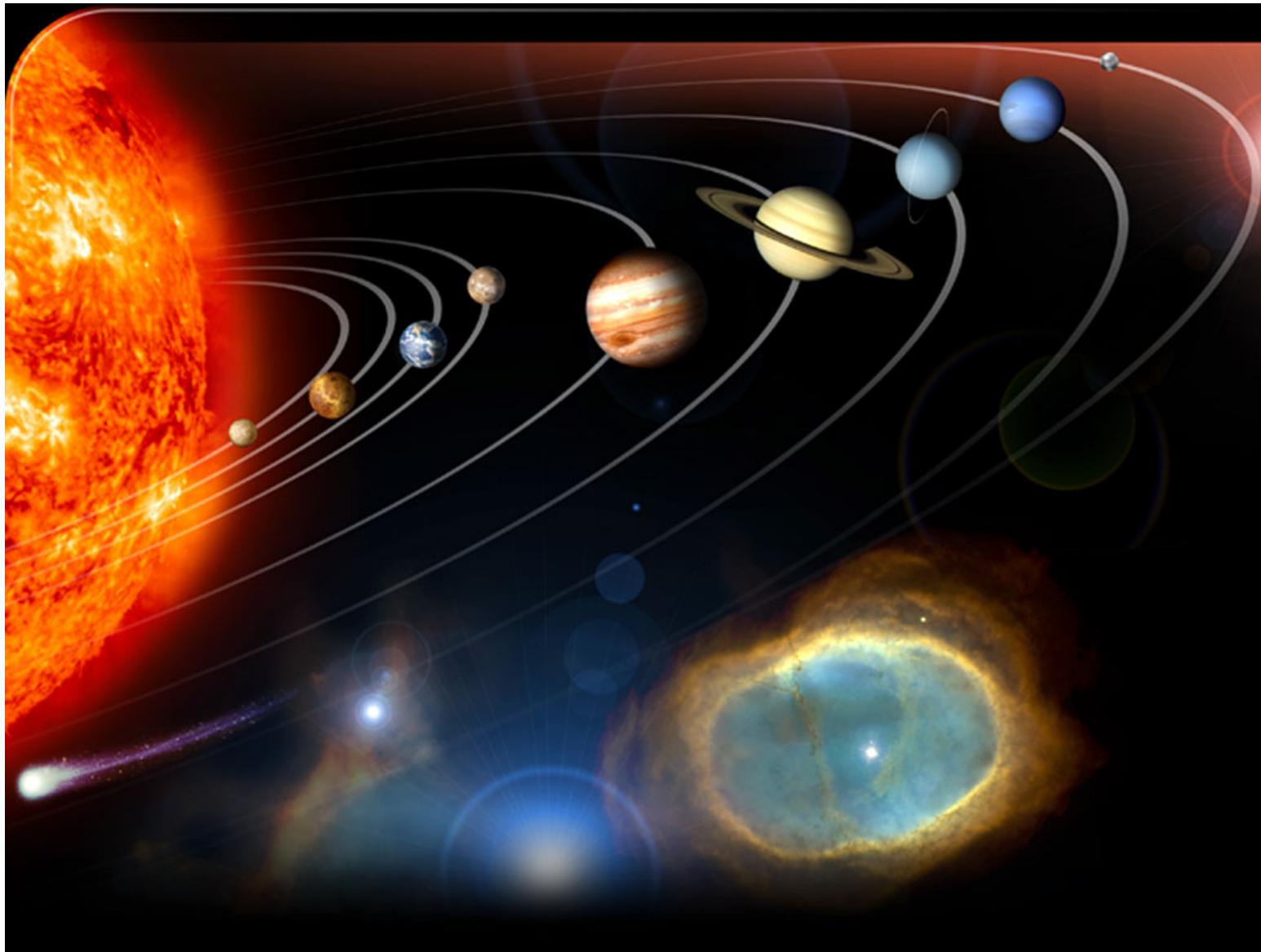
Sushil Atreya

Les Houches, 30 March 2010

map

- *Trace species*: what are they
- *Formation* (origin, source)?
- *Destruction*: how, where, and how fast?
- *Supply*: flux small or big?
- *Meaning*: *Habitability* - tests?
- *Measurements*?

Our Solar System



Mars characteristics from recent observations

- Radius = 3396.2 km (Earth 6398 km)
- Density = 3.933 g/cm³ (Earth 5.4 g/cm³)
- **Semimajor Axis = 1.52 AU**
- Eccentricity = 0.0935
- Inclination = 25.2 degrees
- Orbital period = 686 days (Earth 365 d)
- Rotation period = 24.6229 hrs
- **Surface Gravity = 3.71 m/s²**
- **Internal Magnetic Field: Negligible**
- Average Surface Temperature = 220 K
(range 170-300 K)
- Surface Pressure 6.36 millibars (Earth 1013 mb, or 1 atmosphere)
- Satellites – Phobos and Deimos



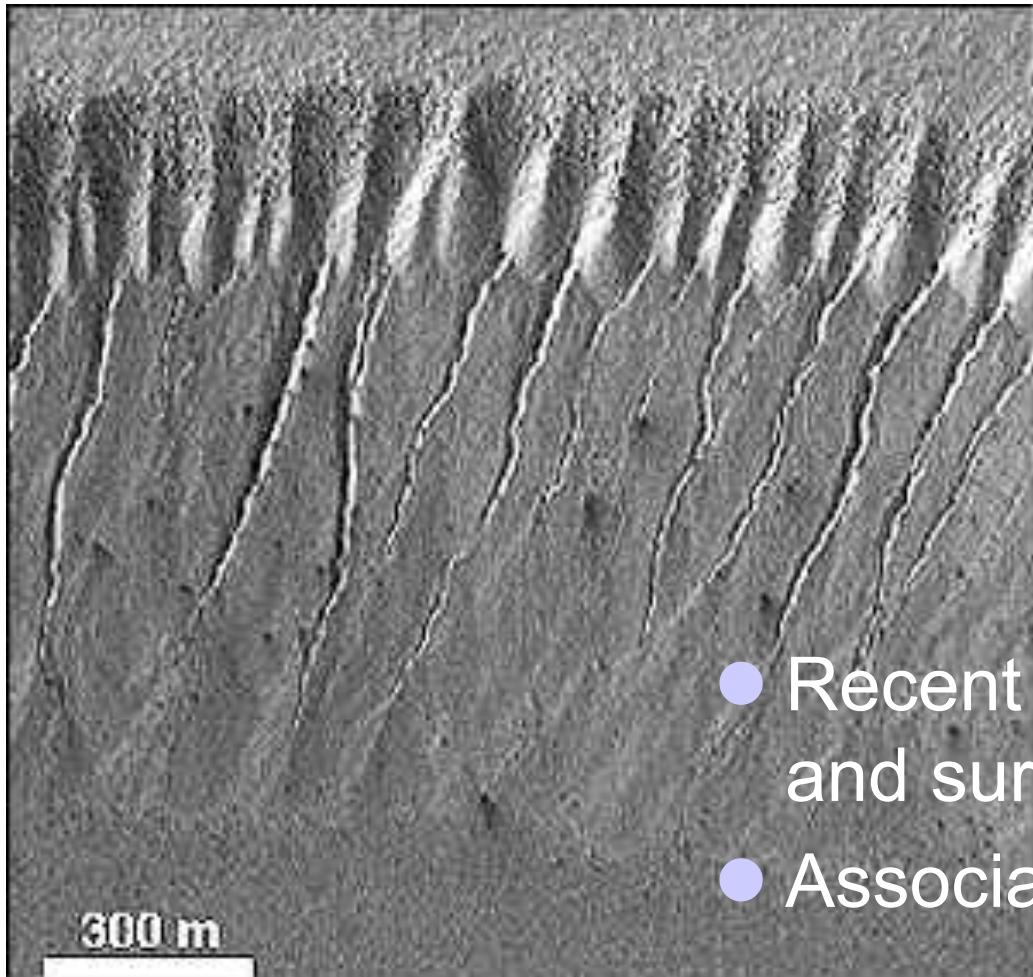
mars atmosphere composition

species	mixing ratio	minor species	<i>upper limit</i>
CO_2	95.32 %	N_2O	1 (-7)
N_2	2.7 %	NO_2	1 (-8)
^{40}Ar	1.6 %	C_2H_2	2 (-9)
O_2	0.13 %	C_2H_4	5 (-7)
CO	0.07 %	C_2H_6	4 (-7)
H_2O	1 - 100 pr- μm	SO_2	1 (-9)
O_3	0.01 – 0.8 ppm	H_2S	1 (-7)
H_2	15 ppm	HCl	2 (-9)
2003 H_2O_2	20-40 ppbv (<i>atmos.</i>)		
2003 CH_4	15 ppbv (global)		surface pressure 6-10 mb

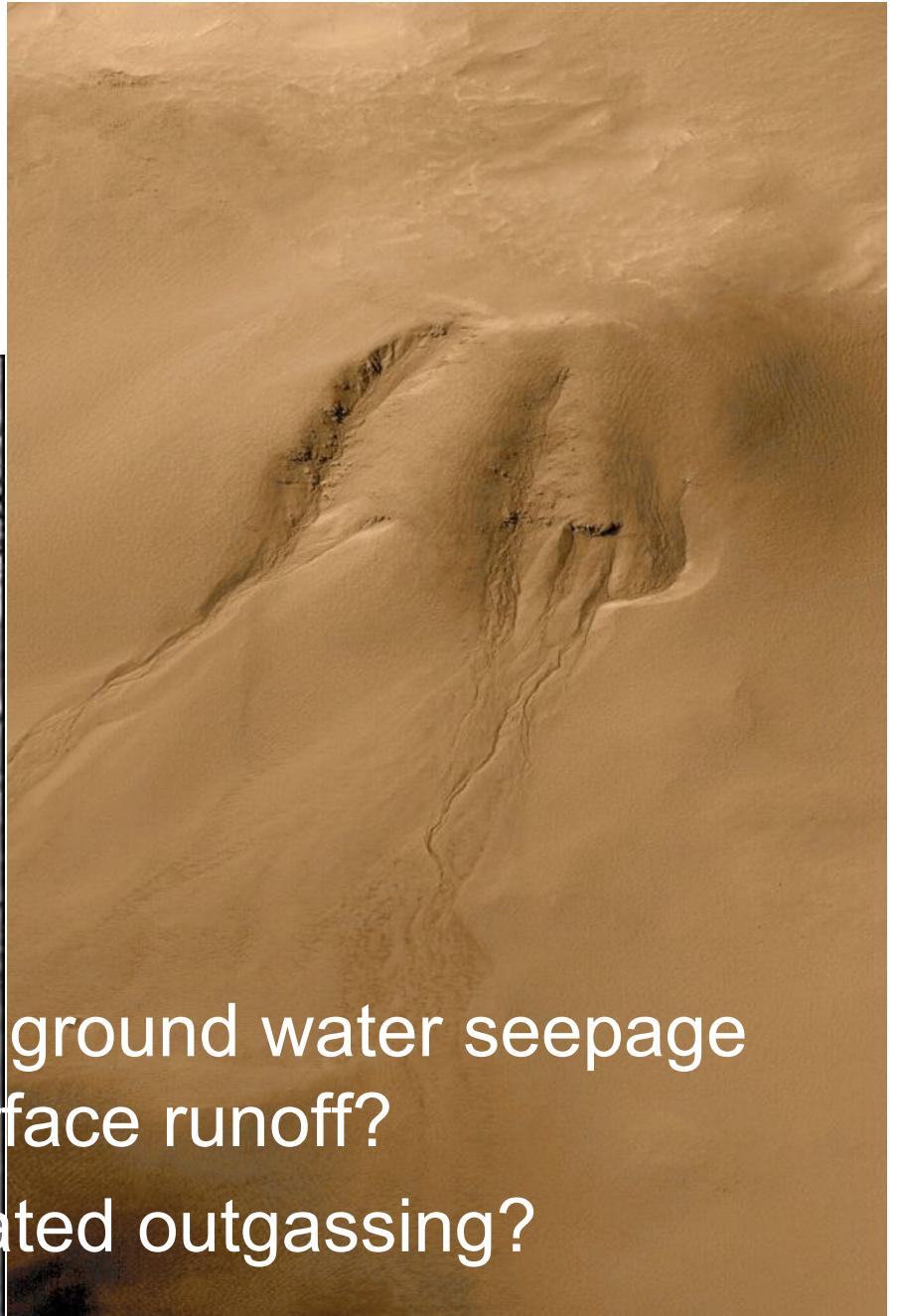
(Earth CFC's ~1 ppbv)

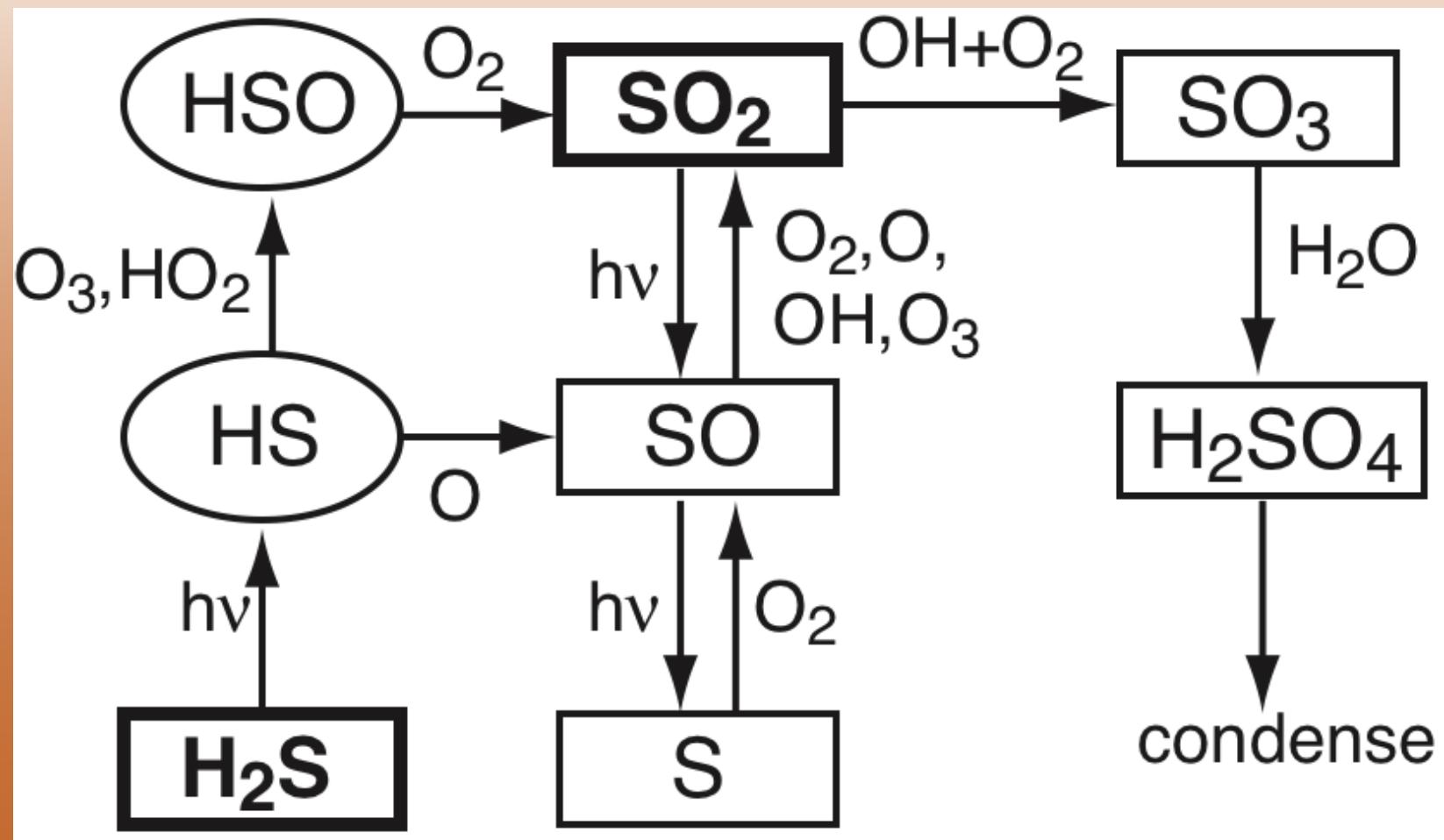
Possible outgassing on Mars?

(Mars Global Surveyor Images, 2000)



- Recent ground water seepage and surface runoff?
- Associated outgassing?

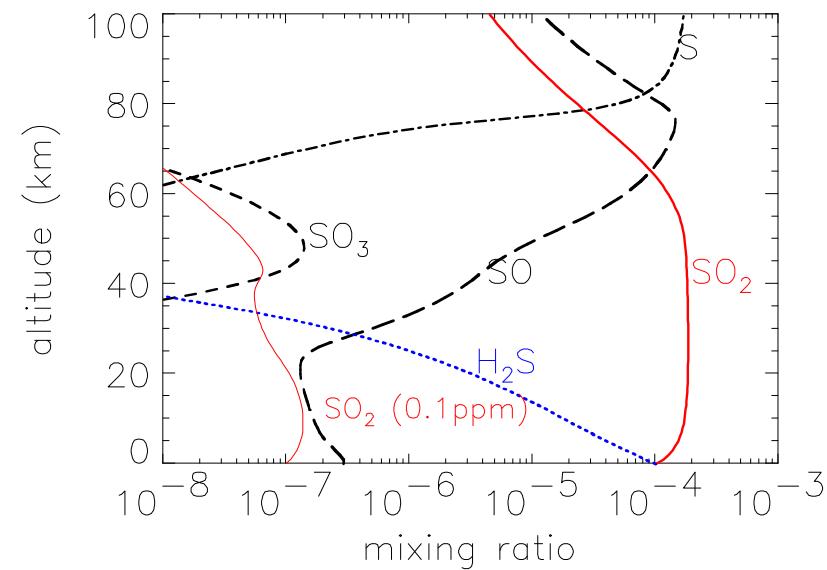
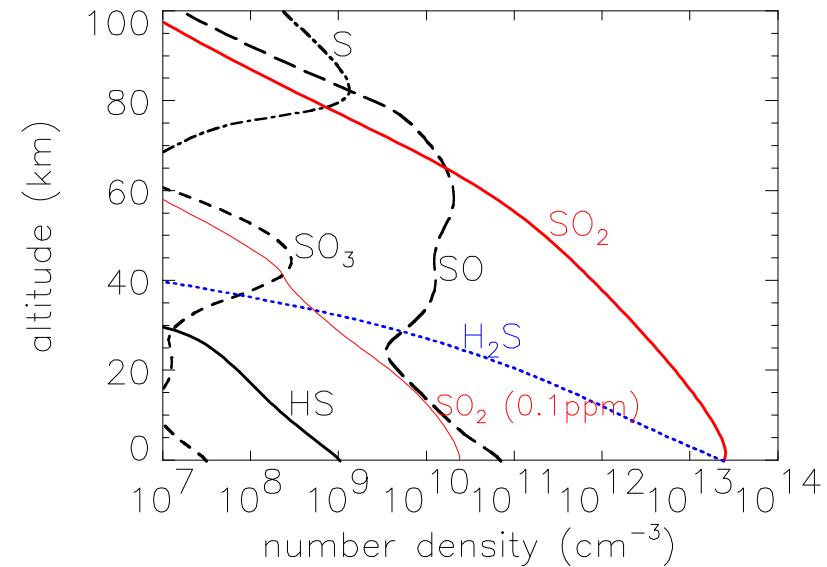
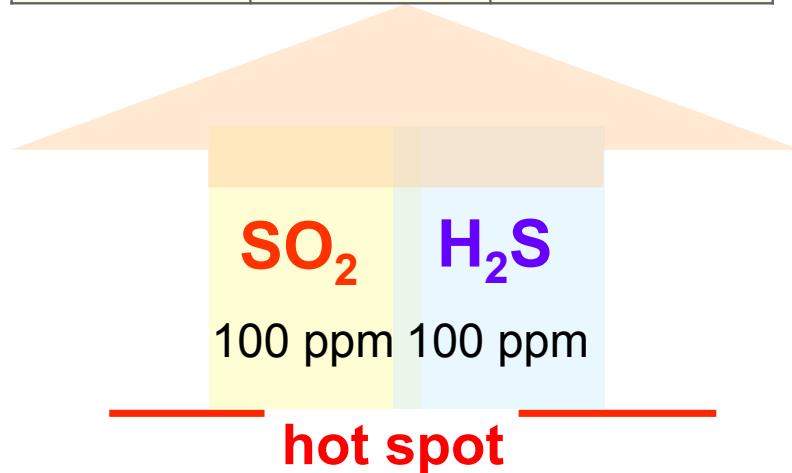




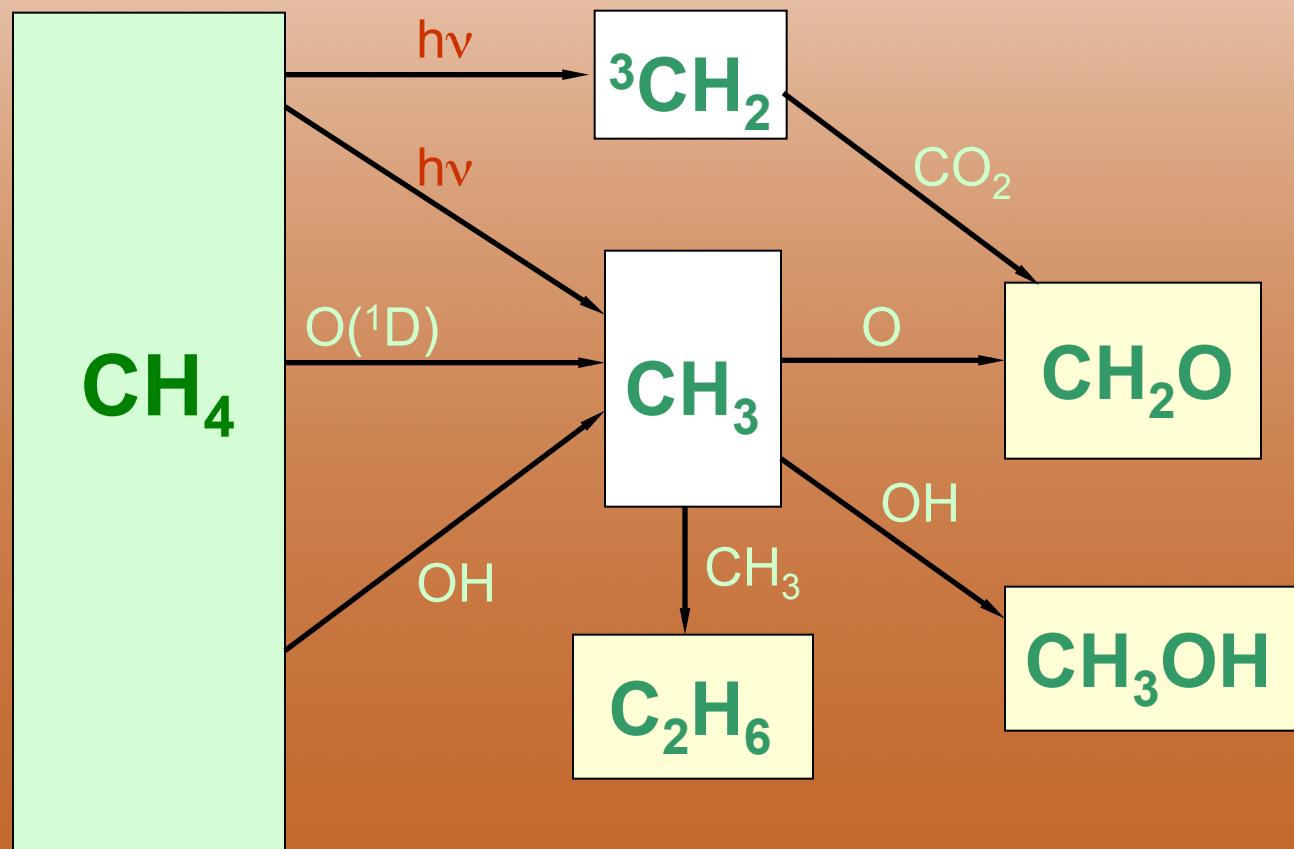
Hotspot sulfur chemistry

Sulfur Species

major species	mixing ratio (10 km)	column (cm^{-2})
SO_2	1.7 (-4)	1.8 (19)
H_2S	1.8 (-5)	7.4 (17)
SO	1.7 (-7)	7.8 (16)



Methane chemistry on Mars

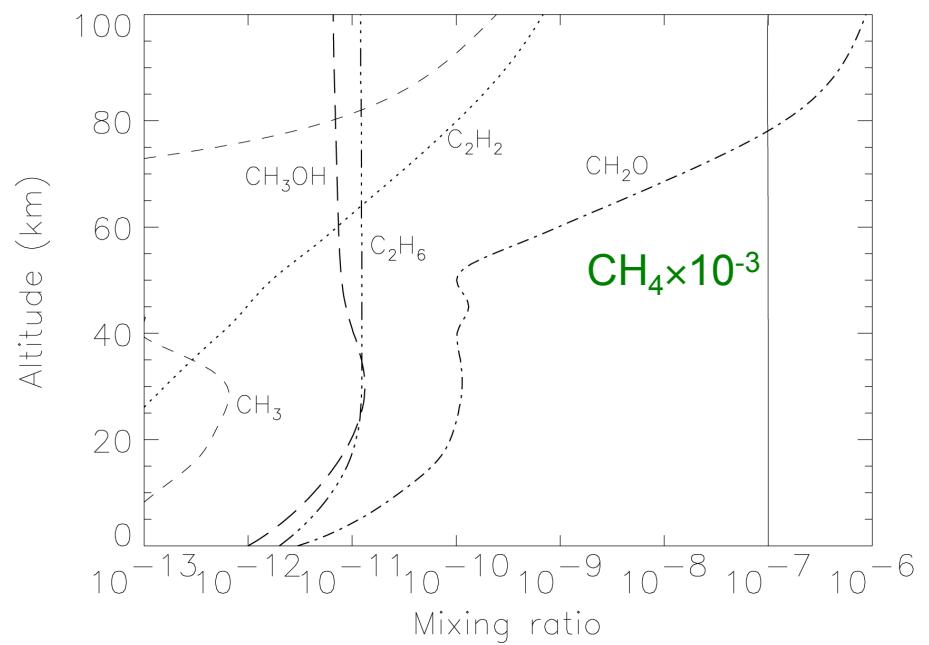
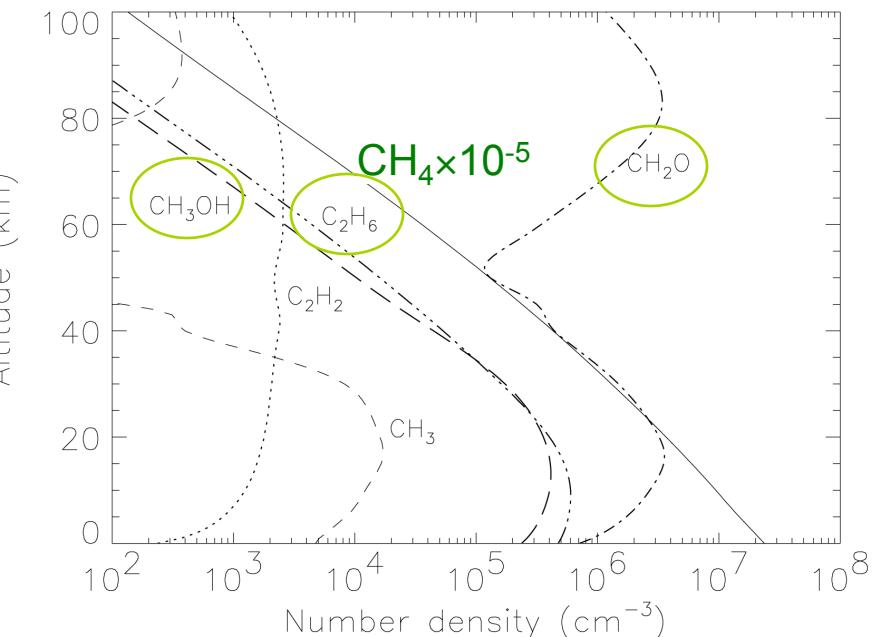
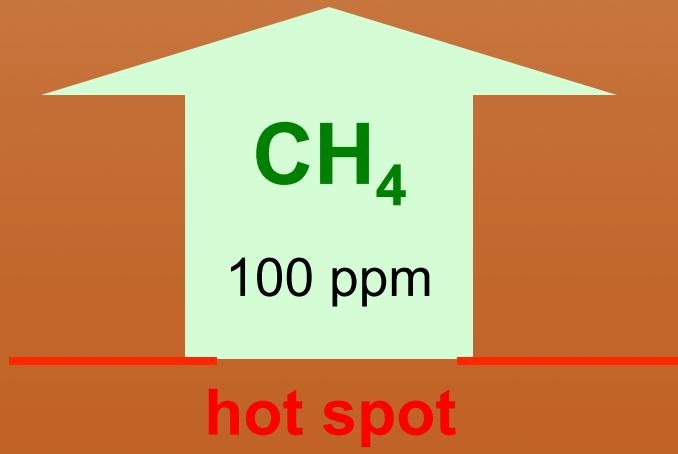


Flux from Mars surface

Photochemical Lifetime of methane 300-600 years

Hydrocarbons

major species	mixing ratio (10 km)	column (cm ⁻²)
CH₄	1.0 (-4)	1.0 (19)
CH ₂ O	2.8 (-11)	1.6 (13)
CH ₃ OH	4.3 (-12)	8.2 (11)
C ₂ H ₆	6.4 (-12)	9.6 (11)

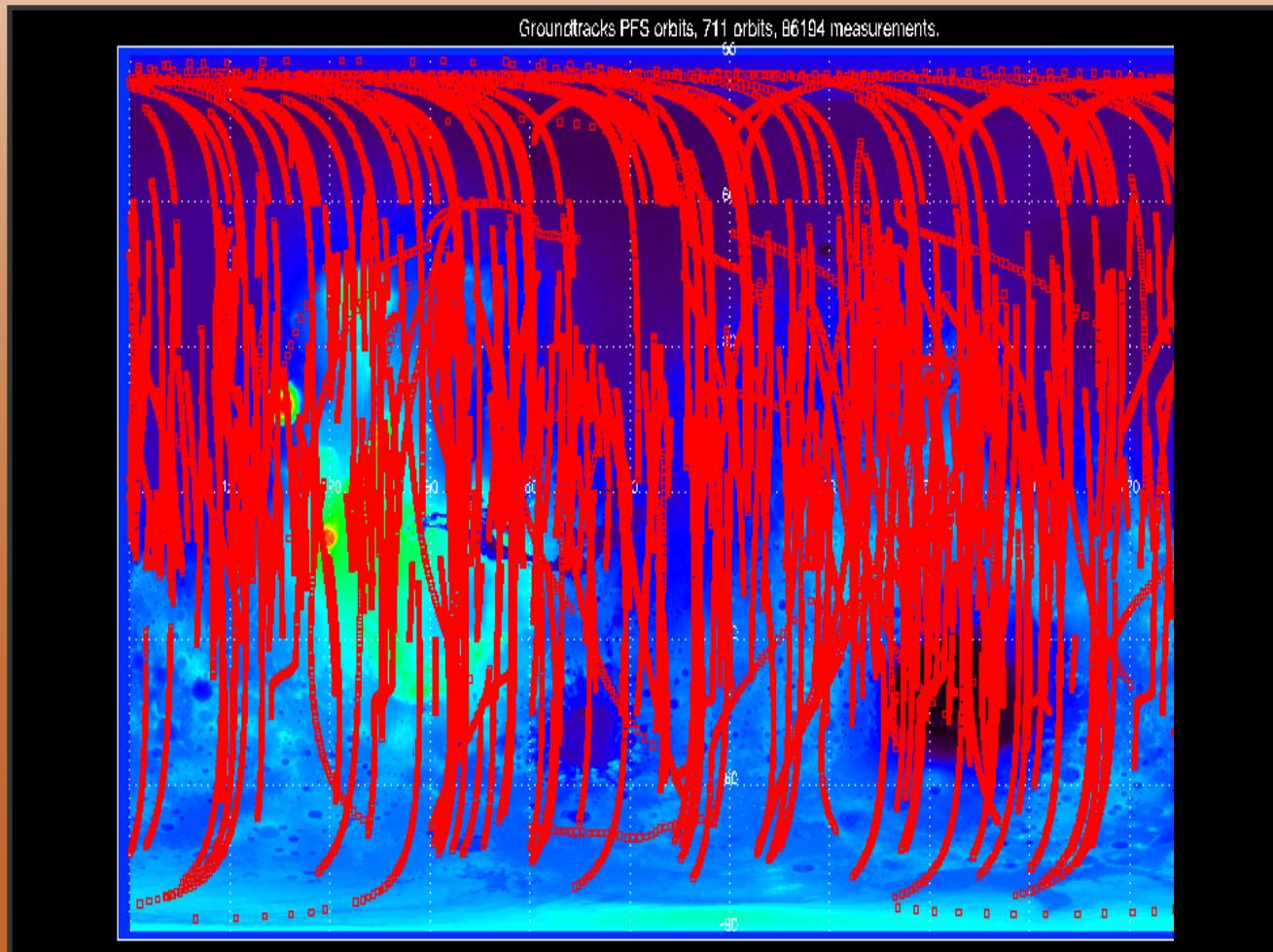


Possible Detectable Species - Model

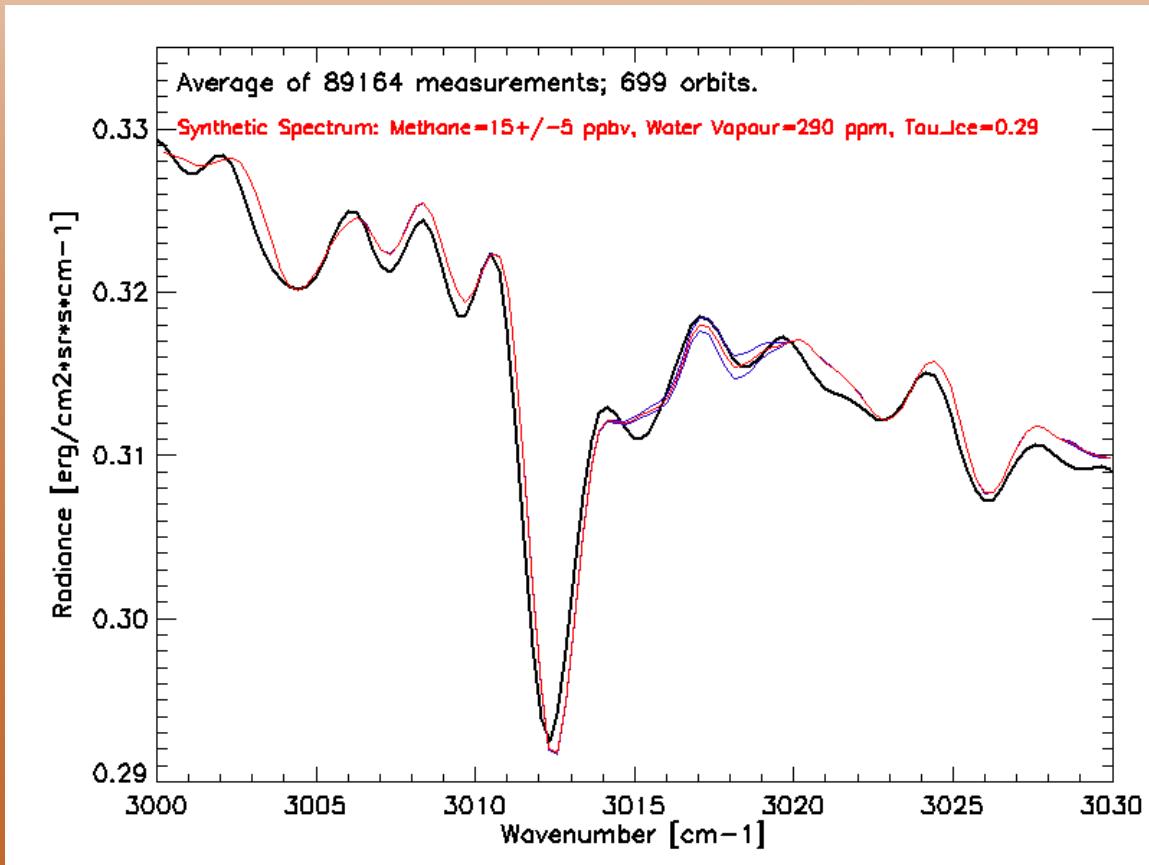
(100 ppm outgassing of SO₂, H₂S, CH₄)

Species	Mixing ratio (10 km)	Column abundance (cm ⁻²)	Photochemical lifetime
SO ₂	1.7 (-4)	1.8 (19)	40 days
H ₂ S	1.8 (-5)	7.4 (17)	9 days
SO	1.7 (-7)	7.8 (16)	4.6 hours
CH ₄	1.0 (-4)	1.0 (19)	670 years
CH ₂ O	2.8 (-11)	1.6 (13)	7.5 hours
CH ₃ OH	4.3 (-12)	8.2 (11)	74 days
C ₂ H ₆	6.4 (-12)	9.6 (11)	25 days

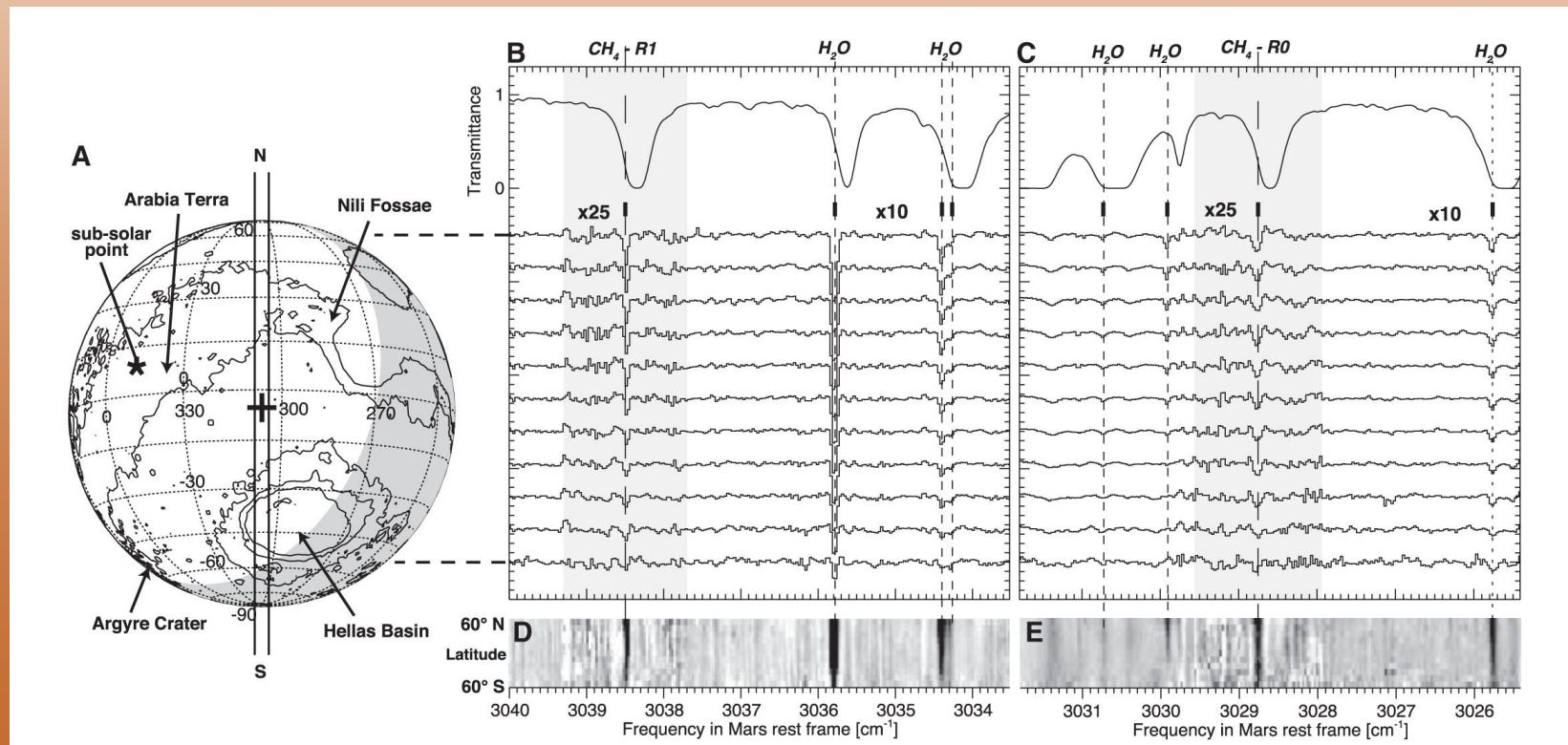
Mars Express PFS orbit ground-tracks



global average methane: 14 ± 5 ppbv



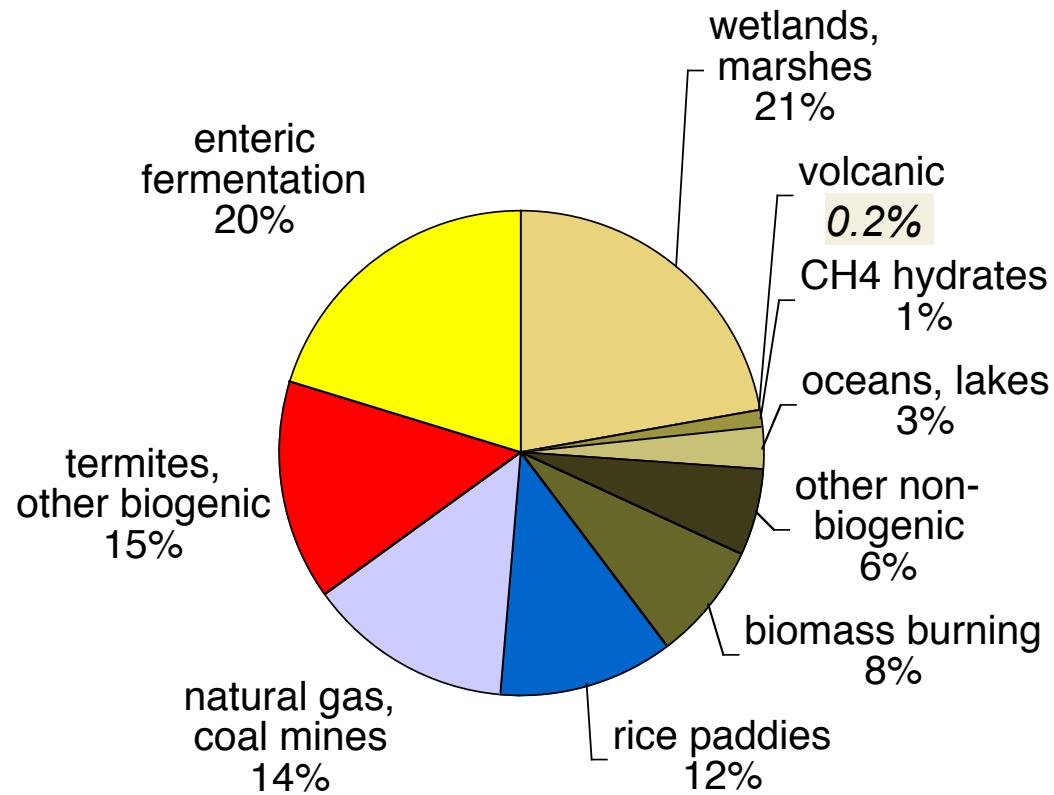
Mars Express PFS (Formisano et al., 2004; Geminale et al., 2007)



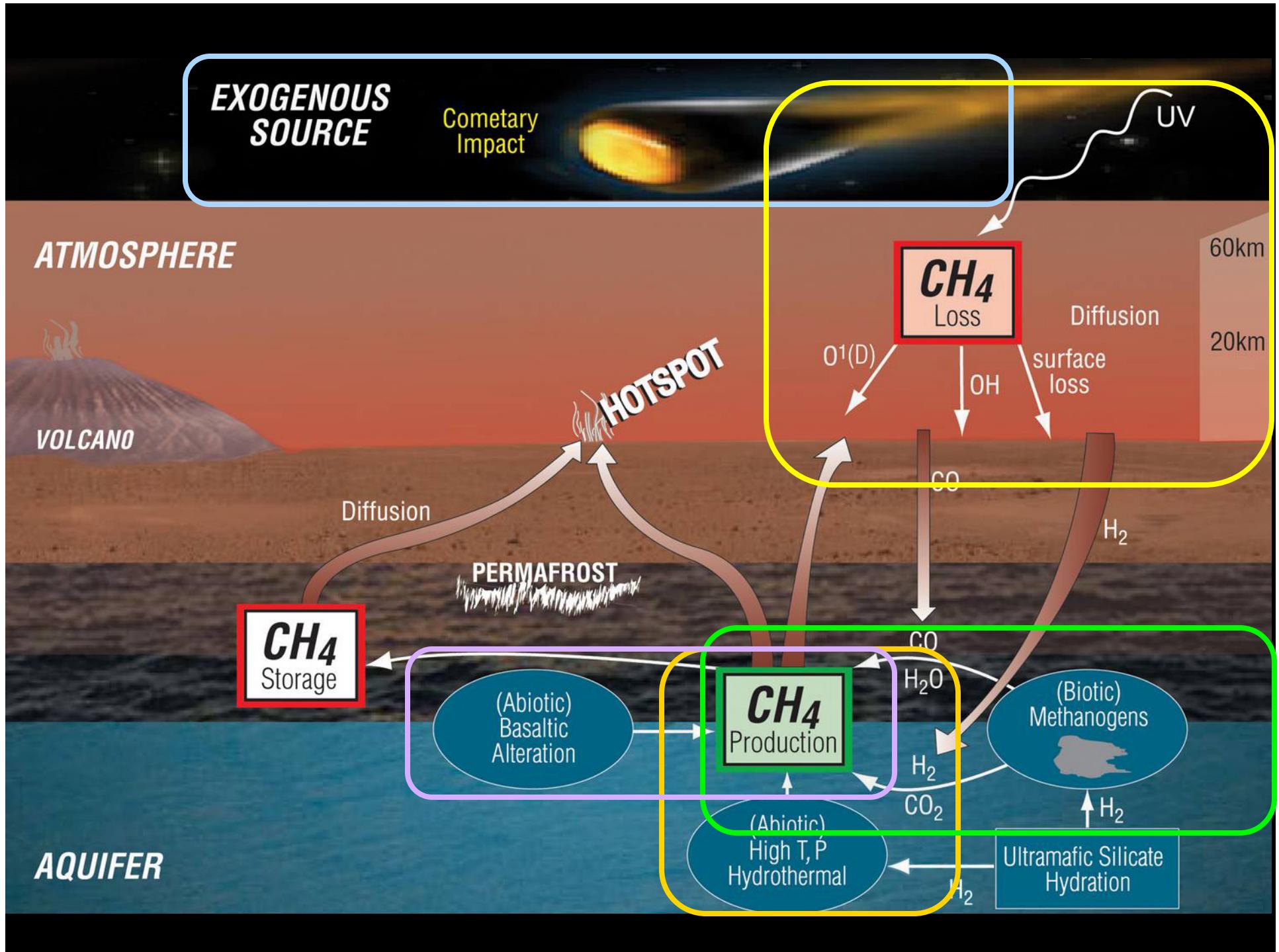
IRTF & Keck 2 (Mumma et al., 2009)

mars methane: significant findings

- MEX/PFS (>500,000 spectra; 3018 cm⁻¹; Formisano et al., 2003-present)
 - average 14 ± 5 ppbv; non uniform; *footprint ≥ 1000 km*
 - max: 35-50 ppbv; *hotspots* - Arabia Terra, Elysium Plenum, N-pole
 - min: <10 ppbv; hints of seasonal variability
- CSHELL/IRTF, NIRSPEC/Keck-2 (R0, R1; Mumma et al., 2003-2006)
 - average 10-20 ppbv; non uniform, *footprint ~ 800 km*
 - max: ~ 35 ppbv; east of Arabia Terra, Nili Fossae, SE Syrtis Major
 - *in plumes* when present; transient
- CFHT/ FTS (3018 cm⁻¹; Krasnopolsky et al., 2004)
 - global average only: 10 ppb (<30 ppbv from later observations)

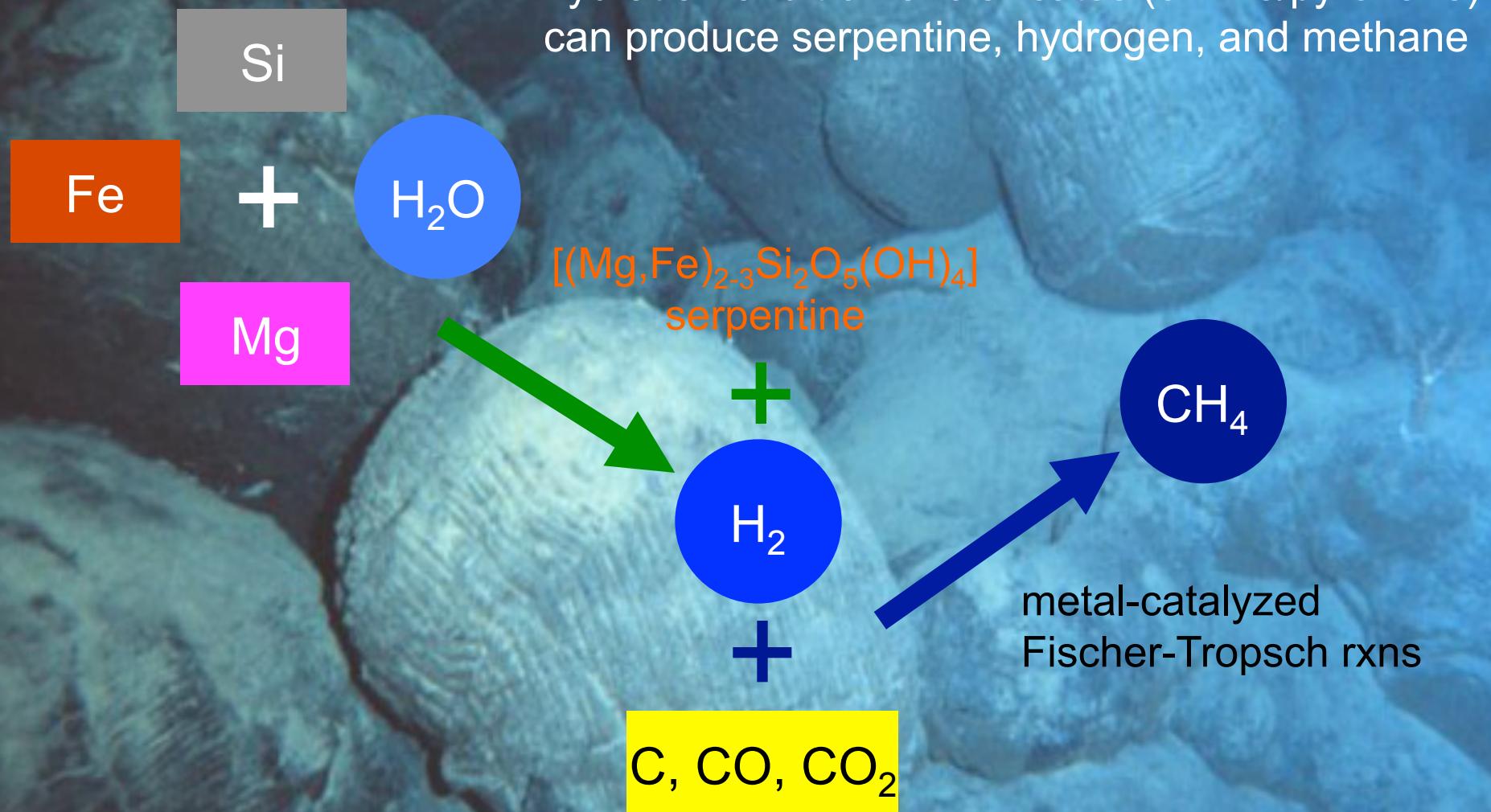


methane sources on Earth (1750 ppbv)

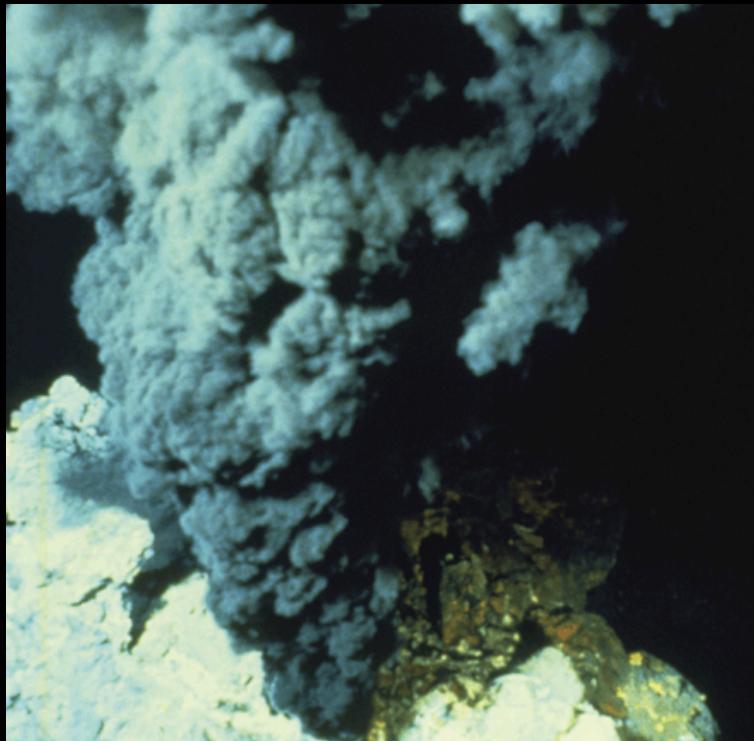


How: serpentinization

liberate hydrogen, mix carbon, make methane (abiotic)



Where (on Earth): Black Smoker hydrothermal vents abiotic CH₄ by *high* temperature serpentinization



Mid Atlantic Ridge

Juan De Fuca Ridge
depth 2222 m
exit temp 342 C
pressure 200 bar
chimney ht. 10 m



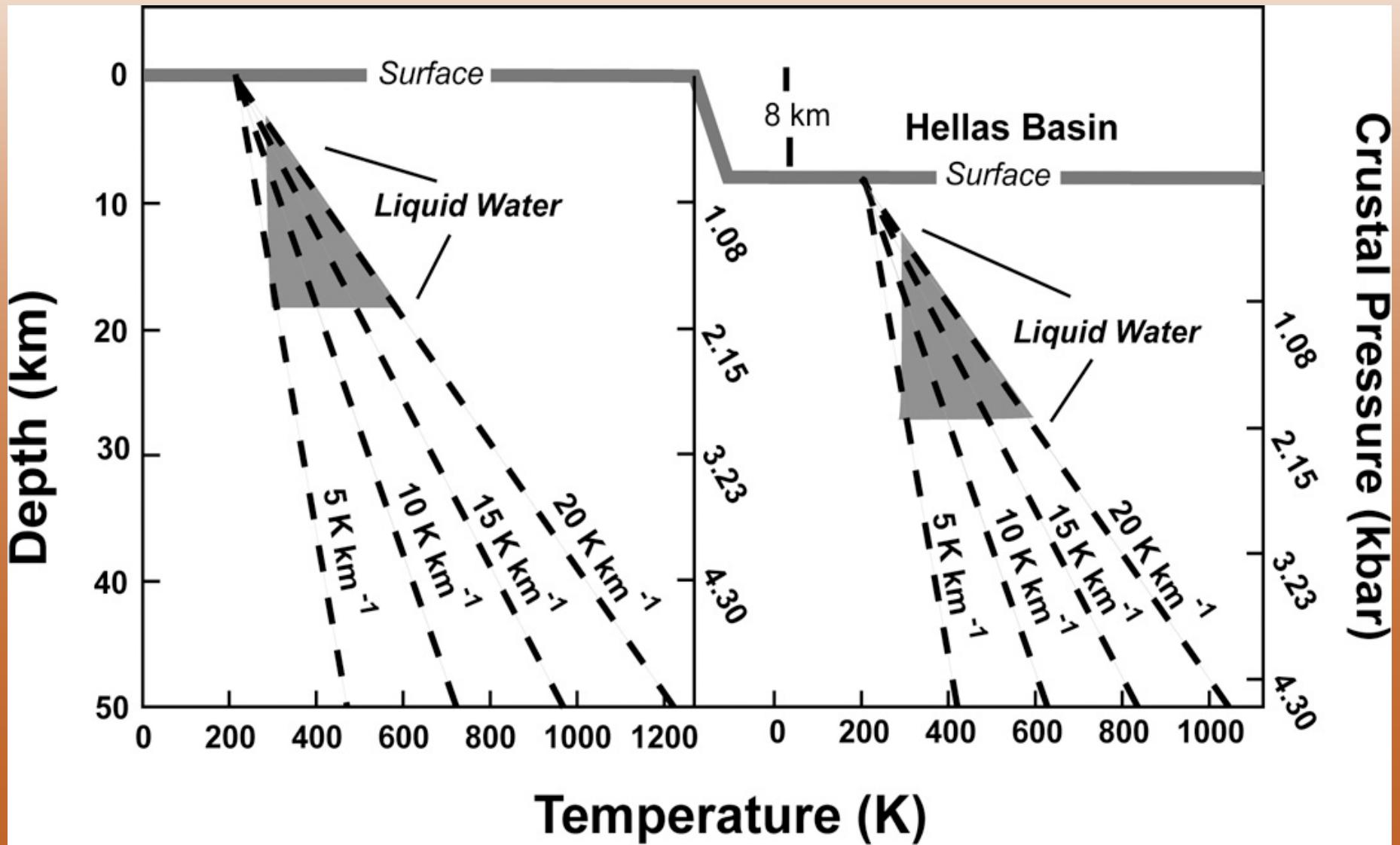
300-500 C, sulfur-rich, highly acidic (~lemon juice)

Where (on Earth): Lost City abiotic CH₄ by *low* temperature serpentinization



- 15 km from Mid-Atlantic Ridge
- 30-90 C
- highly alkaline (~ammonia, milk of magnesia)
- 20m high carbonate towers
- little sulfur minerals

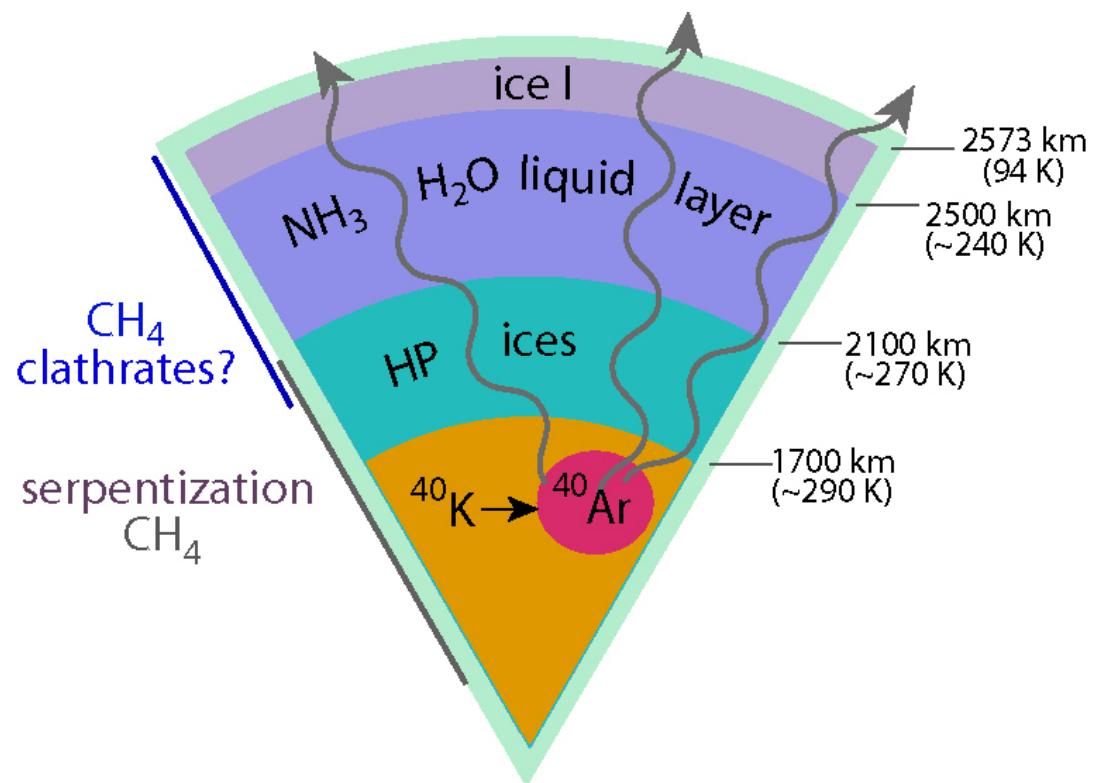
mars geotherms: liquid water stable in 2-25 km, 20-300 C



(Oze, Sharma, 2005)

Outgassing at Titan

- ^{40}Ar detection
 $(^{40}\text{K} \rightarrow ^{40}\text{Ar}$
1.3 Gyr half-life)
 - cryovolcanism
 - CO (VIMS)
- **CH₄ outgassing**



(interior structure, after Grasset et al.)

methane origin, now or in the past

- **Geologic** (or hydrogeochemical)
i.e. water-rock chemistry, involving serpentinization and metal-catalyzed Fischer-Tropsch type reactions, or
- **Biogenic** (microbial methanogenesis)
i.e. by enzymatic reduction of CO_2 , or fermentation of organic matter by microbes in the metabolic process

*Liquid water essential to both
(subsurface aquifers, if CH_4 produced recently)*

source strength: shorter life, bigger flux

- **Photochemical loss:** long life, small source flux
 - 300-600 yrs, ~200 ton/yr → uniform methane
- **Reactive surface:** faster loss, i.e. ~one year lifetime, requiring much bigger source flux
- **Surface loss due to oxidants**
 - **peroxide** (very reactive), i.e.
resulting hydroxyl, hydroperoxy (HO_2), superoxides
 - **perchlorates** (very stable, but reactive with free Cl)

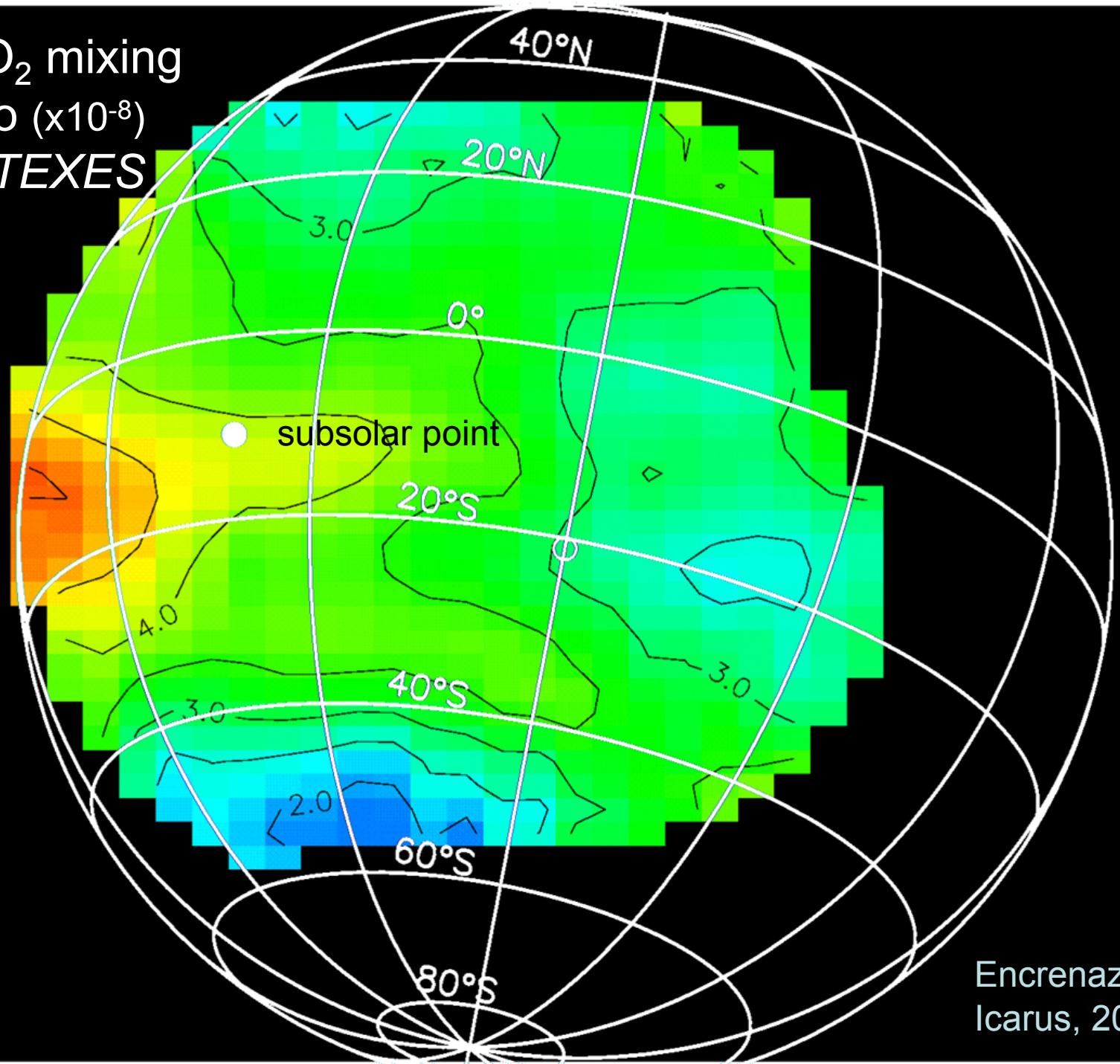
organics and oxidants don't mix

- ▣ **Viking (1976) Life Sciences Experiments (LSE)**
 - ▣ GCMS: no organics found in Martian soil --
 - ▣ Indigenous; or externally delivered by meteorites, IDPs, comets
 - ▣ GEX: nutrients & H₂O added, O₂ released rapidly:
surface oxidant ; required amount **20-250 ppm H₂O₂**
(Oyama, 1977; Hunten, 1979; Huguenin, 1982)
- ▣ In 2003, **20-40 ppb H₂O₂** in **atmosphere** detected, but is **too low** by a factor of 100-1000 compared to LSE requirement

[H₂O₂ detection: submm 362.156 GH,
Clancy et al.; IR (8.04 – 8.13 μm) Encrenaz, et al.]

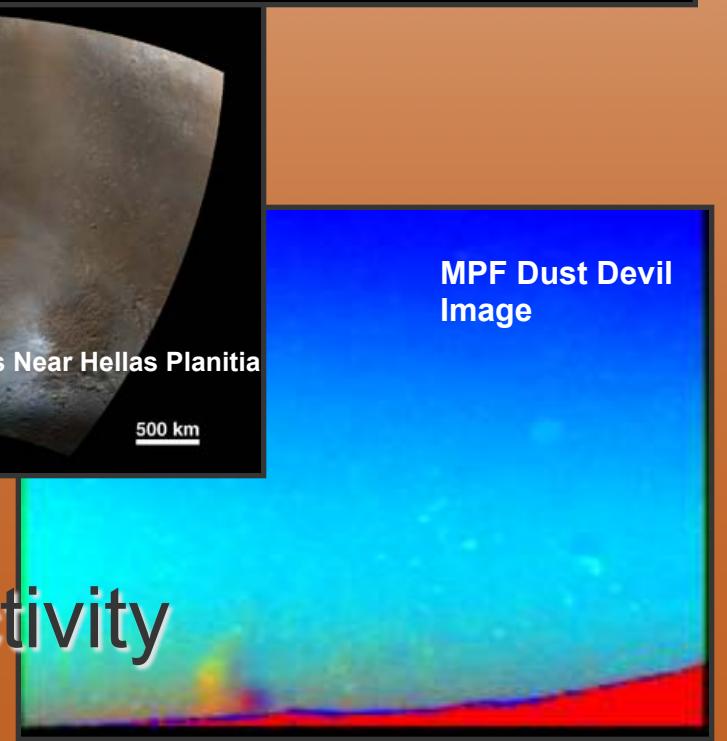
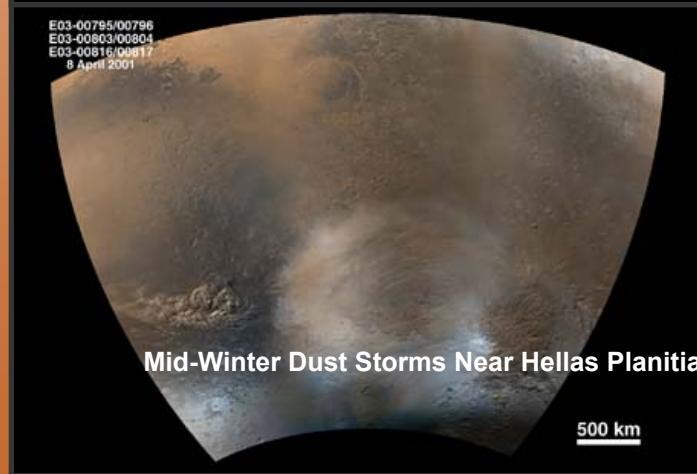
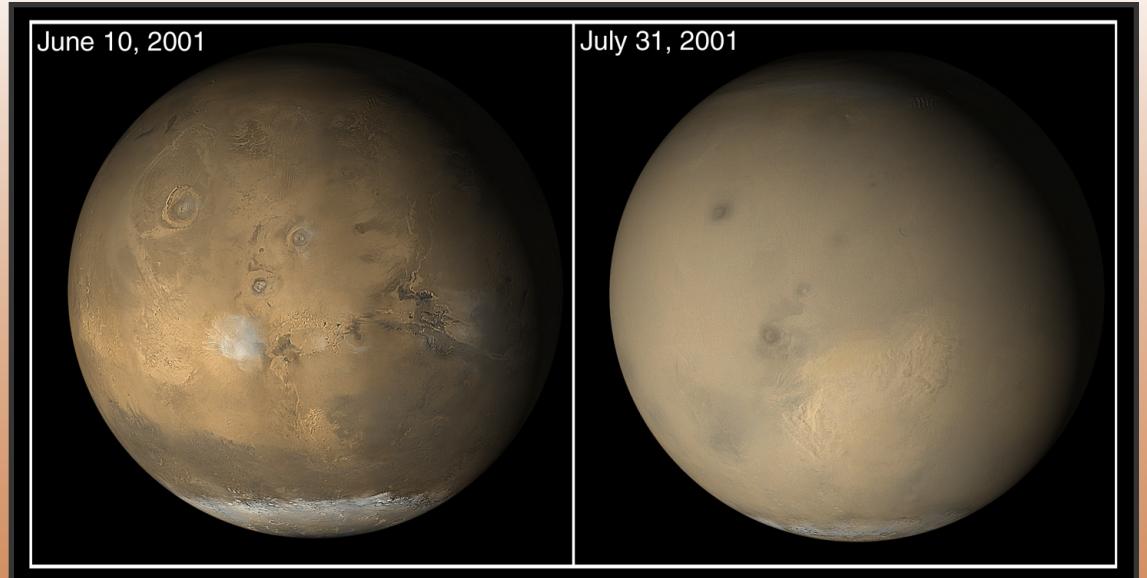
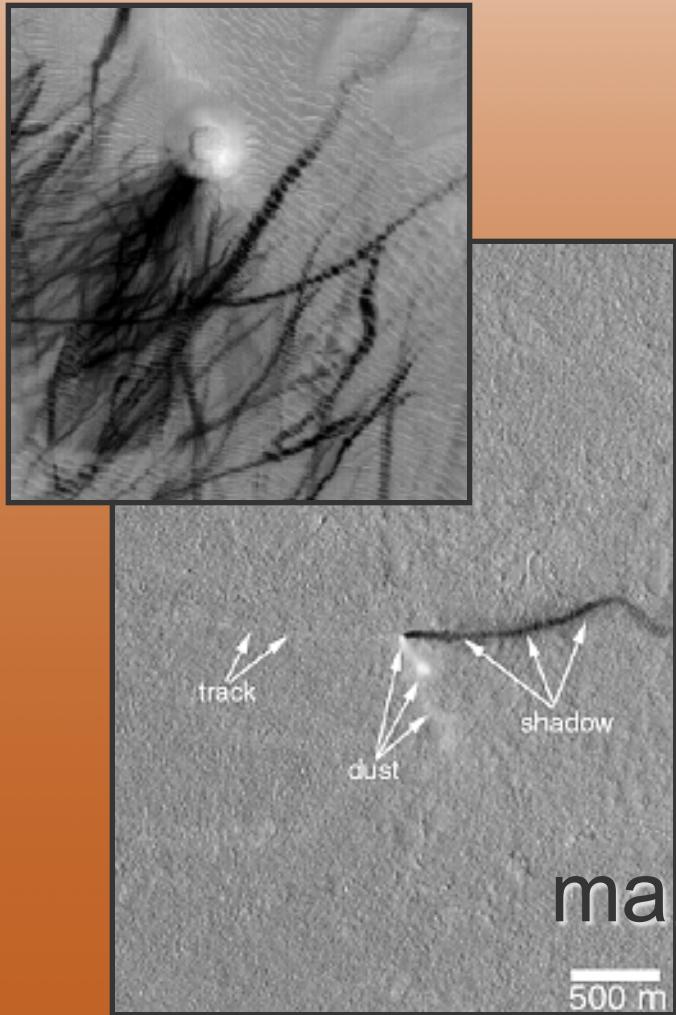


H_2O_2 mixing
ratio ($\times 10^{-8}$)
by *TEXES*



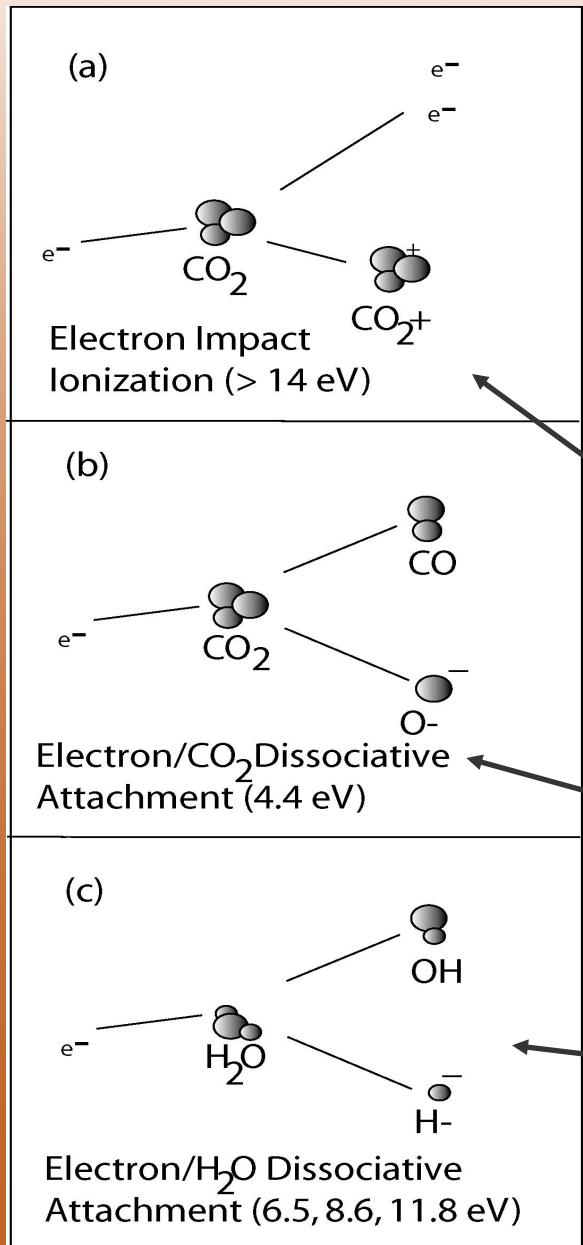
Encrenaz et al.,
Icarus, 2004

aeolian triboelectricity



martian dust activity

enhanced species production in dust storm collisional plasmas



- three new product sets results from the energetic electron population [Delory et al. '06; Farrell et al, '07]
- production of $\text{CO}_2^+ (> 14 \text{ eV})$
$$\text{e} + \text{CO}_2 \rightarrow 2\text{e} + \text{CO}_2^+$$
- production of $\text{CO}/\text{O}^- (\sim 4.4 \text{ eV})$
$$\text{e} + \text{CO}_2 \rightarrow \text{CO} + \text{O}^-$$
- production of $\text{OH}/\text{H}^- (\sim 6.5 \text{ eV})$
$$\text{e} + \text{H}_2\text{O} \rightarrow \text{OH} + \text{H}^-$$

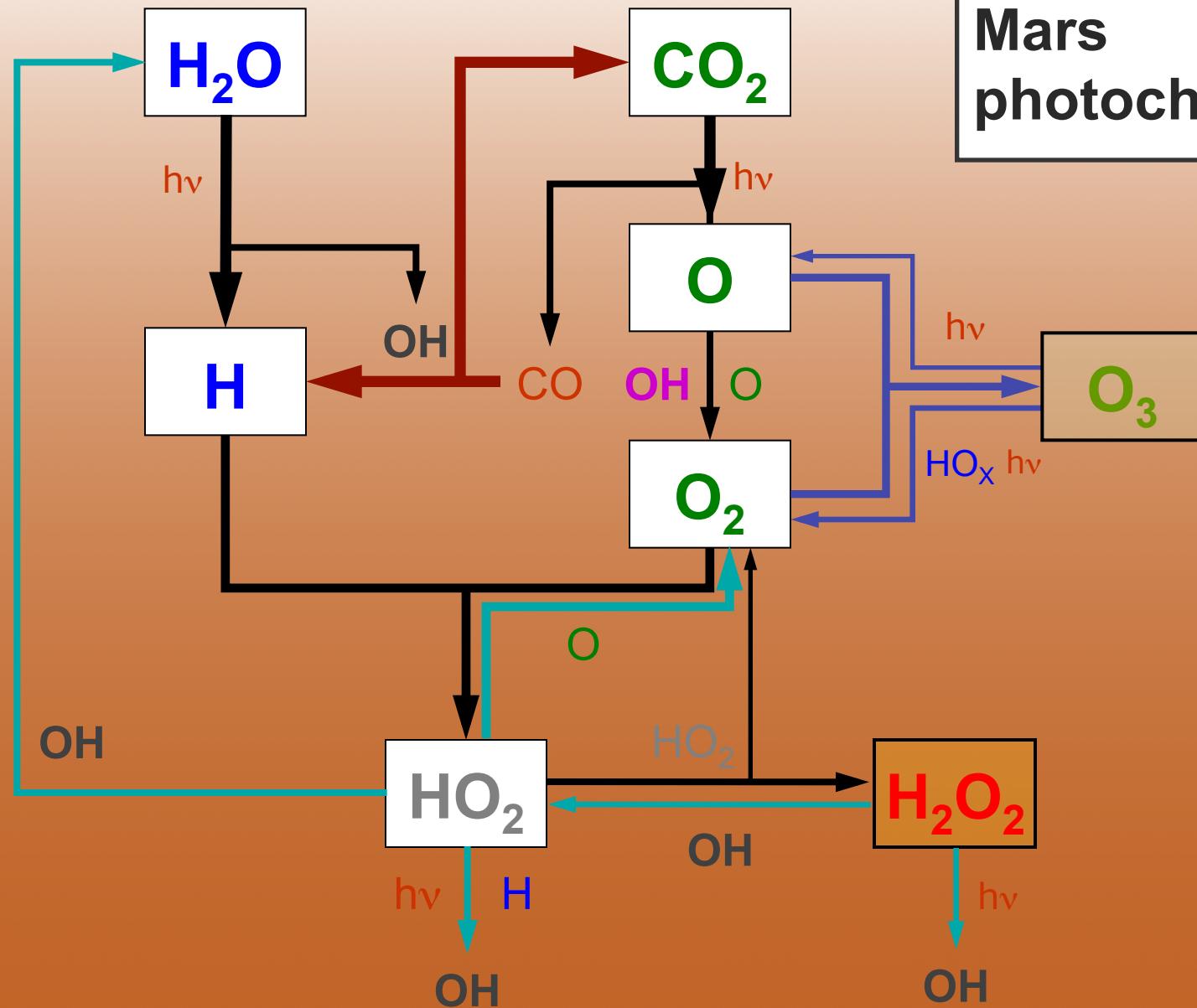
storm electric fields (triboelectricity) change near-surface chemistry

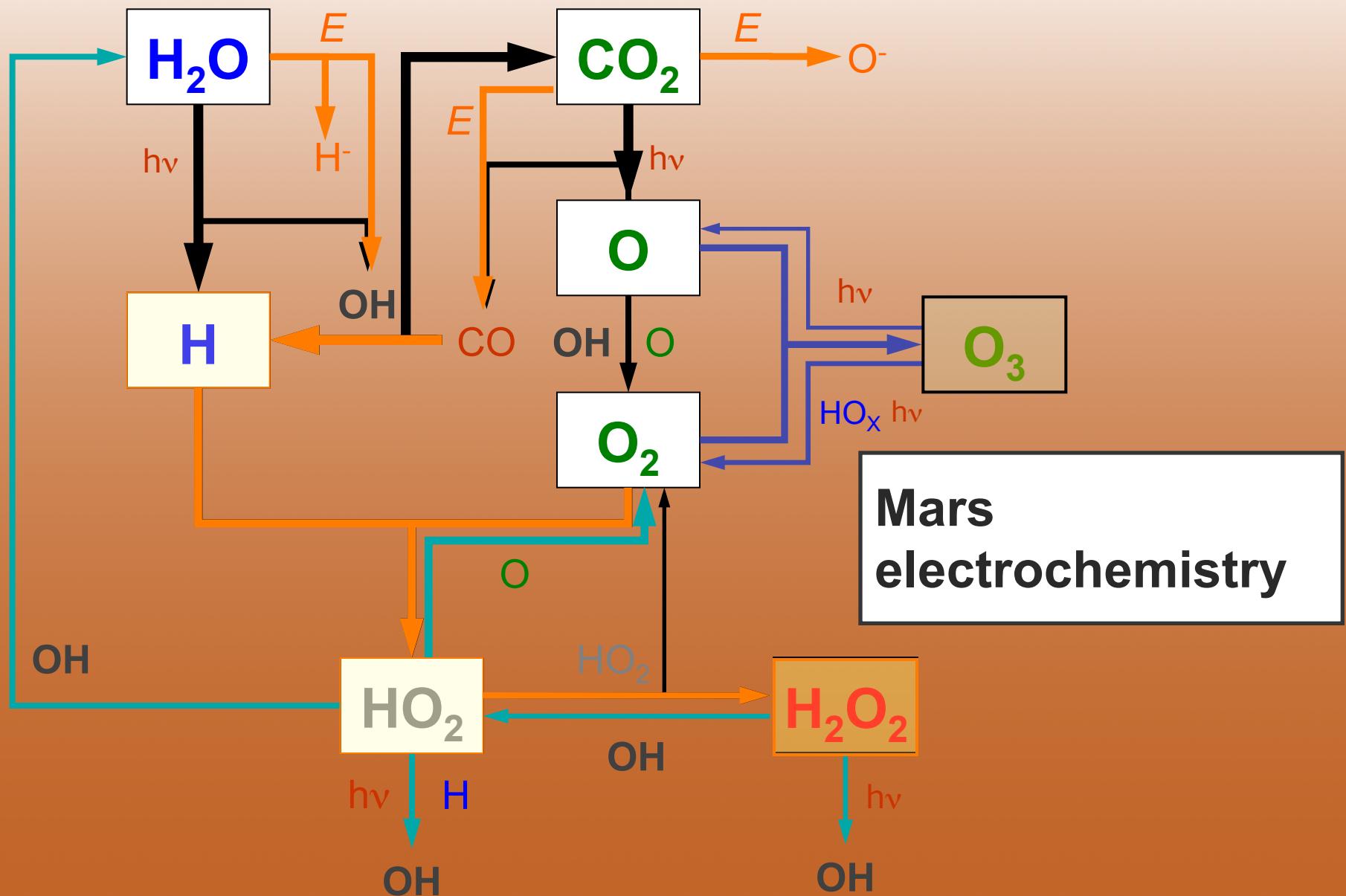
- 5-25 kV/m* electrostatic fields generated in Martian aeolian processes, i.e. in convective dust storms and dust devils
- CO/O-, OH/H- ion pairs* produced→
- H₂O₂ enhanced substantially by electrochemistry over photochemical source**

*Delory, et al. 2006, Farrell et al. 2007, Jackson 2009

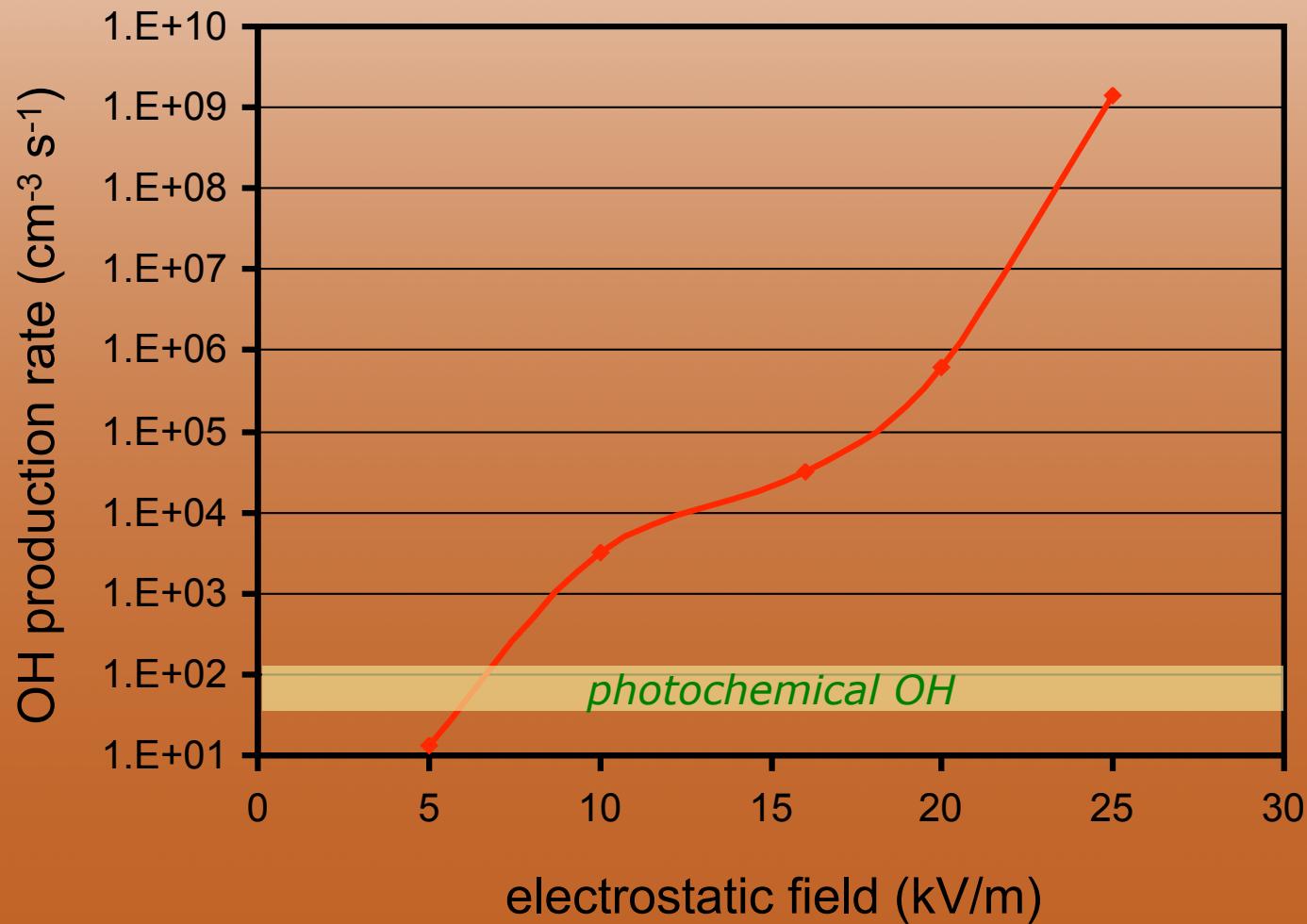
**Atreya, et al. 2006

Mars photochemistry

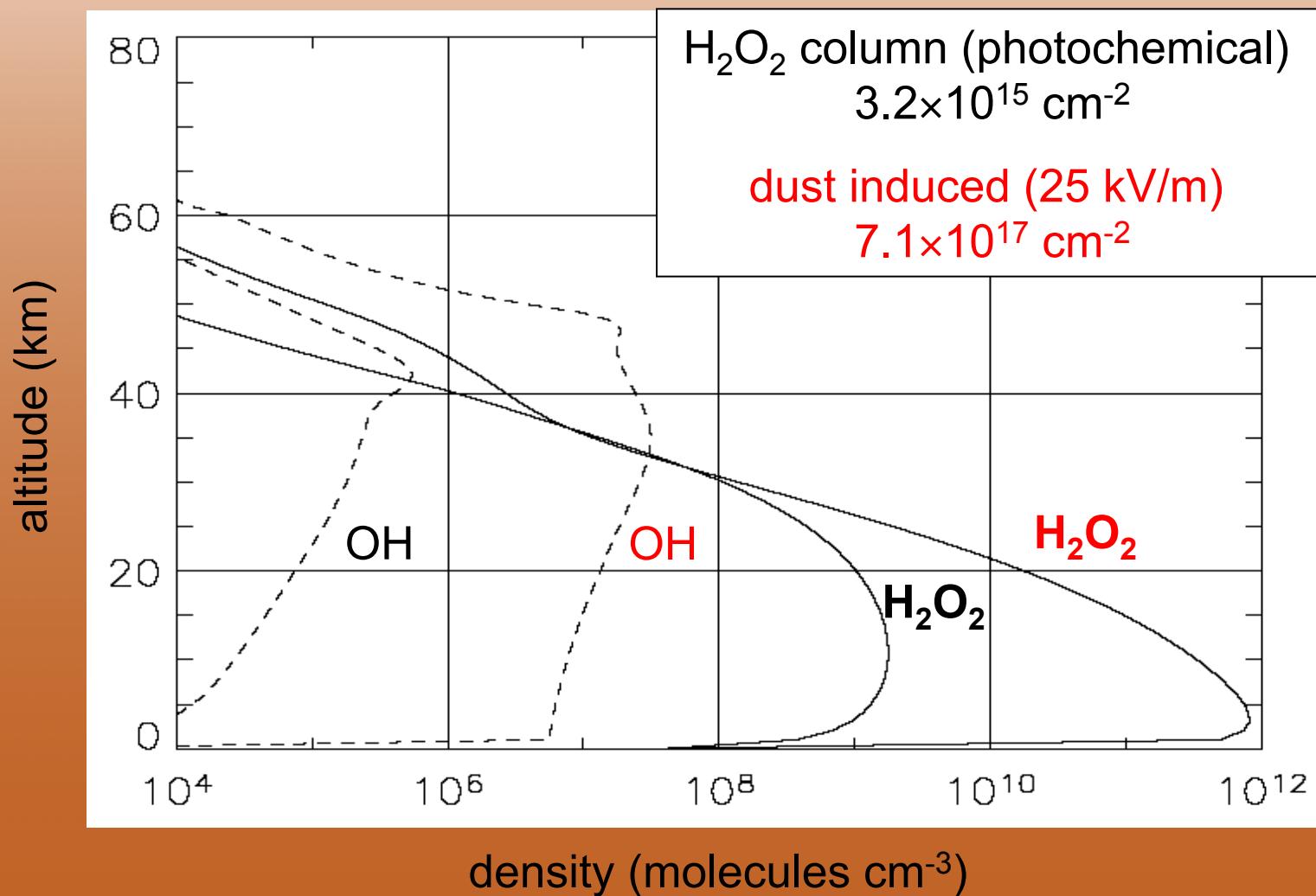




dust-induced OH production



model profiles of OH, H₂O₂



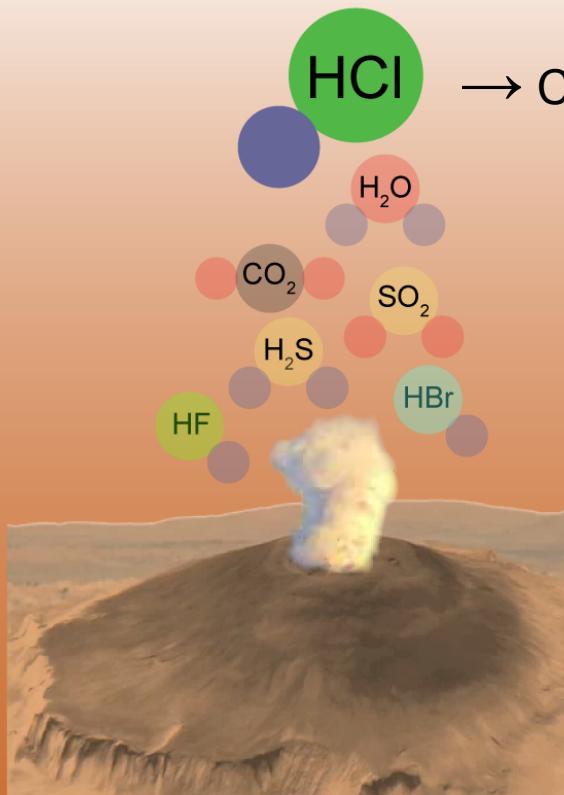
storm electric fields make peroxide; peroxide & products destroy surface organics and methane

- 10,000x enhancement of H_2O_2 over photochemical source ** →
- H_2O_2 “snow” → H_2O_2 into regolith
- H_2O_2 in regolith lives long, up to millions of years (M. Bullock, 1996), compared to *less than one day* in atmosphere
- *superoxides/ hydroxyl/ hydroperoxy (HO_2) generated in soil*
- **rapid destruction of surface organics (directly), and of**
- **methane, by OH , HO_2 and superoxides from H_2O_2**
- **excess CO recycles the CO_2 lost to triboelectricity**

Radiolysis of ice also produces H_2O_2 , as on Europa, Ganymede & Callisto

**Atreya, et al. 2006

perchlorate (ClO_4^-) on Mars and Earth

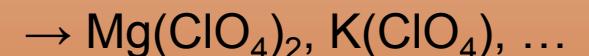
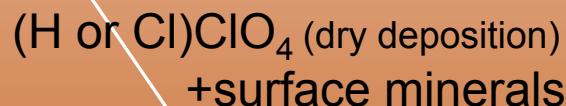
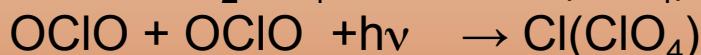


volcano

where
how much
dry period

Mars

north polar region
0.3 – 0.6 wt%
?



brine

Earth

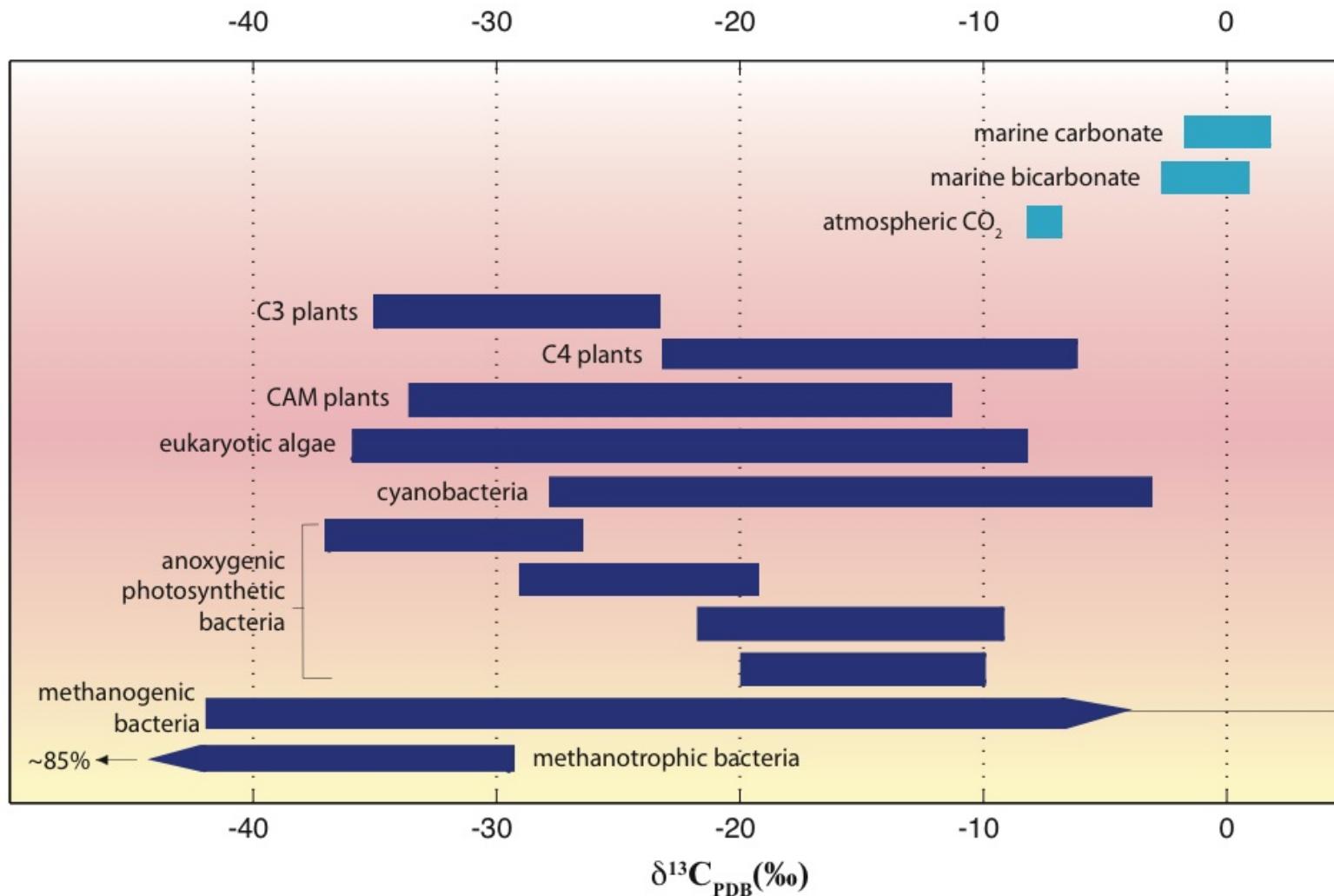
Atacama, Bolivia, Texas
0.03 – 0.6 wt% (Atacama)
5-15 Myr (Atacama)

**Life as we know it produces methane,
but**

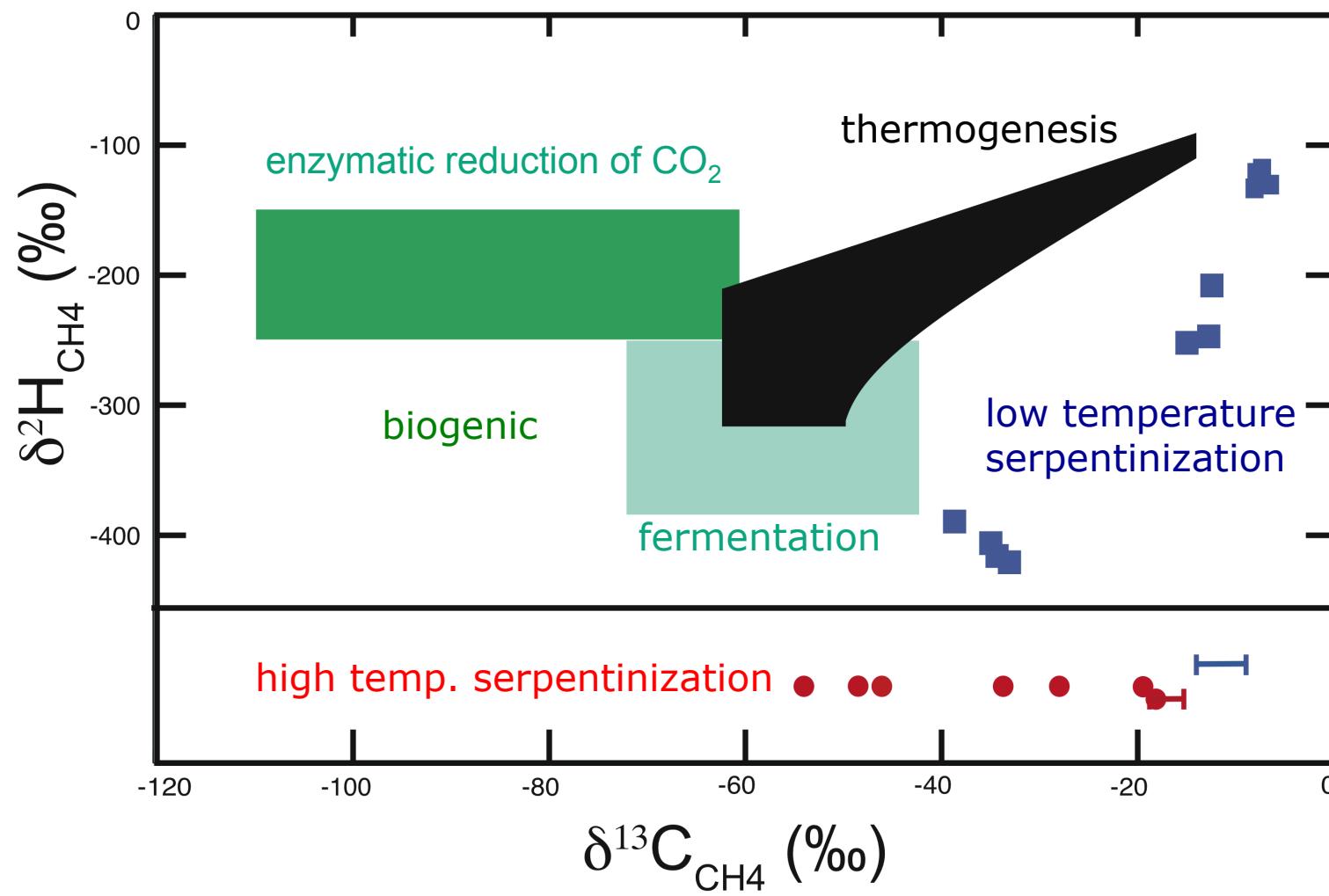
***presence of methane does not by
itself imply presence of life***

Organic (living) vs. Inorganic ^{13}C carbon isotope

V-PDB $^{12}\text{C}/^{13}\text{C}=89.4$

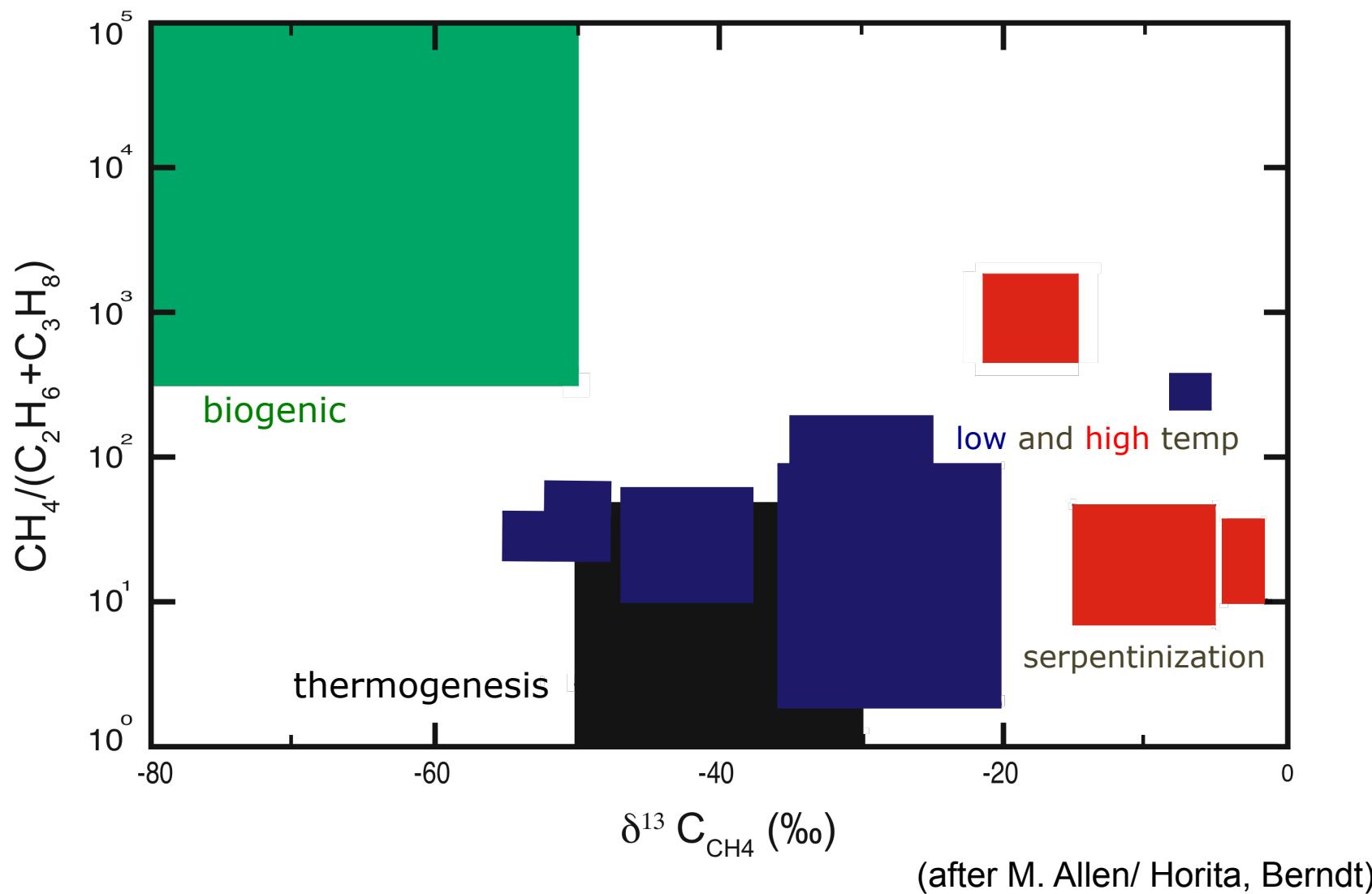


$\delta^2\text{H}$ vs $\delta^{13}\text{C}$ in CH_4

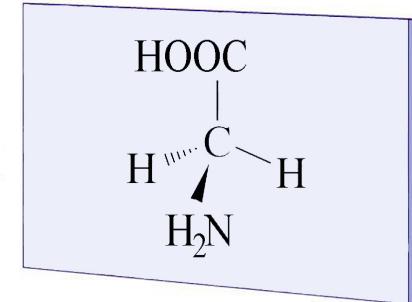
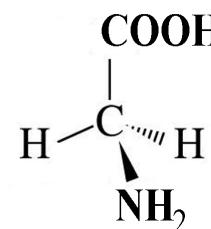
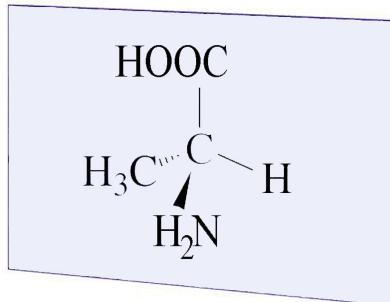
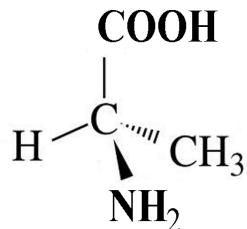


(adapted from M. Allen/ T. Onstott/ Sherwood-Lollar)

ratio of methane/(ethane + propane) with $\delta^{13}\text{C}$ (methane)



alanine

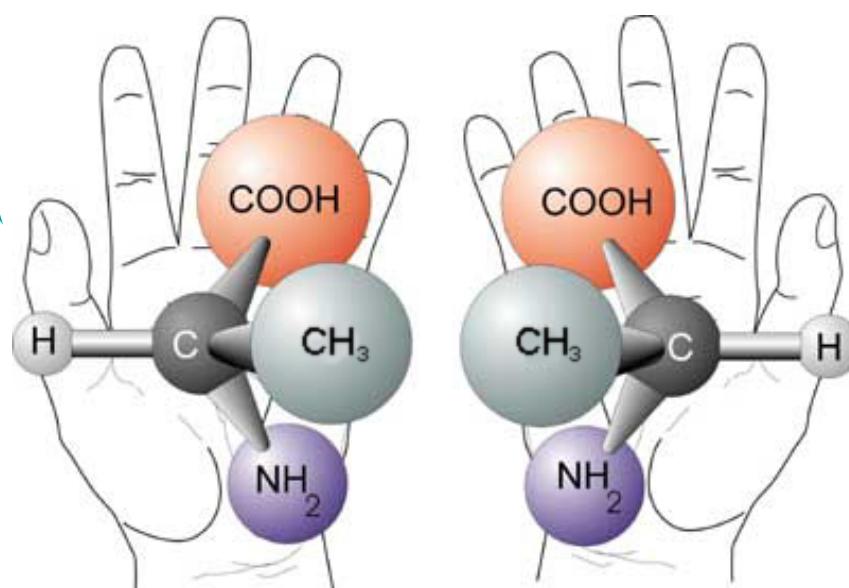


chiral molecule

nonsuperimposable mirror image

achiral molecule

superimposable mirror image



glycine

CHIRALITY

MER, Sojourner, MSL at JPL Science Fair



SAM Suite top assembly

Solid sample inlets penetrate through MSL top deck

Atmospheric inlets and vents located on side of SAM box and penetrate +Y face of MSL WEB

Tunable Laser Spectrometer

Wide Range Pump

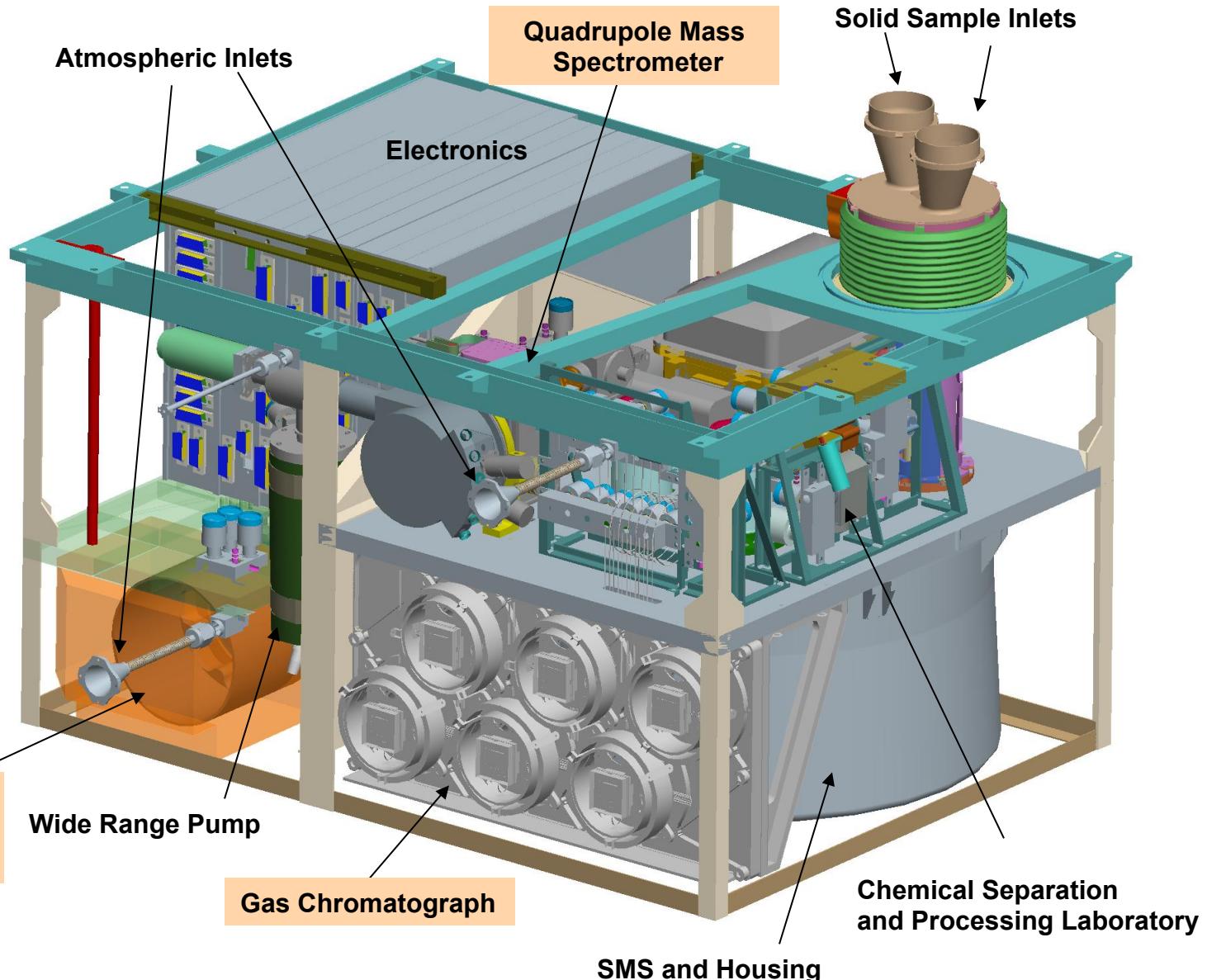
Gas Chromatograph

Solid Sample Inlets

Quadrupole Mass Spectrometer

SMS and Housing

Chemical Separation and Processing Laboratory



take home

- ***Map of Trace Constituents*** in solid & gas samples, including methane, but not limited to it, essential to address key questions of Mars, such as habitability, and extinct or extant life
- ***Find the Organics***
- ***Isotopes of C, S, O, etc.*** also essential
- ***Context: Mineralogical, Geologic, Environmental*** essential
- ***??Surface oxidation state, surface-atmosphere-interior coupling over geologic time, electric fields at surface and in atmosphere, are some of the big unknowns, that still need addressing***

MSL and ExoMars will address some of the above

?

- atreya@umich.edu
- pdf's of pubs: my website (google sushil atreya)
- Scientific American article, “Mystery of Methane on Mars and Titan”: ask or email me