

Overarching Theme: The emergence of habitable worlds around gas giant planets



Addresses two Cosmic Vision Themes:

- What are the conditions for planet formation and the emergence of life?
- How does the Solar System work?

Presentations by:

Michele Dougherty (Imperial College London)

Introduction and Summary

Olivier Grasset (Nantes University)

Habitability of the icy moons

Emma Bunce (Leicester University)

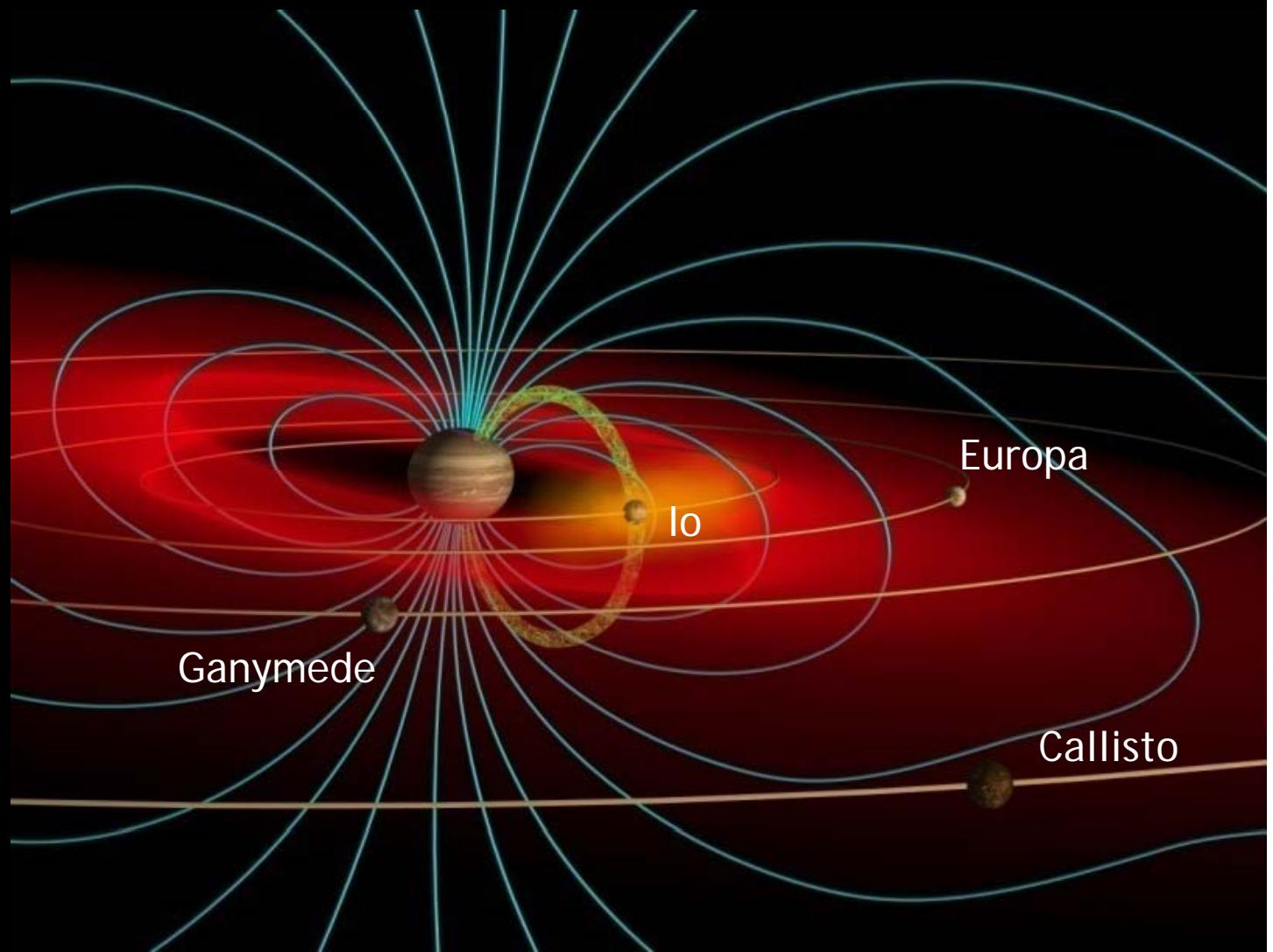
Explore the Jupiter System as an archetype for gas giants

Christian Erd (ESA)

Mission Implementation

Introduction : The diverse Jupiter System

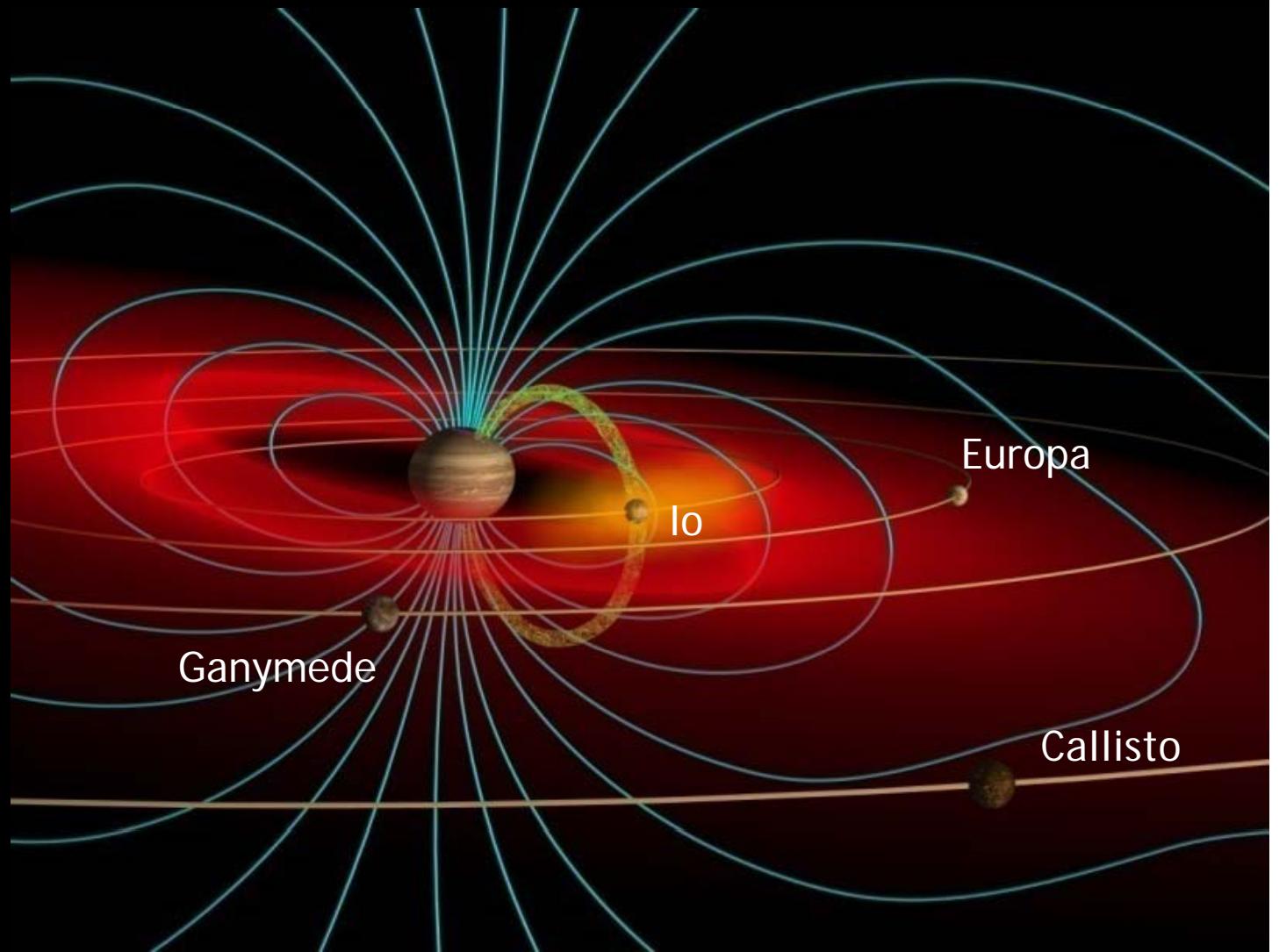
EJSM-Laplace



Introduction : The diverse Jupiter System

EJSM-Laplace

Largest planet
with the largest
magnetic field

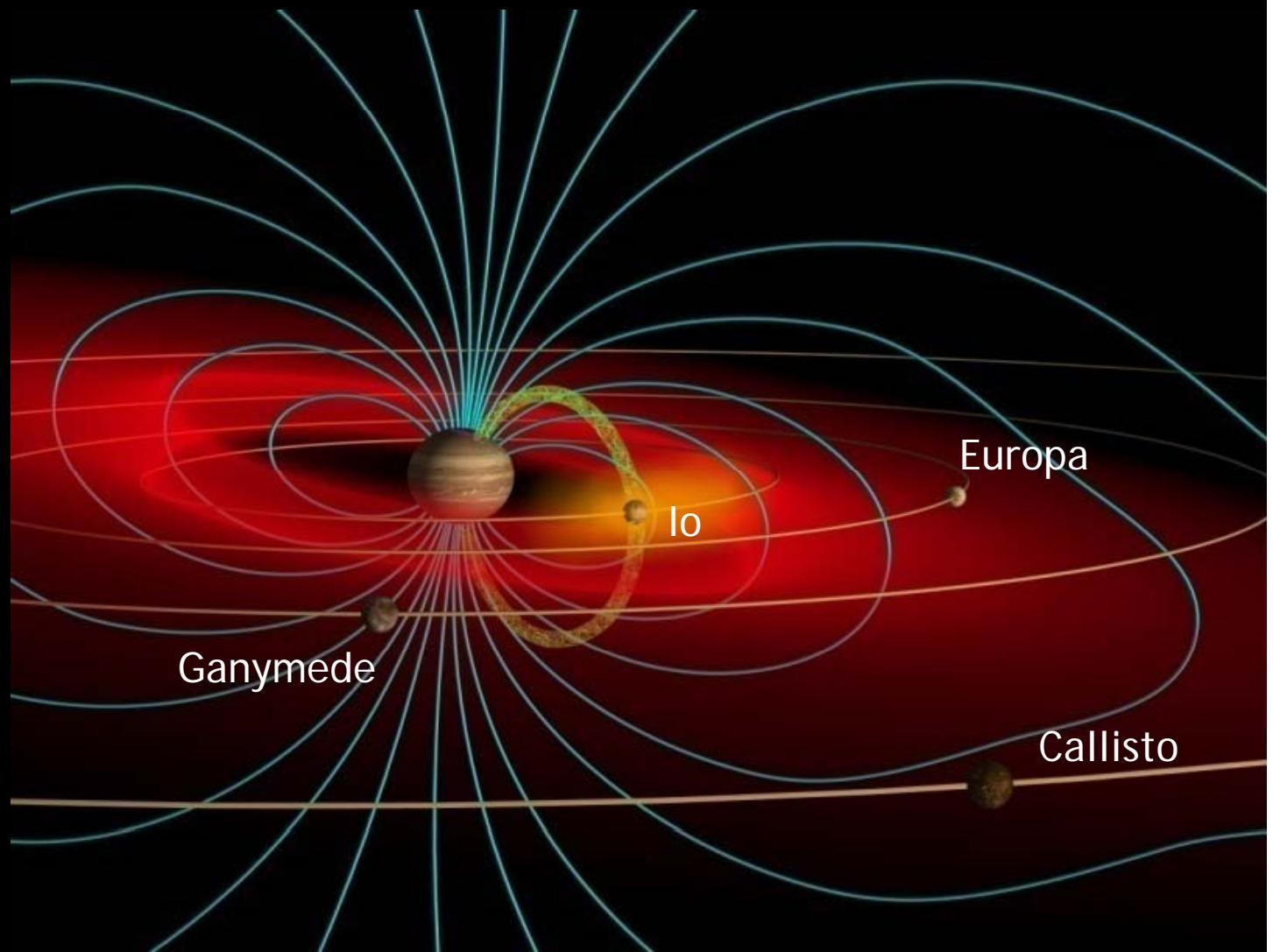


Introduction : The diverse Jupiter System

EJSM-Laplace

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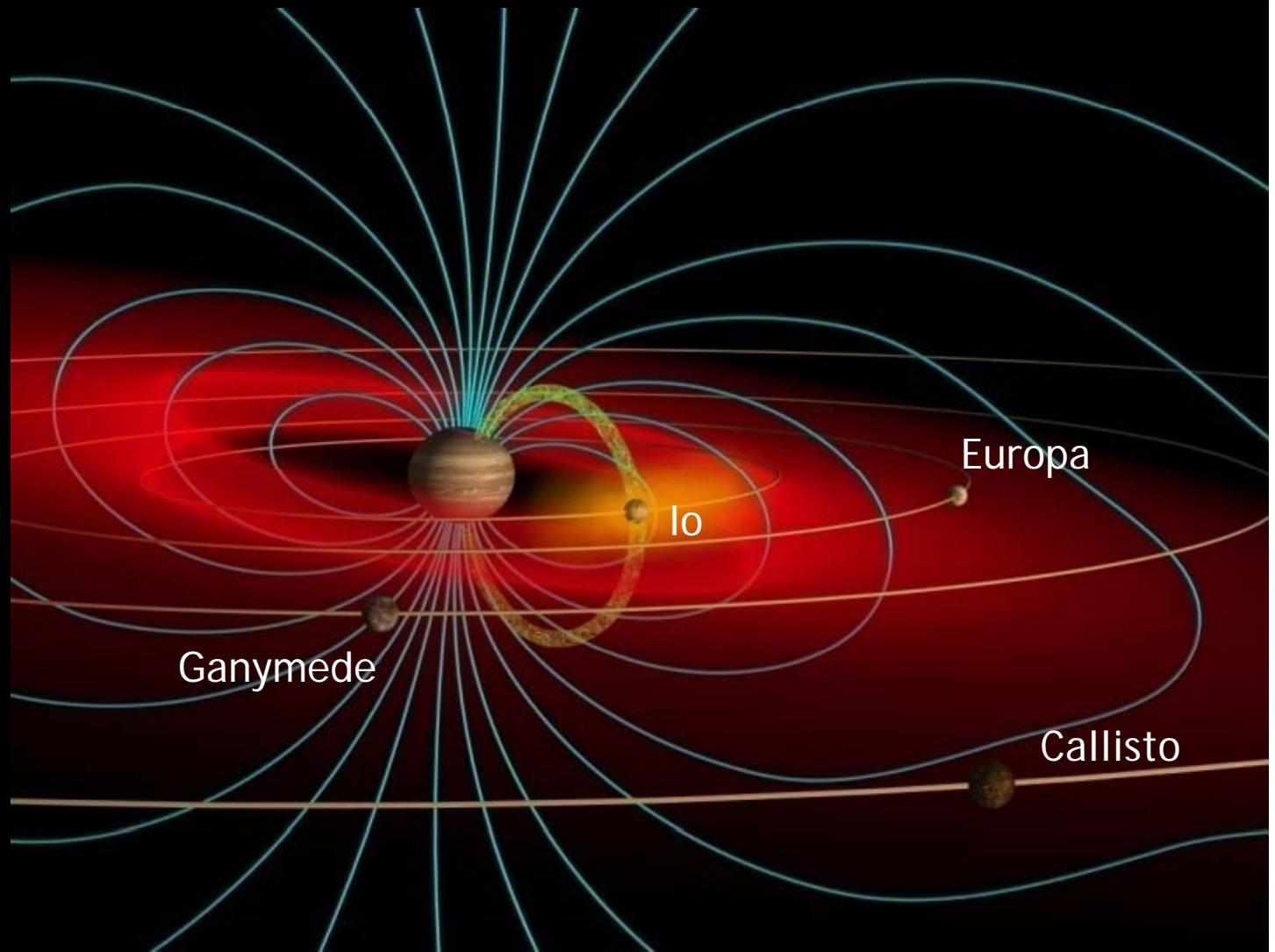
Most active
atmosphere



Largest planet
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Most active
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Most powerful
and largest
magnetosphere



Introduction : The diverse Jupiter System

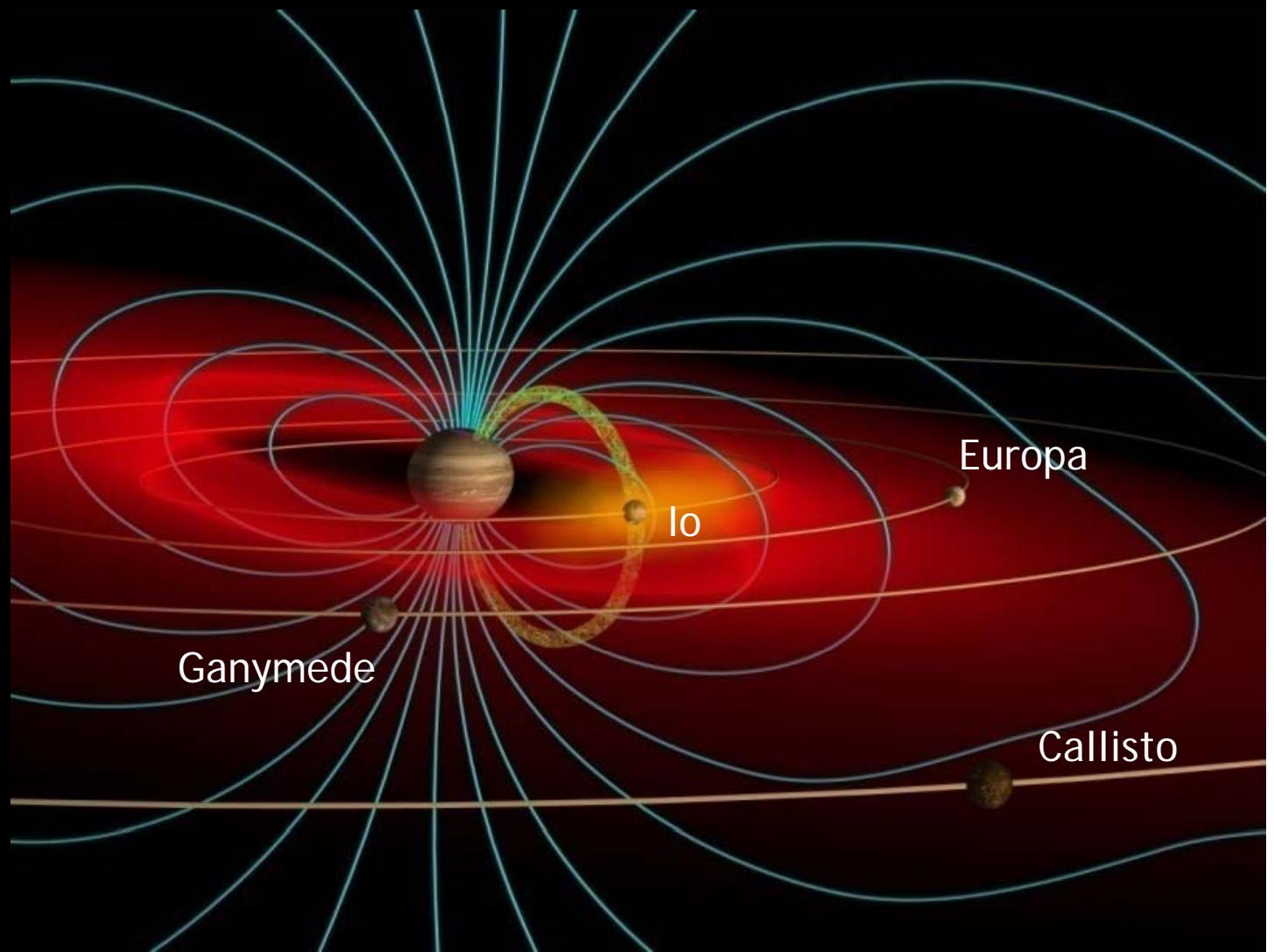
EJSM-Laplace

Largest planet
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Most active
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Most powerful
and largest
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63 moons, 4
large Galilean
moons



Introduction : The diverse Jupiter System

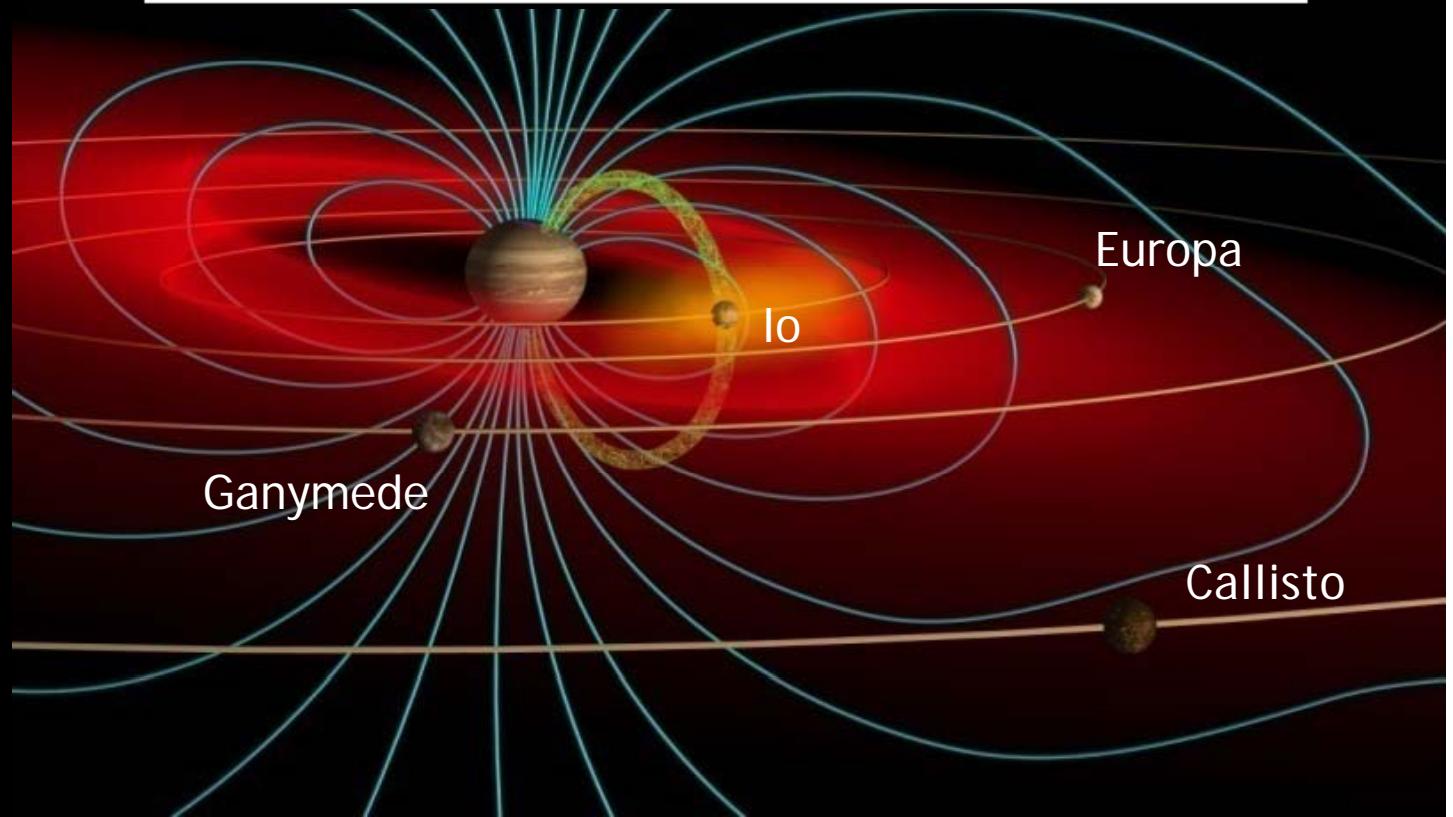
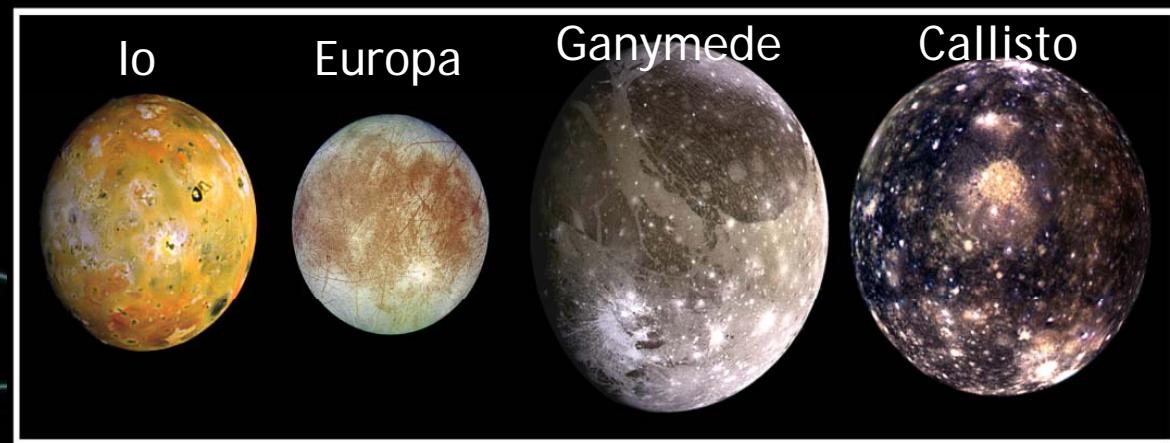
EJSM-Laplace

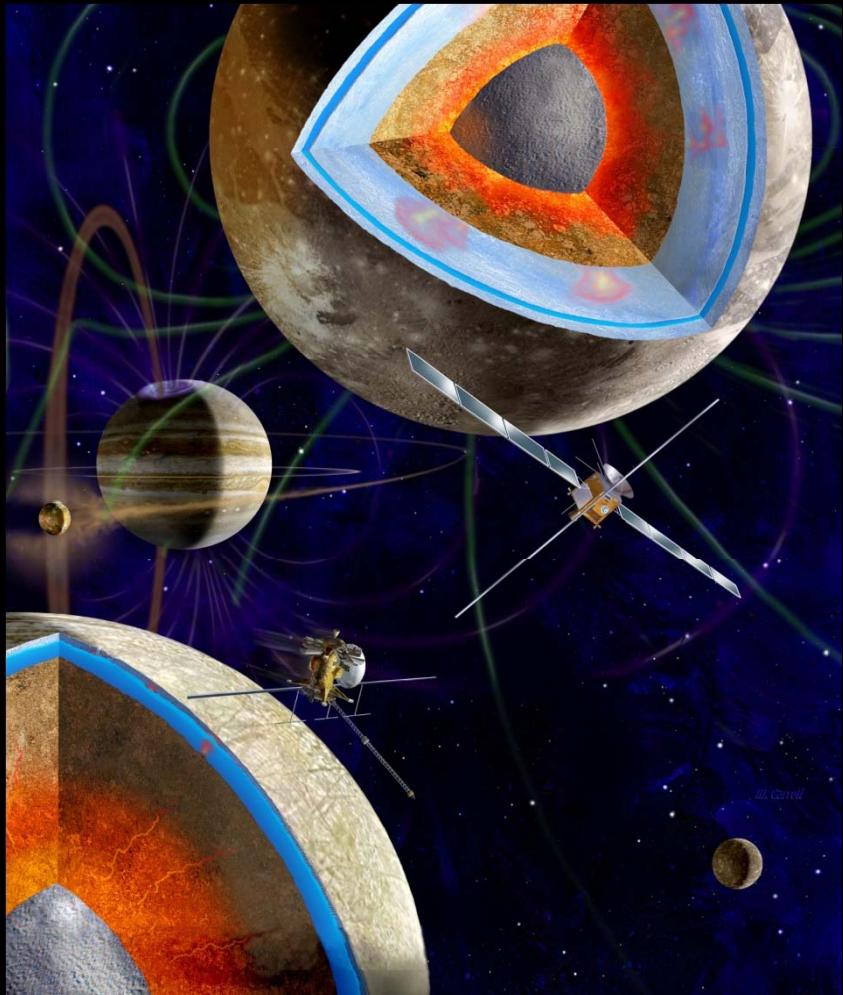
Largest planet
with the largest
magnetic field

Most active
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Most powerful
and largest
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63 moons, 4
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Previous studies:

- *ESA's Laplace proposal*
- *2 NASA studies*

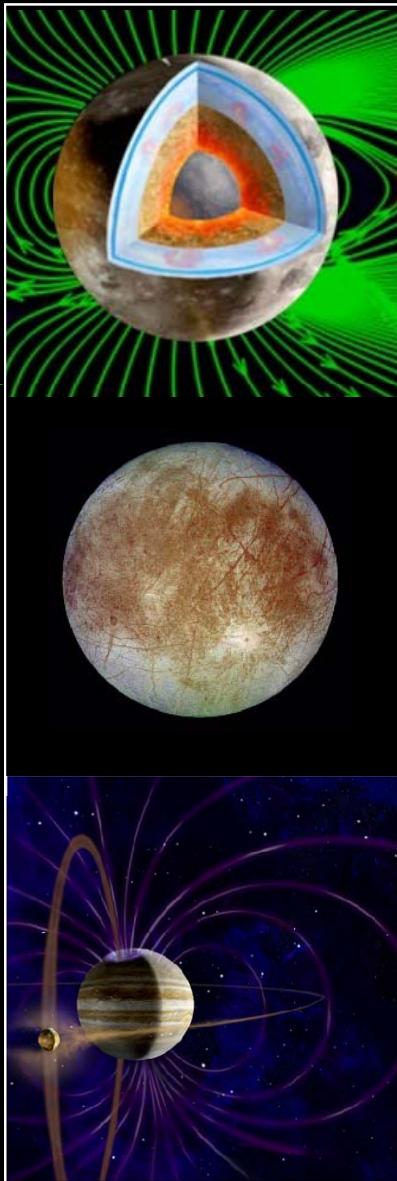
Key Questions:

- Does the Jupiter system harbor habitable worlds?
- What are the processes operating within the Jupiter system?

EJSM-Laplace mission:

- Jupiter Ganymede Orbiter (JGO)
- Jupiter Europa Orbiter (JEO)

The emergence of habitable worlds around gas giants

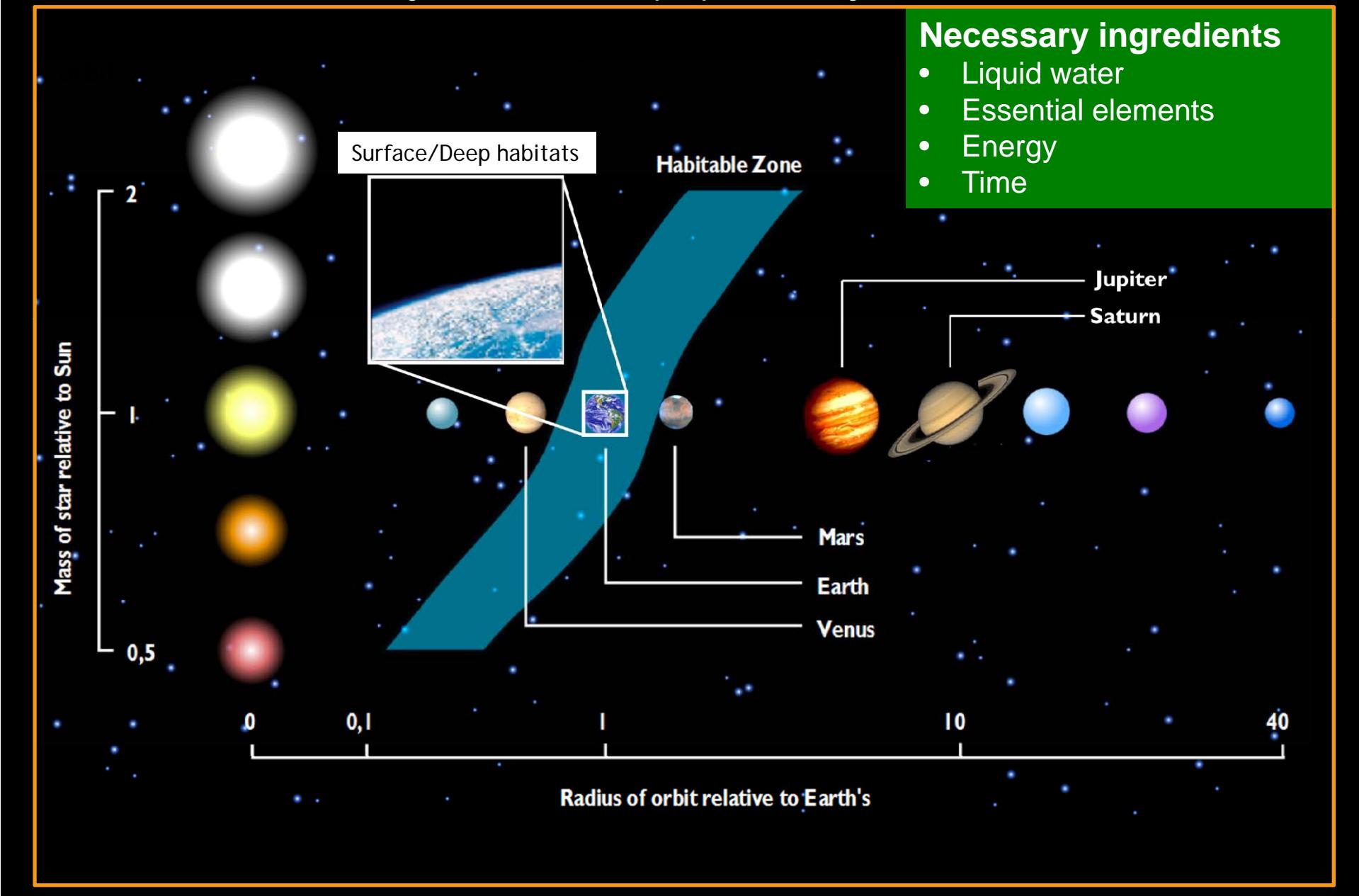


- **Ganymede (JGO-focus):**
Characterise Ganymede as a planetary object including its potential habitability
- **Europa (JEO-focus):**
Explore Europa to investigate its habitability
- **Jupiter System (JGO + JEO):**
Explore the Jupiter system as an archetype for gas giants

Introduction : Habitability of the icy moons

EJSM-Laplace

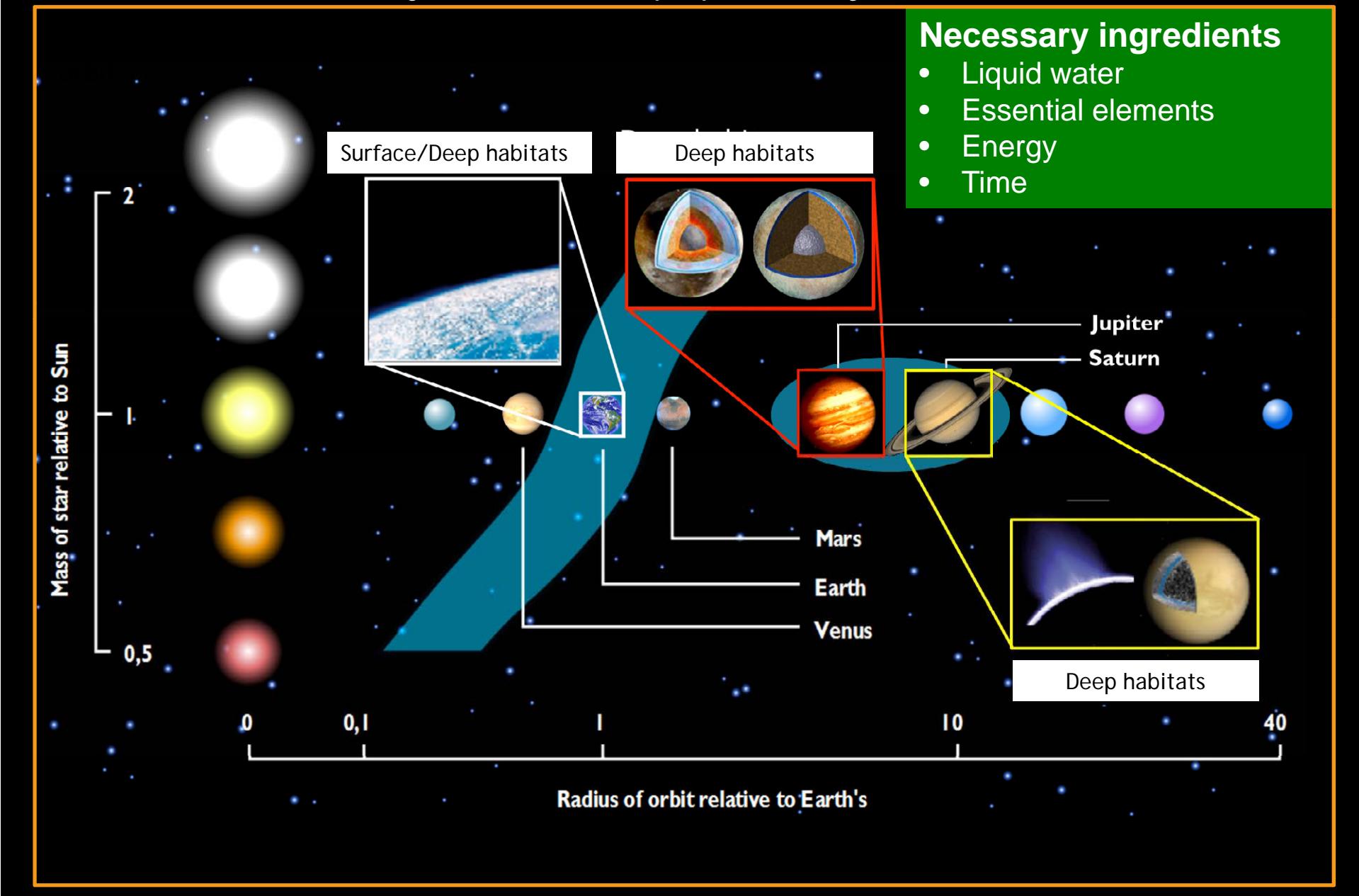
Are Ganymede and Europa potentially habitable worlds ?



Introduction : Habitability of the icy moons

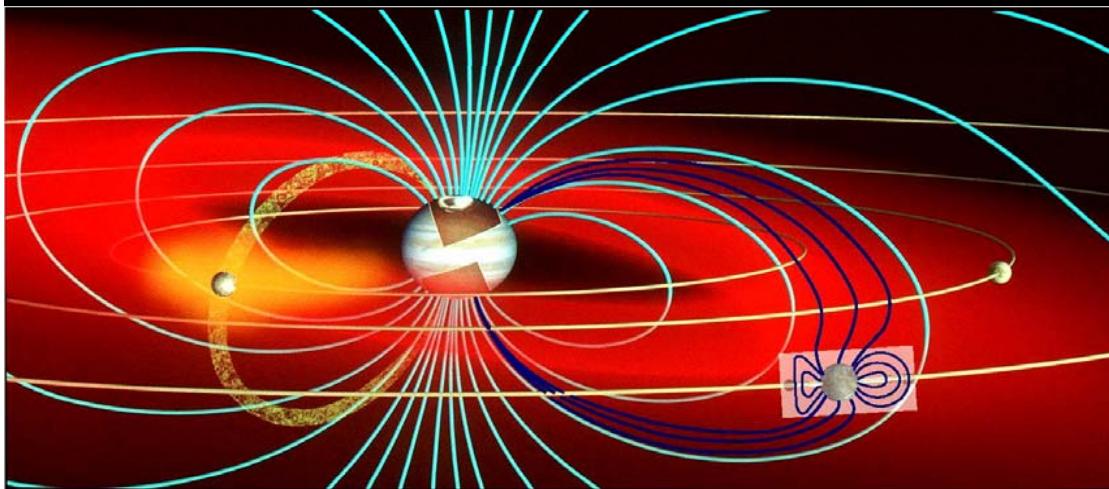
EJSM-Laplace

Are Ganymede and Europa potentially habitable worlds ?



Introduction: Exploration of the Jupiter system

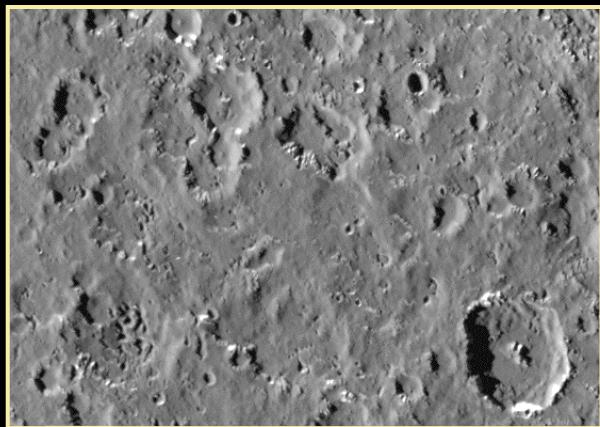
EJSM-Laplace



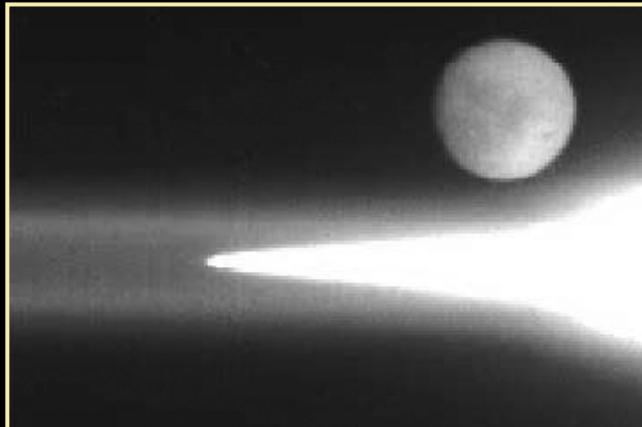
Jovian magnetosphere



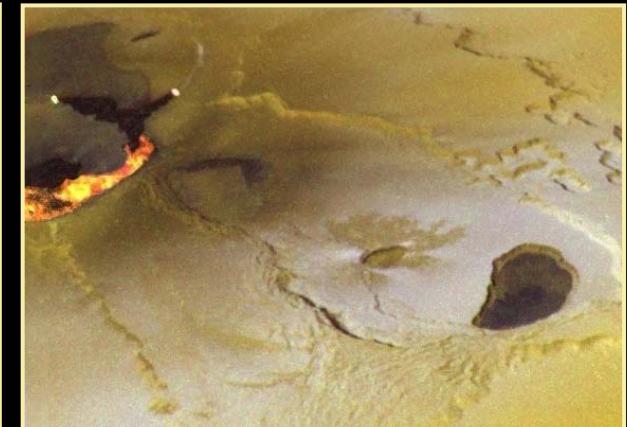
Jovian atmosphere



Callisto



Rings & small satellites



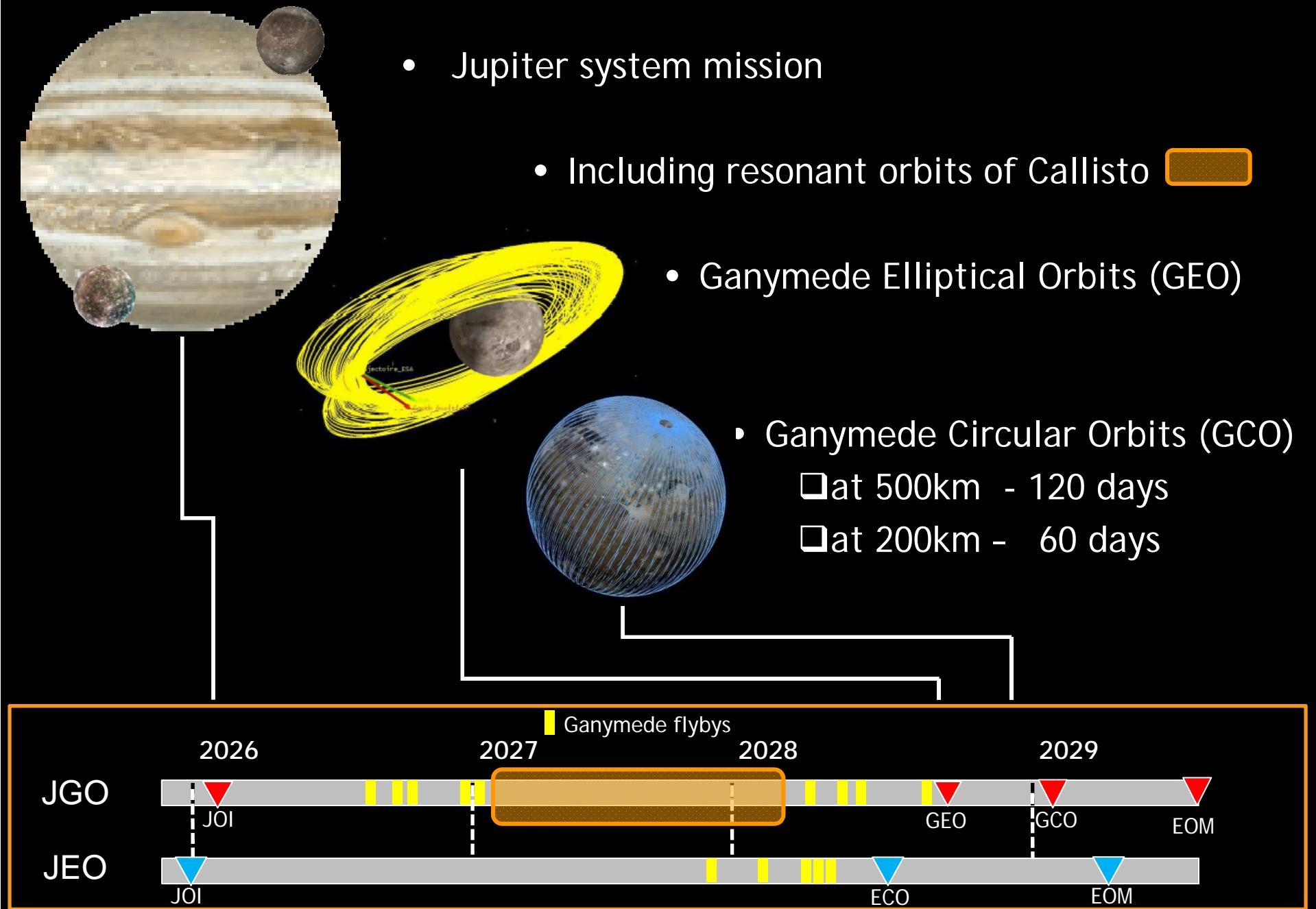
Io

The Jupiter system is rich in dynamic and coupled processes

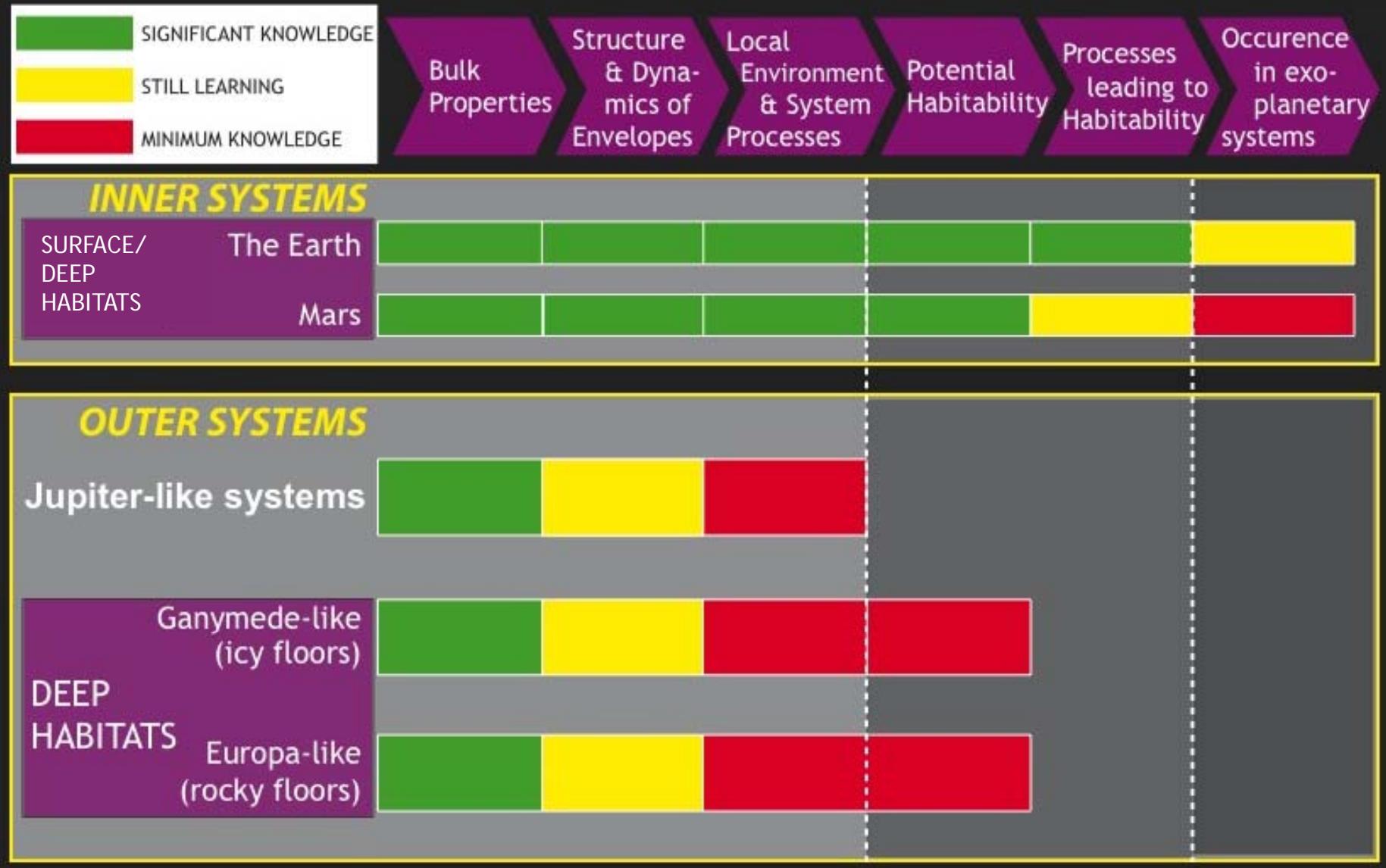
Introduction: Mission Timeline

EJSM-Laplace

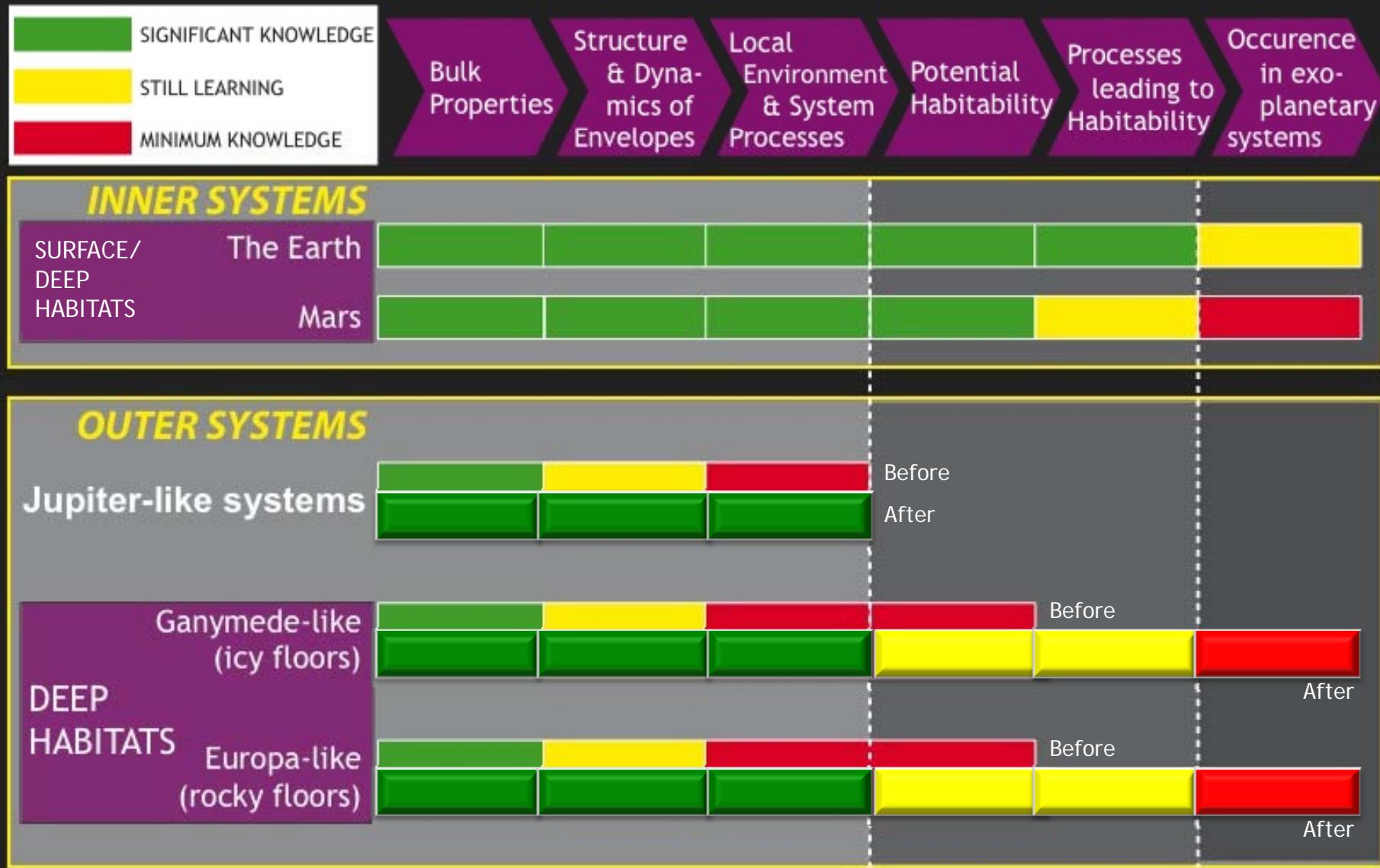
- Jupiter system mission
 - Including resonant orbits of Callisto
- Ganymede Elliptical Orbits (GEO)
- Ganymede Circular Orbits (GCO)
 - at 500km - 120 days
 - at 200km - 60 days



The Emergence of Habitable Worlds: Different destinations, the same steps...



The Emergence of Habitable Worlds: Different destinations, the same steps...



Ganymede & Europa exploration

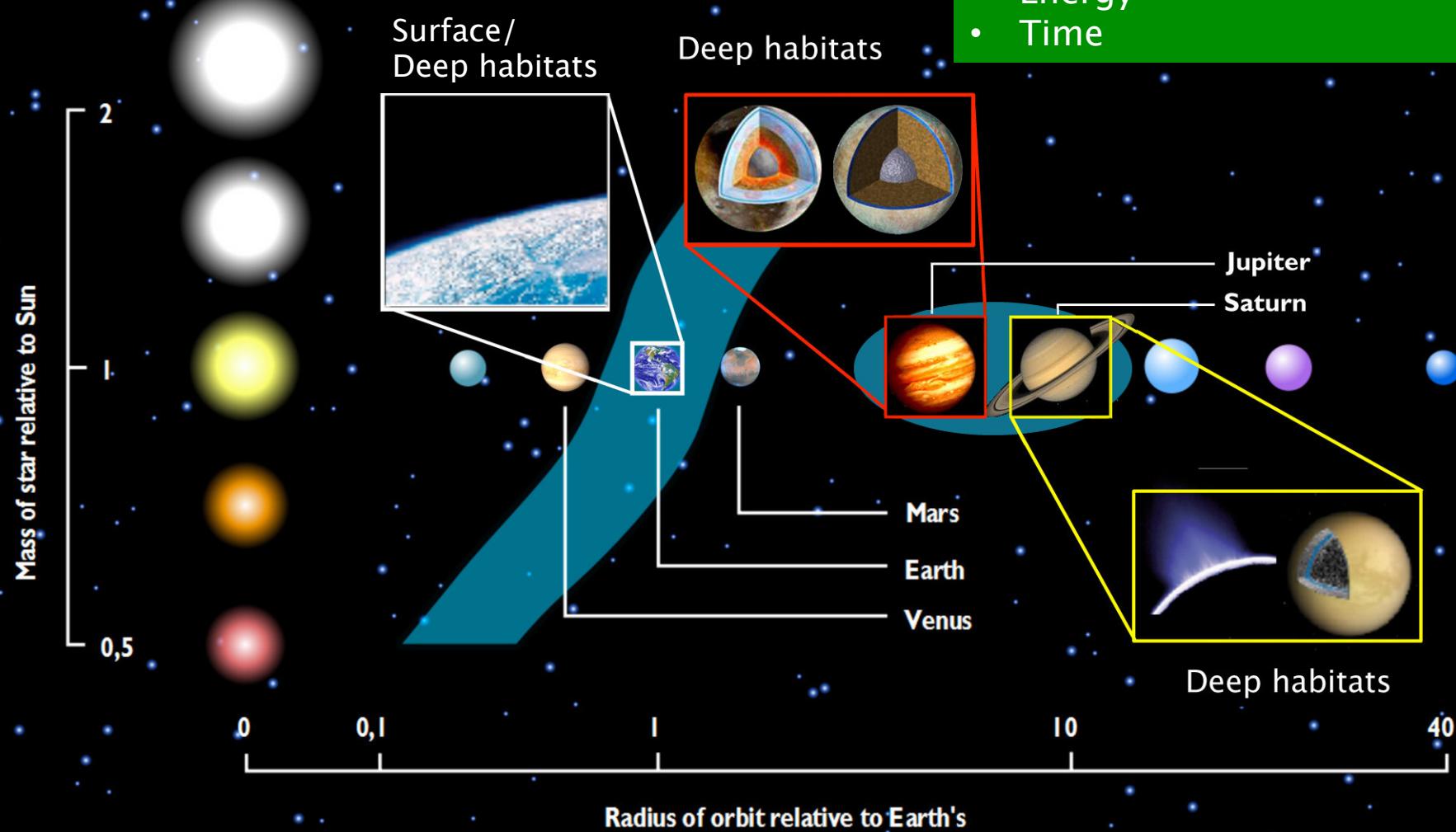
Olivier Grasset

Are Ganymede and Europa potentially habitable worlds ?

The habitable zone is not restricted to the Earth's orbit...

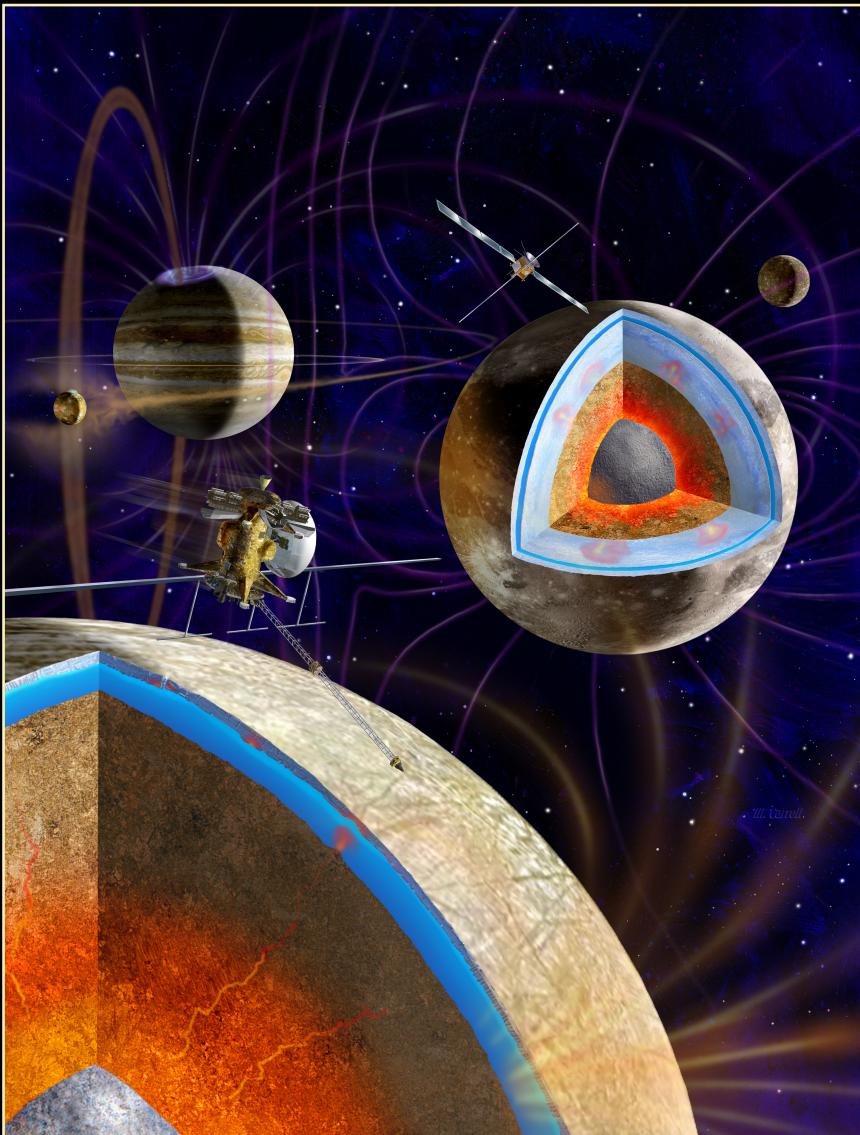
Necessary ingredients

- Liquid water
- Essential elements
- Energy
- Time



Are Ganymede and Europa potentially habitable worlds ?

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Necessary ingredients

- Liquid water
- Essential elements
- Energy
- Time

Ganymede

- Tenuous atmosphere
- Icy surface
- **Two icy layers**
- One liquid layer
- Silicates
- Iron core

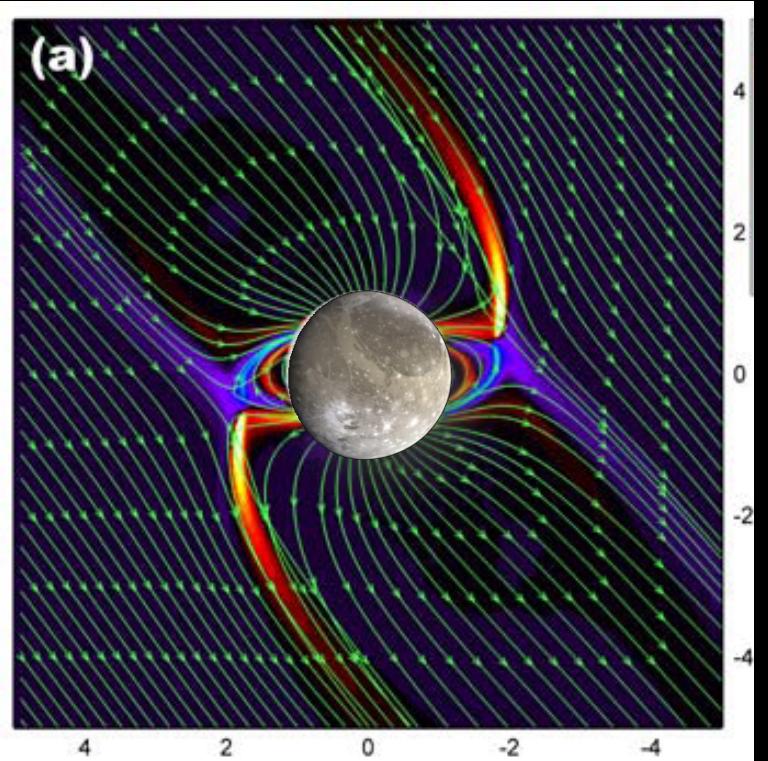
Europa

- Tenuous atmosphere
- Icy surface
- **One icy crust**
- One liquid layer
- Silicates
- Iron core

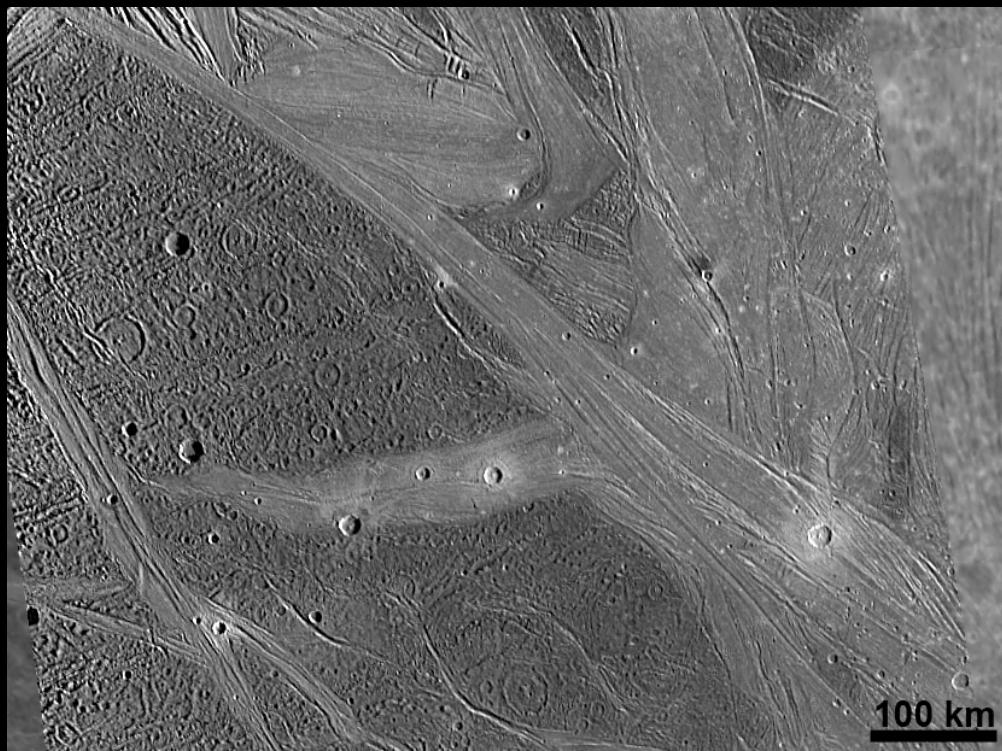
Galileo evidences of liquid layers

Induced magnetic field

Marginally demonstrated but more data needed



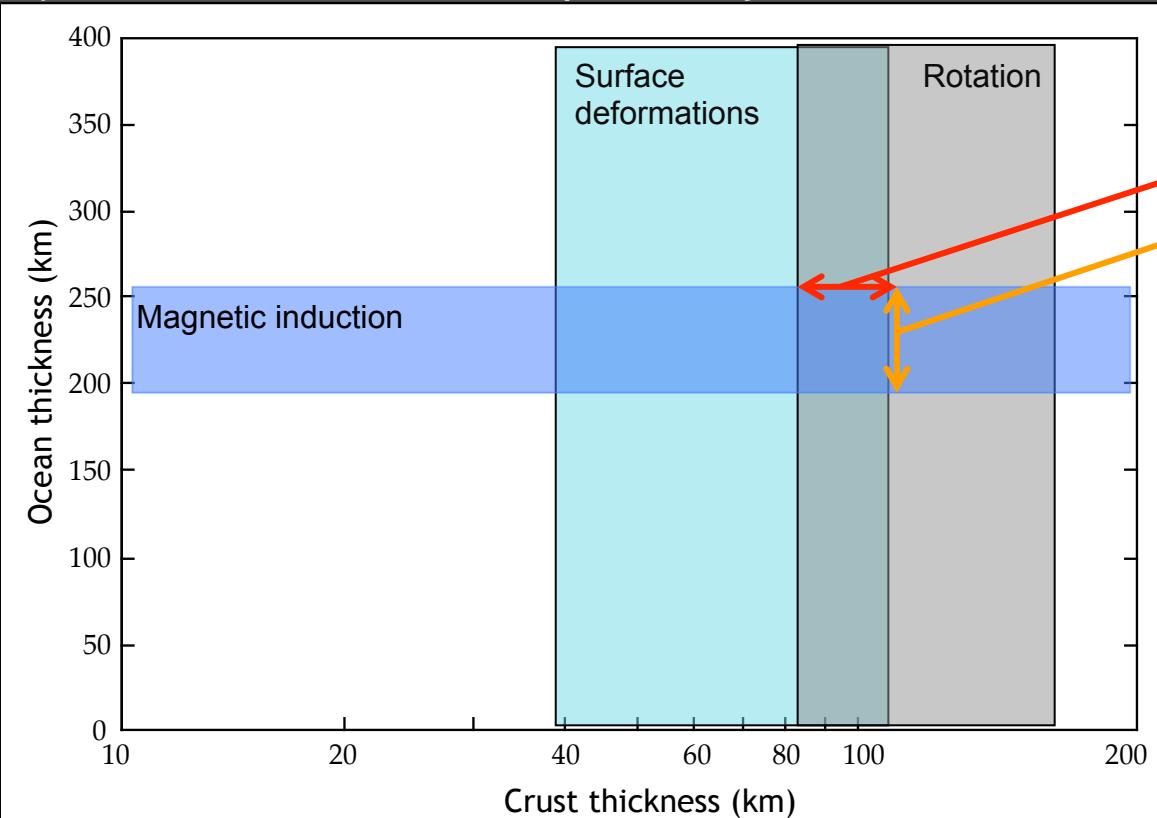
Geologic activity



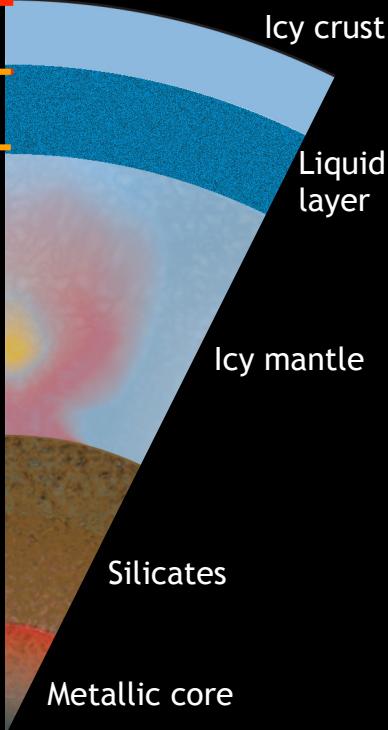
Science questions

- ✧ Existence of a liquid layer
- ✧ Location and size

Ganymede's ocean: EJSM-Laplace objectives



Internal structure



JGO measurements

- Surface deformations
- Rotation
- Magnetic induction

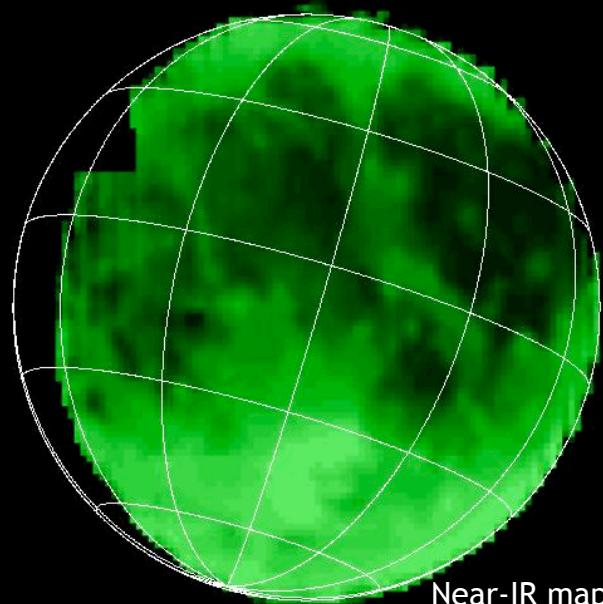
Instrument Packages

- In situ Fields and Particles
- Imaging
- Sounders and Radio Science

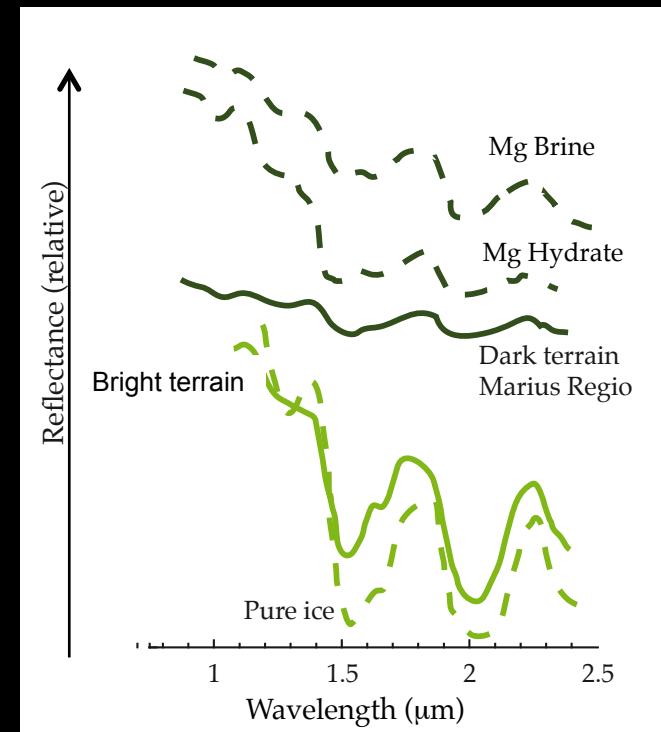


Galileo evidences of an outstanding complexity

Complex chemistry



Near-IR map of Ganymede

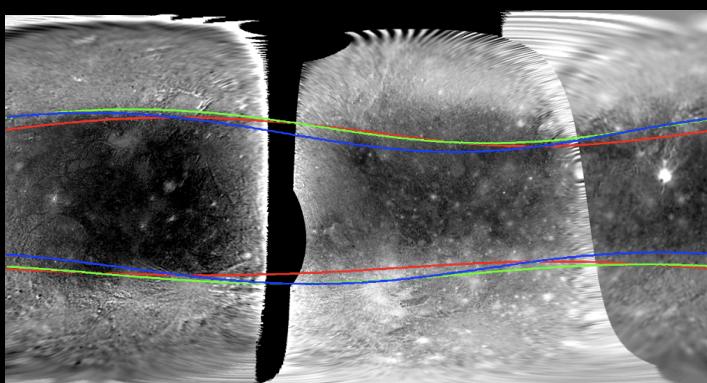


Frozen brines and hydrates ?

Complex dynamics

Specific albedo distribution

Important exogenic contribution

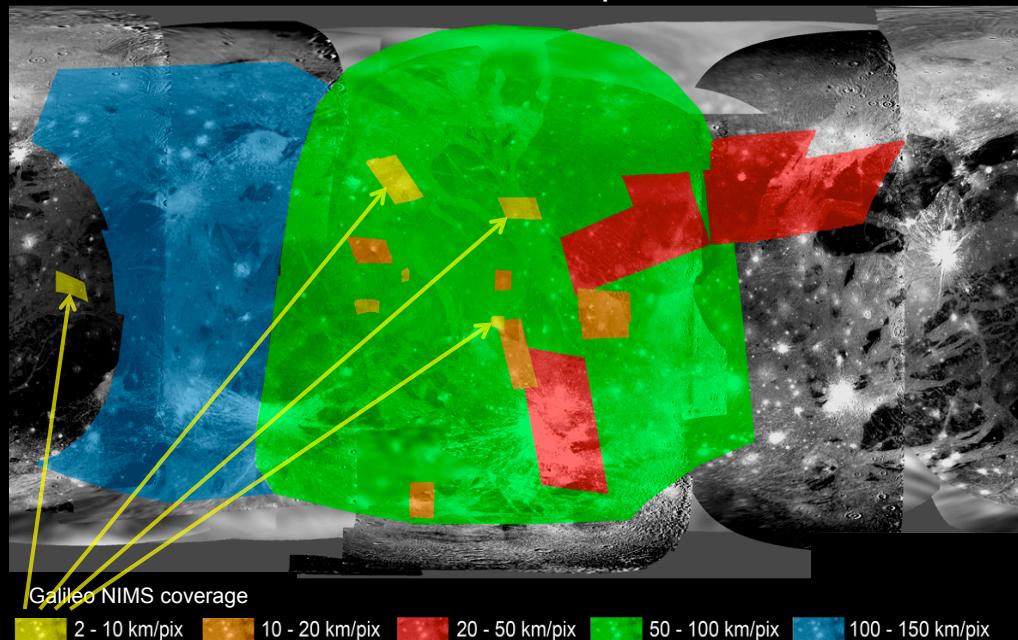


Science questions

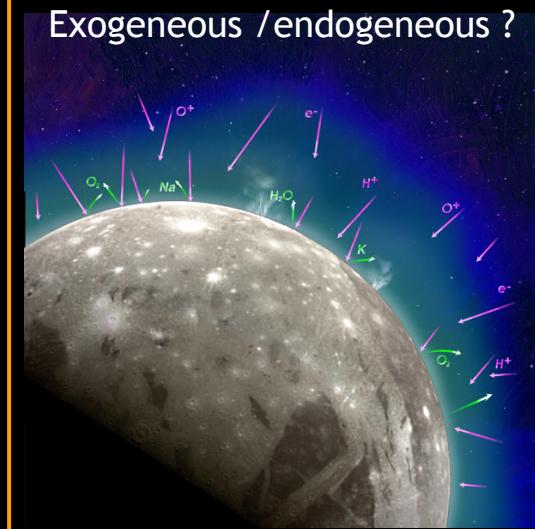
- ❖ Nature of non-ice compounds
- ❖ Surface / sub-surface correlation
- ❖ Endogenous/exogenous contribution

Surface composition: EJSM-Laplace objectives

What are the surface chemical compounds ?



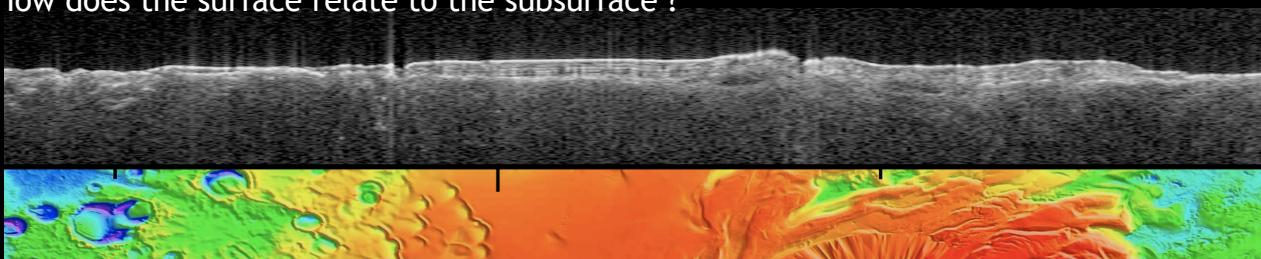
Exogeneous / endogeneous ?



Volatiles

Ions and Neutrals

How does the surface relate to the subsurface ?



Instrument Packages

- Spectroscopy
- Imaging
- In situ Fields and Particles
- Sounders

JGO

2026

2027

2028

2029

JOI

Ganymede flybys

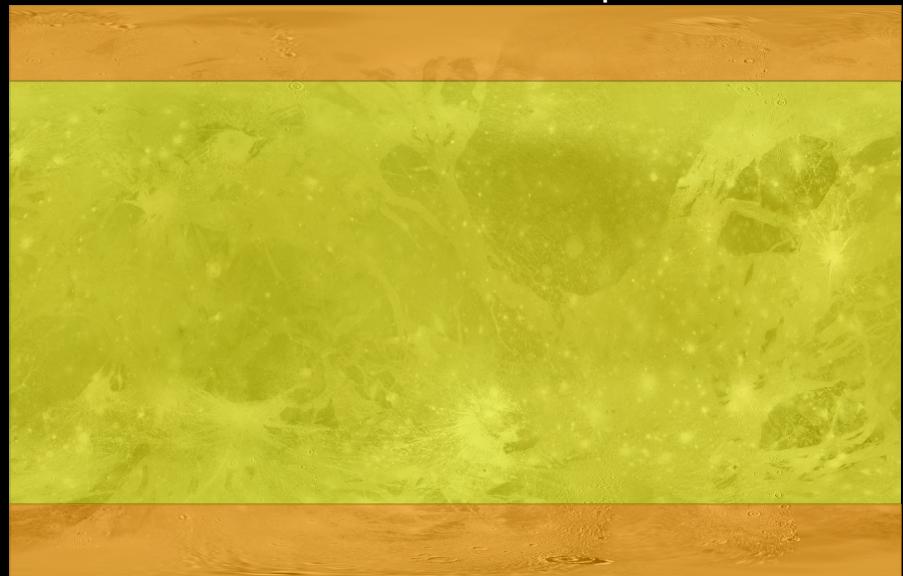
GEO

GCO

EOM

Surface composition: EJSM-Laplace objectives

What are the surface chemical compounds ?



Galileo NIMS coverage

2 - 10 km/pix 10 - 20 km/pix 20 - 50 km/pix 50 - 100 km/pix 100 - 200 km/pix

Imaging spectroscopy

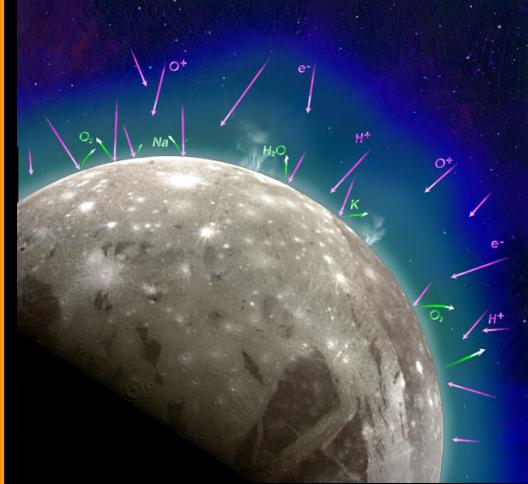
Spatial coverage

- Global at a few km/px
- 100 m/px on a few %
- 10 m/px where needed

Spectral coverage

- > 5 times better than Galileo NIMS
- Close to lab data quality when needed

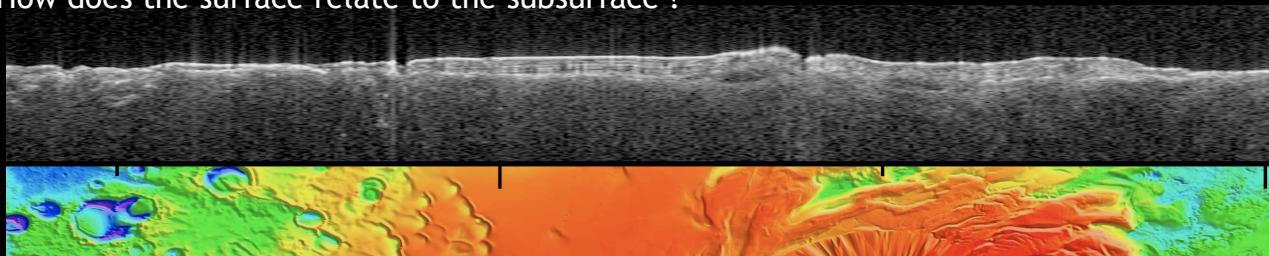
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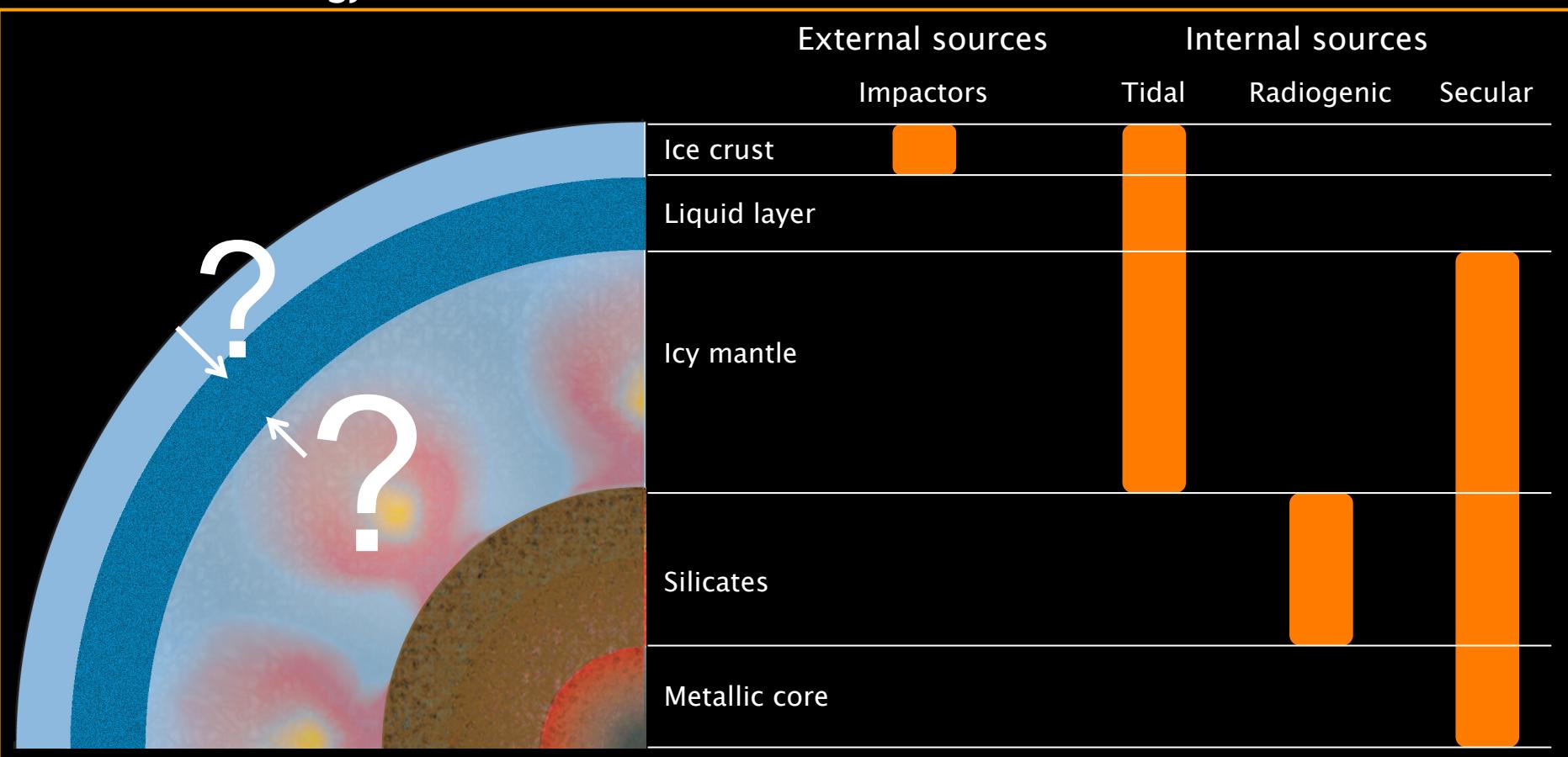
Ganymede flybys

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What are the energy sources?

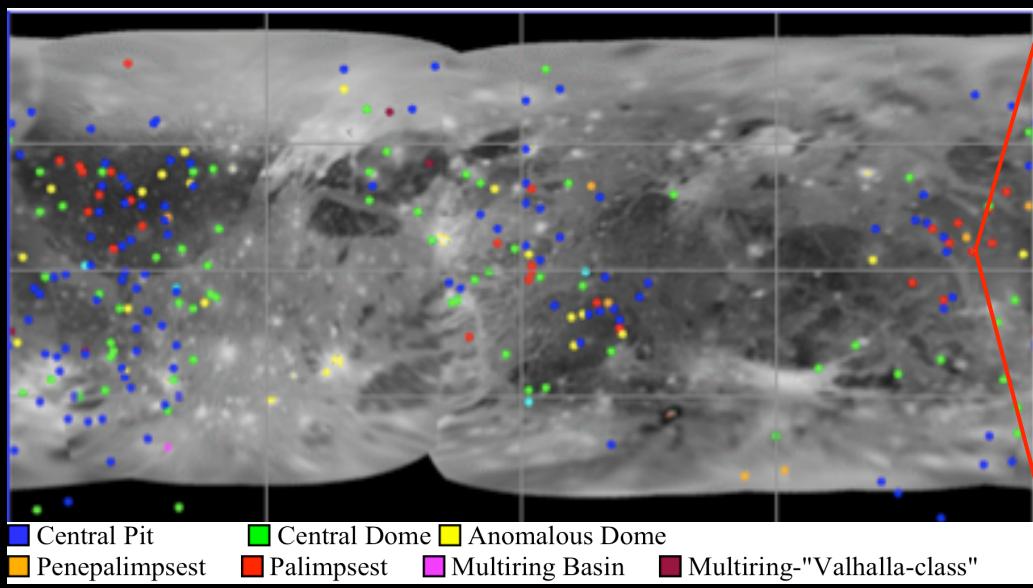


Science questions

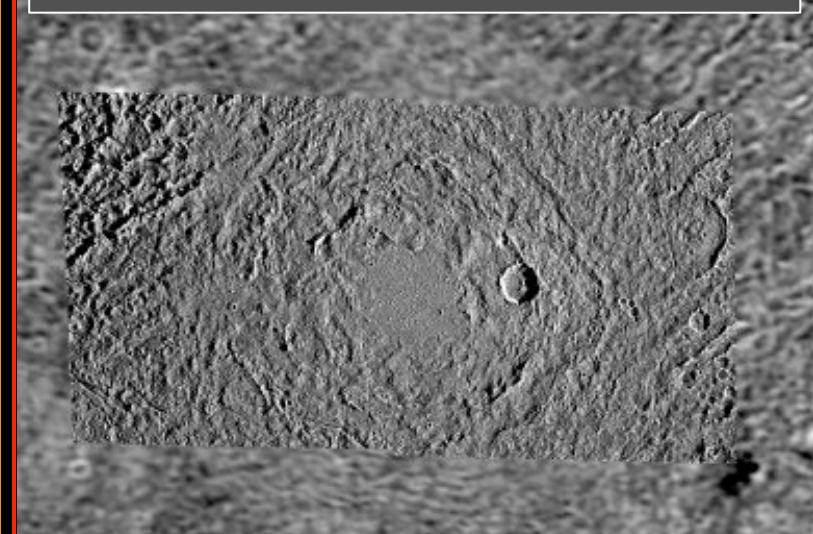
- ✧ Impacts: description, processes, chronology
- ✧ Internal structure: investigation, implications
- ✧ Energy transfer to the habitable zone

Impact cratering: description, processes, chronology

Ganymede possesses the widest range in crater morphology



Analysis requires high-resolution imaging



JGO measurements

- Global imaging at 200-400 m/px
- High Resolution target areas (5-50 m/px)
- Compositional relationships
- Subsurface exploration

Instrument Packages

- Imaging
- Spectroscopy
- Sounders

JGO

2026

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Energy for deep habitats ?

Induced magnetic field

Surface motions

Icy crust and sub-surface ocean

Degree 2 g-field + tidal motions

Density profile and equilibrium state

High degree g-field

Density anomalies

Intrinsic magnetic field

Metallic core properties



Instrument Packages

- In situ Fields and Particles
- Imaging
- Sounders & Radio Science

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JGO

JOI

Ganymede flybys

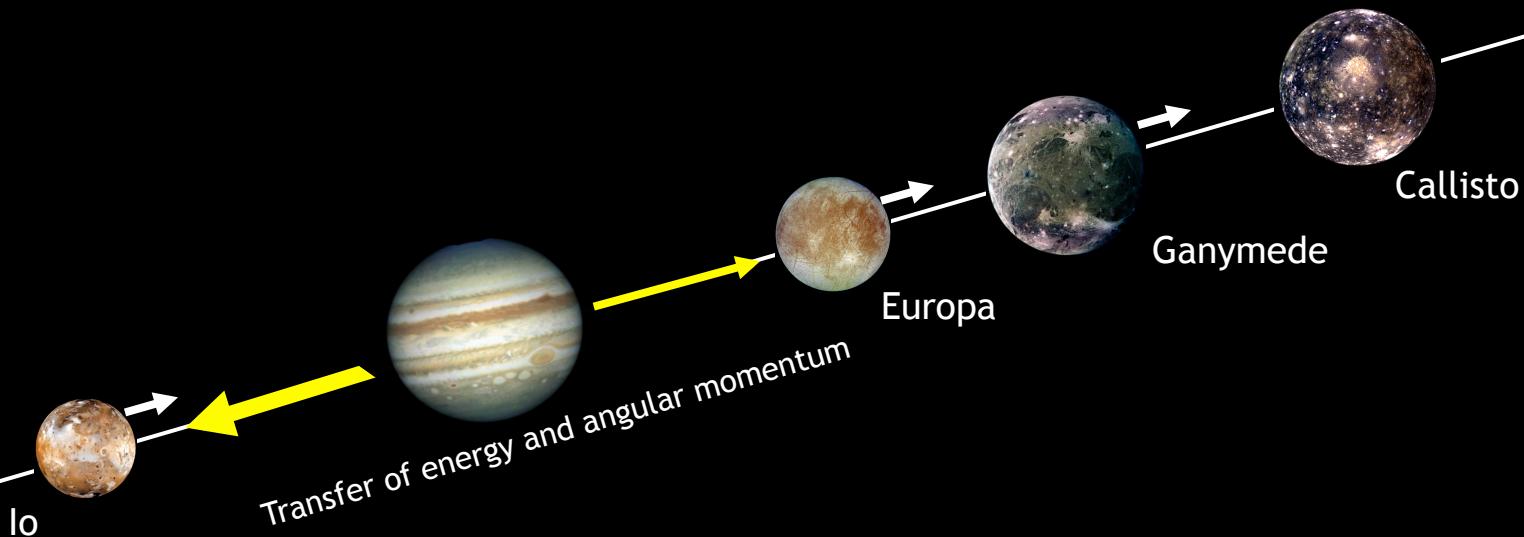
GEO

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How stable are the present habitats ?

Laplace resonance evolution



Science questions

- ✧ System stability: Laplace resonance

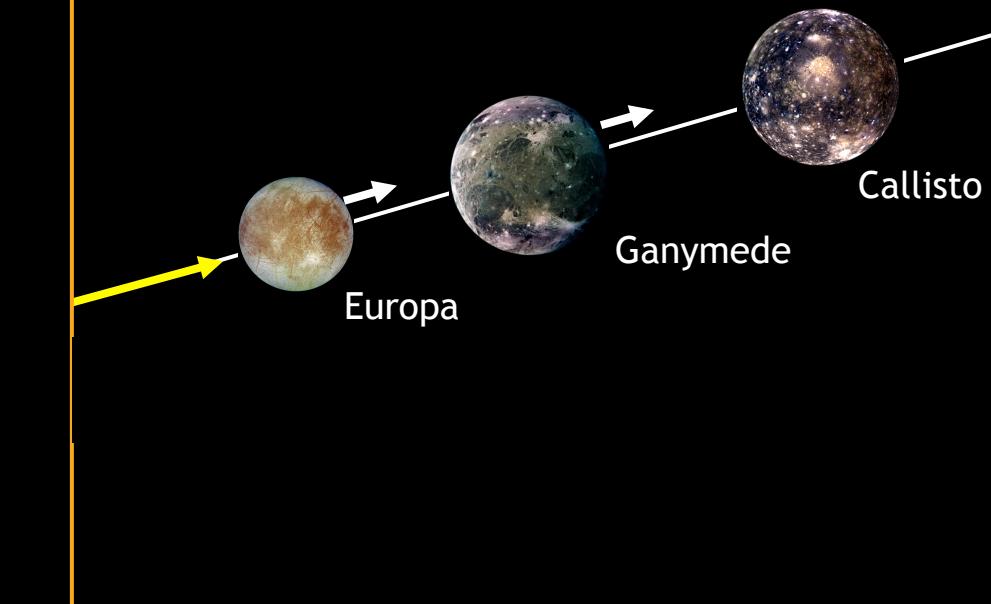


How stable are the present habitats ?

Geology as a witness of moon's evolution



Laplace resonance evolution

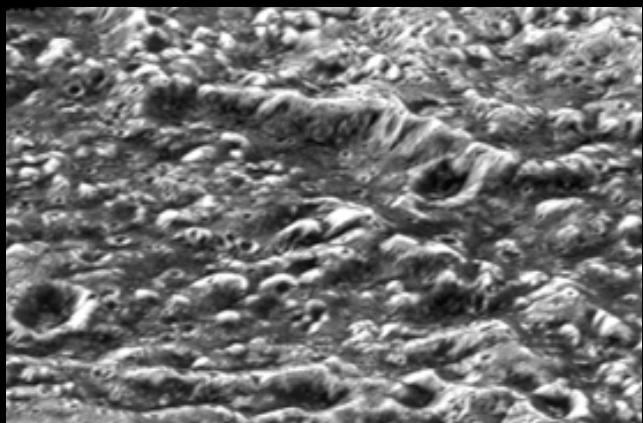


Science questions

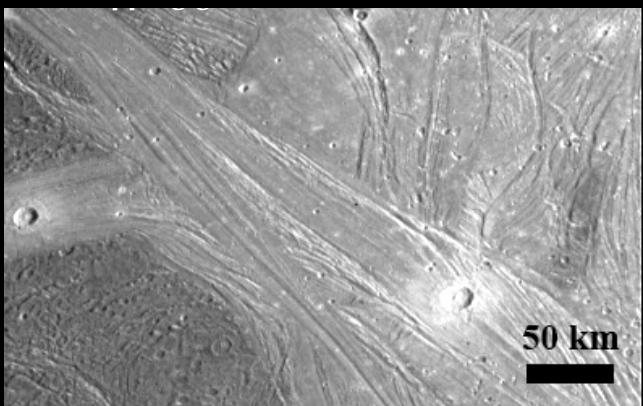
- ✧ System stability: Laplace resonance
- ✧ Moons' histories: Geologic evidences



Stability of the habitats: geology as a witness of moon's activity



Dark terrain: ancient surface; furrow tectonics



Bright terrain: Rift-like tectonism, icy volcanism



JGO measurements

- Global imaging at 200-400 m/px
- High Resolution target areas
- Topography/ morphology
- Subsurface exploration
- Compositional relationships

Instrument Packages

- Imaging
- Spectroscopy
- Sounders

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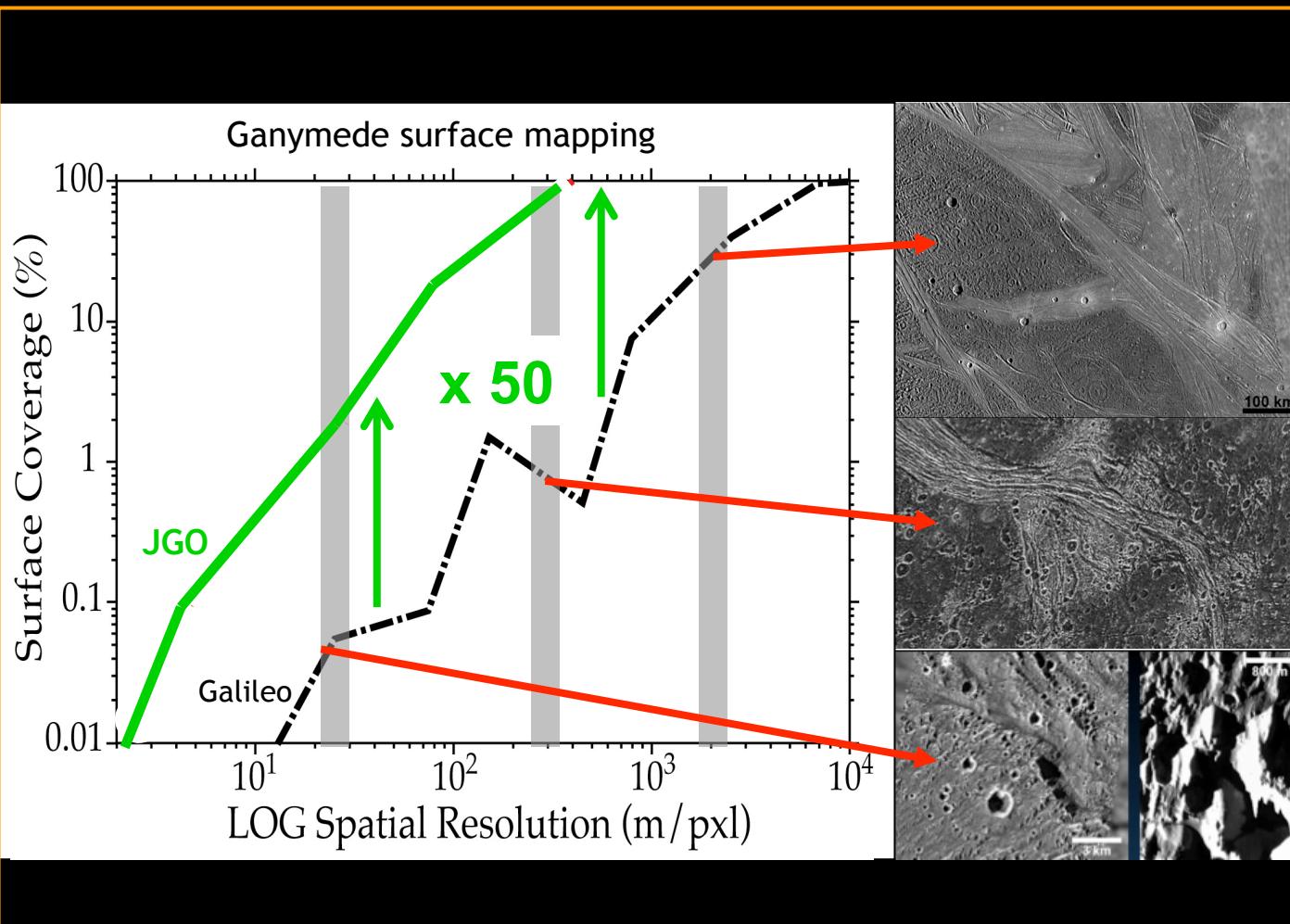


GCO



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Stability of the habitats: geology as a witness of moon's activity



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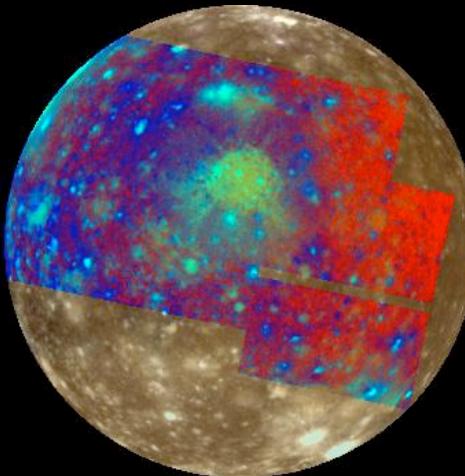


Callisto: a witness of early ages

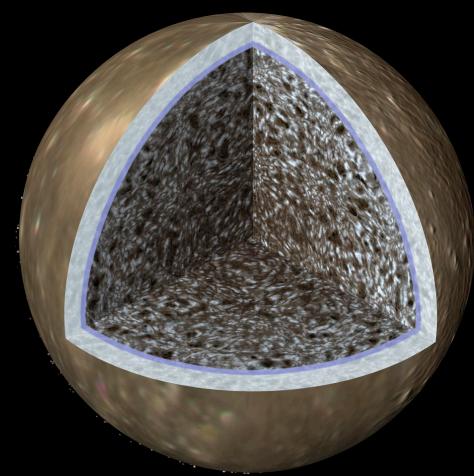
Oldest surface of the system



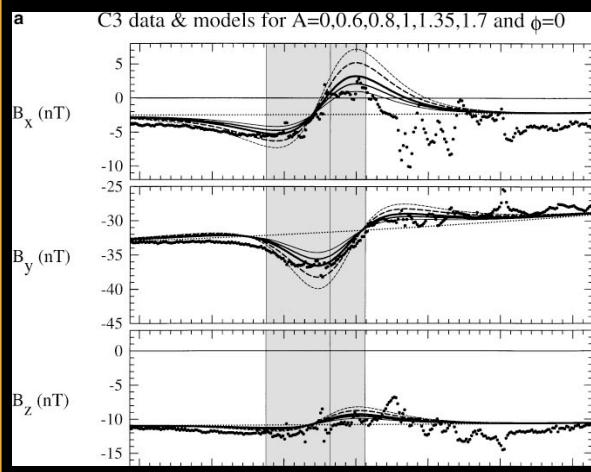
Complex chemistry



Not totally differentiated



An induced magnetic field

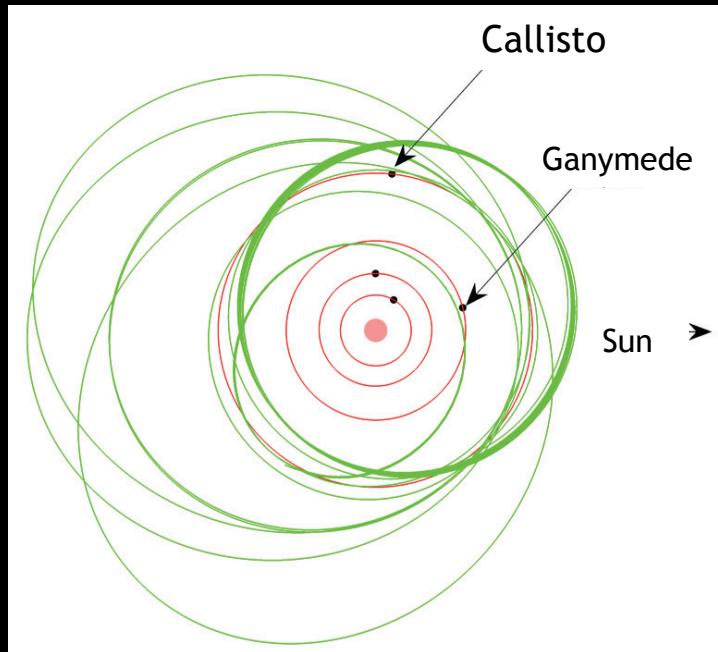


Science questions

- ❖ Presence and extent of a subsurface ocean
- ❖ Degree of differentiation
- ❖ Cratering record and early geological history
- ❖ Surface composition including organics and carbon dioxide

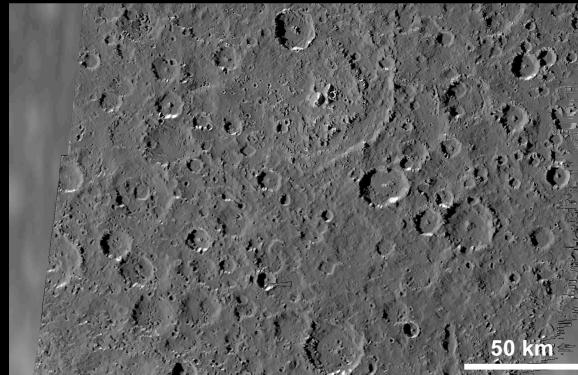
The oldest surface of the jovian system will be deciphered

Callisto phase

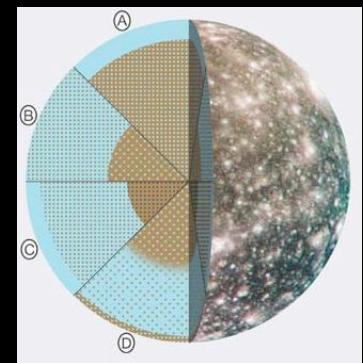


Science objectives

Cratering record and early geological history

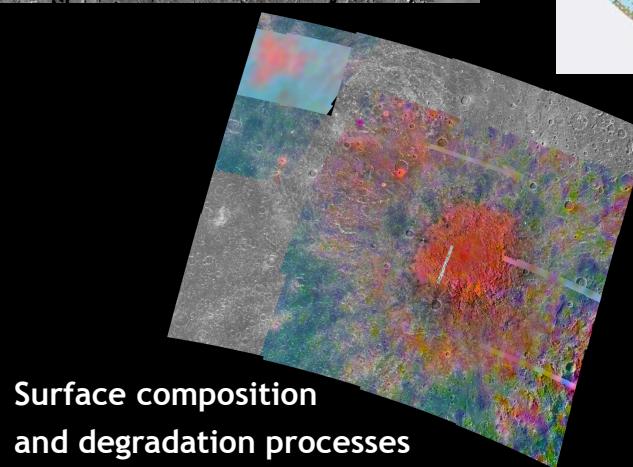


Structure and internal differentiation

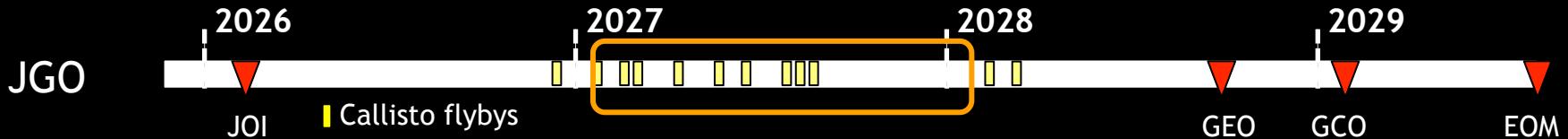


Mission profiles

- Dedicated close flybys for JGO
- Close flybys opportunities for JEO
- JGO / JEO complementarities



Surface composition
and degradation processes

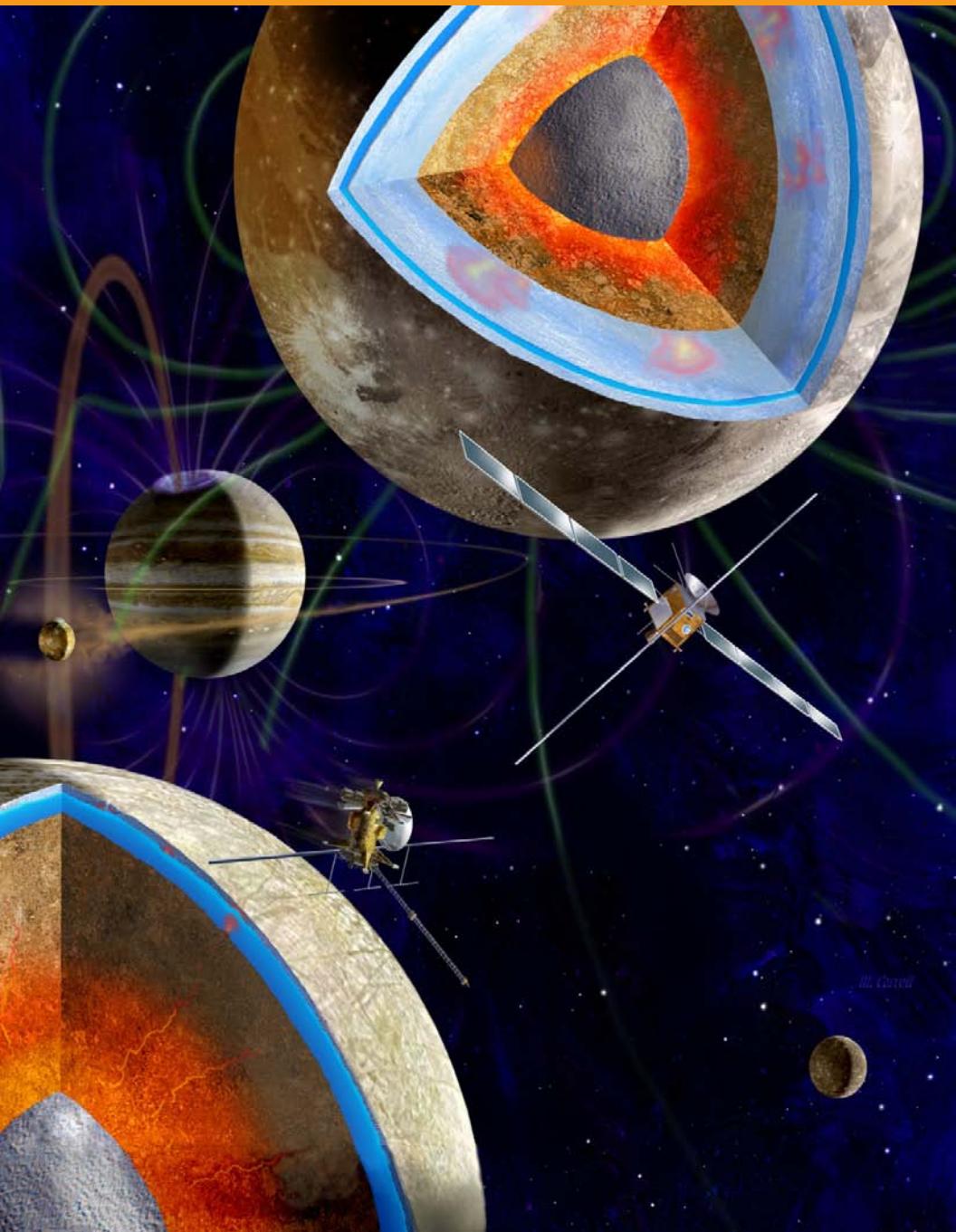


From the objectives to the measurements

	INSTRUMENT	LIQUID	CHEMISTRY	ENERGY	TIME
In situ Fields and Particles	Magnetometer (MAG)	X	X		
	Radio and Plasma Wave Instr. (RPWI)	X	X		
	Particle and Plasma Instr.– Ion Neutral Mass Spectr. (PPI-INMS)	X	X		
Imaging	Narrow Angle Camera (NAC)	X	X	X	X
	Wide Angle Camera (WAC)	X	X	X	X
Spectroscopy	Visible Infrared Hyperspectral Imaging Spectrometer (VIRHIS)		X	X	X
	UV Imaging Spectrometer (UVIS)		X	X	
	Sub-mm Wave Instrument (SWI)		X		X
Sounders & Radio Science	Laser Altimeter (LA)	X		X	X
	Ice Penetrating Radar (IPR)		X	X	X
	Radio Science Instrument (JRST+USO)	X		X	X

Jupiter System

Emma Bunce



The Jupiter system

Archetype for gas giants

Solar system in miniature

Paradigm for exoplanetary systems

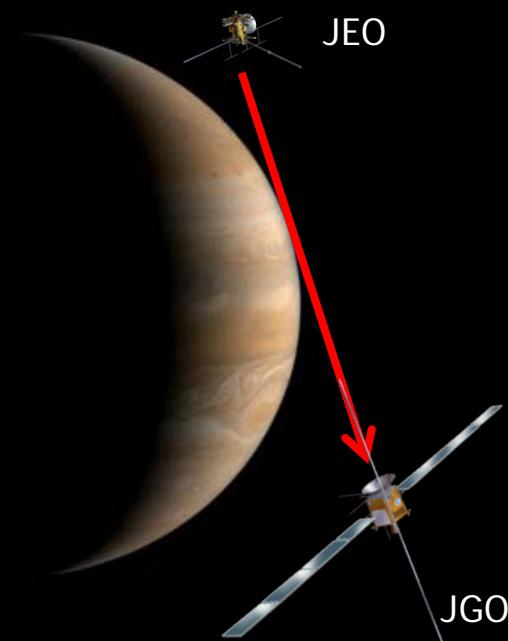
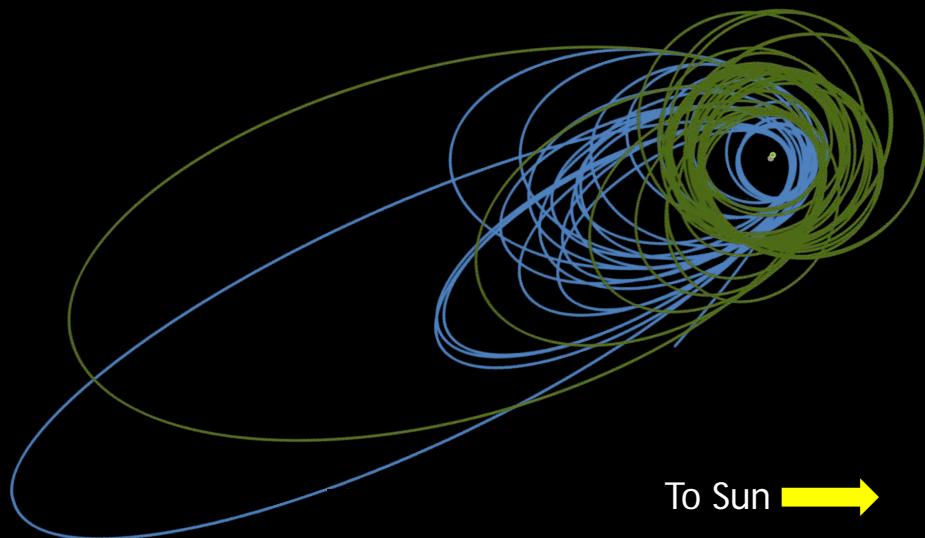
EJSM-Laplace represents...

FIRST two-spacecraft mission to
a gas giant system

Significant advance in
our knowledge

A major opportunity for
complementary and synergistic science

Complementary and Synergistic Science: Where JEO and JGO unite



Complementary - two spacecraft extend the coverage of the system

- ✓ two spacecraft in orbit around Jupiter and/or a moon at the SAME TIME
- ✓ trajectory of two spacecraft extends coverage in LOCAL TIME/LATITUDE

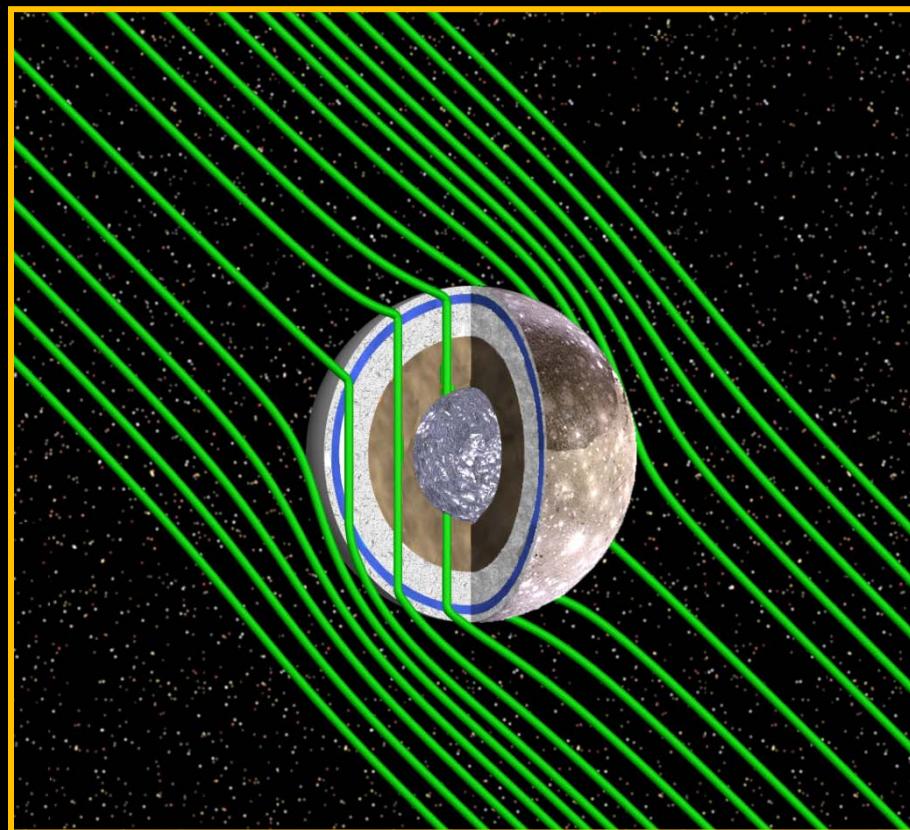
Synergistic - two spacecraft together provide new opportunities

- ✓ spacecraft-to-spacecraft link provided by receiving capability on JGO
- ✓ dual observations of spatially and temporally varying phenomena

GANYMEDE
ENVIRONMENT

JUPITER
MAGNETOSPHERE

JUPITER
ATMOSPHERE



Previous Exploration

- ✓ *Induced* magnetic field → sub-surface (conducting) ocean?
- ✓ *Intrinsic* dipole magnetic field
- ✓ *Mini-magnetosphere* is created
- ✓ Interaction is complex & highly variable



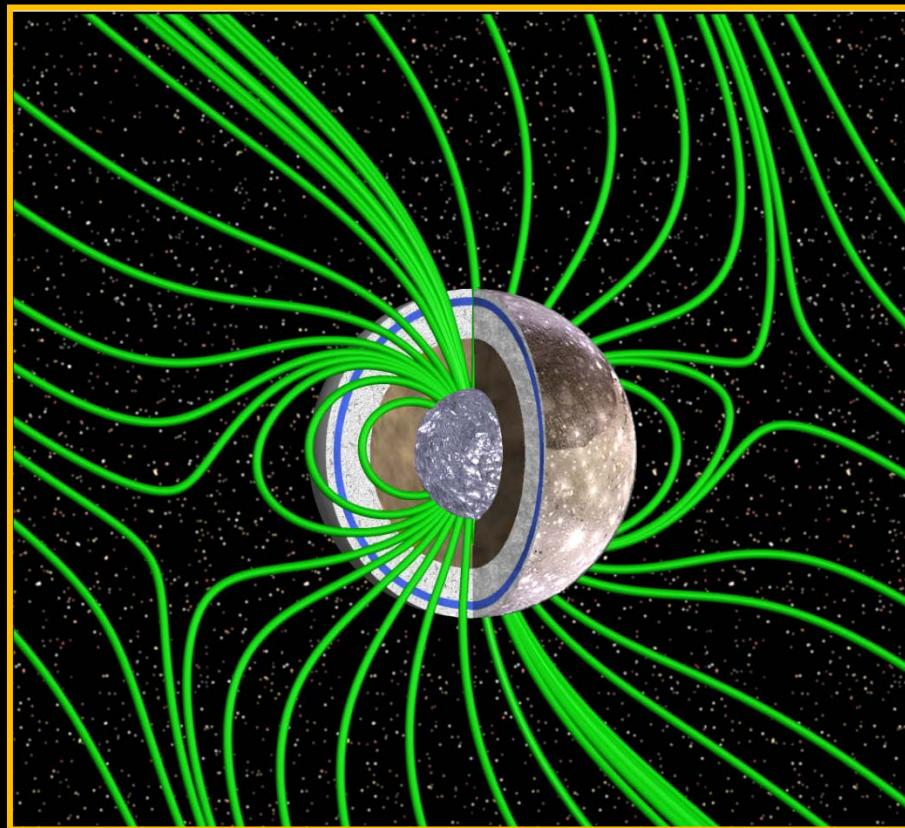
Science Topics:

Ganymede's dipole magnetic field and mini-magnetosphere
Coupling to Jupiter's magnetosphere
Local plasma environment effects

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ENVIRONMENT

JUPITER
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JUPITER
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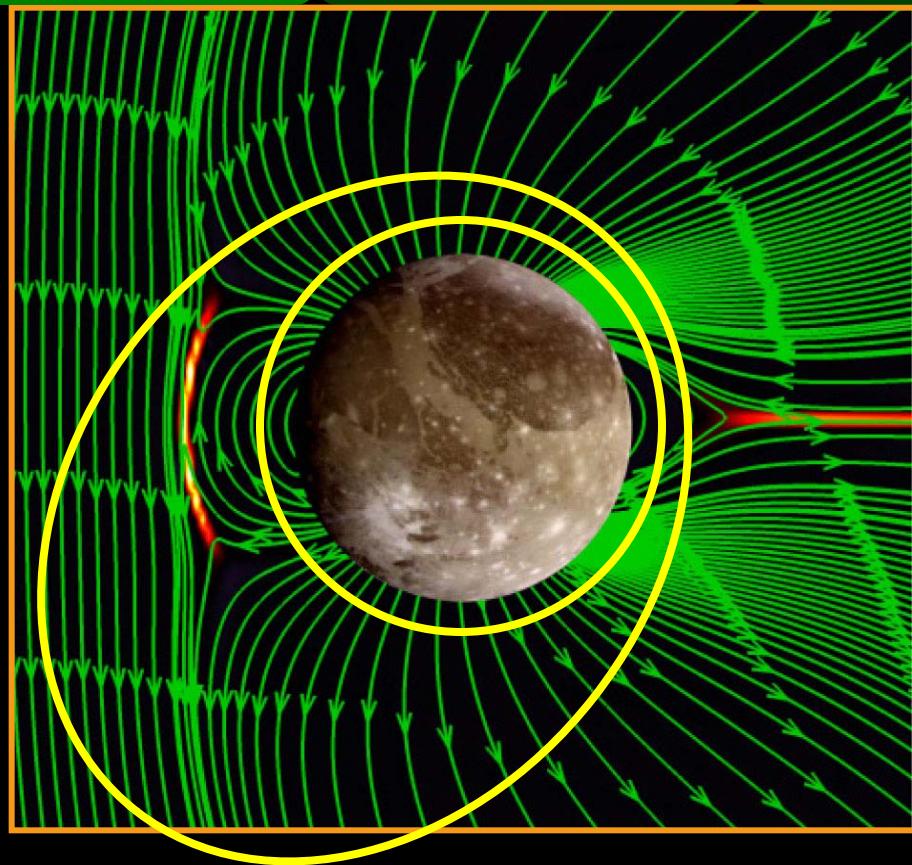
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JUPITER
MAGNETOSPHERE

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Ganymede's dipole magnetic field and mini-magnetosphere
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Local plasma environment effects

The Jupiter System

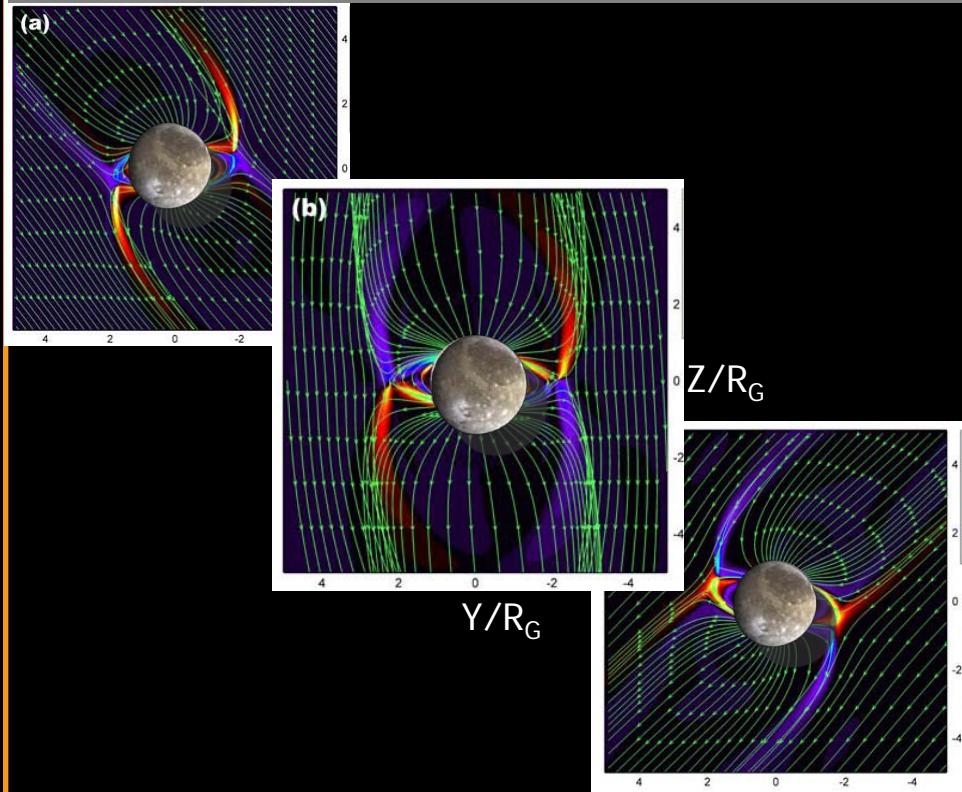
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ENVIRONMENT

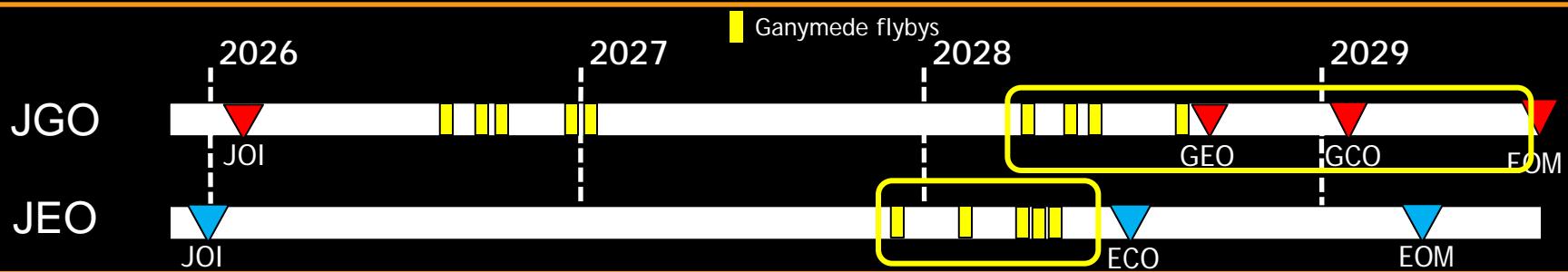
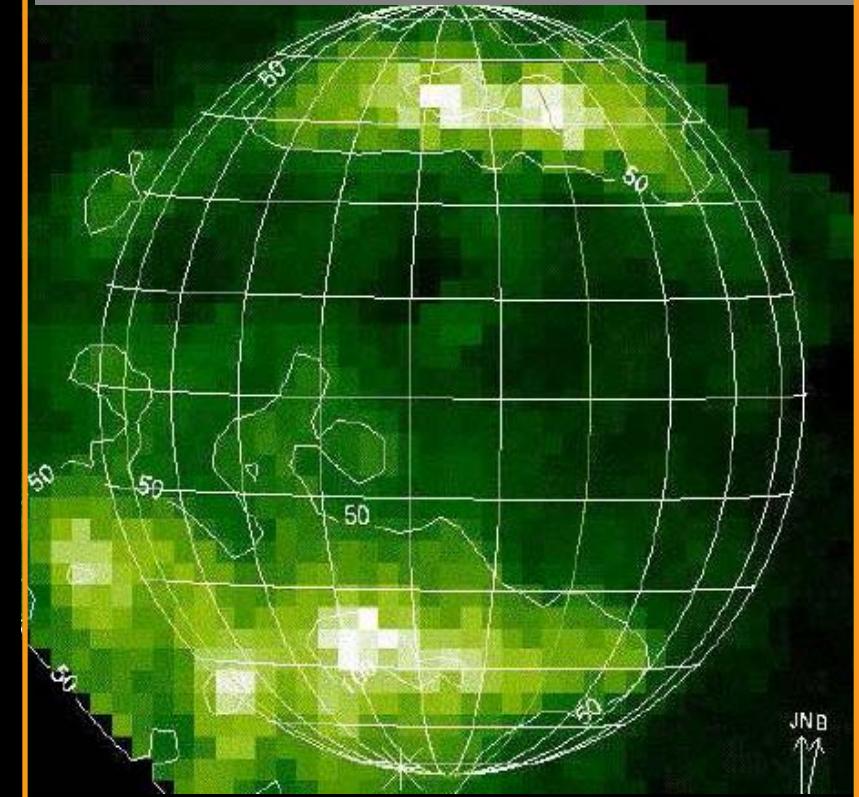
JUPITER
MAGNETOSPHERE

JUPITER
ATMOSPHERE

Dipole magnetic field and mini-magnetosphere



Coupling to Jupiter's magnetosphere



The Jupiter System

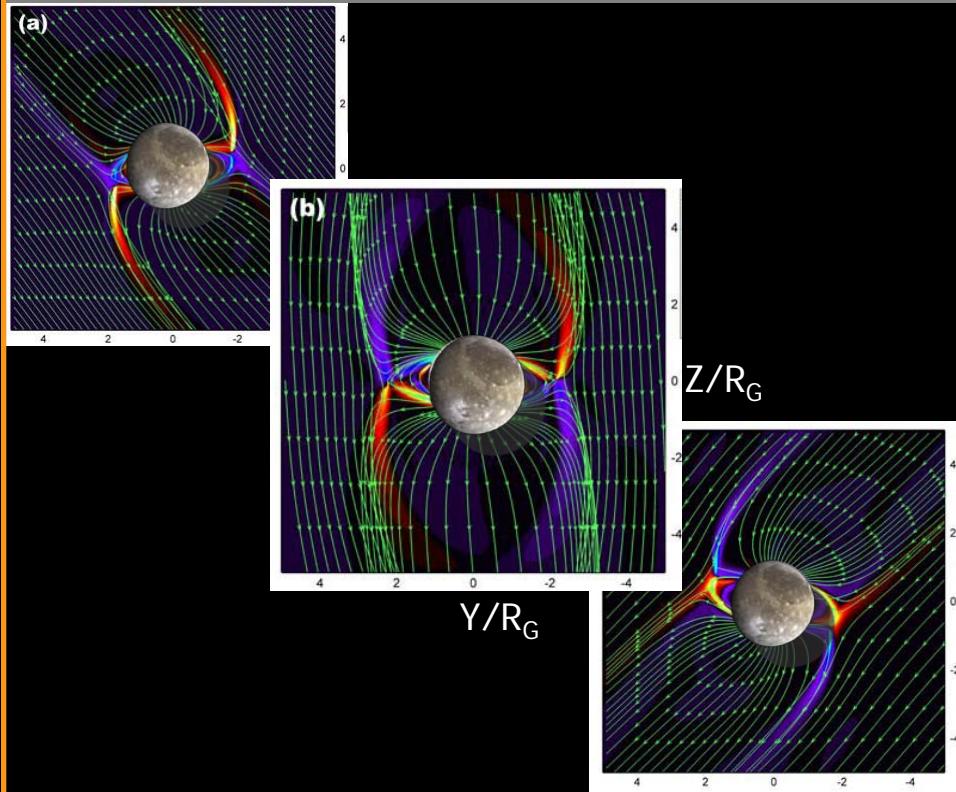
EJSM-Laplace

GANYMEDE
ENVIRONMENT

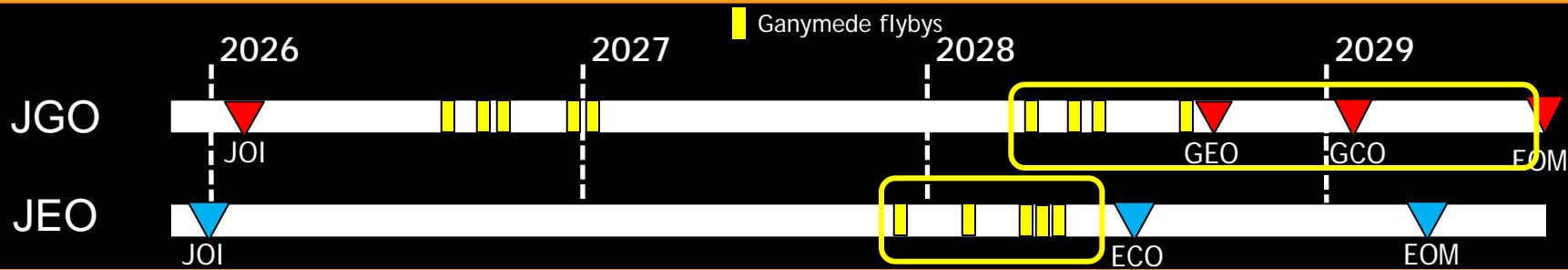
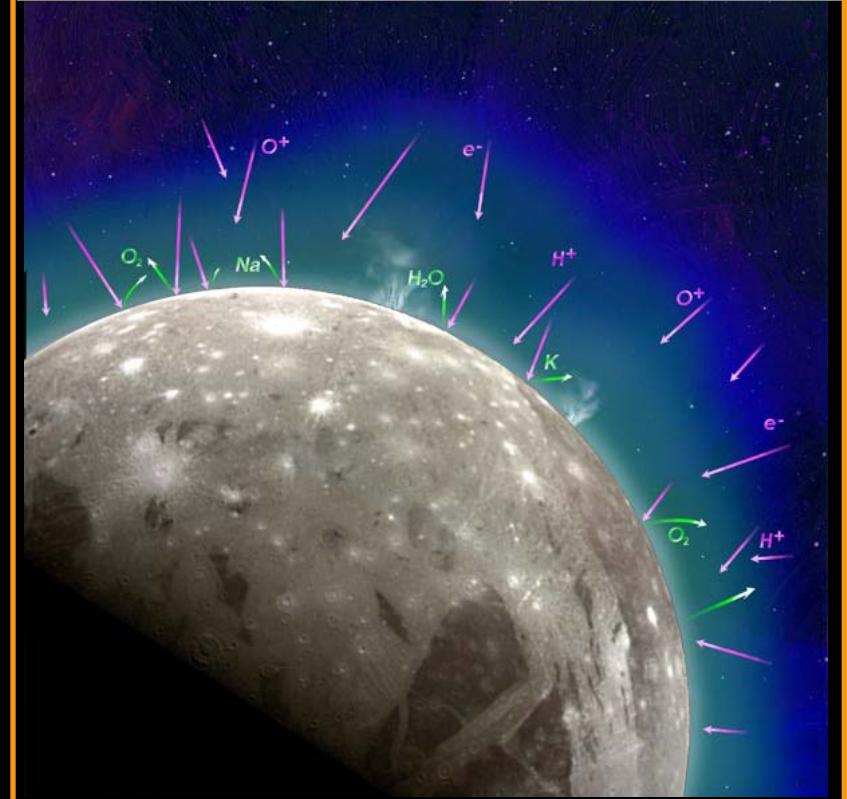
JUPITER
MAGNETOSPHERE

JUPITER
ATMOSPHERE

Dipole magnetic field and mini-magnetosphere



Sources and sinks of atmosphere



The Jupiter System

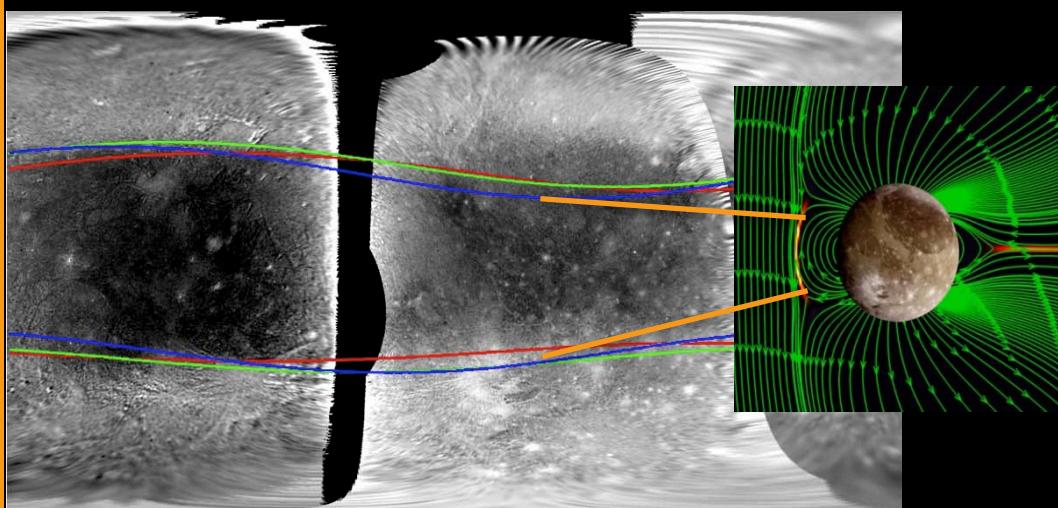
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GANYMEDE
ENVIRONMENT

JUPITER
MAGNETOSPHERE

JUPITER
ATMOSPHERE

Local plasma environment effects



INSTRUMENT PACKAGES

In situ Fields and Particles

Imaging

Spectroscopy

Radio science

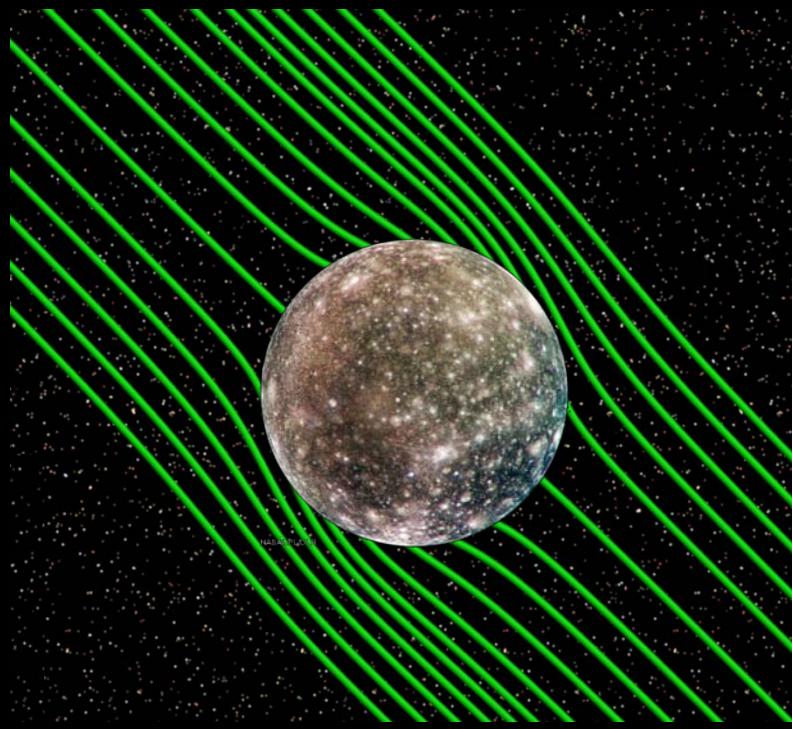


CALLISTO
ENVIRONMENT

JUPITER
MAGNETOSPHERE

JUPITER
ATMOSPHERE

Local plasma environment effects



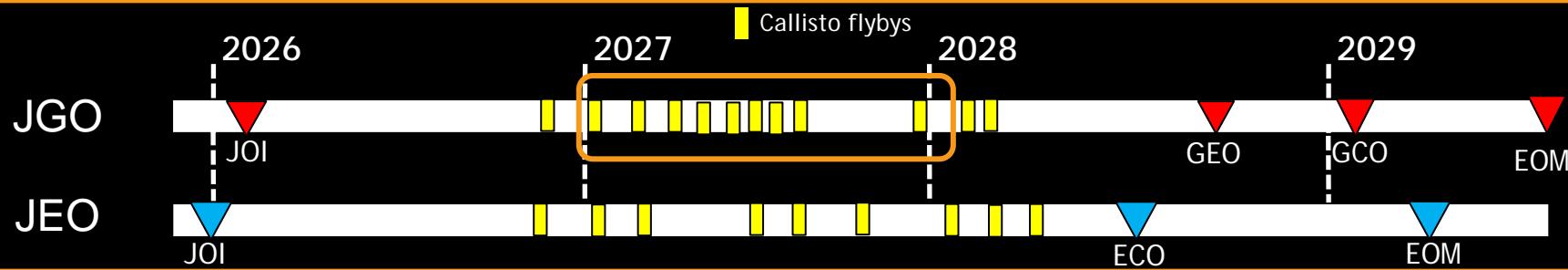
INSTRUMENT PACKAGES

In situ Fields and Particles

Imaging

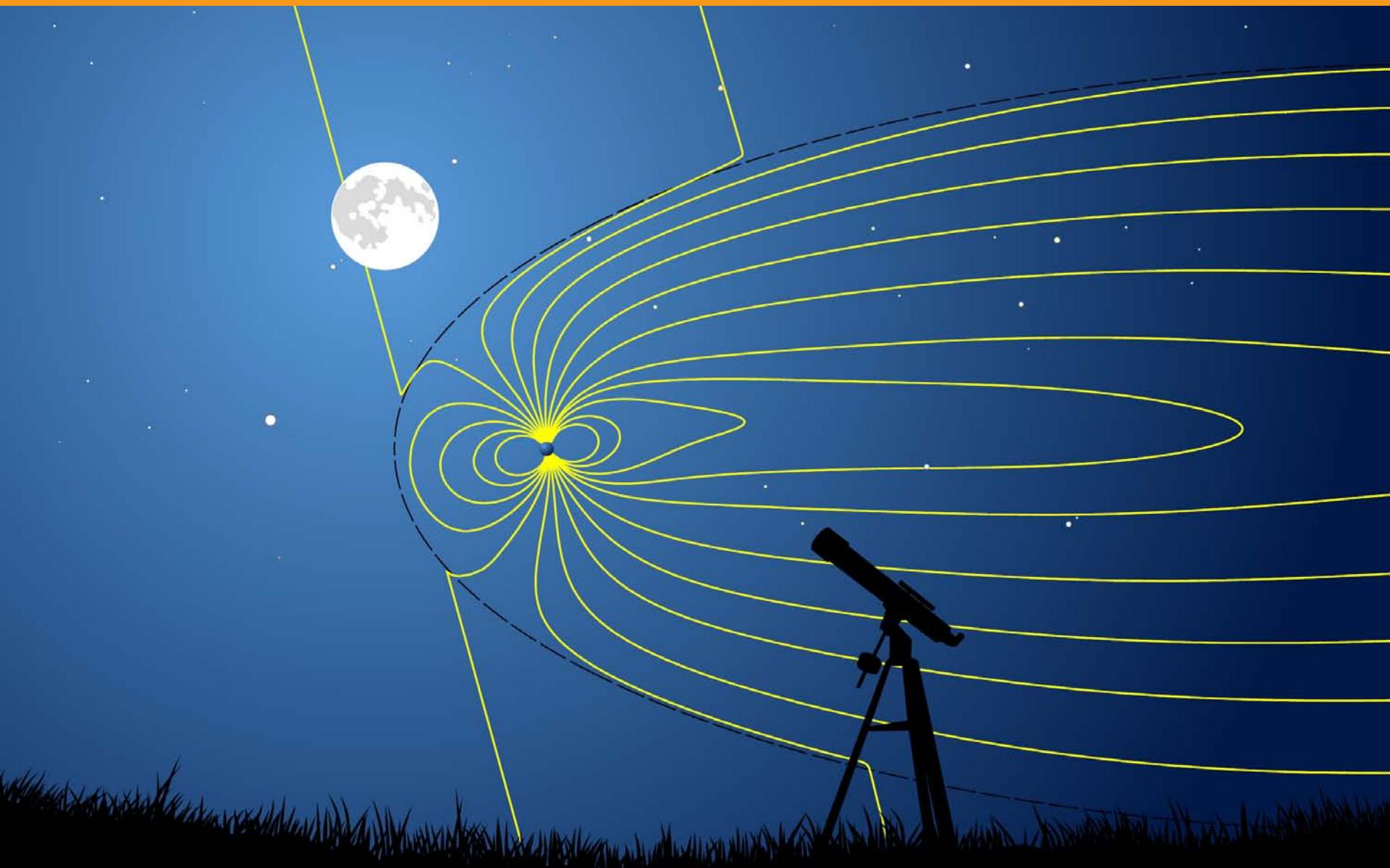
Spectroscopy

Radio science



The Jupiter System

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MOON
ENVIRONMENTS

JUPITER
MAGNETOSPHERE

JUPITER
ATMOSPHERE



Previous missions (1979-2007)

- Brief fly-through observations
- Galileo completed a survey

Previous
Exploration

Future
Exploration



Complementary to Juno's (2016+) investigations

- Close mapping of Jovian internal field
- Polar magnetosphere

Science Topics:

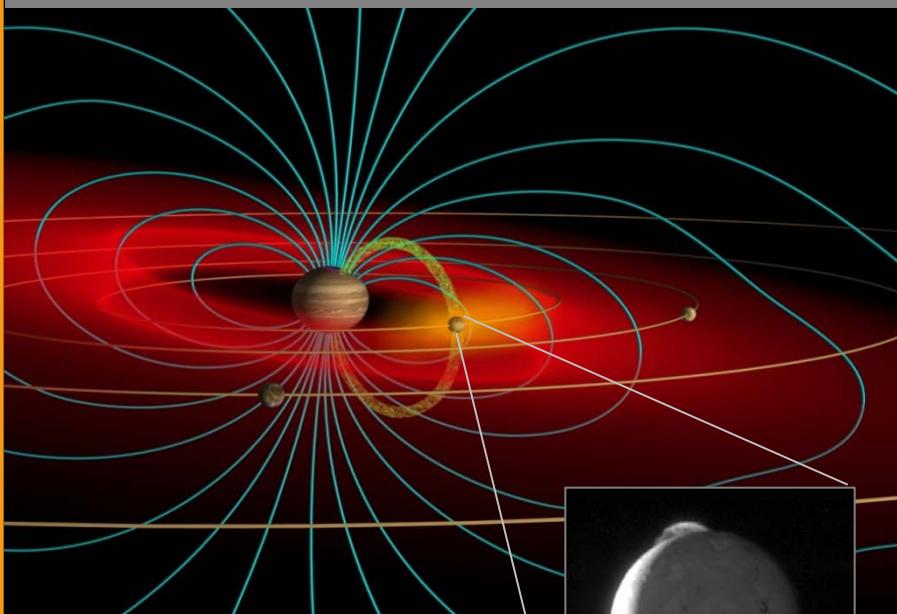
Properties of the magnetodisc
Magnetospheric driving; Energy & Momentum transfer
Particle acceleration and loss
Moon-magnetosphere interactions

MOON
ENVIRONMENTS

JUPITER
MAGNETOSPHERE

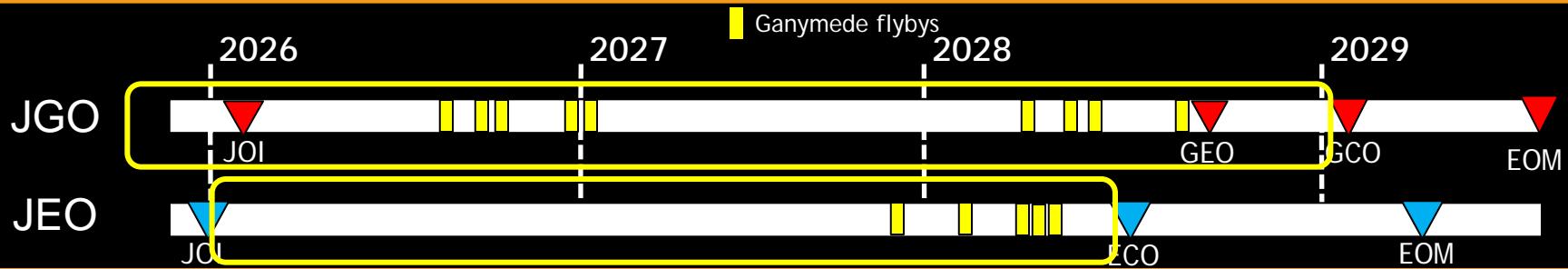
JUPITER
ATMOSPHERE

Properties of the magnetodisc



Energy transfer in the coupled system

UV aurora - main oval and moon footprints

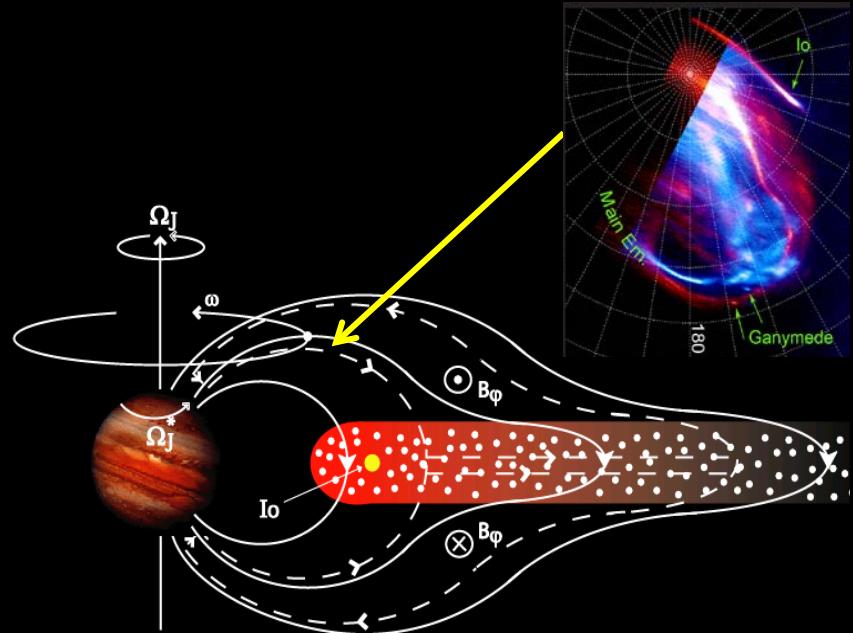


MOON
ENVIRONMENTS

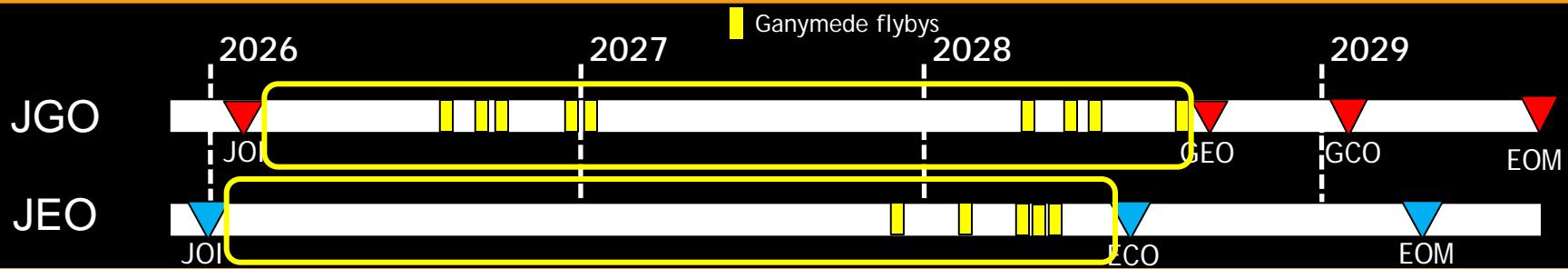
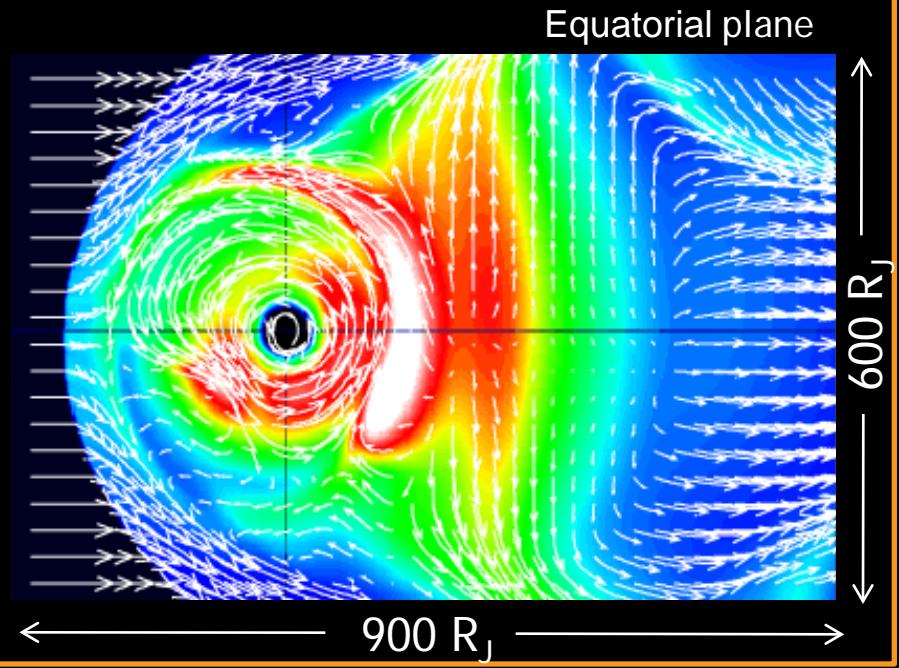
JUPITER
MAGNETOSPHERE

JUPITER
ATMOSPHERE

Particle acceleration



Mass loss processes

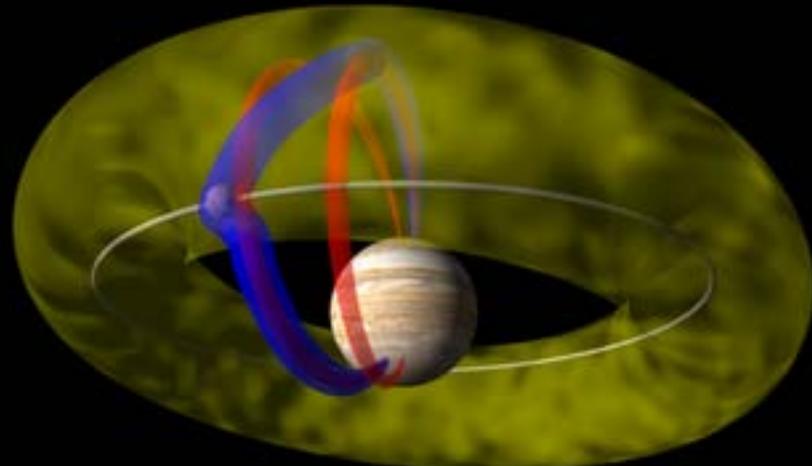


MOON
ENVIRONMENTS

JUPITER
MAGNETOSPHERE

JUPITER
ATMOSPHERE

Moon-magnetosphere interactions



Io-Jupiter interaction

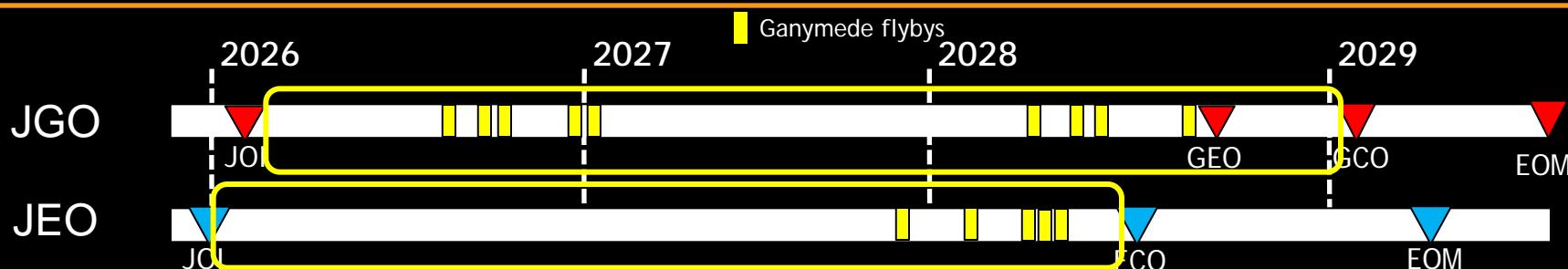
INSTRUMENT PACKAGES

In situ Fields and Particles

Imaging

Spectroscopy

Radio science



MOON
ENVIRONMENTS

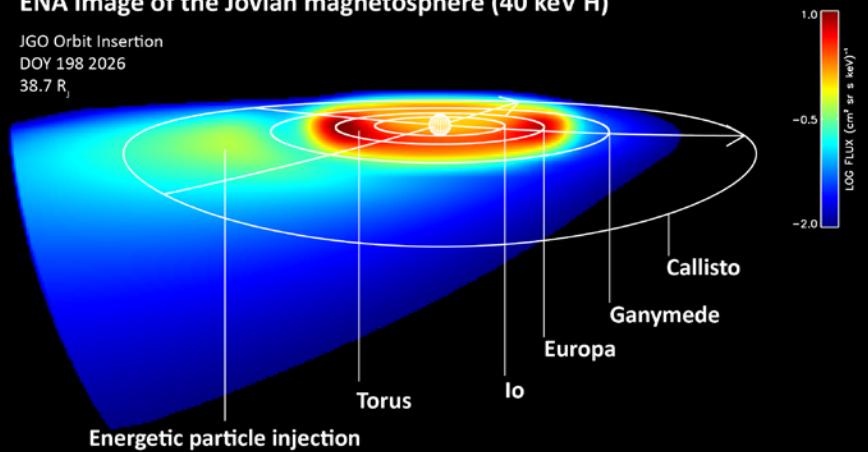
JUPITER
MAGNETOSPHERE

JUPITER
ATMOSPHERE

Moon-magnetosphere interactions

ENA image of the Jovian magnetosphere (40 keV H)

JGO Orbit Insertion
DOY 198 2026
38.7 R_J



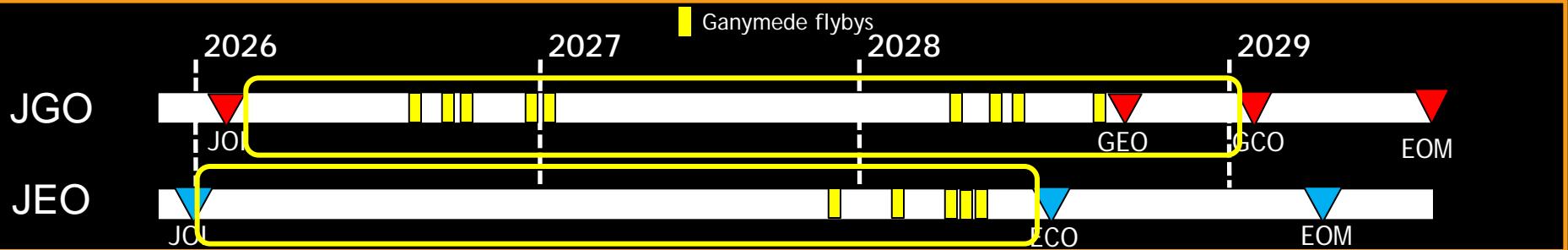
INSTRUMENT PACKAGES

In situ Fields and Particles

Imaging

Spectroscopy

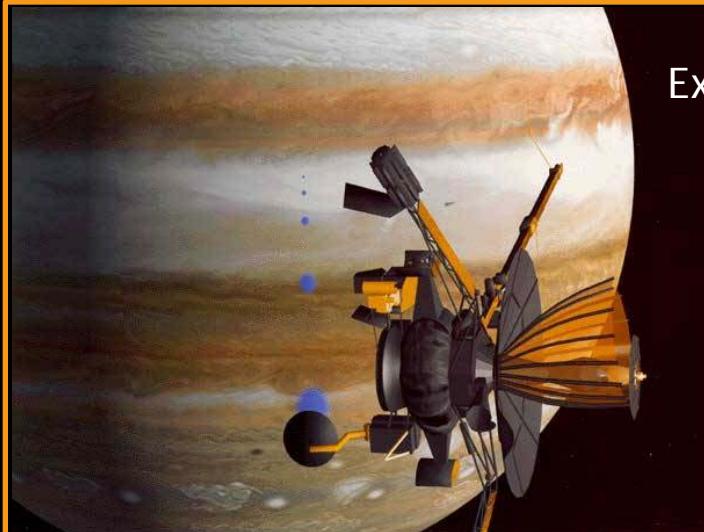
Radio science



MOON
ENVIRONMENTS

JUPITER
MAGNETOSPHERE

JUPITER
ATMOSPHERE

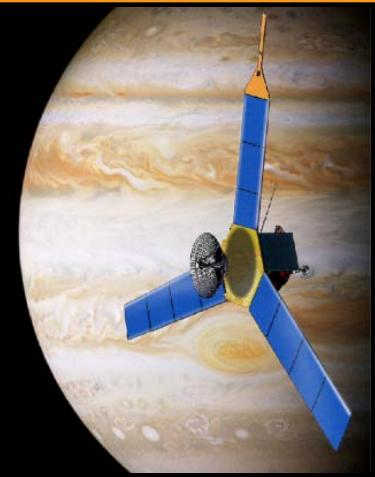


Previous
Exploration

Previous missions (1979-2007)

- Snap-shot observations
- Probe findings (local measurements)

Future
Exploration



Complement and extend Juno's (2016+) investigation of deep interior

- Microwave studies of volatiles in deep interior
- Gravity mapping of Jupiter

Science Topics:

Global atmosphere dynamics and variability
Jupiter's bulk composition
Transport of energy and material tracers
Vertical coupling

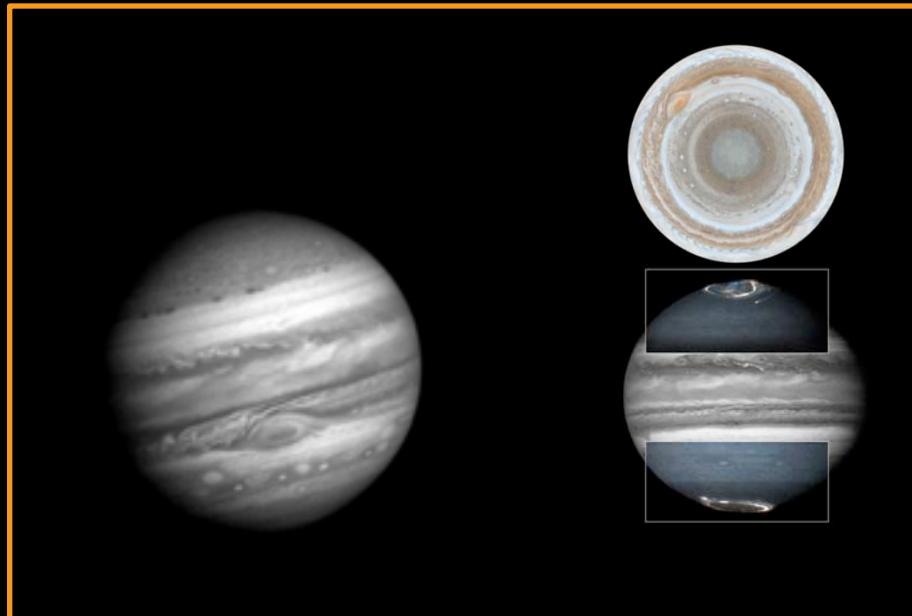
MOON
ENVIRONMENTS

JUPITER
MAGNETOSPHERE

JUPITER
ATMOSPHERE

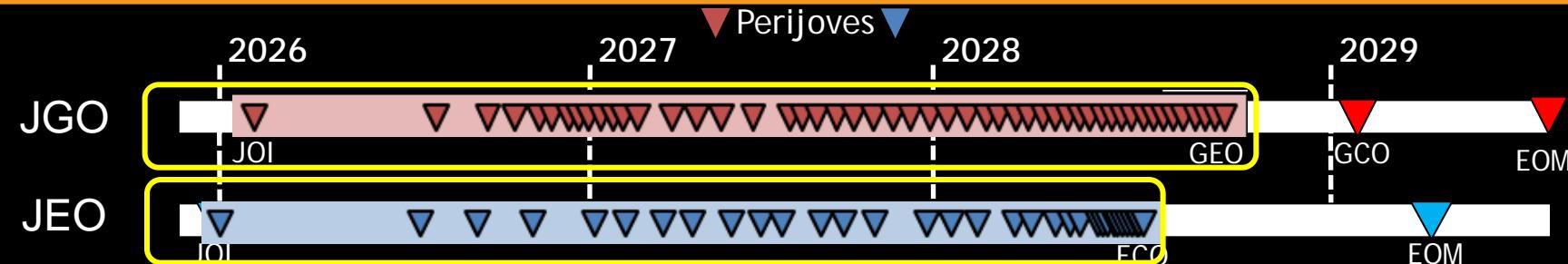
Dynamics and Circulation

- Winds and thermodynamics
- Lightning studies
- Wave propagation
- Vertical coupling
- Storms, vortices, and instabilities
- Variability of circulation



INSTRUMENT PACKAGES

- Imaging
- Spectroscopy
- Radio Science



MOON
ENVIRONMENTS

JUPITER
MAGNETOSPHERE

JUPITER
ATMOSPHERE

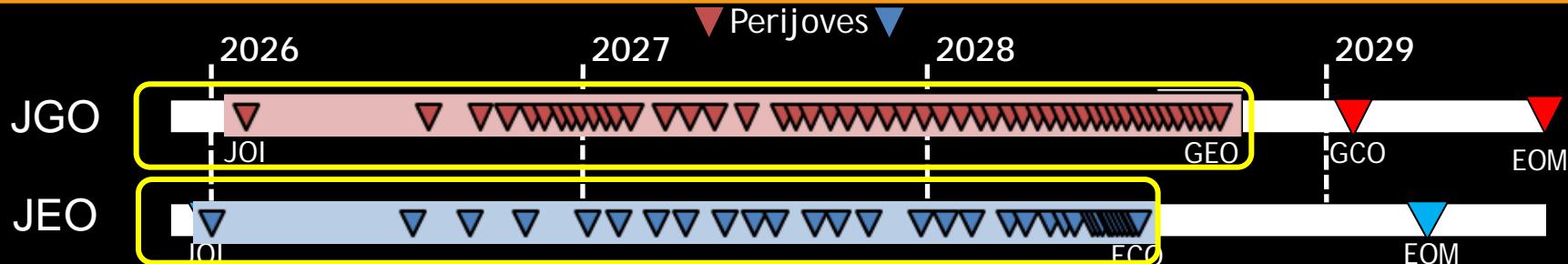


Composition and Chemistry

Distribution of condensable volatiles
Tracers of weather-layer dynamics
Hydrocarbon chemistry
Stratospheric trace species

INSTRUMENT PACKAGES

Imaging
Spectroscopy
Radio Science



MOON
ENVIRONMENTS

JUPITER
MAGNETOSPHERE

JUPITER
ATMOSPHERE

Gravity waves in Jupiter's atmosphere

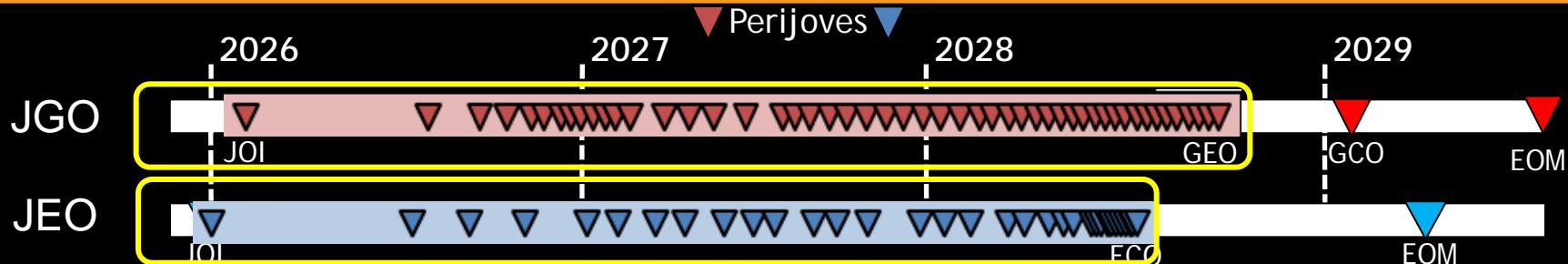


Vertical Structure and Interior

Properties of cloud and haze layers
 Vertical aerosol distribution
 Coupling of atmosphere to deep interior and external environment

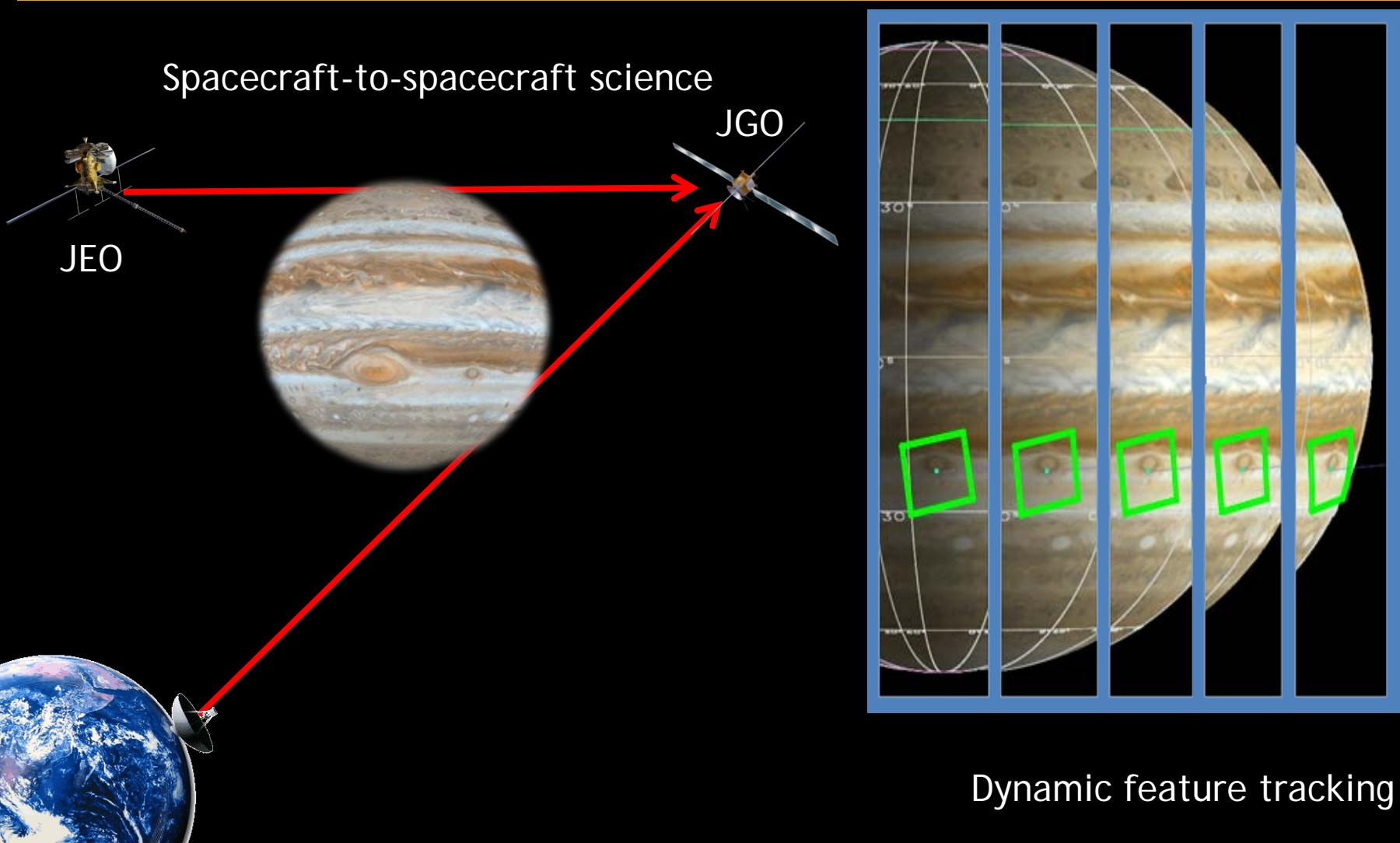
INSTRUMENT PACKAGES

Imaging
 Spectroscopy
 Radio Science



Complementary and synergistic atmosphere studies

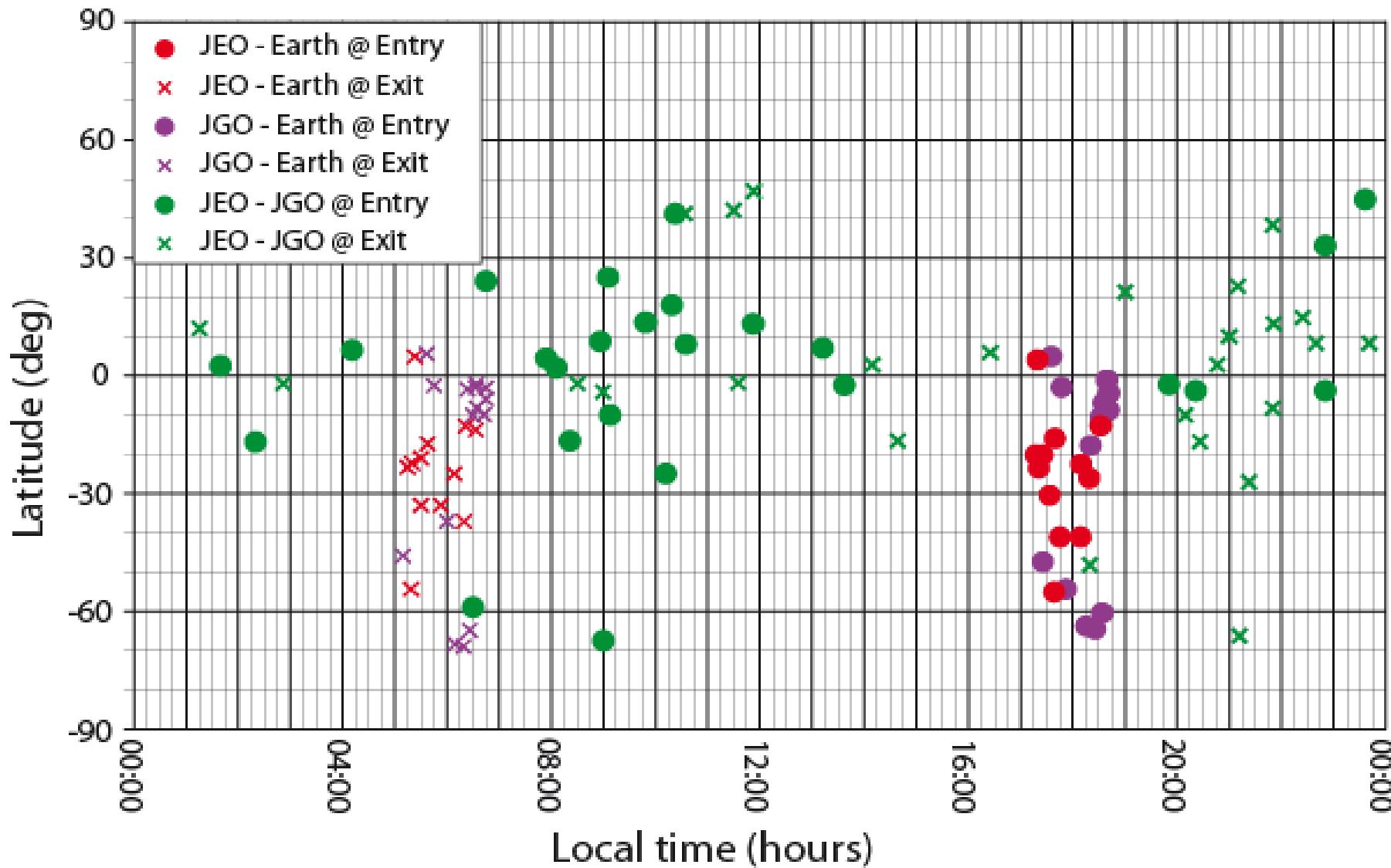
- ✓ Radio science occultations of Jupiter/satellite atmospheres = Increased opportunities
- ✓ Dual observations of atmospheric phenomena
- ✓ Enhanced flexibility with specialised instrumentation and wavelength coverage



Complementary and synergistic atmosphere studies

- ✓ Radio science occultations of Jupiter/satellite atmospheres = Increased opportunities

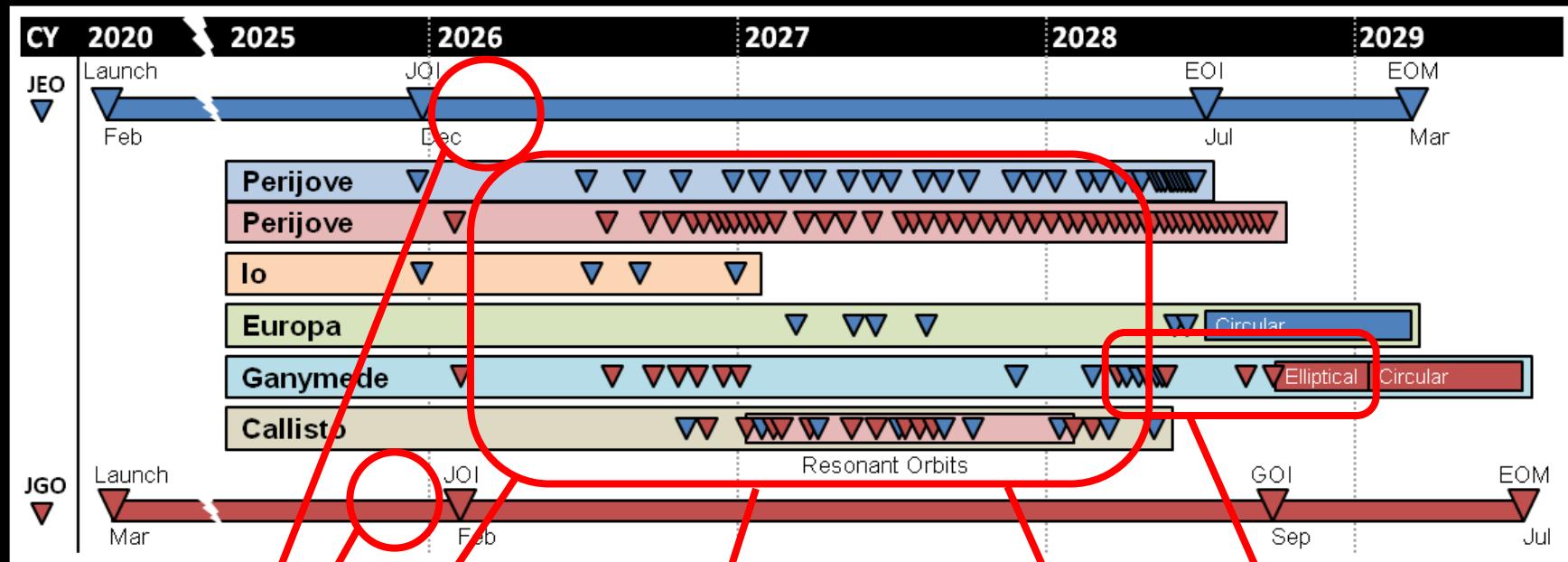
Spacecraft-to-spacecraft Radio Science Opportunities



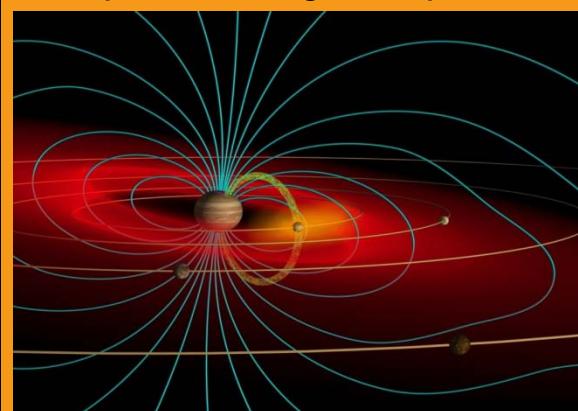
The Jupiter System

EJSM-Laplace

Complementary and Synergistic Science: Where JEO and JGO unite



Jupiter's Magnetosphere



Jupiter Science



Satellite Science



Ganymede Magnetosphere Studies



From the objectives to the measurements

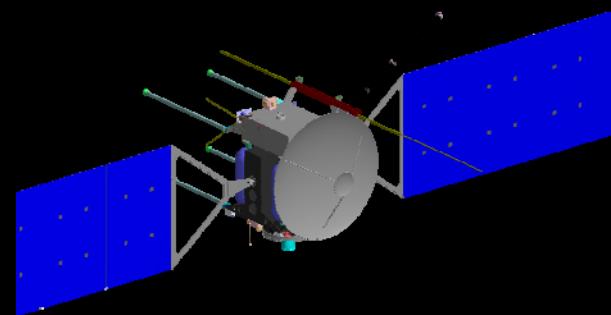
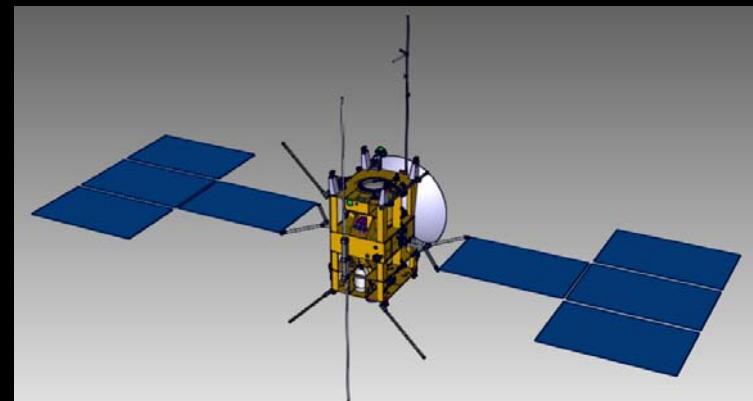
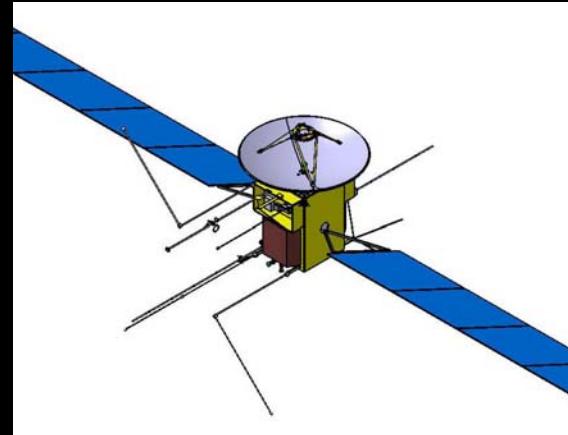
	INSTRUMENT	MOON ENVIRONMENTS	JUPITER MAGNETOSPHERE	JUPITER ATMOSPHERE
In situ Fields and Particles	Magnetometer (MAG)	X	X	
	Radio and Plasma Wave Instr. (RPWI)	X	X	
	Particle and Plasma Instr.— Ion Neutral Mass Spectr. (PPI-INMS)	X	X	
Imaging	Narrow Angle Camera (NAC)	X	X	X
	Wide Angle Camera (WAC)	X	X	X
Spectroscopy	Visible Infrared Hyperspectral Imaging Spectrometer (VIRHIS)	X	X	X
	UV Imaging Spectrometer (UVIS)	X	X	X
	Sub-mm Wave Instrument (SWI)	X		X
Sounders & Radio Science	Laser Altimeter (LA)			
	Ice Penetrating Radar (IPR)			
	Radio Science Instrument (JRST+USO)	X		X

JGO Mission Implementation

Christian Erd

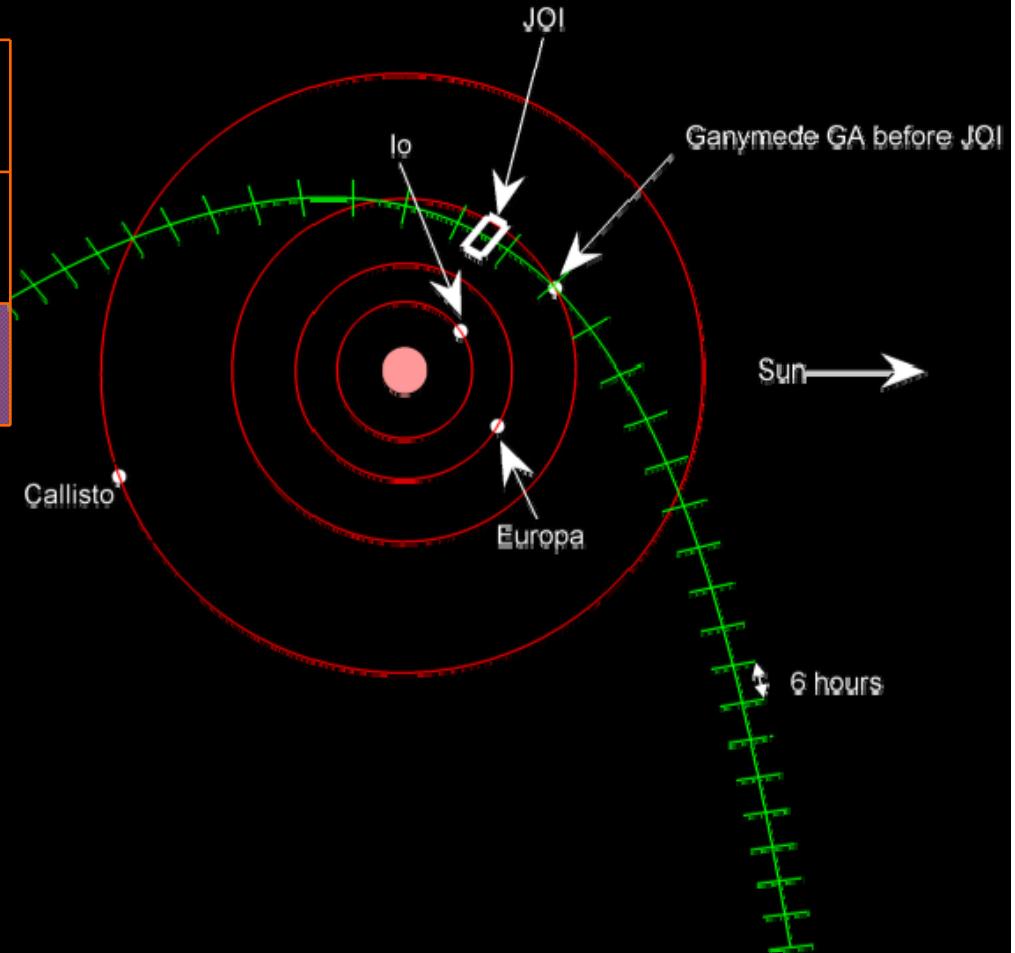
JGO Spacecraft Salient Properties

- Dry mass ~1700 kg, propellant mass ~2500 kg
- High Δv required: 2350 m/s
- Model payload 104 kg, ~120 – 150 W
- 3-axis stabilized s/c
- Power: solar array 60 – 70 m², 640 – 700 W
- HGA: >3 m, fixed to body, X & Ka-band
- Data return >1 Gb per 8 h pass (1 ground station)



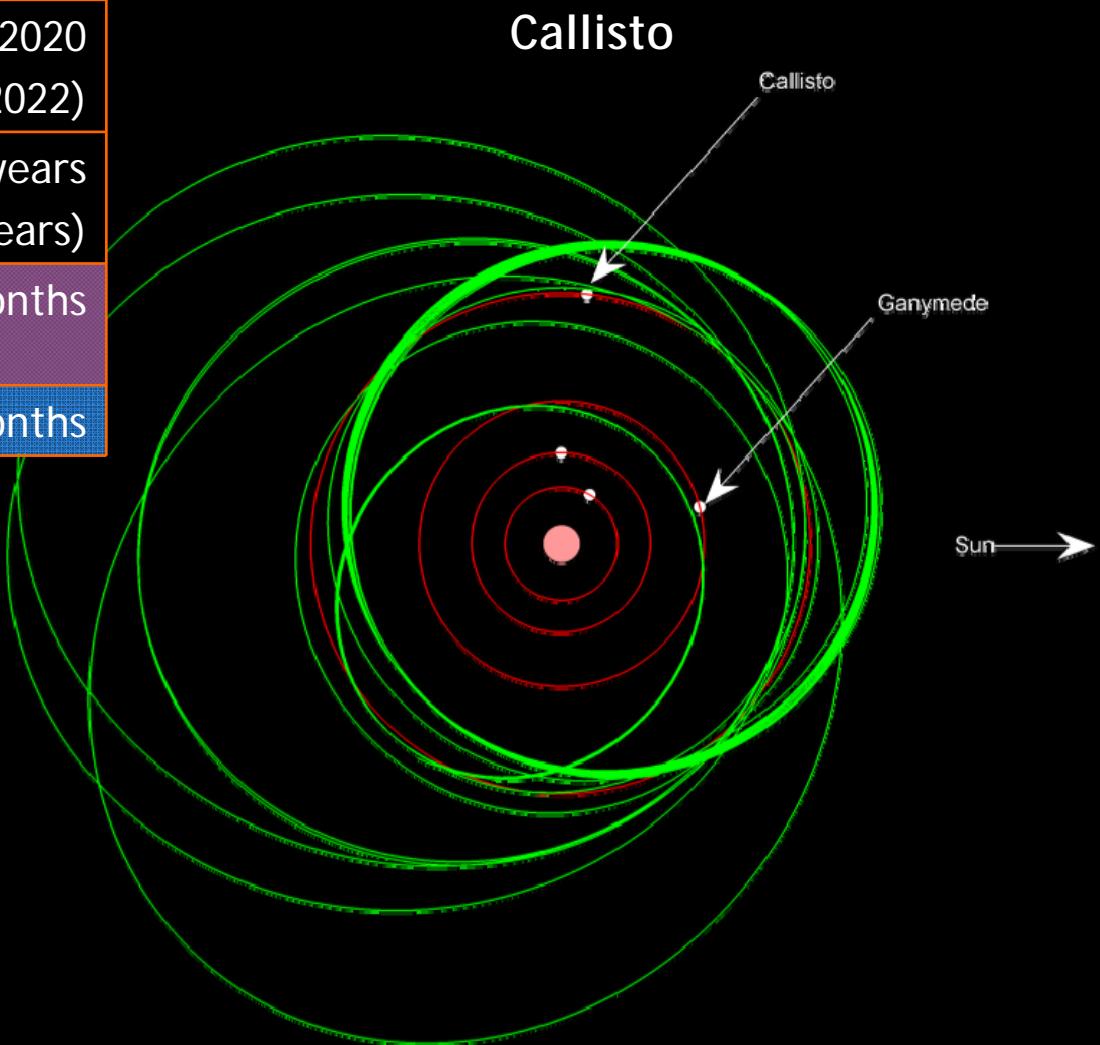
Mission Trajectory

Launch	March 2020 (June 2022)
Interplanetary Transfer	5.9 years (7.3 years)
Jupiter Orbit Insertion to Callisto	~11 months



Mission Trajectory

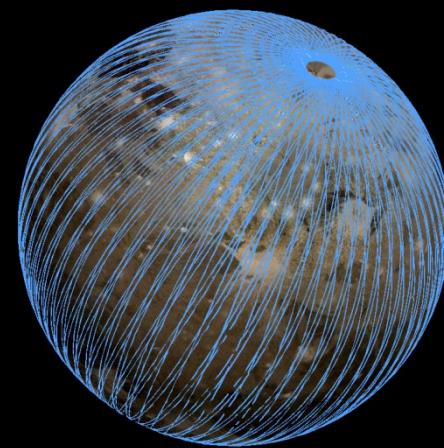
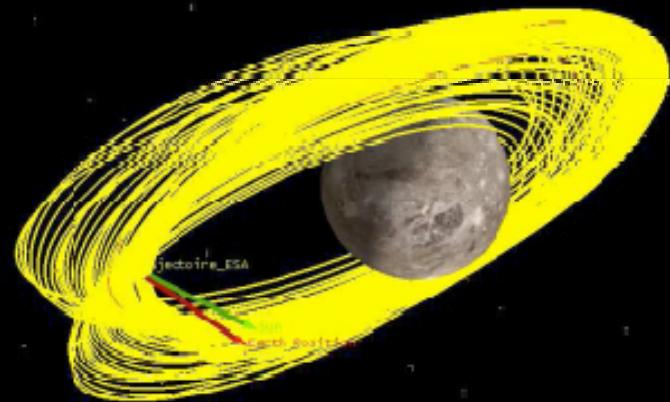
Launch	March 2020 (June 2022)
Interplanetary Transfer	5.9 years (7.3 years)
Jupiter Orbit Insertion to Callisto	~11 months
Callisto flybys	~13 months



Mission Trajectory

Launch	March 2020 (June 2022)
Interplanetary Transfer	5.9 years (7.3 years)
Jupiter Orbit Insertion to Callisto	~11 months
Callisto flybys	~13 months
Callisto to Ganymede	6 months
Ganymede (polar) 10,000x200 km & 5000 km 500 km circular 200 km circular	120 days 120 days 60 days
Total mission duration	9.2 years (10.6 years)

Ganymede



Model Payload

Total mass: 104 kg

Imaging	
Narrow Angle Camera (NAC)	10 kg
Wide Angle Camera (WAC)	4.5 kg

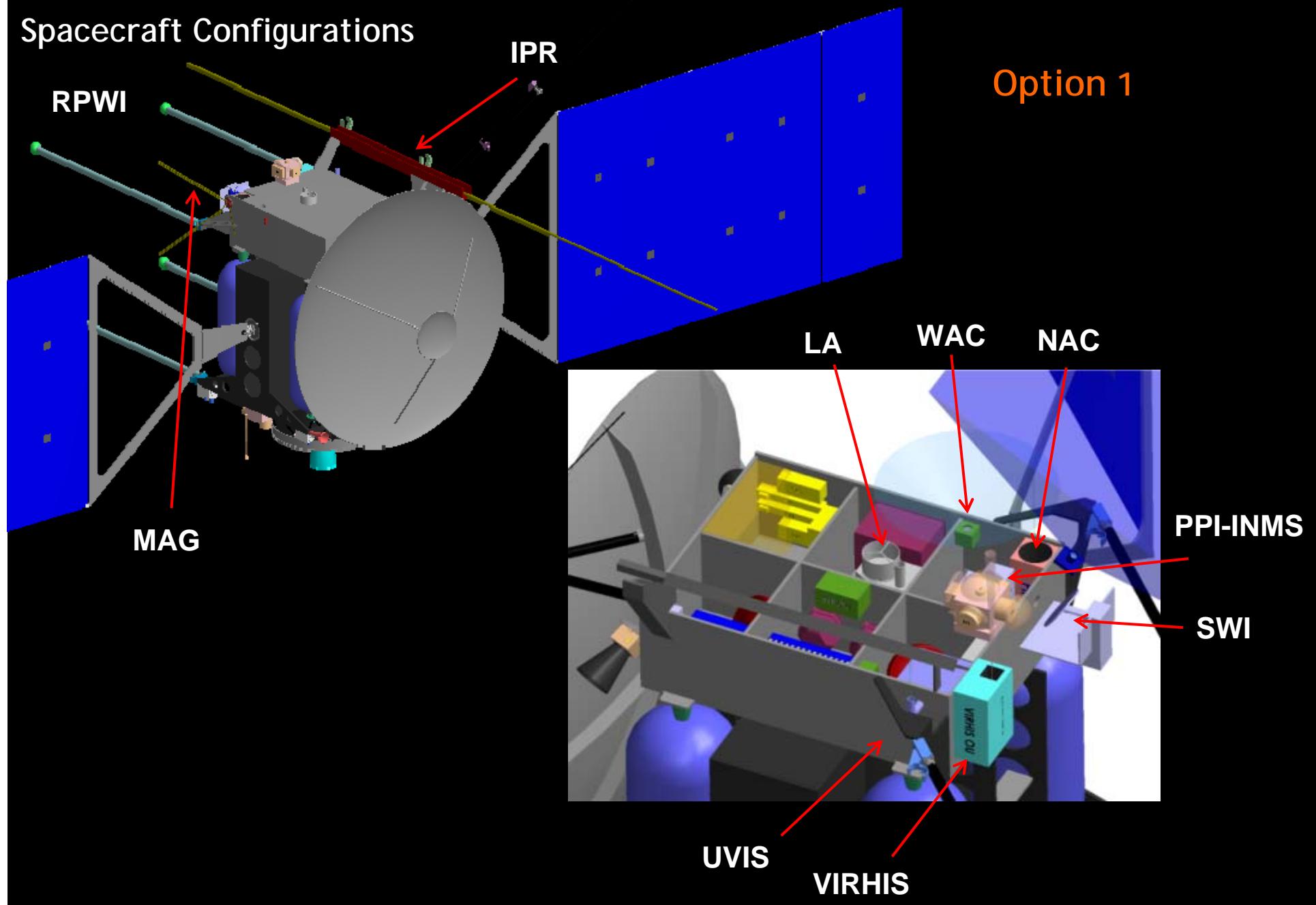
In situ Fields and Particles	
Magnetometer (MAG)	1.8 kg
Radio and Plasma Wave Instr. (RPWI)	11.2 kg
Particle and Plasma Instr.— Ion Neutral Mass Spectr. (PPI-INMS)	18.2 kg

Spectroscopy	
Visible Infrared Hyperspectral Imaging Spectrometer (VIRHIS)	17 kg
UV Imaging Spectrometer (UVIS)	6.5 kg
Sub-mm Wave Instrument (SWI)	9.7 kg

Sounders & Radio Science	
Laser Altimeter (LA)	11 kg
Ice Penetrating Radar (IPR)	10 kg
Radio Science Instrument (JRST+USO)	4 kg

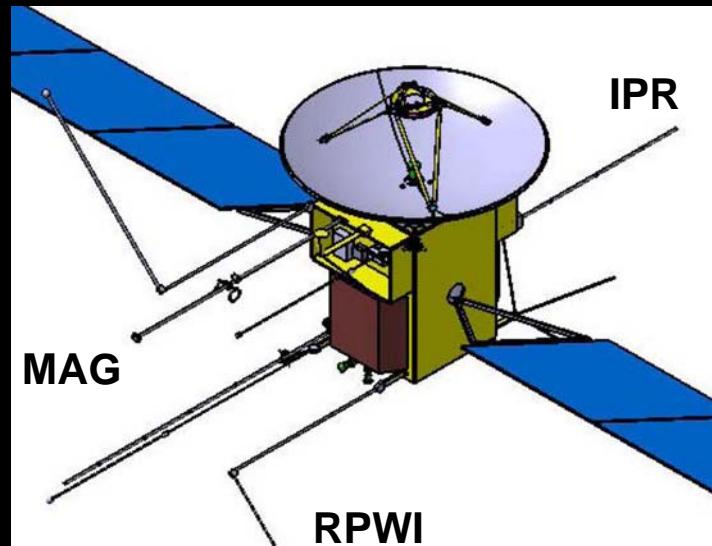
Model payload is based on heritage: BepiColombo, Juno, Mars Express, Double Star, Venus Express, Rosetta, Dawn, Cassini, etc.

Spacecraft Configurations

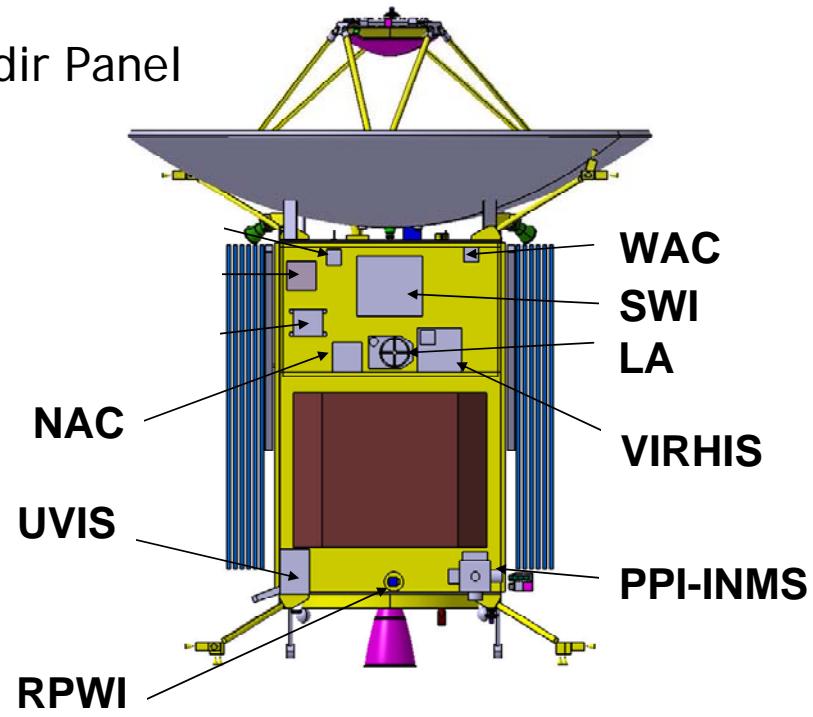


Spacecraft Configurations

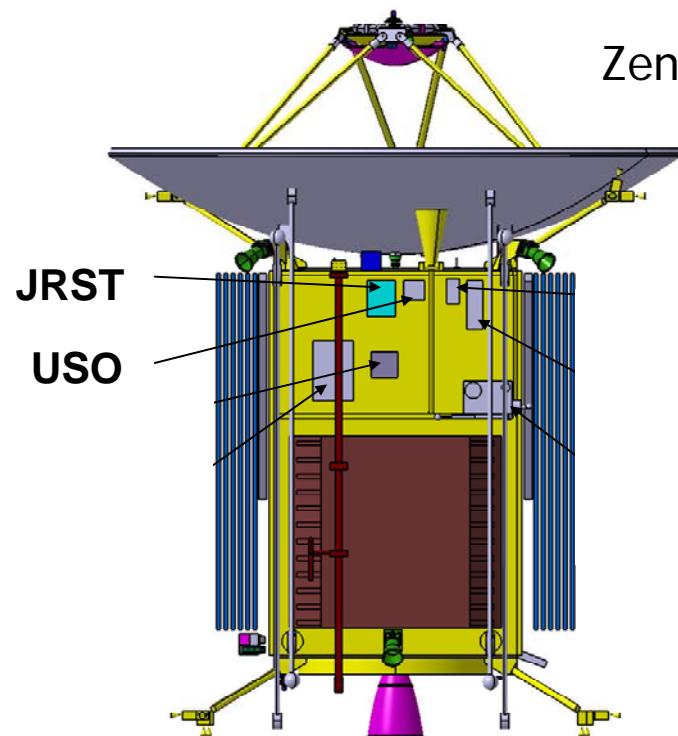
Option 2



Nadir Panel

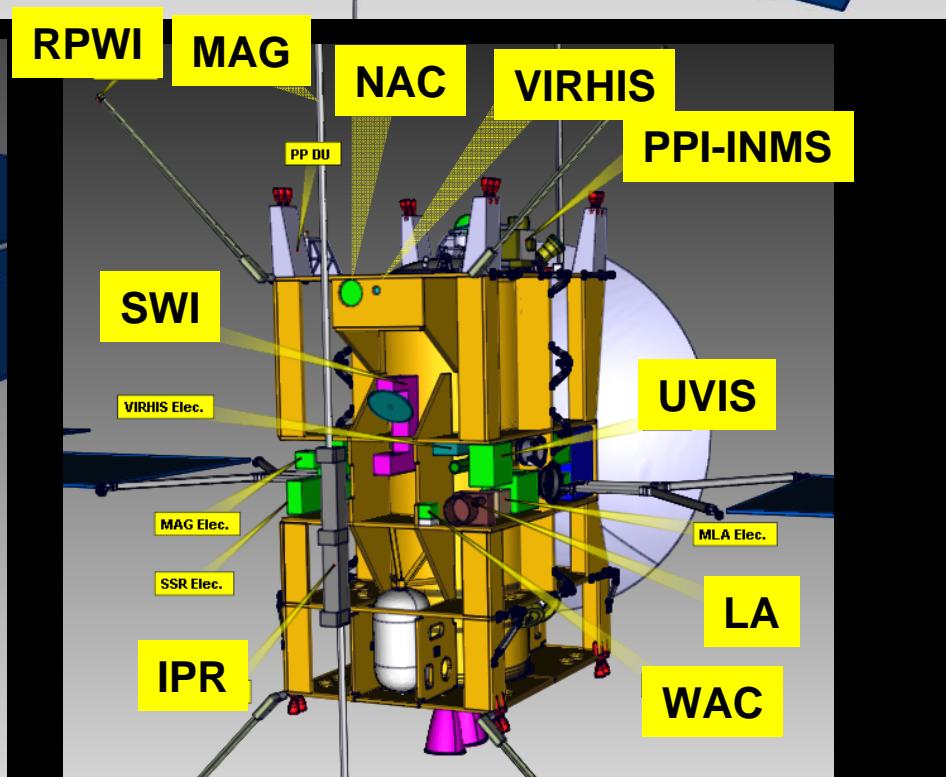
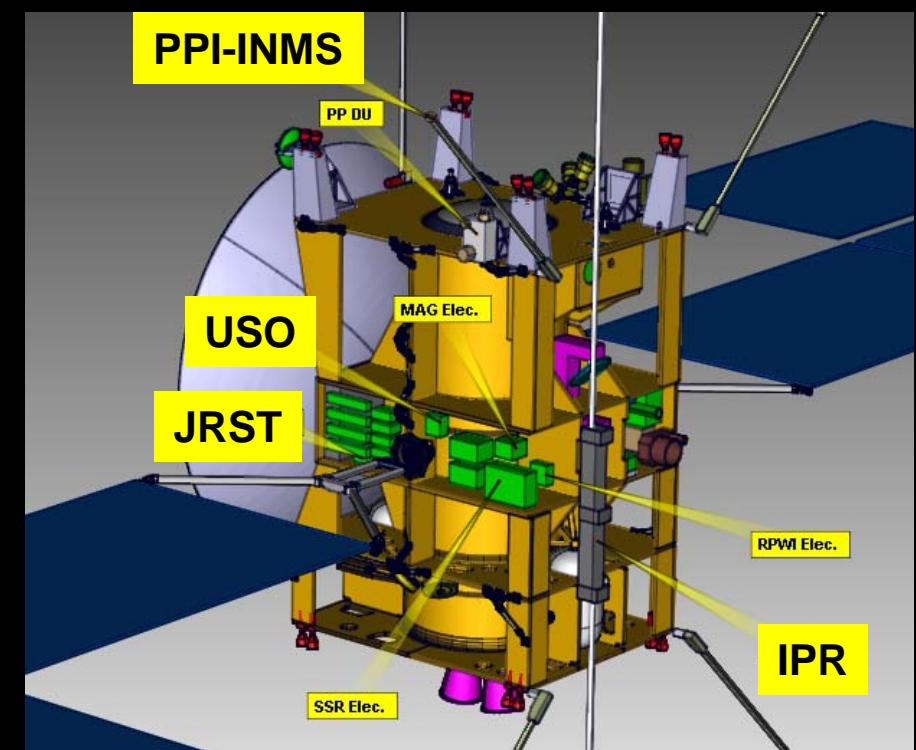
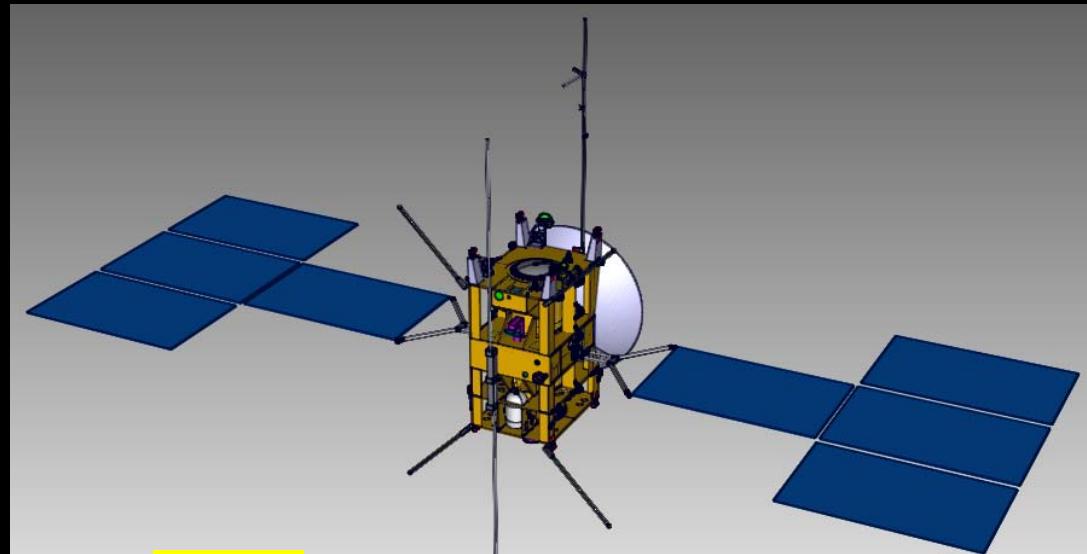


Zenith Panel



Spacecraft Configurations

Option 3



JGO Technology

- Radiation
 - 85 krad inside 10 mm Al
 - Layered shielding strategy achieving 25 krad
 - Shielded equipment compartment (“vault”)
 - Equipment housing
 - Spot shielding of components
 - Sufficient reduction, components available with tolerance of 50 krad
- Solar arrays
 - Low Intensity Low Temperature technology confirmed
 - Enhancements expected from ongoing developments
- Planetary Protection easily achievable



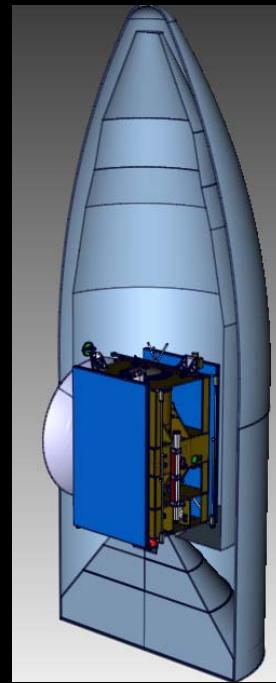
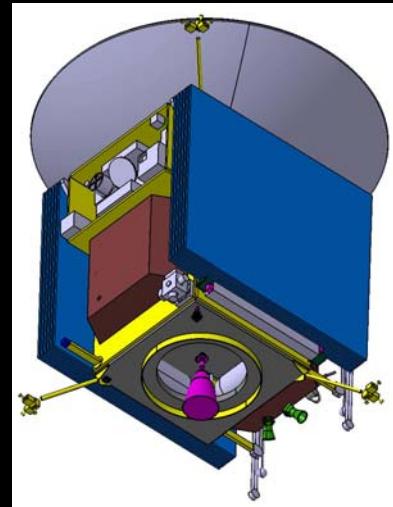
No need for critical technologies

ESA Technical Review: Summary of Risk Assessment – Spacecraft

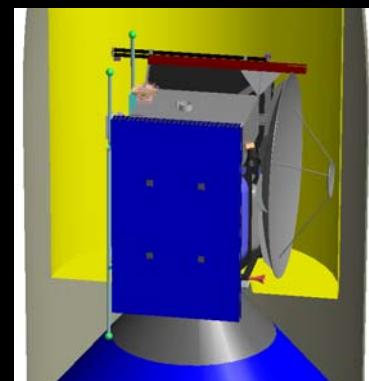
Spacecraft Risk Area	Low	Low-Medium	Medium	Medium-High	High
Mass					
Solar array					
Radiation tolerance					
Materials					
Electrostatic charging					
Magnetic cleanliness					
Communications					
Planetary Protection					
Autonomy and ops					
Technology					
Schedule				2020 launch	

ESA Technical Review: Follow-up and way forward

- **Mass: way forward**
 - Interplanetary transfer: Δv , duration
 - Launcher performance confirmed
 - Optimization of shielding feasible
 - Required system margin (20%) feasible



- **Schedule: way forward**
 - 2020 launch is feasible, but is tight and requires efficient decision process
 - 2021 launch being looked at
 - 2022 can be met comfortably



ESA Technical Review: Summary of Risk Assessment – Model Payload

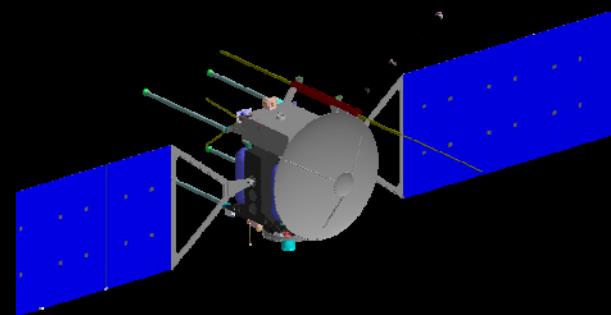
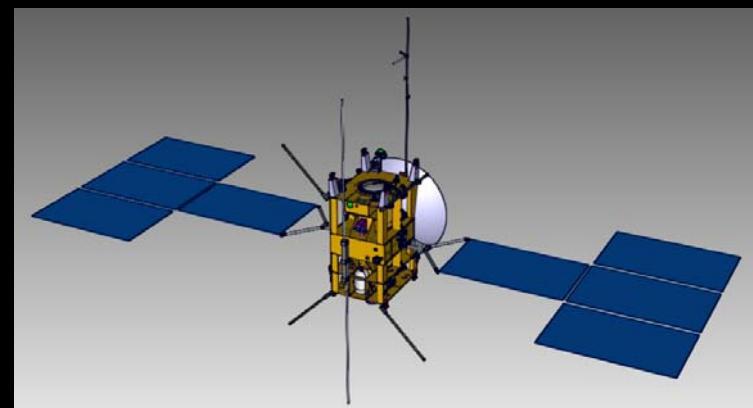
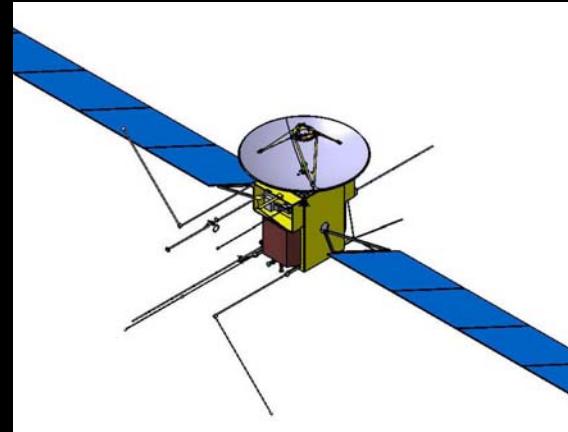
- Instrument teams should avoid single point failures (mechanisms)
- Recommendation for resource sharing, in particular for redundancy:
 - E.g. to consider sharing of p/I processors

Model Payload Risk Area	Low	Low-Medium	Medium	Medium-High	High
Instrumentation for JGO					
New developments for JGO (sensors & optics)					
Schedule					

European Instrumentation for JEO					
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Conclusions

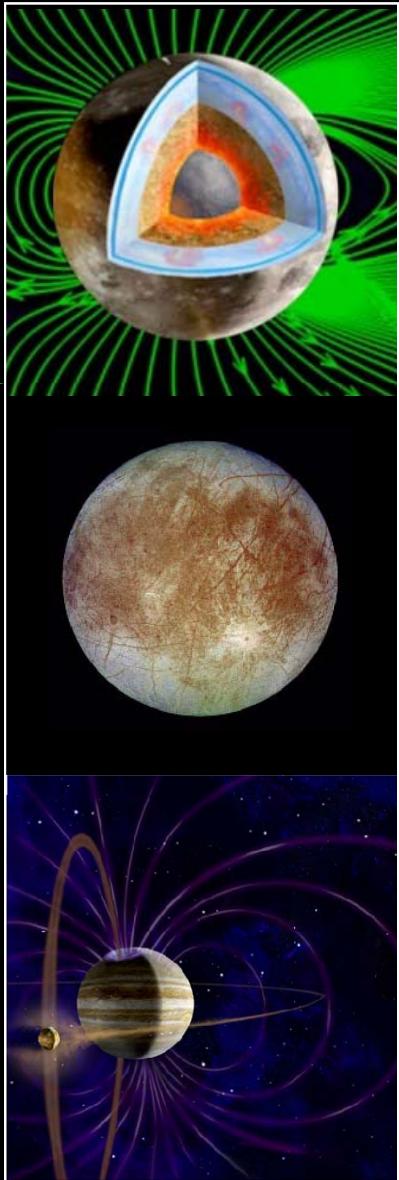
- JGO is the ESA spacecraft within the EJSM-Laplace mission
- JGO is technically feasible
- JGO is ready to go: No need for critical technology identified
- JGO is compatible with programme constraints



Summary

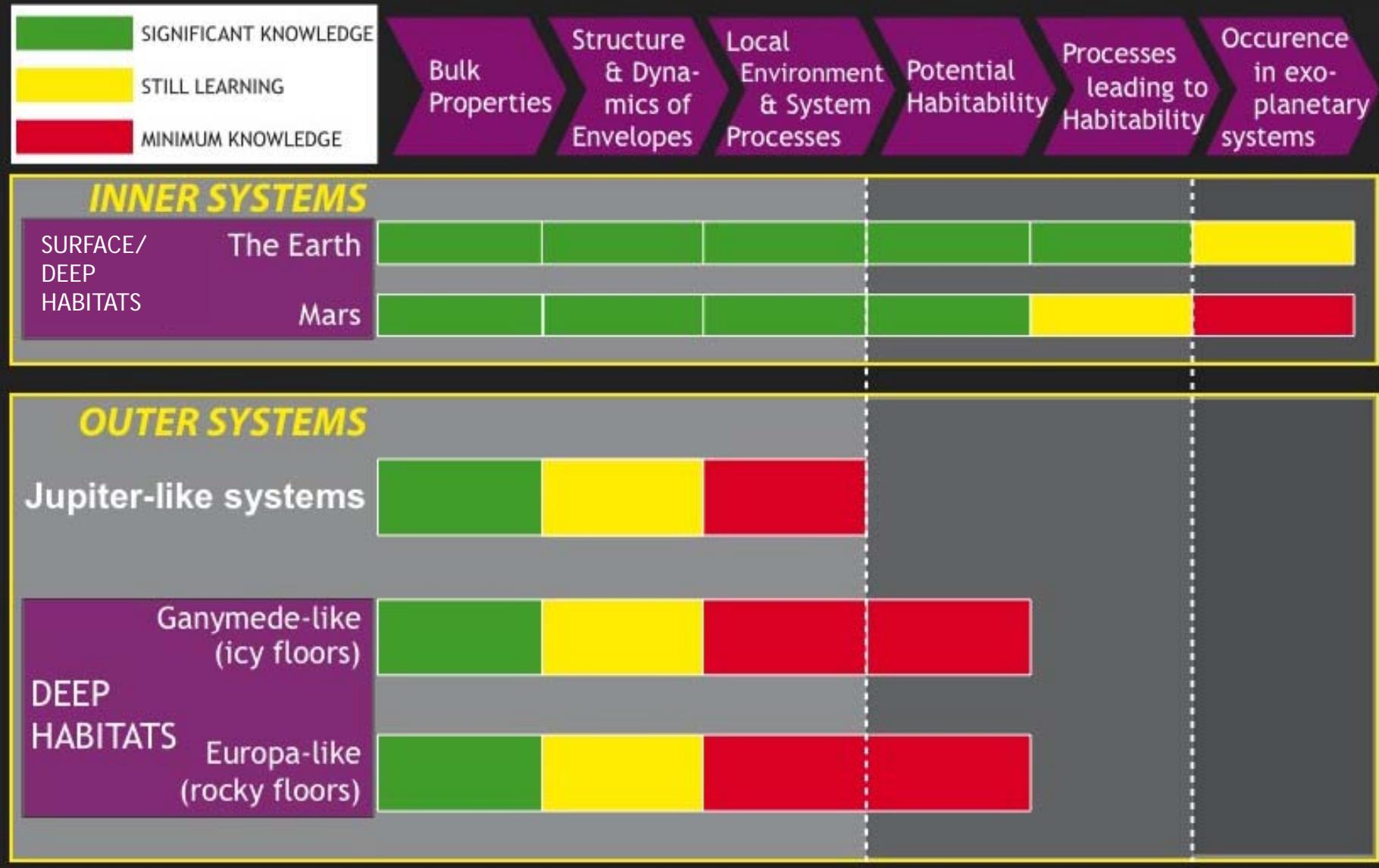
Michele Dougherty

The emergence of habitable worlds around gas giants

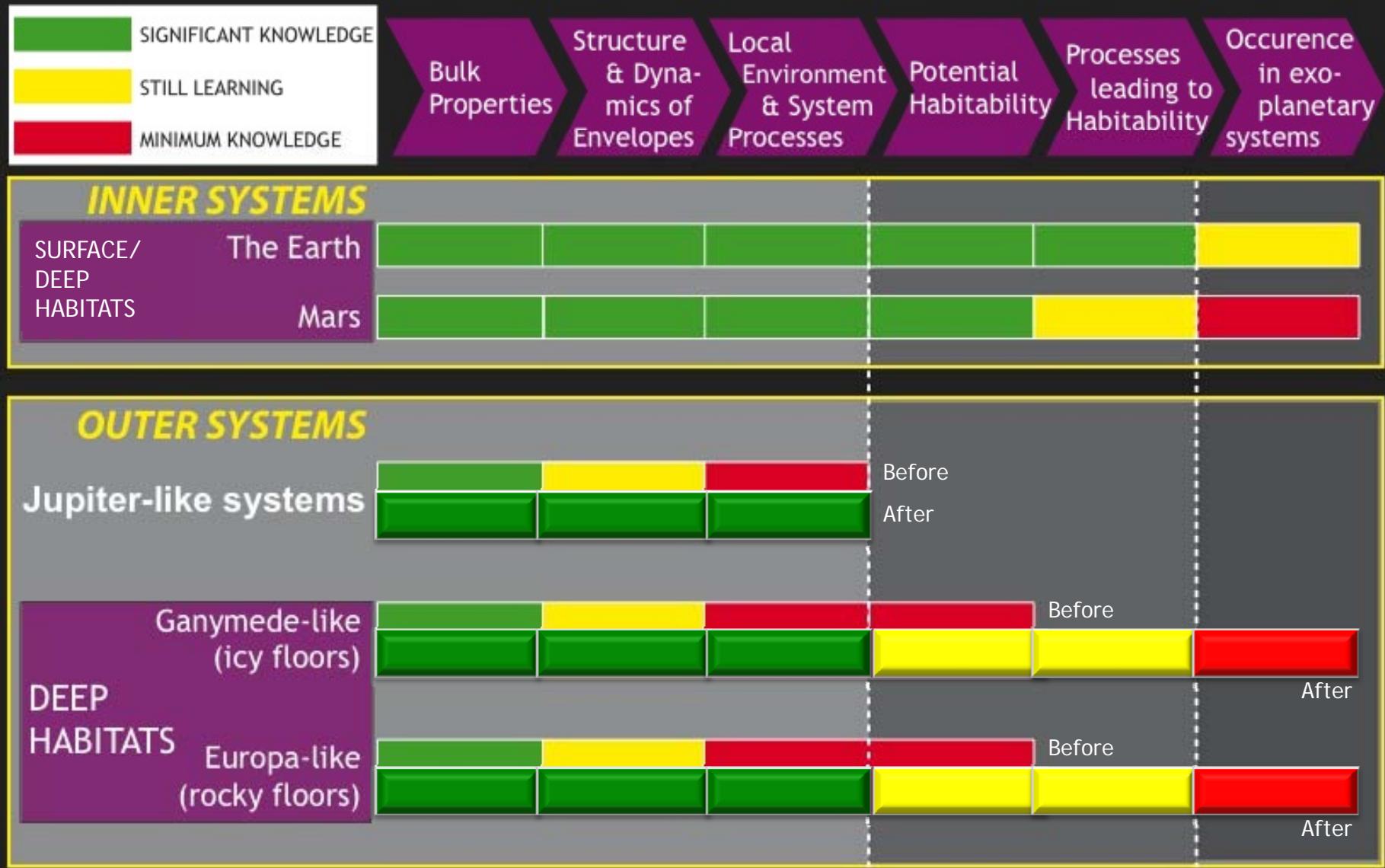


- **Ganymede (JGO-focus):**
Characterise Ganymede as a planetary object including its potential habitability
- **Europa (JEO-focus):**
Explore Europa to investigate its habitability
- **Jupiter System (JGO + JEO):**
Explore the Jupiter system as an archetype for gas giants

The Emergence of Habitable Worlds: Different destinations, the same steps...



The Emergence of Habitable Worlds: Different destinations, the same steps...



Thanks to the Joint Science Definition Team

EJSM-Laplace

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3.02.2011) ESA

Dima Titov (Study Scientist 3.02.2011 -) ESA

Robert Pappalardo (US Study Scientist) JPL

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Louise Prockter (US Deputy Study Scientist)
Johns Hopkins University—APL USA

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Mark Showalter (USA)

Amy Simon-Miller (USA)

Mitch Sogin (USA)

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Takeshi Takashima (Jp)

Paolo Tortora (It)

Federico Tosi (It)

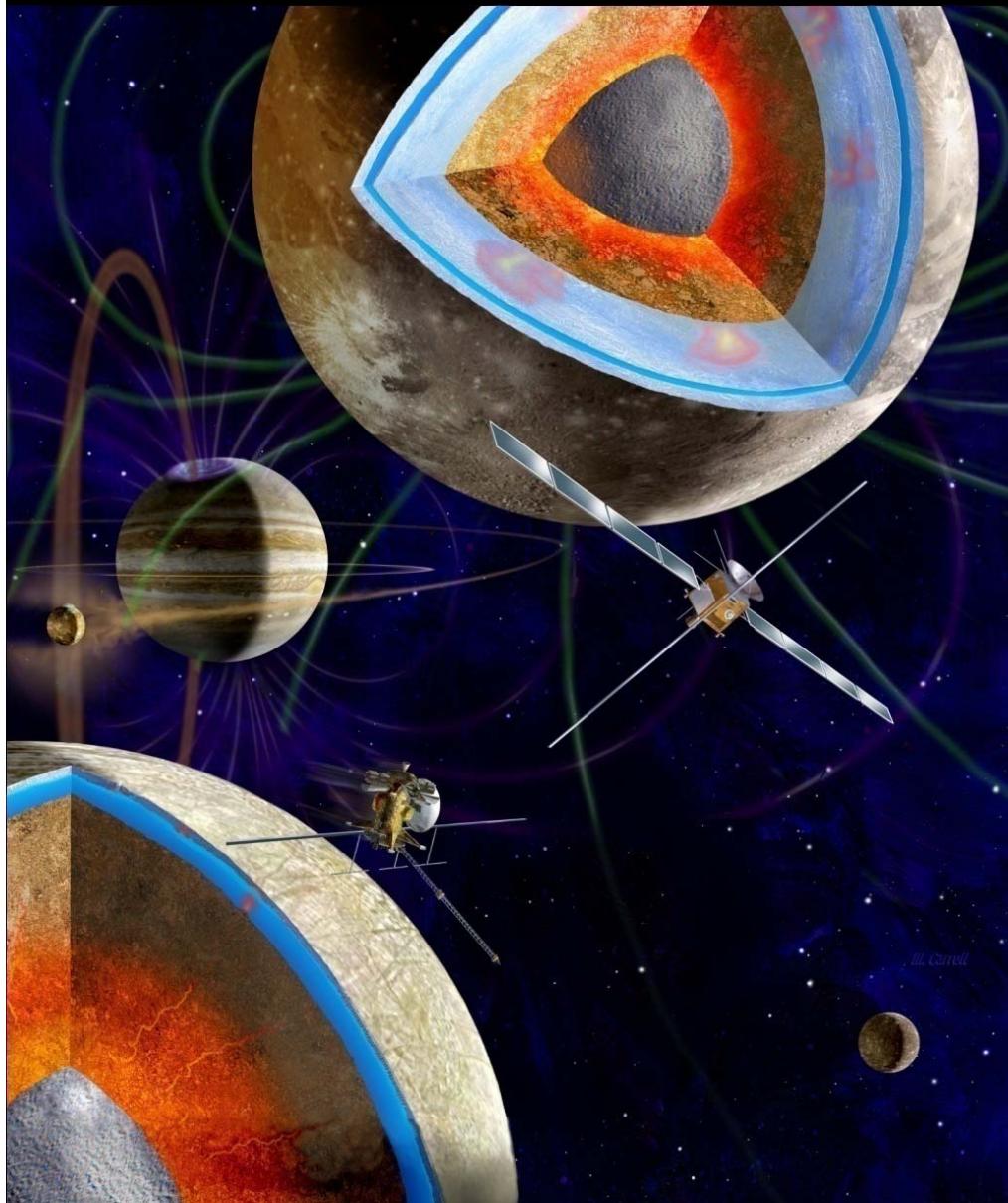
Elizabeth Turtle (USA)

Tim Van Hoolst (Bel)

Hunter Waite (USA)

Summary: L-class Cosmic Vision Science

EJSM-Laplace



- Scientific goals and objectives are well-defined:
 - Broad appeal to different science communities
 - Flows from community priorities
 - Mature science coupled with exploration opportunities
- International cooperation:
 - JGO and JEO will offer unique scientific synergies
 - JGO and JEO each address key portions of EJSM-Laplace science
 - JGO and JEO could each stand on their own to achieve compelling icy world and Jupiter system science
- Mission is judged feasible with no need for critical technologies
- Many opportunities for public outreach

The emergence of habitable worlds around gas giant planets

