Cross-Scale Multi-scale Coupling in Space Plasmas

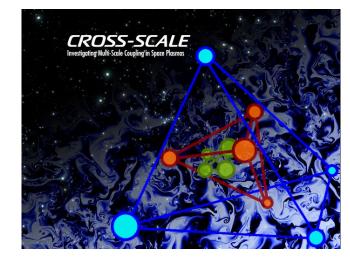
Steve Schwartz & Peter Falkner On behalf of the Cross-Scale Science Study Team

1 December 2009



Outline

- Science
- Payload
- Operations
- Mission
- Development Approach



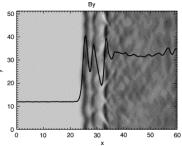


Cross-Scale

Universal Plasma Processes

How do shocks accelerate and heat particles?





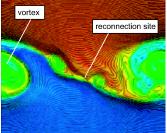
CV Q: "How does the Solar System [and other plasma systems] work?"



Universal Plasma Processes

- How do shocks accelerate and heat particles?
- How does reconnection convert magnetic energy?





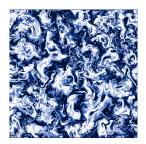
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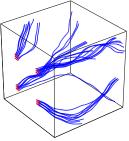


Cross-Scale

Universal Plasma Processes

- How do shocks accelerate and heat particles?
- How does reconnection convert magnetic energy?
- How does turbulence control transport in plasmas?





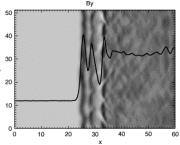
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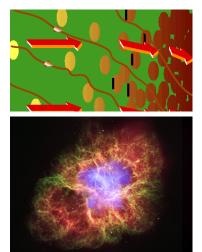
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Questions:

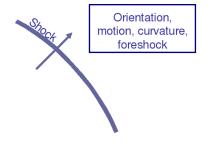
- Acceleration?
- Energy partition?
- Variability?

Universality:

- Cosmic ray acceleration
- CR injection at sub-shock with $M \sim 5 10$
- Variability/efficiency of electron acceleration? (synchrotron radiation used as proxy for ion acceleration)

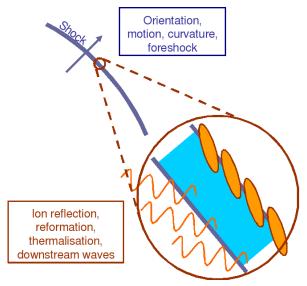


Cross-Scale Coupling at Shocks



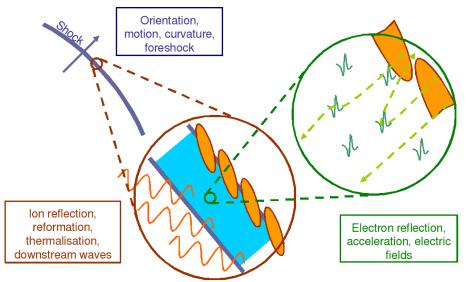


Cross-Scale Coupling at Shocks





Cross-Scale Coupling at Shocks

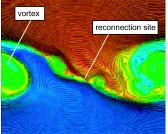




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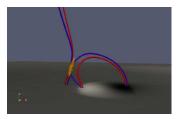
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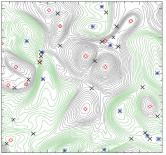
Questions:

- Initiation?
- Magnetic topology?
- · Particle acceleration?

Universality:

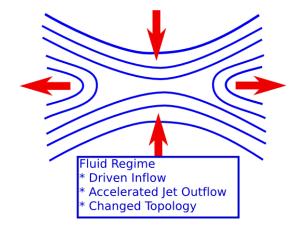
- 3D reconnection in solar flares
- Turbulent reconnection
- Relaxation of galactic dynamo to large-scale field
- Disruption events in laboratory plasmas







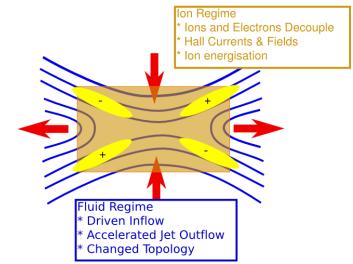
Cross-Scale Coupling and Reconnection





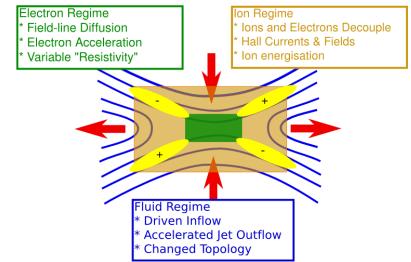
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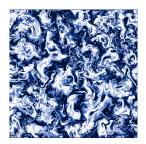
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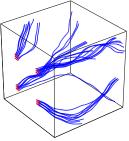




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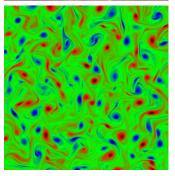


How does turbulence control transport in plasmas?

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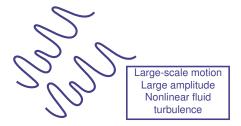
- Energy cascade?
- Anisotropies?
- Coherent structures?
- Universality:
 - · Non-local & inverse cascade
 - Field-line wandering/CR modulation & propagation
 - Heating, acceleration, kinetic modes, ...







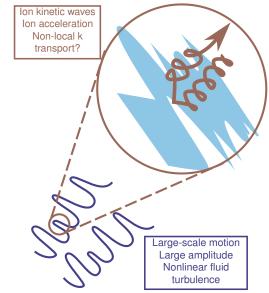
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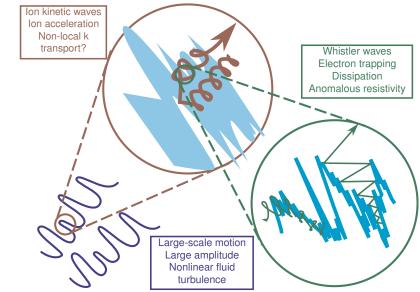
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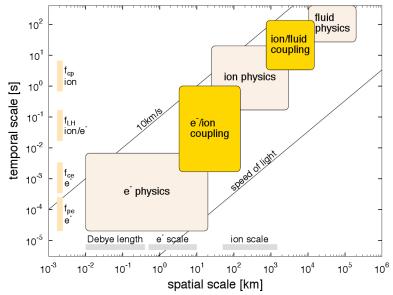


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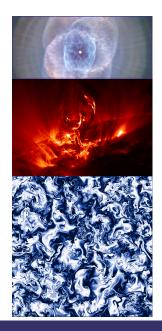


Science Requirements: Conspiracy of Scales





- Time-varying
- Fundamentally 3D
- 3 spatial & temporal scales electron, ion & fluid
- Nonlinearly coupled ⟨Behaviour⟩ ≠ Behaviour of ⟨⟩
- · Simultaneous multi-scale
- Focus: coupling between scales
- Collisionless \Rightarrow Kinetic
- Plasma ⇔ EM fields
- Near-Earth: unique *in situ* plasma laboratory





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10 km; 10 ms: comprehensive P/L



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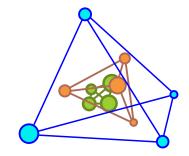


10 km; 10 ms: comprehensive P/L 500 km; 1-5 s: targetted P/L



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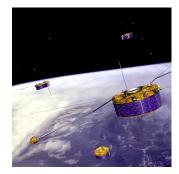
10 km; 10 ms: comprehensive P/L 500 km; 1-5 s: targetted P/L 5000 km; 30 s: context P/L



State of the Art

Cluster

- · First 3D plasma measurements
- · One scale at a time
- Results highlighted need for simultaneous multi-scale
- Success of mission extension also confirming need for all 3 scales in 3D





Cross-Scale

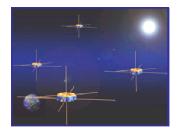
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MMS

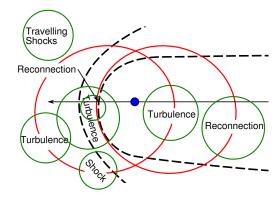
- · Access to electron scales in 3D
- · Only one scale at a time





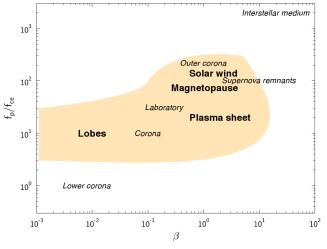
Science Orbit

- Near-Earth space contains
 - Shocks
 - Reconnection
 - Turbulence
- 7 ESA: dual-scale coupling
- Modular instrument
 accommodation
- Scale-specific minimal payload
- Standardised bus
- Shared objectives and orbit
 with JAXA/SCOPE





Science Requirements: Parameter Space



Note wide range of conditions in near-Earth space Overlap with other plasmas

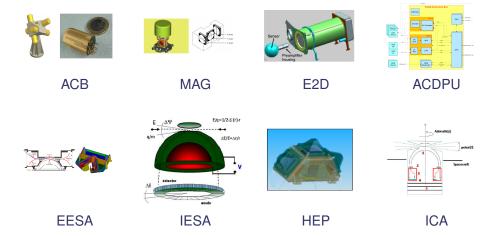


Measurement Requirements

Abbrev	Instrument	Measurement *						
EM Fields								
ACB	AC Magnetic Field	Hi-f waveform & spectra						
MAG	DC Magnetometer	Magnetic field to 128 Hz						
E2D	Spin-plane electric field wire booms	DC & AC electric field						
ACDPU	Fields Electronics Box	Density sounder, wave anal-						
		yser, fields control, processing,						
		power						
	3D $f(\vec{v})$ Partic	les						
EESA	Electron analyser	3D electron $f(\vec{v})$ to 30 keV						
IESA	lon analyser	3D ion $f(\vec{v})$ to 30 keV						
HEP	High energy particle detector	3D particles to 1 MeV						
ICA	lon composition analyser	3D ions to 40 keV/q by mass						
CPP	Common particle processor	control, data processing						
ASP	Active S/C Potential control	ion current						
* See Yellow Book & Science Req. Doc for performance, resolution, etc.								



Instrumentation - Strong Heritage



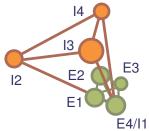


Cross-Scale

Payload Deployment

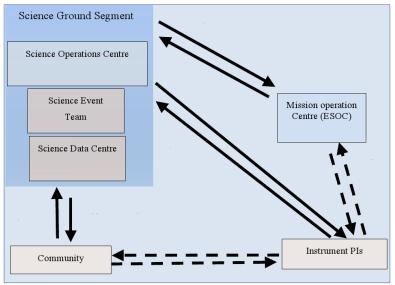
E1	E2	E3	E4/I1	12	13	14
ACB	ACB	ACB	ACB	ACB	ACB	ACB
MAGx2	MAGx2	MAGx2	MAGx2	MAGx2	MAGx2	MAGx2
E2Dinclx4	E2Dx4	E2Dinclx4	E2Dx4	E2Dx4	E2Dx4	E2Dx4
ACDPU	ACDPU	ACDPU	ACDPU	ACDPU	ACDPU	ACDPU
EESAx4	EESAx4		EESAx2	EESAx2	EESAx2	EESAx2
ASP	ASP		IESAx2	IESAx4	IESAx2	IESAx2
			HEP		HEP	ICA
CPP	CPP		CPP	CPP	CPP	CPP

- Fields suite common
- · Particle instruments targetted & balanced
- Initial ion-electron coupling configuration
- · Reconfigure or join international partners





Science Ops - Early planning, Good data access





Cross-Scale

Science Data Selection

- Summary Data
 - 100% recovered
 - \sim 10% telemetry
 - Comprehensive; science quality
- Full-resolution Data
 - Taken continuously
 - Stored onboard

S/C 1	$t_1 1$	t ₁ 2	t ₁ 3	t14	t ₁ 5	t ₁ 6	t ₁ 7	t ₁ 8	 t ₁ (i-1) t ₁ i
S/C 2	t-1	t ₂ 2	t.3	t24	t ₂ 5	t26	t ₂ 7	t28	 t ₂ (i-1) t ₂ i
	2	-	-	2	-	-	-	-	2() 2
S/C 3	t ₃ 1	t ₃ 2	t ₃ 3	t ₃ 4	t ₃ 5	t ₃ 6	t ₃ 7	t ₃ 8	 t ₃ (i-1) t ₃ i
S/C 4	t ₄ 1	t ₄ 2	t ₄ 3	t ₄ 4	t ₄ 5	t46	t ₄ 7	t ₄ 8	 t4(i-1) t4i
							_		
S/C 5	t51	t ₅ 2	t ₅ 3	t ₅ 4	t ₅ 5	t56	t ₅ 7	t ₅ 8	 t ₅ (i-1) t ₅ i
S/C 6	t ₆ 1	t ₆ 2	t ₆ 3	t ₆ 4	t ₆ 5	t ₆ 6	t ₆ 7	t ₆ 8	 t ₆ (i-1) t ₆ i
a (a =									
S/C 7	t ₇ 1	t ₇ 2	t ₇ 3	t ₇ 4	t ₇ 5	t ₇ 6	t ₇ 7	t ₇ 8	 t7(i-1) t7i



Science Data Selection

- Summary Data
 - 100% recovered
 - \sim 10% telemetry
 - Comprehensive; science quality
- Full-resolution Data
 - Taken continuously
 - Stored onboard
 - Science Event List selects subsets
- Science Event List part of operations

S/C 1 t ₁ 1	t ₁ 2 t ₁ 3	14	t ₁ 5	t ₁ 6	t ₁ 7	t ₁ 8		t ₁ (i-1)	t ₁ i
S/C 2 t ₂ 1	t ₂ 2 t ₂ 3	24	t ₂ 5	t ₂ 6	t ₂ 7	t ₂ 8		t ₂ (i-1)	t ₂ i
S/C 3 t ₃ 1	t ₃ 2 t ₃ 3	34	t ₃ 5	t36	t₃7	t ₃ 8		t ₃ (i-1)	t3i
S/C 4 t ₄ 1	t ₄ 2 t ₄ 3	44	t ₄ 5	t46	t₄7	t ₄ 8		t ₄ (i-1)	t4i
S/C 5 t ₅ 1	t ₅ 2 t ₅ 3	54	t ₅ 5	t56	t₅7	t ₅ 8		t ₅ (i-1)	t5i
S/C 6 t ₆ 1	t ₆ 2 t ₆ 3	64	t ₆ 5	t ₆ 6	t ₆ 7	t ₆ 8		t ₆ (i-1)	t ₆ i
S/C 7 t ₇ 1	t ₇ 2 t ₇ 3	:74	t ₇ 5	t ₇ 6	t ₇ 7	t ₇ 8		t7(i-1)	t7i
FR Data to be telemetred									



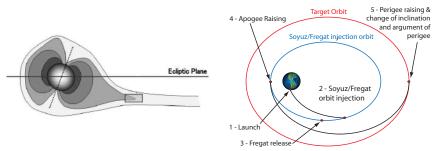
Mission Overview: Launch & Insertion

- · 3 year mission, 7 Spacecraft & 1 transfer stage
- Single Soyuz-Fregat launch from Kourou, October 2017
- Launch mass: 3570 3703 kg
- Insertion into:

200 km x 5.3 (5.8) R_E orbit

1630 (1420) m/s

- Transfer by chemical propulsion module (optional lunar resonance)
- Transfer ΔV :
- Target Orbit: 10 x 25 R_E, $i = 14^{\circ}$, $\omega = 205^{\circ}$, T = 4.3 days





Mission Overview: Configuration & Environment

- Configuration ΔV : 14 200 m/s
- Scales: electron: 2-100 km, ion: 50-2000 km, fluid: 3000-15000 km
- · Electrical power per spacecraft:
- · Mass per spacecraft:
- Radiation Environment:

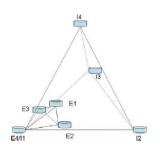
160 kg to 210 kg (dry)

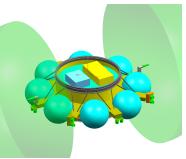
220 – 240 W

33 krad (1.5mm), 9 krad (4 mm)

(54 krad (1.5 mm), 11 krad (4 mm))

Operation done by ESOC and ESAC



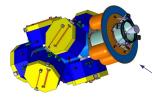




Mission Concept

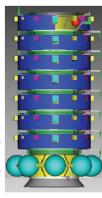
- Propulsion module (PM) performs insertion of the stack
 - PM brings stack into 10 by 25 Re orbit, then discarded
 - PM controlled by spacecraft on top of it
 - Analysis of collision between PM and spacecraft: no risk
- · 7 identical spacecraft
- · Inter-spacecraft ranging on electron scale spacecraft
- · Each spacecraft has a direct link with Earth
- 15 rpm spinning spacecraft (after separation)
- · Simplified propulsion system for science spacecraft





S/C in control of PM

Propulsion module

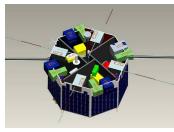


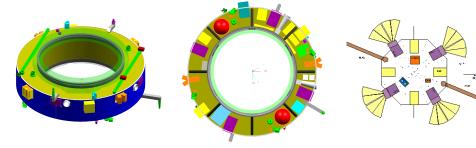
Dimension of spacecraft driven by power demand

Dimension of spacecraft driven by structural stiffness

Model Payload

- 5 different payload configurations
- P/L Mass range:
- P/L Power range:
- Average data rate:
- Total 93 sub-units
- · Up to 1 Tbit mass memory per spacecraft
- · Spacecraft configuration allows variation on P/L





15 – 33 kg

15 - 60 W

800 kbps

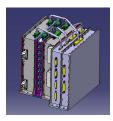


Cross-Scale

Development Approach

- A: 2 x Proto-Flight Model + 5 recurring flight models
- B: EQM + 7 recurring flight models
- · Parallel AIV for the flight model via island approach
- · Requires seamless arrival of instruments in time
- Schedule for 2017 launch is challenging
- Minimal technology upgrade activities required
 - 15 rpm star mapper
 - Low gain antenna improvement
 - Inter-spacecraft link adaptation
 - Mass memory optimisation





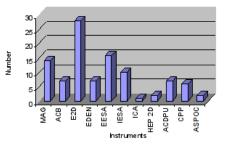


Cross-Scale

Most Critical Items

Feasible mission concept with following critical items:

- Cost at Completion
- Schedule (7+1 spacecraft need to be made) for launch 2017
- · Number of Instrument sub-units (93 FM) & availability in time
- Number of Spacecraft and resulting AIT/AIV
- EMC requirements driving AIV/AIT

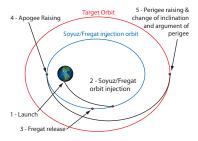


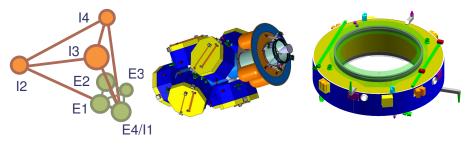


Technology Summary

Feasible mission concept

- Good industrial solutions
- Strong heritage for instruments & space sector
- Clearly-identified but manageable risks



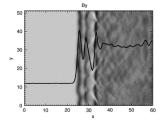




Cross-Scale

Cross-Scale addresses fundamental questions about how solar system astrophysical, and laboratory plasmas work by *simultaneously* measuring the coupling between electron, ion, and fluid scales.

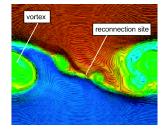
· How do shocks accelerate particles?





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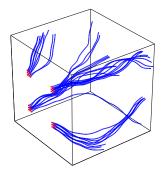
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- How does turbulence control transport in plasmas?





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- · How do shocks accelerate particles?
- How does reconnection convert magnetic energy?
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The mission

- · Mature concept within ESA and member states
- Strong synergies with JAXA/SCOPE including CSA

