

Selecting an image sensor for the *EJSM VIS/NIR* camera systems

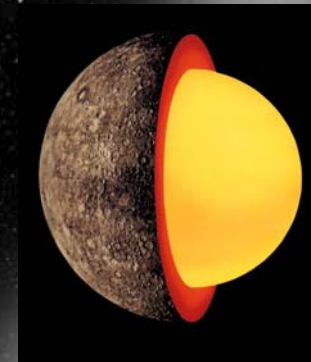
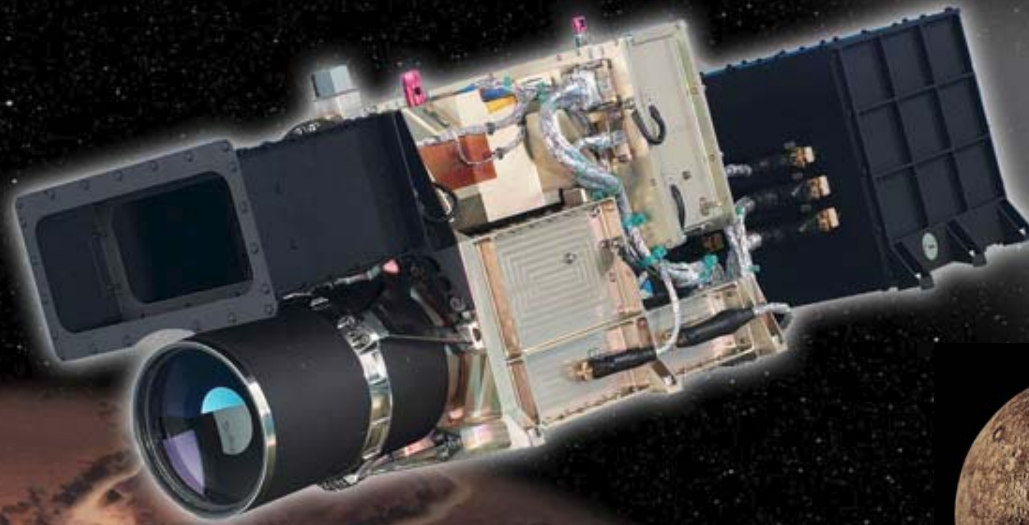
presented by

Harald Michaelis (DLR-PF)

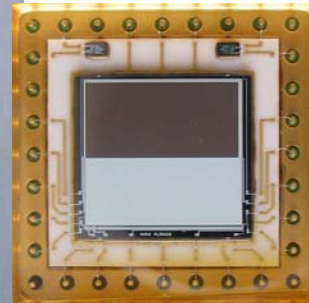
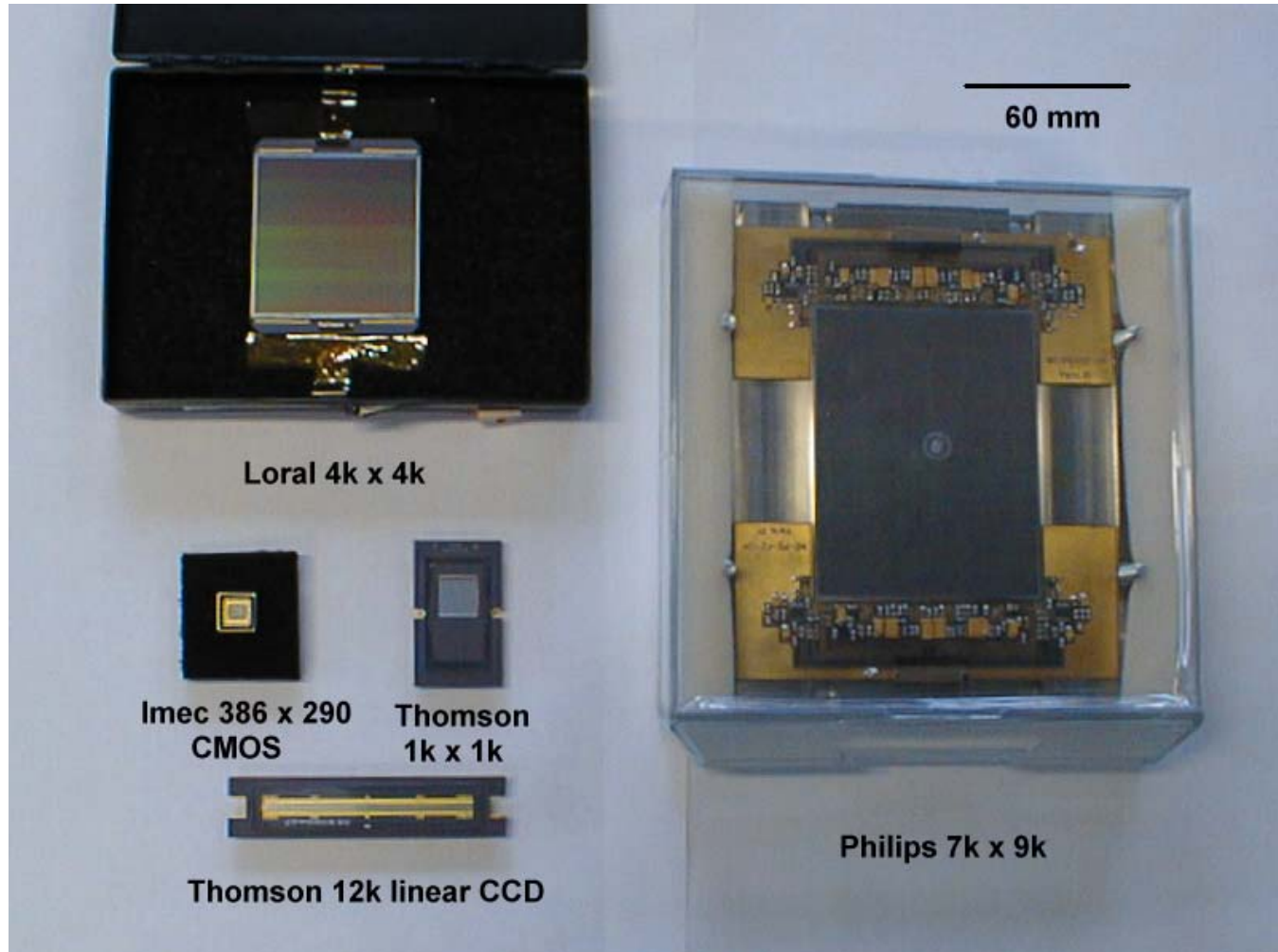


Deutsches Zentrum
für Luft- und Raumfahrt e.V.
in der Helmholtz-Gemeinschaft

What for a detector/sensor we shall chose for EJSM?



Used Linear CCD, CCD Arrays and CMOS Detectors at DLR



Cassini-Huygens- DISR
MPS; 512x512



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Contents

- Introduction
- **General performance requirements**
- Environment
- **Draft analysis and estimations**
- **The current baseline sensor candidates**
- Alternatives
- Conclusions and outlook



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Introduction

- What are the performance requirements for the **high resolution camera**
- What is the radiation environment and the resulting effects on the detector performance
- What are the current detector- candidates and what for a performance they provide

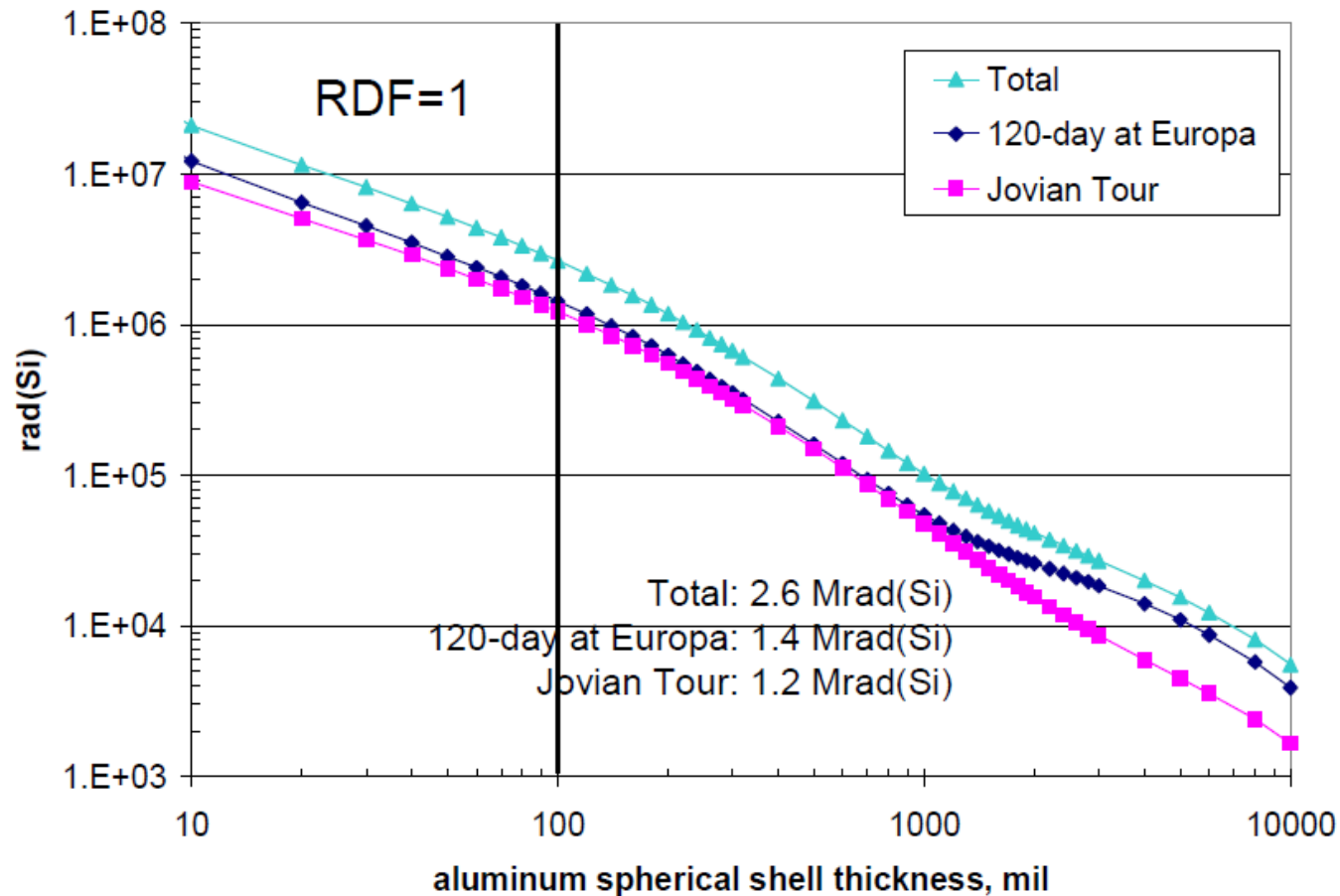


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General performance requirements (HRC)

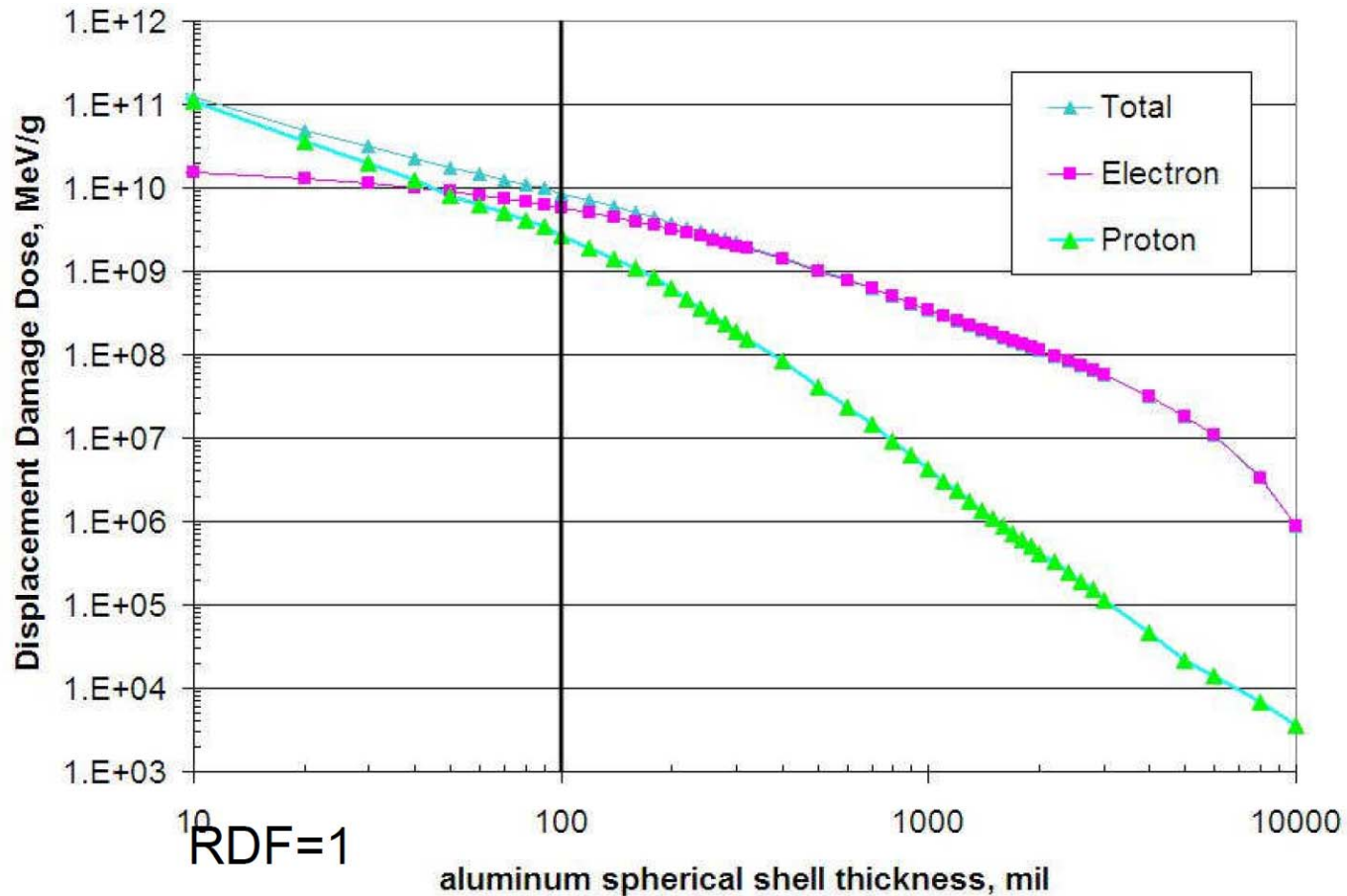
- **Resolution (IFOV): 1m/px...5m @ 200km**
- **swath width: ~1km ...10km) → >1000px**
- **Spectral range (channels): 350nm ...1050nm; filter channels: PAN, TBD**
- **SNR: ~100**
- **(mapping capability)**

EE2007 Mission Dose-Depth Curve by Mission Segment



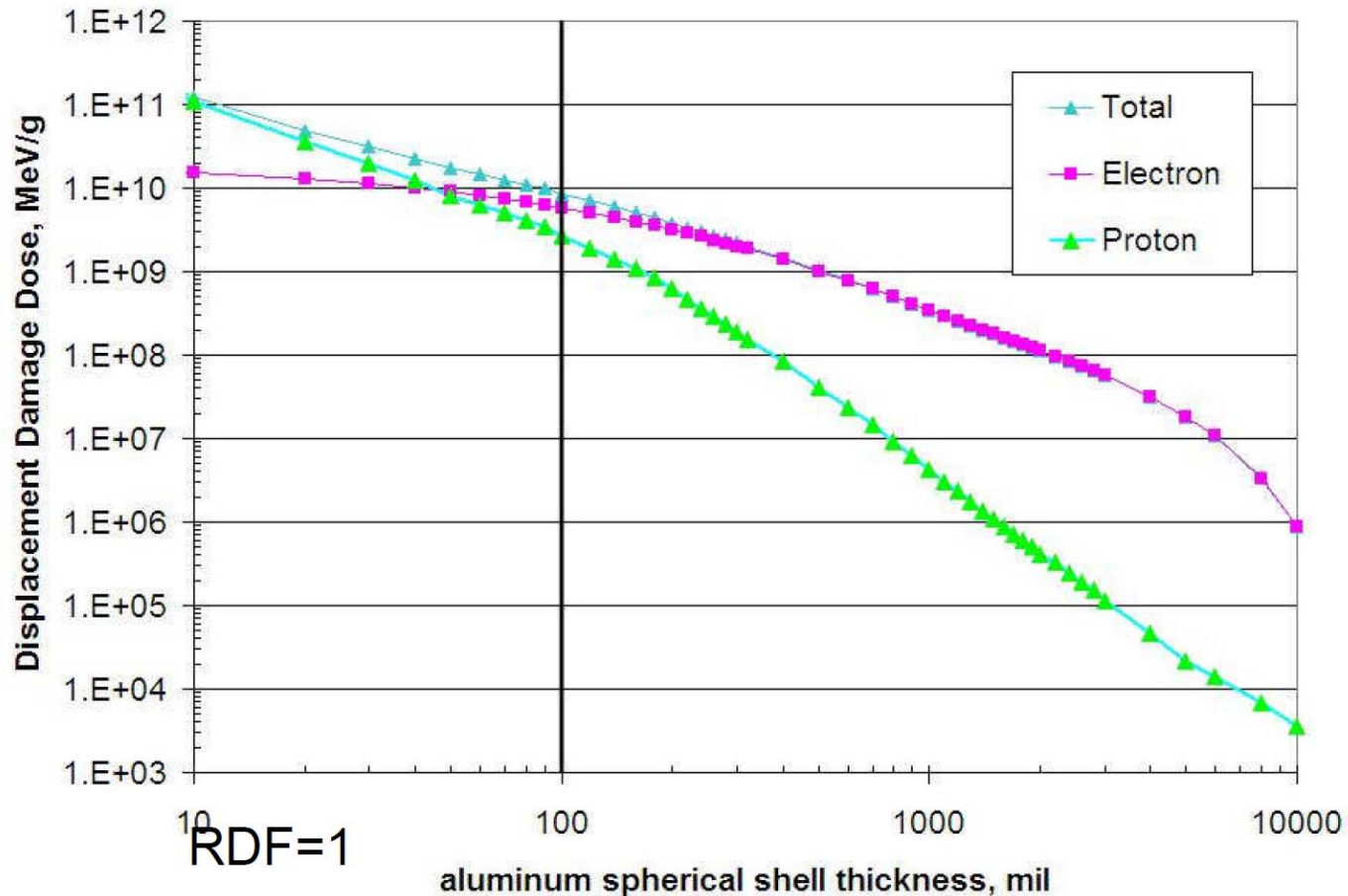
- Copy from presentation 'Characterization of Radiation Environments- Europa Orbiter';
Insoo Jun; June 2008; JPL

EE2007 Displacement Damage Dose [2]



- Copy from presentation 'Characterization of Radiation Environments- Europa Orbiter';
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EE2007 Displacement Damage Dose [2]



- This shows the contribution of different particle species to the damage dose. As shown, the electron is the dominant contributor to the damage dose. -



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Radiation Environment - Europa / Ganymed

- radiation belt with 100KeV...50MeV protons and electrons
- deep penetrating electrons cause high TID

~1MRad(Si) @ 5mm - JEO

~0.1MRad(Si) @ 5mm - JGO

-High energetic particles (ions, protons, neutrons, e⁻) cause DDD

~1.3 10⁰⁸ MeV/g (behind 1cm Ta shielding - JEO)

~3 10⁰⁸ MeV/g (behind 1cm Al shielding - JGO)





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Radiation effects on the detector and impact on instrument performance

- Increase of dark current and dark current shot noise
 - Hot spots ↑; transients↑ (dark spikes)
 - Decrease of CTE (CCDs)
 - Degradation of QE
 - Latch up effects (CMOS)
- Big impact on noise and signal level
- **SNR will suffer**
- Danger of total damage



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Draft analysis and estimations

' What is the signal level and the SNR that I get with the detector'?



Impact on sensor performance requirements and sensor design architecture





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Draft analysis and estimations

1. Assumptions

- IFOV: 1m and 5m/px @200km
- 10 μ m pixel size
- Focal length: ~2m...0.4m
- aperture: 0.1...0.2m (max.; because of mass)
- F#: 2...10
- vg: ~1870m/s ; Tdwell: 0.5...2.5ms; Tint: $\frac{1}{4}$ Tdwell
- Albedo: 0.43
- Spectral range: PAN (550nm...800nm); FWHM: 250nm
- QE: ~80%; optical efficiency: 0,8
- Detector noise: 2, 10, 100e rms

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Draft signal analysis and SNR estimations

- 1. Signal level @ 1m/px and 5m/px

F#:	2	4	10
signal[ke] @ 1m	11	2,8	0,4
@ 5m	56	14	2,3

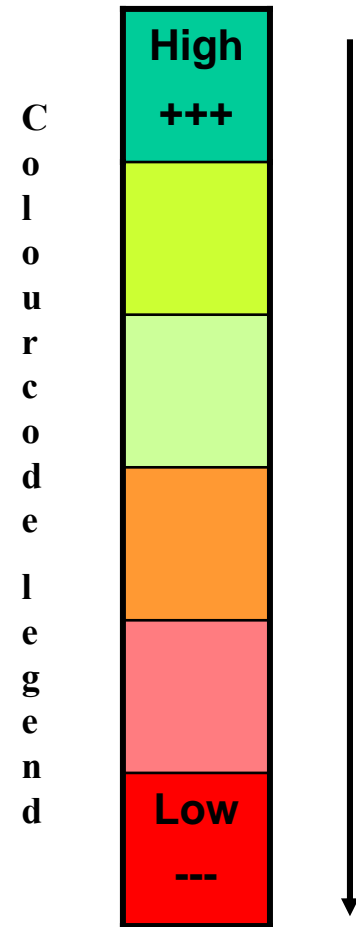


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Draft analysis and estimations

1. SNR estimations @ 1m

F# /:	2	4	10
Noise [e rms]			
2	106	53	21
10	105	52	19
100	77	25	4

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Draft analysis and estimations

1. SNR estimations @ 5m

F# /:	2	4	10
Noise [e rms]			
2	237	118	47
10	237	118	46
100	218	90	20



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Draft analysis and estimations – First Conclusions

If 1m resolution is required then

→ a very low noise sensor with high QE and fast optics
(~F/2)

or



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Draft analysis and estimations – First Conclusions

If 5m resolution is sufficient then

→ a medium fast optics ($\sim F/4$)

→ Readout noise sensor of $\leq 50e$ rms is sufficient

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SYMBIO-SYS and STAR1000 sensor characteristics

1. 'SYMBIO-SYS' (Data TBC)

Heritage	Bepi Colombo
Type	Si-PIN-CMOS (Hybrid)
Format and pixel size	2kx2k; 10 μ m
QE	>80%
Noise [e rms]	<100 (TBC)
Radiation Tolerance	~100 krad (TBC)
others	off-chip ADC; global shutter, no TDI

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SYMBIO-SYS and STAR1000 sensor characteristics

1. STAR 1000

Heritage	ROKVISS, many star trackers
Type	CMOS- APS
Format and pixel size	1kx1k; 15µm
QE	~30%
Noise [e rms]* *: temporal readout noise	~40e
Radiation Tolerance	>230 krad;
others	on-chip ADC; rolling shutter , no TDI

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HRC with STAR1000 sensor → instrument characteristics

Goal	SNR: >100
Focallength needed:	3m
F#	~1.2
Total aperture:	>2.5m
Mass*:	>500kg
Others:	<i>rolling shutter is not applicable</i>

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HRC with 'SYMBIO-SYS' sensor → instrument characteristics

Goal	SNR: >100
Focallength needed:	2m
F#	~2 (...4)
Total aperture:	~50-100cm*
Mass**:	>100kg
Others:	
	*: depends on assumed detector noise level **: 20cm telescopes:>10kg



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Conclusion II (1m- HRC)

- STAR1000 is **NOT** applicable *for 1m- HRC*
- SYMBIO-SYS: appears as difficult or not applicable (only applicable with trade-offs - FWHM, Tint, IFOV);
 - It would be much better to have a **TDI- sensor architecture** and (high QE, low noise-)



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Conclusion II - continued

- Either we can relax the 1m- resolution requirement
- or
- we have to apply **TDI** or motion compensation

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TDI and its effect on needed telescope aperture (SNR=100)

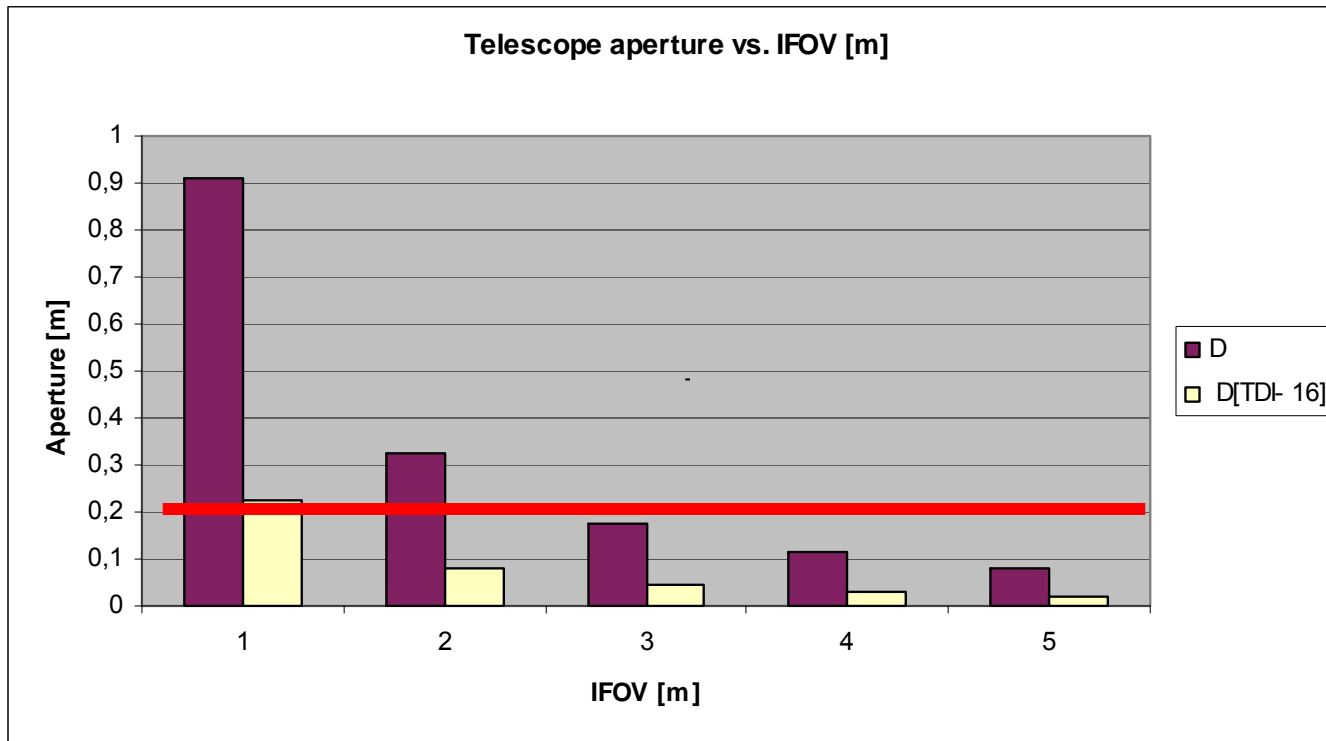


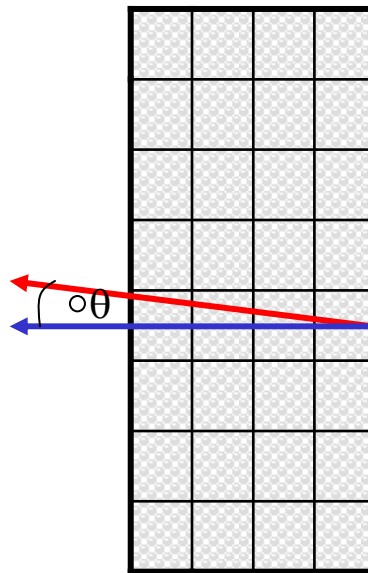
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Conclusion II - continued

TDI or motion compensation requires that the **S/C motion vector** is parallel to the **pixel (charge- transport-) vector**

$$\theta \leq 0.5/n_{\text{TDI}}$$

$$n_{\text{TDI}} = 8 \dots 64$$



$$f_{\text{readout}} > n_{\text{pxl}}/T_{\text{dwell}}$$

f_{readout} : detector readout frequency

n_{pxl} : number of pixel per line

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CCD vs CIS –few words

- Both are silicon detectors with capacitors and transistor(s) -

CCDs
Capacitors are the dominant components
Ionizing radiation → oxid charging → flatband voltage shift
DDD have big impact on CTE
Insensitive to SE-latch-up
High QE is standard
Low noise at low readout rates
TDI is straightforward
High performance customized design and manufacturing is rather easy

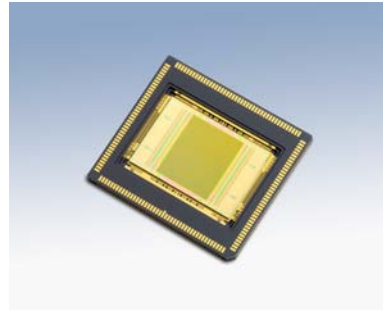
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CCD vs CIS –few words

CIS
transistors are the dominant components
Complex electronics can be integrated in customized designs
No CTE issue
Sensitive to SE-latch-up
High QE only for customized components
Low noise event at high readout rates is feasible
TDI is feasible but not straightforward
High performance customized design and manufacturing is very difficult

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CIS – some additional remarks (s-CMOS)



- ☺ High performance science CMOS- sensors are now reality
- ☺ QE: ~90%
- ☺ Readout noise: <2 e rms @ readout frequency of 30fps (150MHz!!!)
- ☺ on-chip readout electronics and ADCs
- ☹ **'ESA-Science' is a too small customer for the big CMOS-players**



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Conclusions and further study actions

- (Re-)definition of HRC resolution requirement (IFOV) + trade-offs
- If 1m: Investigate motion compensation vs. TDI
- Analyze CMOS- TDI capabilities and performance
- Comparison of CCD-TDI with CMOS- TDI
- Analyze radiation environment
- Analyze (and test) CCD radiation hardening options and limits
- Investigate procurement possibilities of a customized CIS (access to high performance CIS- process)

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Example: Focal Plate with two sensor types

