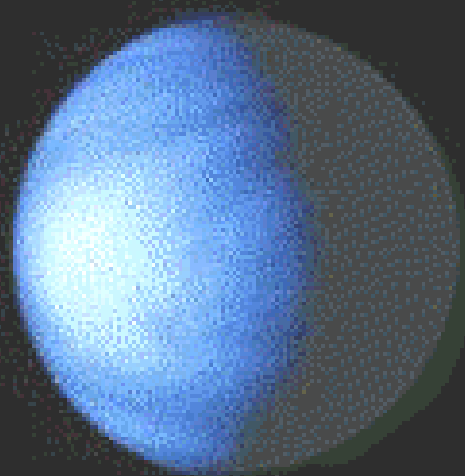
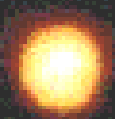
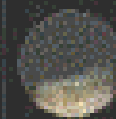
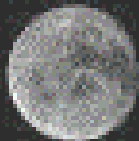
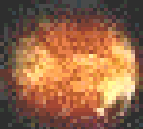
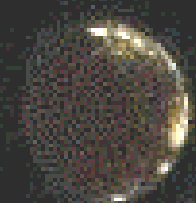
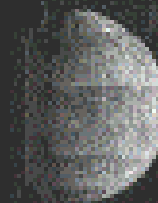


HABITABILITY IN THE SOLAR SYSTEM AND BEYOND



Frances Westall



Centre de Biophysique Moléculaire, CNRS, Orléans, France
westall@cns-orleans.fr



→ **WHAT IS HABITABILITY ?**

→ **EARTH**

→ **MARS**

→ **EUROPA**

→ **(VENUS?)**

→ **OTHER?**

WHAT IS HABITABILITY?

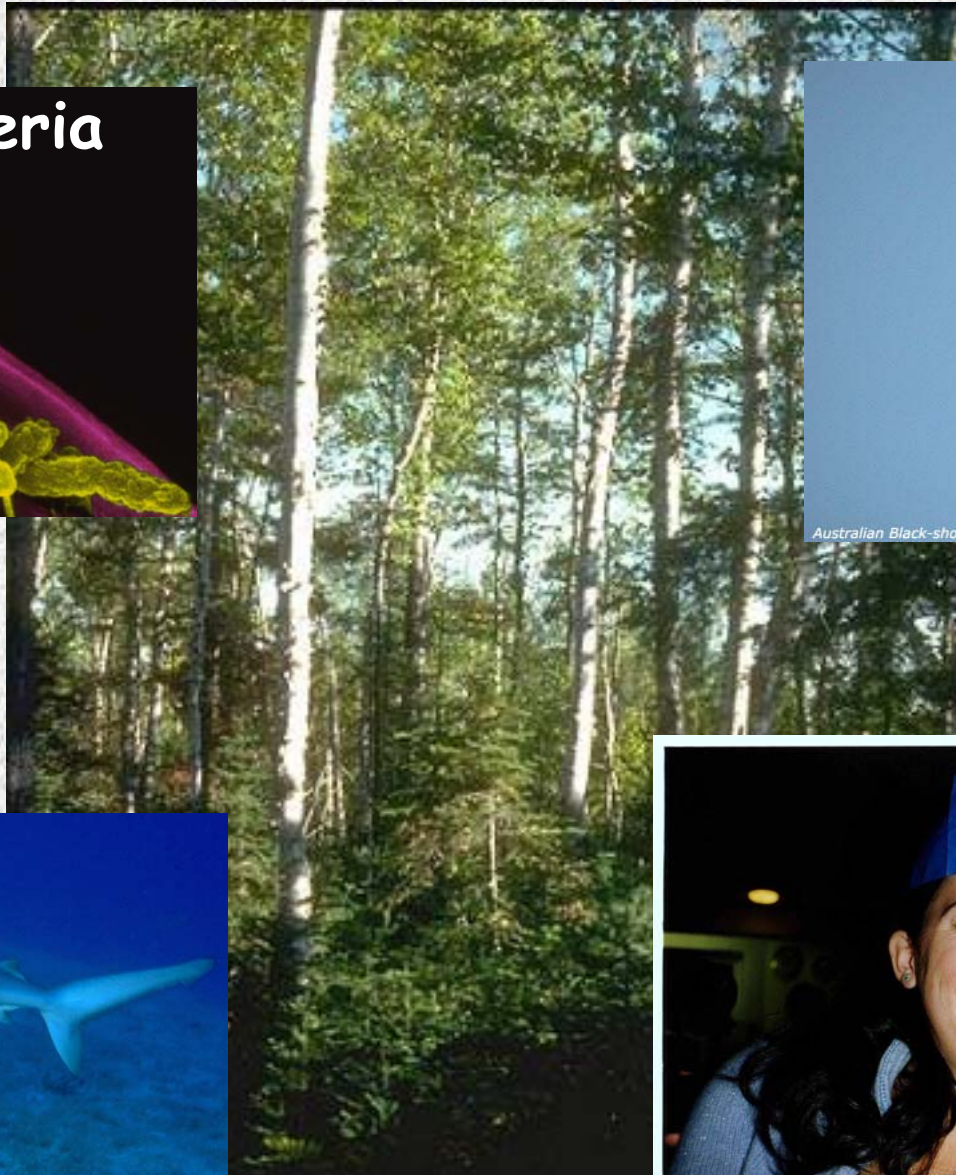
Habitable = fit to be lived in



FOR WHAT KIND OF LIFE?

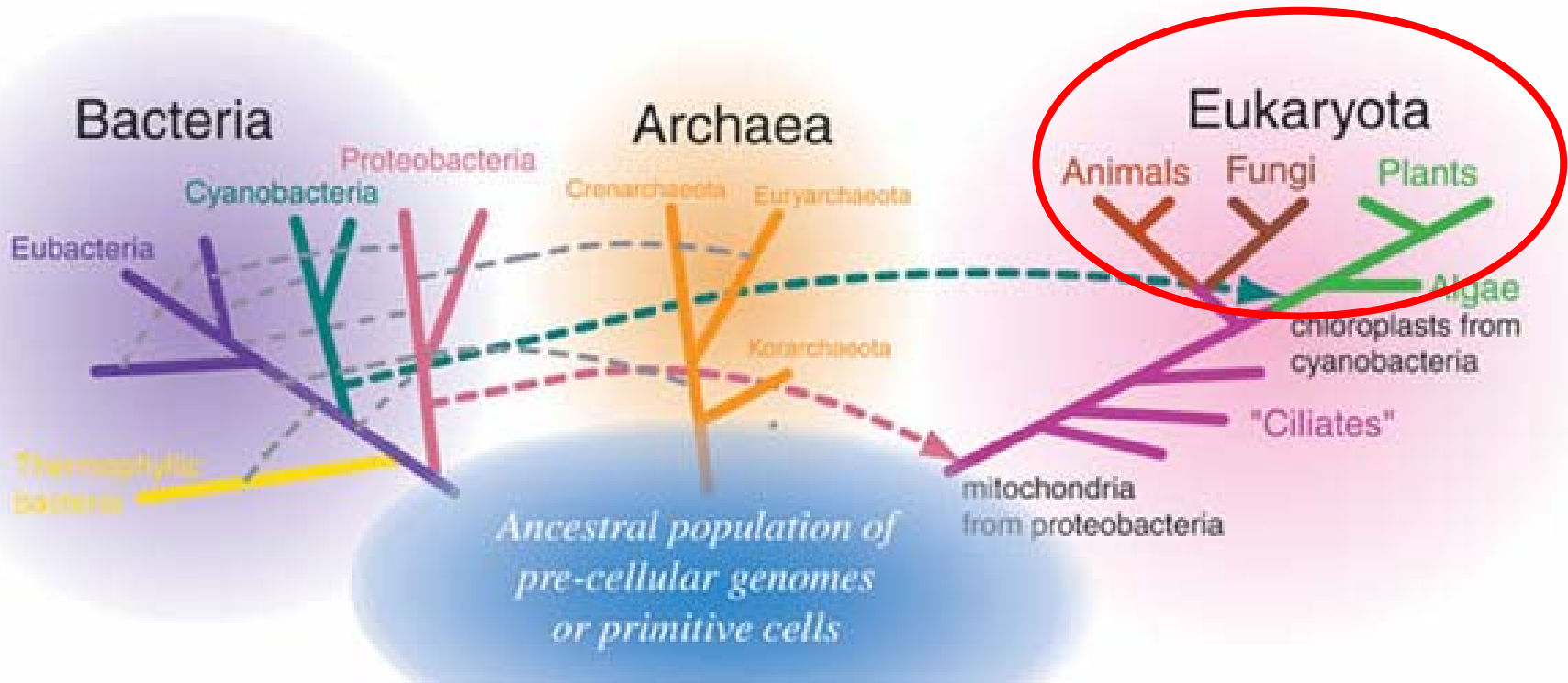
What kind of life?

bacteria



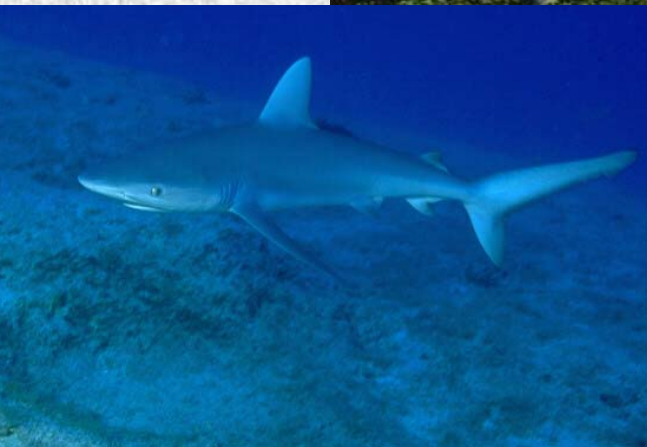
KINGDOMS OF LIFE

Most visible surface life



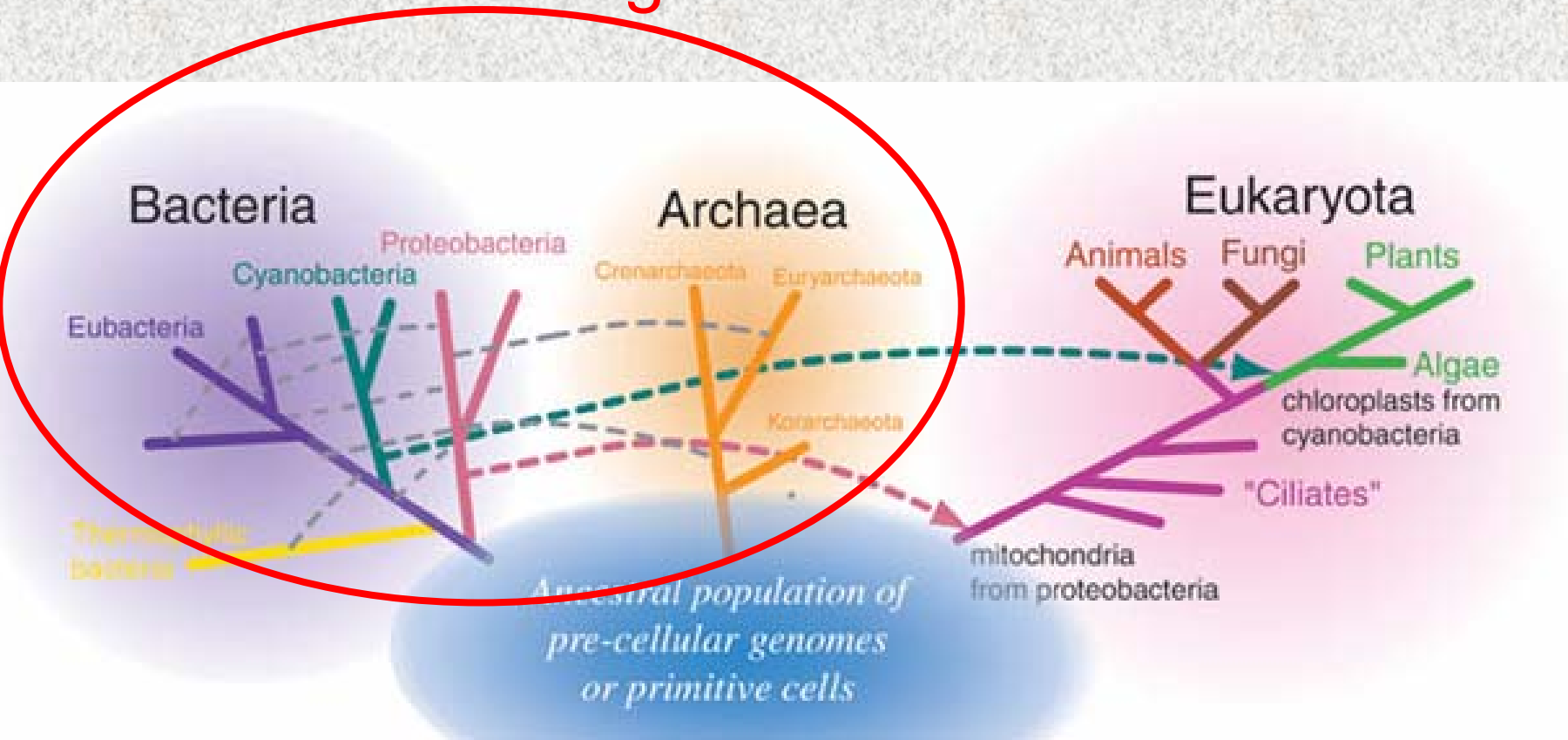
After Woese

Eucaryotes are the product of long co-evolution with a geologically active planet



KINGDOMS OF LIFE

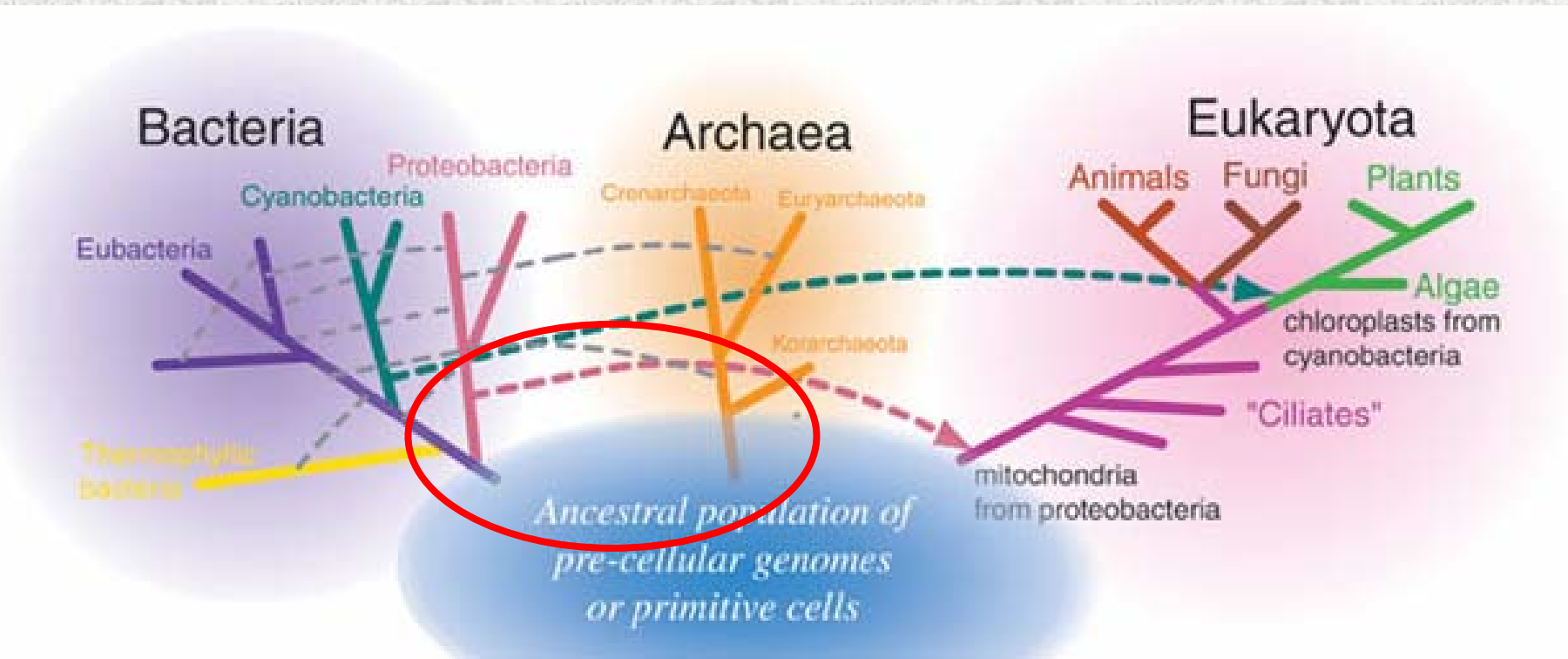
Largest biomass



After Woese

KINGDOMS OF LIFE

Earliest life and probable level of ET in the SS



After Woese

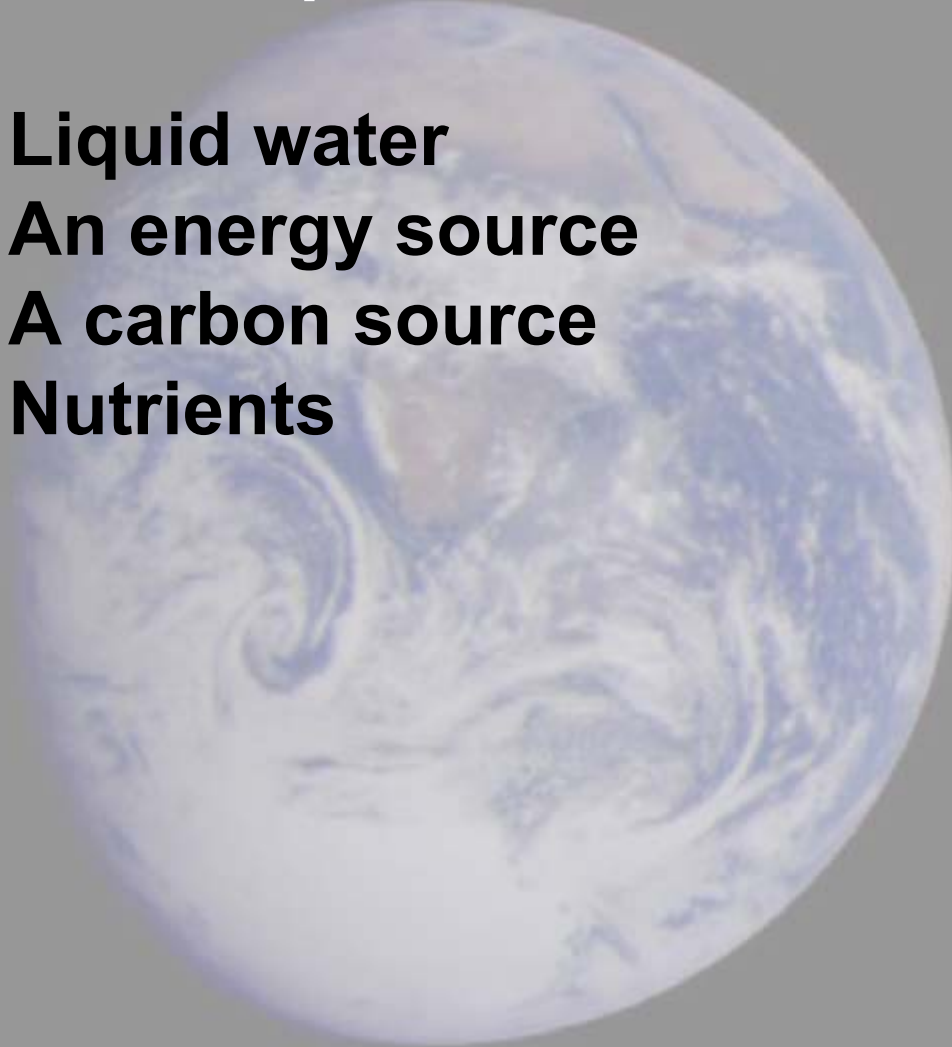
Life needs :

- **Liquid water**
- **An energy source**
- **A carbon source**
- **Nutrients**

→ Conditions suitable for spontaneous generation and reproduction of life

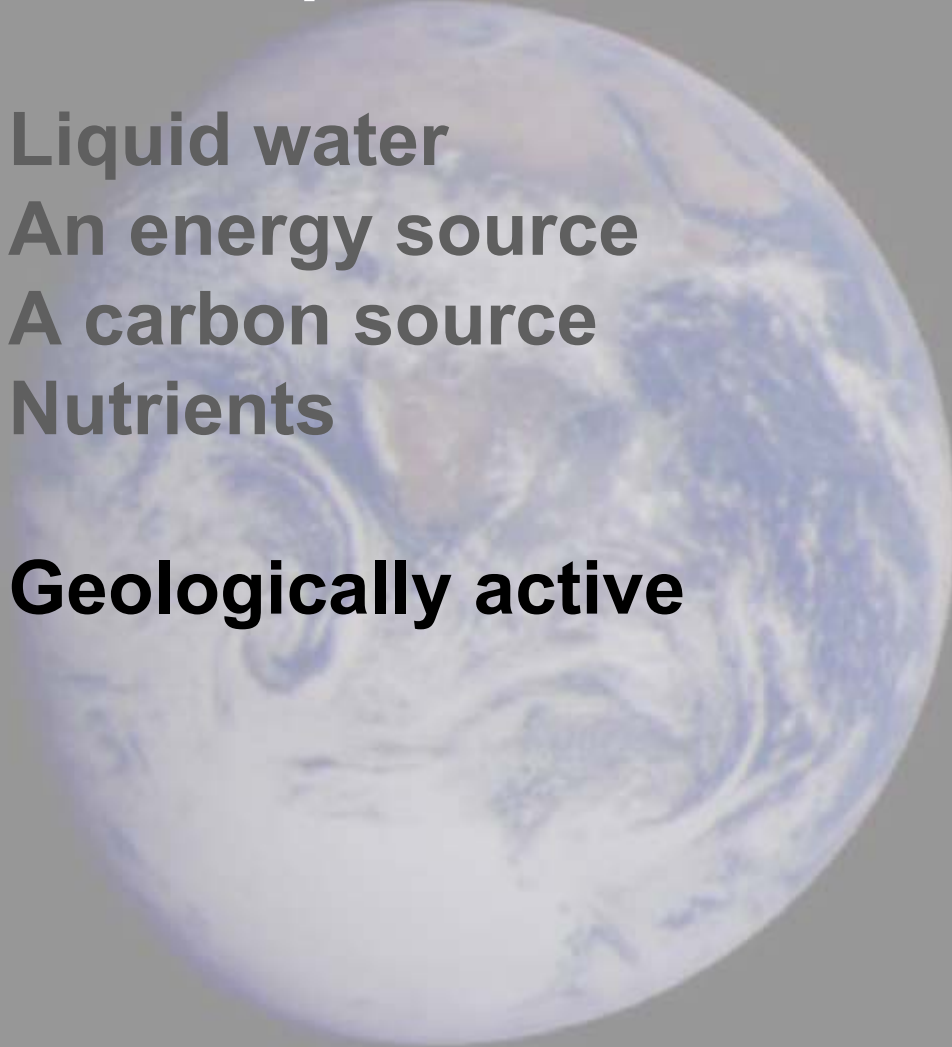
An habitable planet should have :

- **Liquid water**
- **An energy source**
- **A carbon source**
- **Nutrients**



An habitable planet should be :

- Liquid water
- An energy source
- A carbon source
- Nutrients
- **Geologically active**



Survival conditions

**Life can do without liquid water,
energy and carbon for a certain
period of time**

HABITABILITY IN TIME

TODAY

EARTH

MARS ?

EUROPA ?



HABITABILITY IN TIME

TODAY

EARTH
MARS ?
EUROPA ?

4 b.y. AGO

EARTH
MARS ?
VENUS ?
EUROPA ?



THE HABITAT OF LIFE ON EARTH TODAY

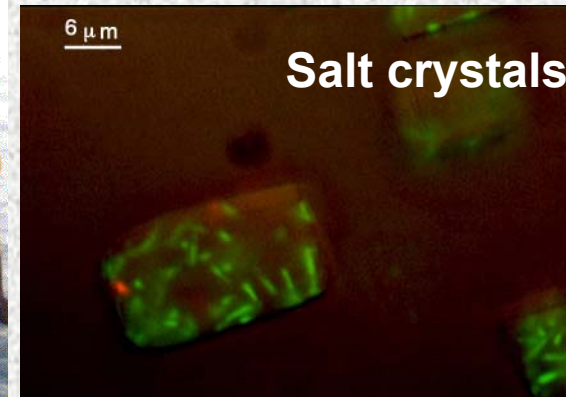
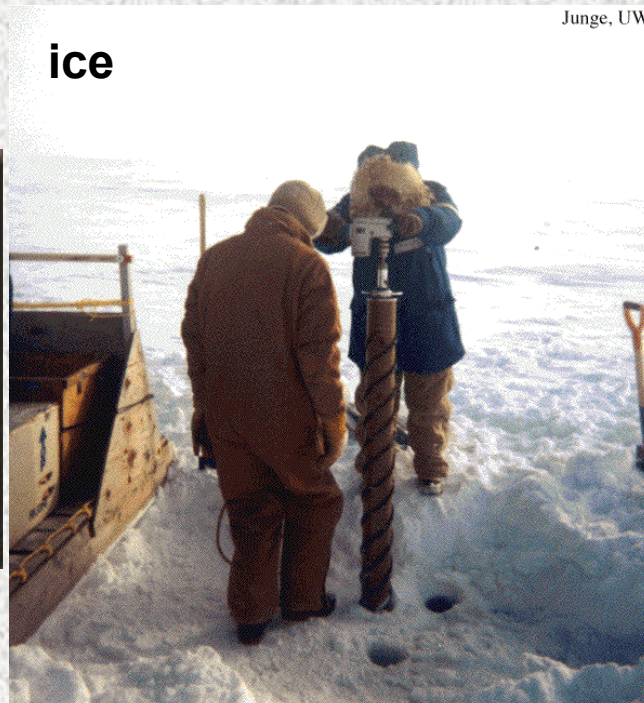
Average environmental conditions:

- **Temperature** **15°C**
- **pH** **7.2-7.4**
- **Atmosphere** **21% O₂**
- **Radiation** **1W/m²**
(DNA-weighted UV)

LIFE ON EARTH TODAY

Life in extreme conditions

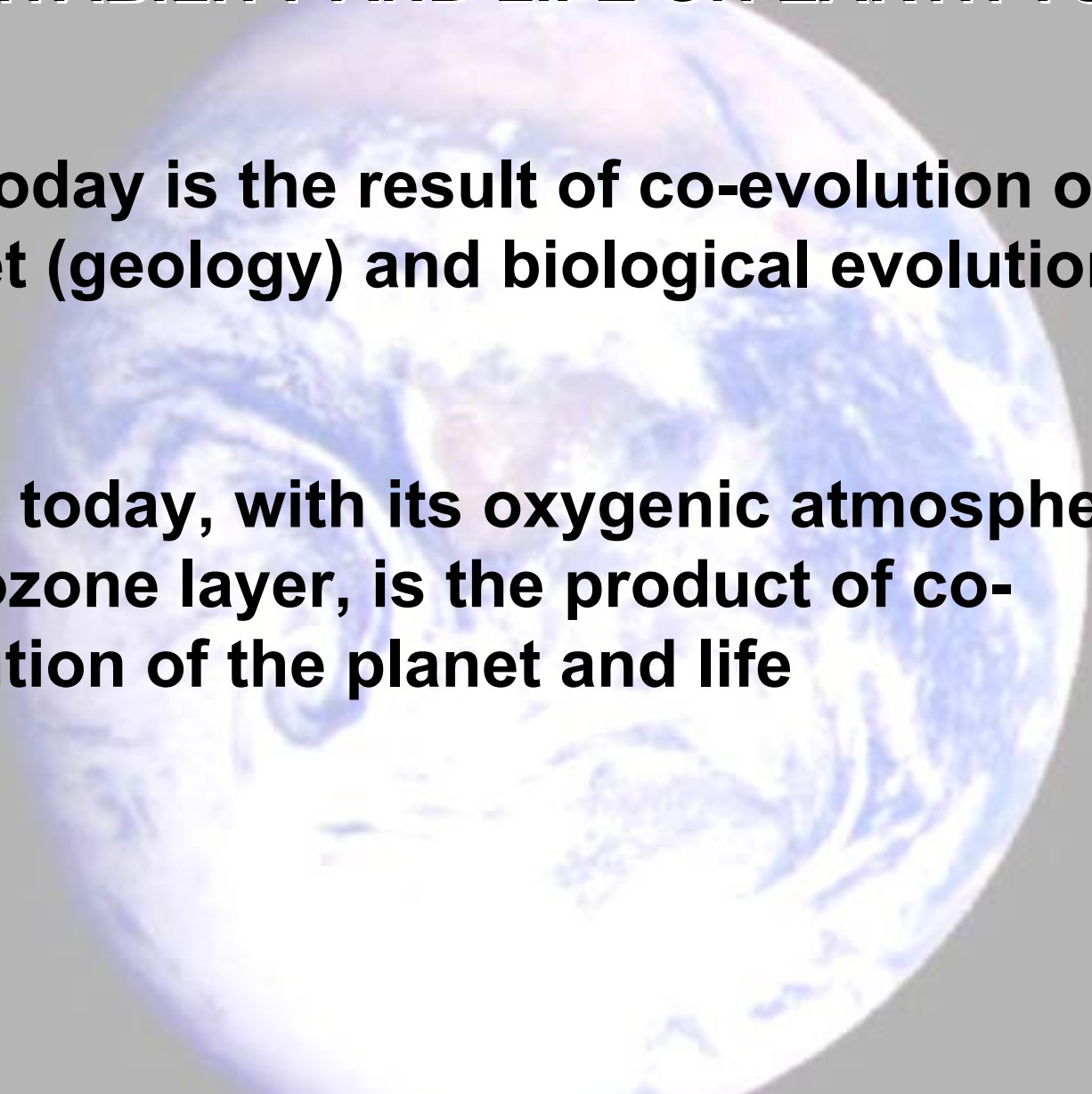
- Life can resist almost all extreme conditions
 - except high temperatures $> \sim 120^{\circ}\text{C}$
 - not for very extended periods of time ($> 10^6$ y?).



HABITABILITY AND LIFE ON EARTH TODAY

Life today is the result of co-evolution of the planet (geology) and biological evolution.

Earth today, with its oxygenic atmosphere and ozone layer, is the product of co-evolution of the planet and life



HABITABILITY AND LIFE ON EARTH TODAY



Life today is the result of co-evolution of the planet (geology) and biological evolution.

Earth today, with its oxygenic atmosphere and ozone layer, is the product of co-evolution of the planet and life

**Understanding geological evolution
is very important**

A satellite view of Earth showing swirling cloud patterns over the continents. The image is centered on the Atlantic Ocean, with North and South America visible on the left and Europe and Africa on the right. The clouds are white and blue, creating a complex, swirling pattern. The text is overlaid in the center of the image.

**HABITABILITY AND
GEOLOGICAL EVOLUTION
ON EARTH**

HABITABILITY AND LIFE ON EARTH 3.5 GA



Only 2 areas in the world that contain ancient, well-preserved rocks:

Barberton, South Africa

The Pilbara, Australia

→ Problem of lack of data points

HABITABILITY AND LIFE ON EARTH 3.5 GA

average environmental conditions:

	3.5 Ga	TODAY
Temperature	> 50°C	15°C
pH	5-6	7.2-7.4
Atmosphere	<0.2% O₂	21% O₂
Radiation	54 W/m²	1W/m²
	(DNA-weighted UV)	

HABITABILITY AND LIFE ON EARTH 3.5 GA

Signatures of life ?

- **Extant – extinct**
- **Structural**
- **Biogeochemical**

HABITABILITY AND LIFE ON EARTH 3.5 GA

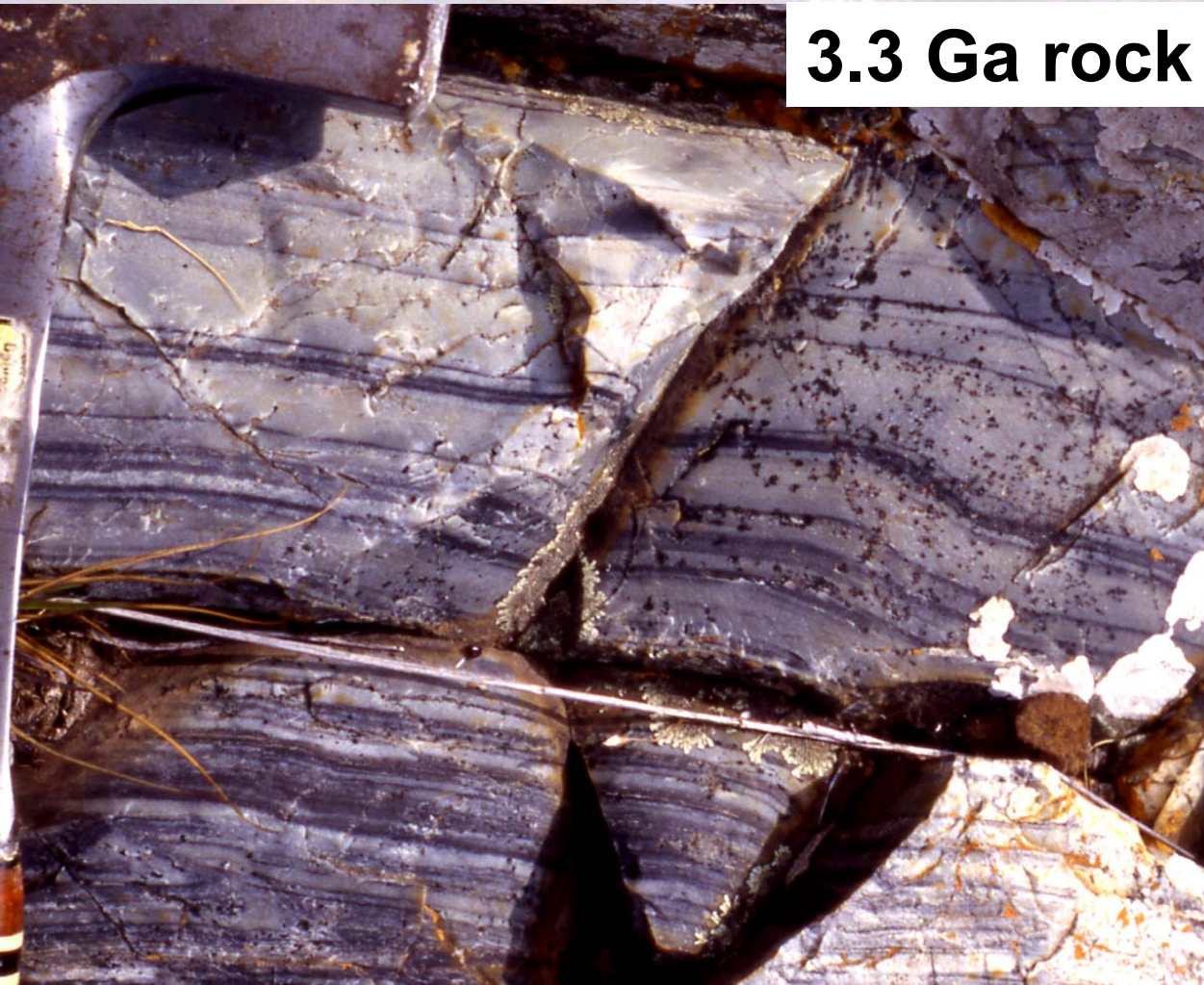
Abundance of life:

➤ Far lower biomass than on present Earth
(more primitive type of metabolism)

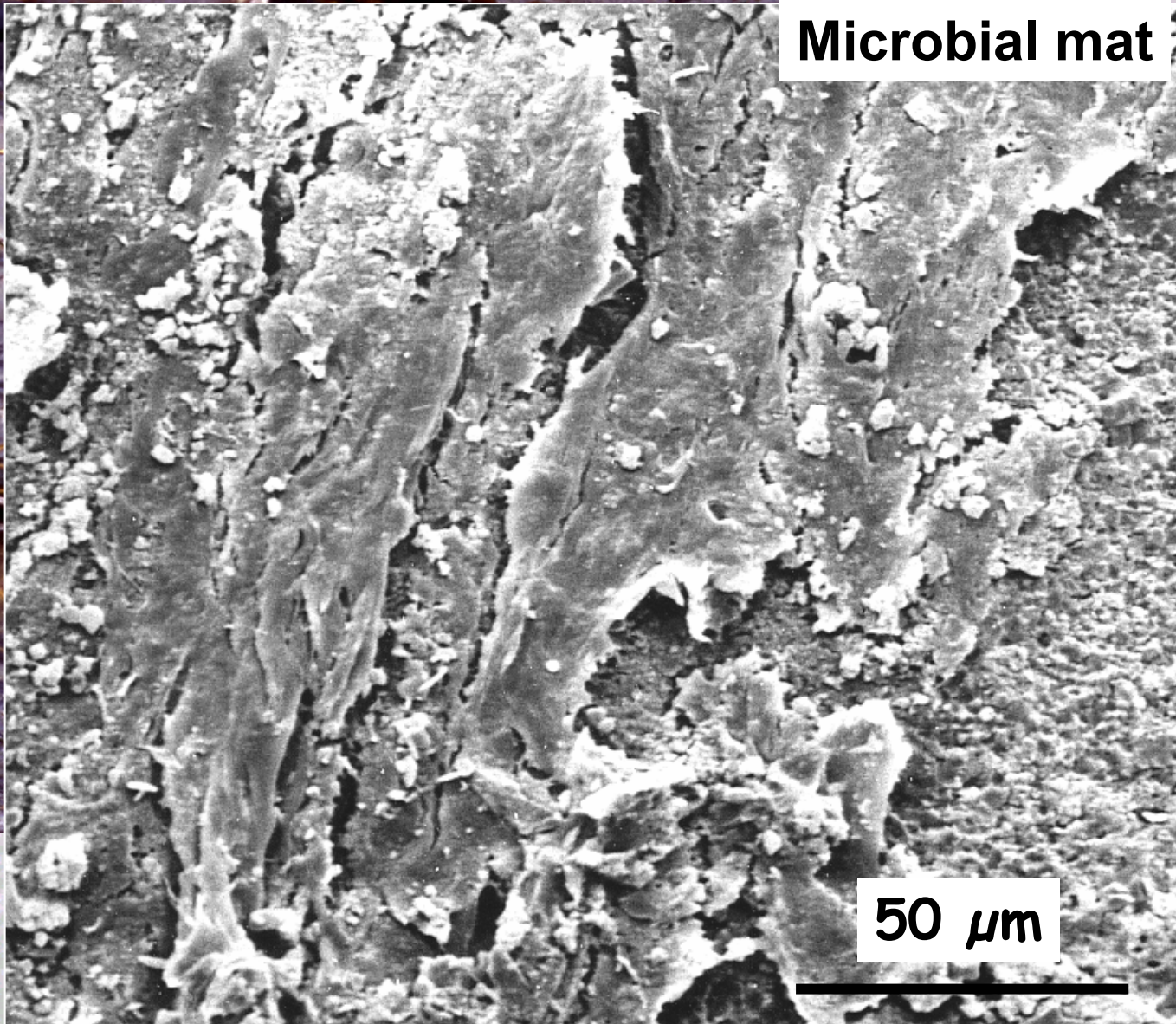
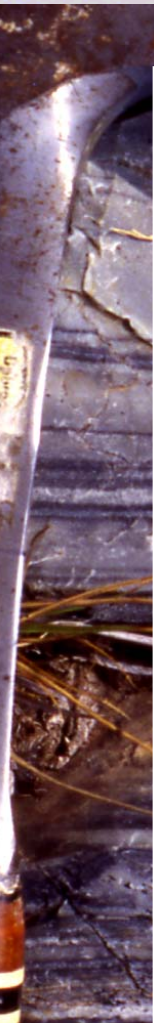
➔ **Manifestations of life far fewer and more subtle**

HABITABILITY AND LIFE ON EARTH 3.5 GA

3.3 Ga rock

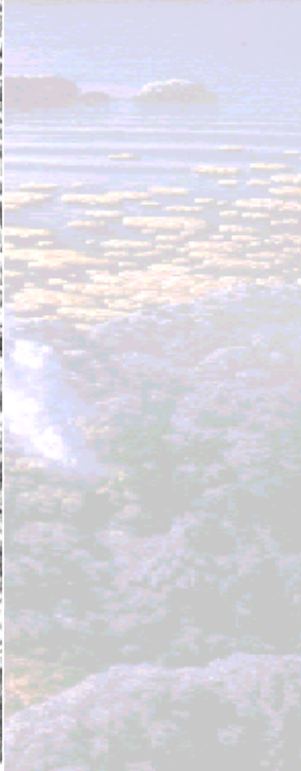


HABITABILITY AND LIFE ON EARTH 3.5 GA

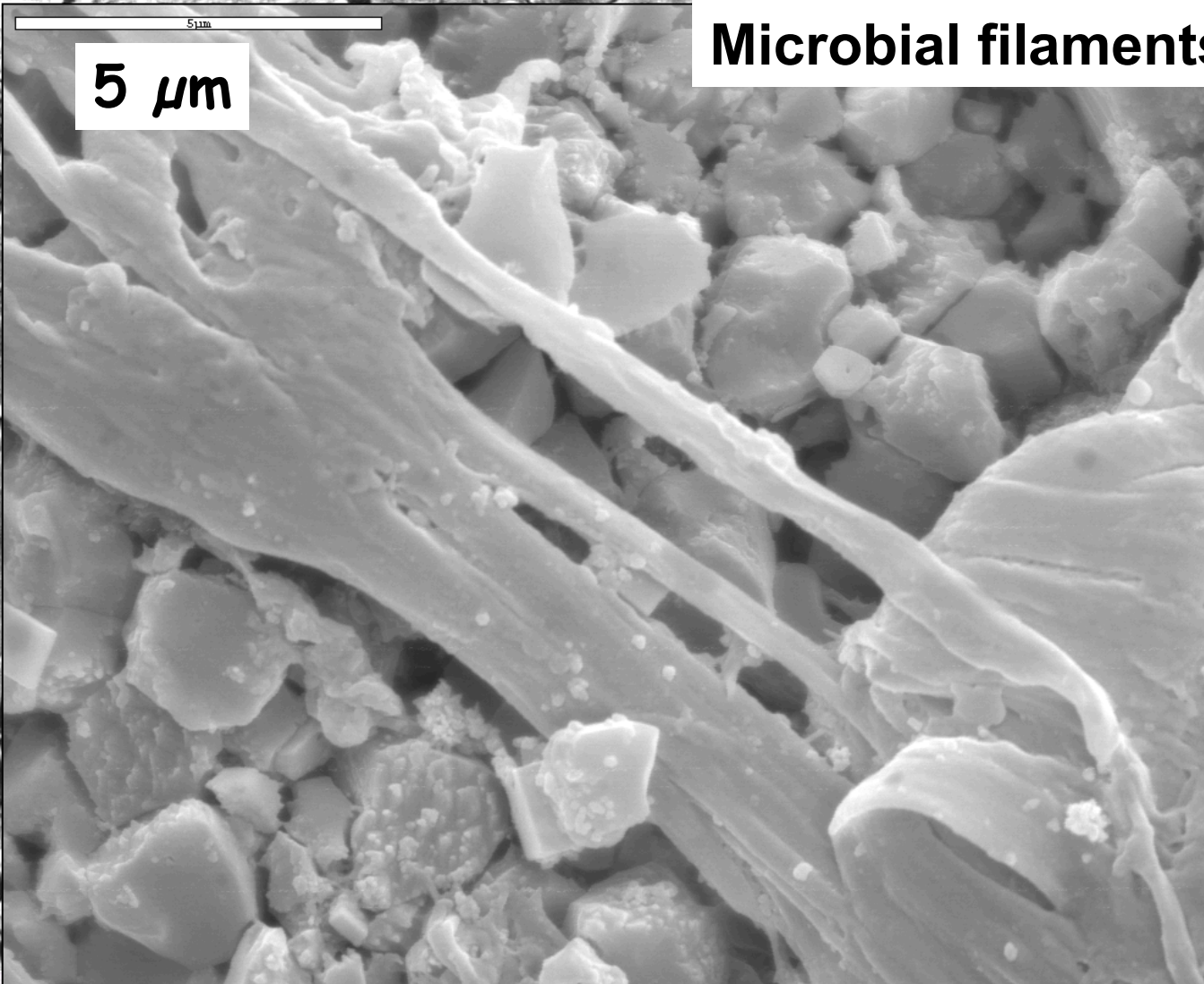
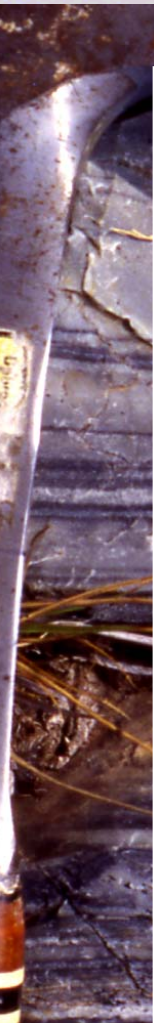


Microbial mat

50 μm

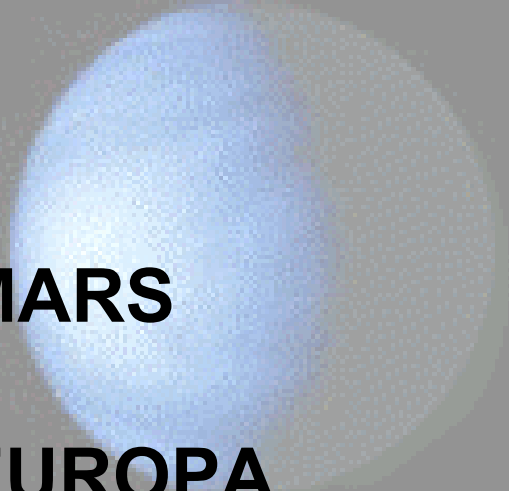


HABITABILITY AND LIFE ON EARTH 3.5 GA



5 μm

Microbial filaments

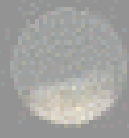
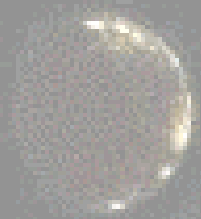
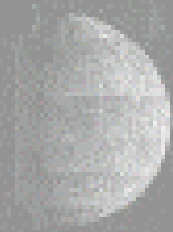
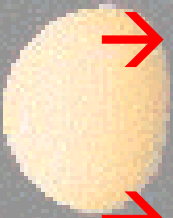


→ **MARS**

→ **EUROPA**

→ **(VENUS?)**

→ **OTHER?**



A large, reddish-orange planet, Mars, is shown against a black background. The planet's surface is textured with various shades of red, orange, and brown, indicating different geological features. The word "Mars" is written in a bold, red, sans-serif font with a white outline, centered on the planet's surface.

Mars

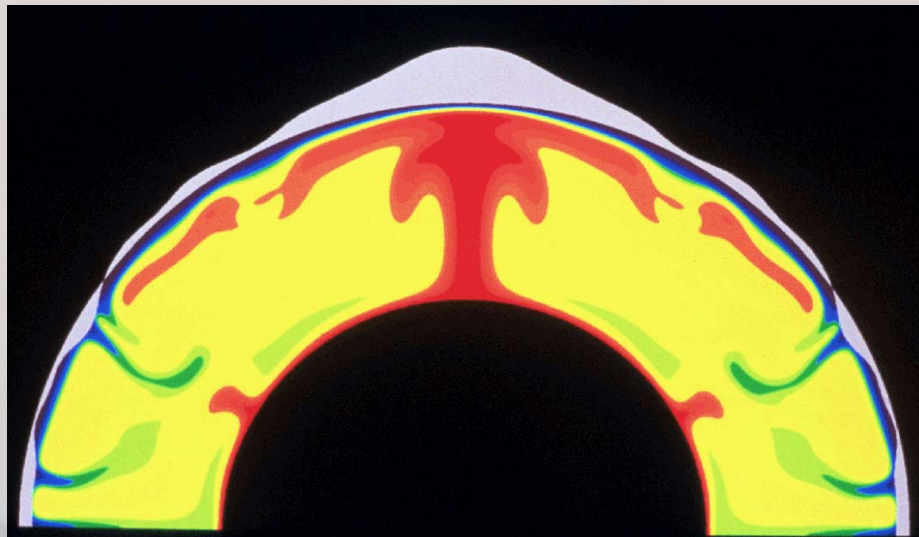
HABITABILITY OF A PLANET PAST AND PRESENT?

- 1. The geological characteristics of the planet (structure, dynamics, composition) throughout its history**
- 2. History of changing environmental conditions**
- 3. Was there/is there life on the planet (spontaneous generation or transported from elsewhere)?**

HABITABILITY OF MARS PAST AND PRESENT?

1. STRUCTURE

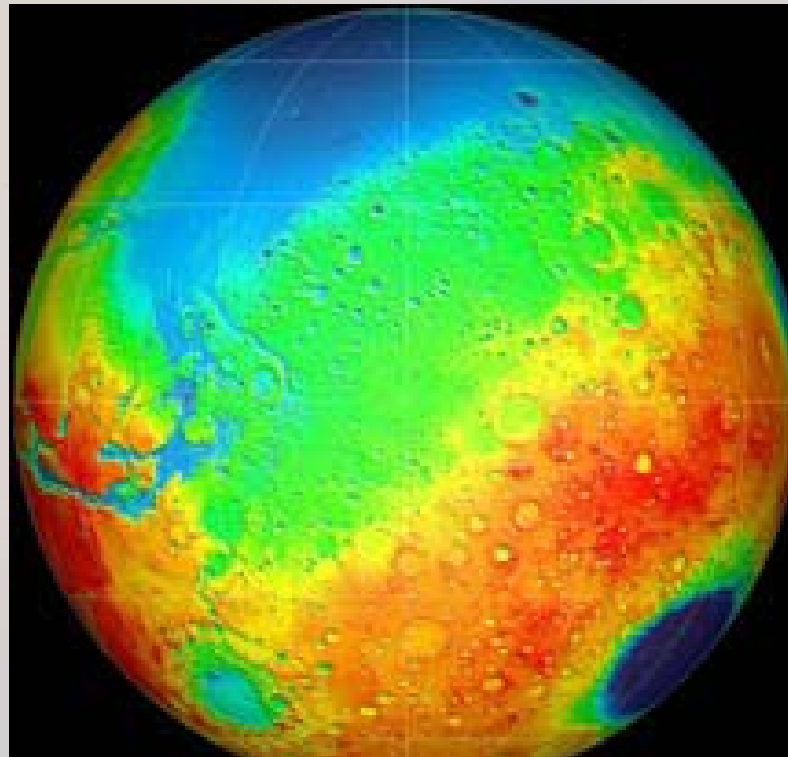
- Core, mantle, crust structure
- One plate planet



HABITABILITY OF MARS PAST AND PRESENT?

1. STRUCTURE

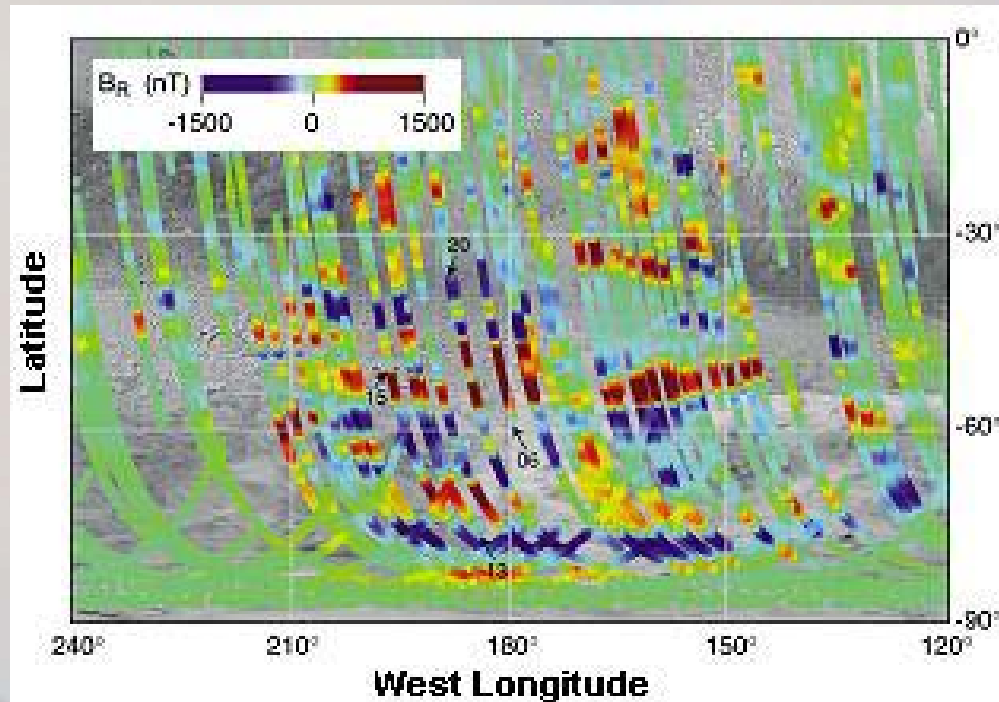
- N/S crustal dichotomy



HABITABILITY OF MARS PAST AND PRESENT?

2. DYNAMICS

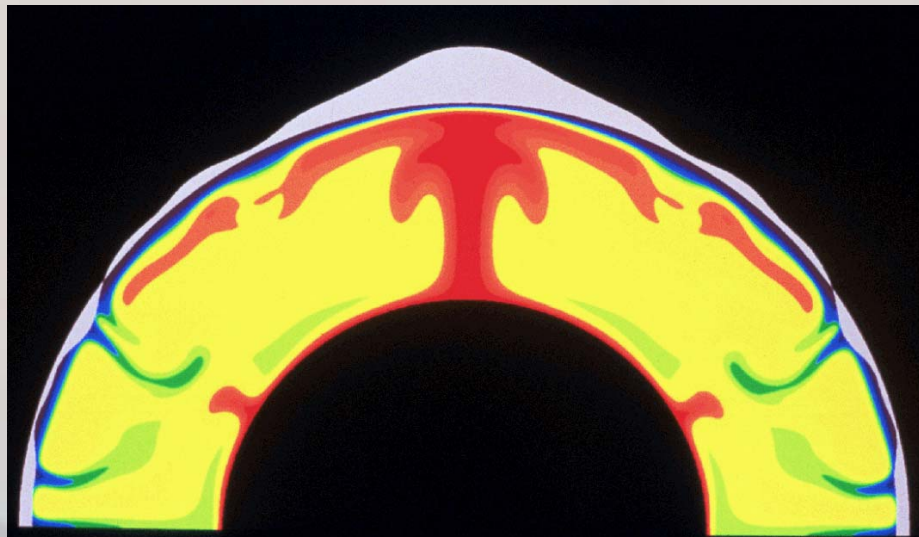
- Pre 4 Ga magnetic field



HABITABILITY OF MARS PAST AND PRESENT?

2. DYNAMICS

- No lateral plate tectonics
- Volcanism (up to recent)



HABITABILITY OF MARS PAST AND PRESENT?

3. COMPOSITION

- Fe-rich mantle
- Basaltic crustal rocks



HABITABILITY OF MARS PAST AND PRESENT?

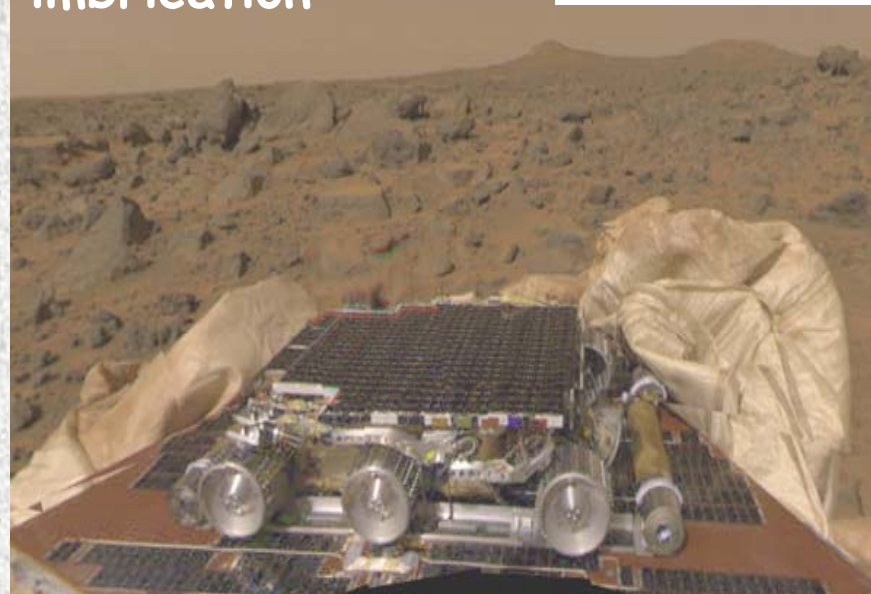
ENVIRONMENTAL CONDITIONS

- **Early Mars was water-rich**
- **Carbon**
- **Energy sources**
(geochemical, hydrothermal, solar)
- **Nutrients**

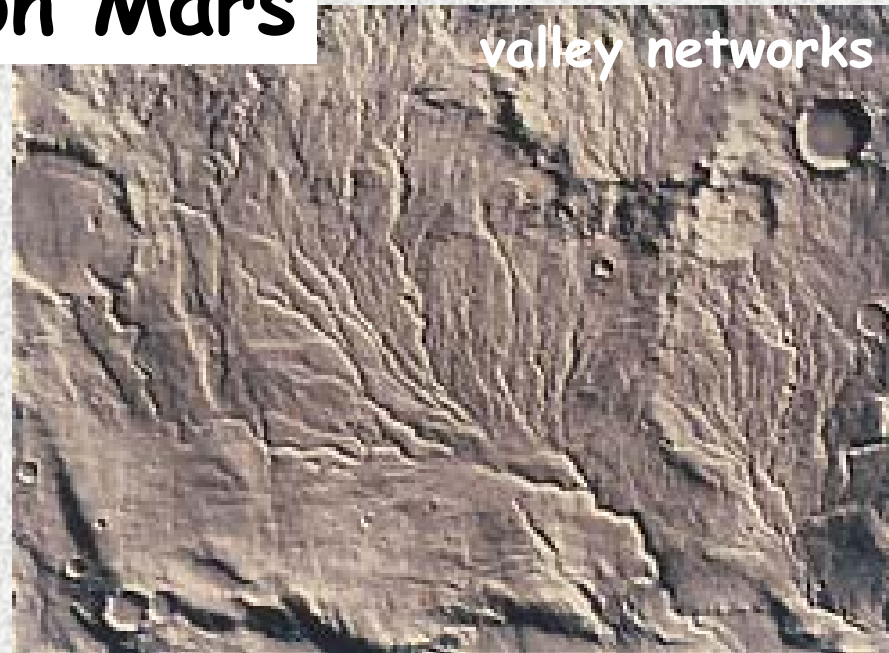
(radiation $\sim 54\text{W/m}^2$)

Water on Mars

imbrication



valley networks



outflow channel & craters



streamlined islands



HABITABILITY OF MARS PAST AND PRESENT?

**EARLY MARS WAS
HABITABLE**

HABITABILITY OF MARS PAST AND PRESENT?

EARLY MARS WAS HABITABLE

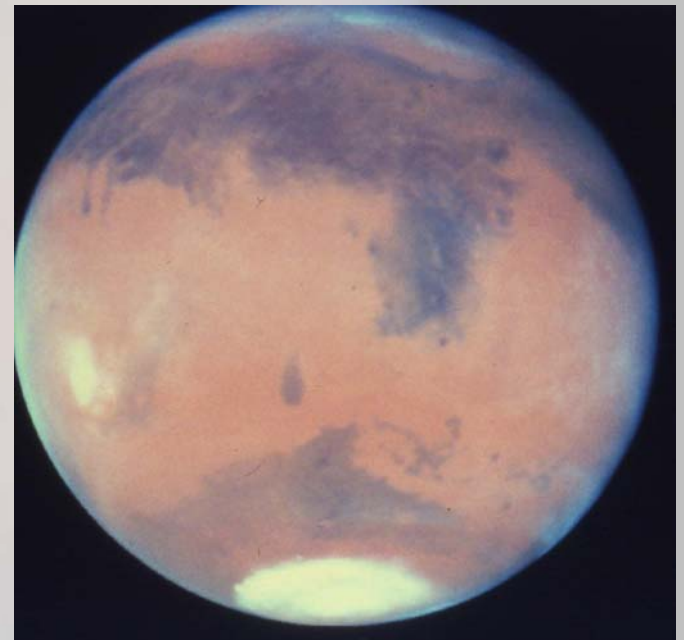
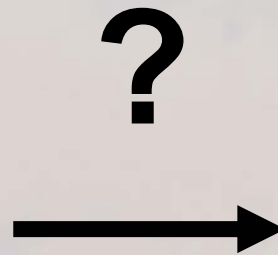
- On the surface and subsurface**
- By primitive microorganisms**
- Signatures of life will be subtle**

CHANGING ENVIRONMENTAL CONDITIONS

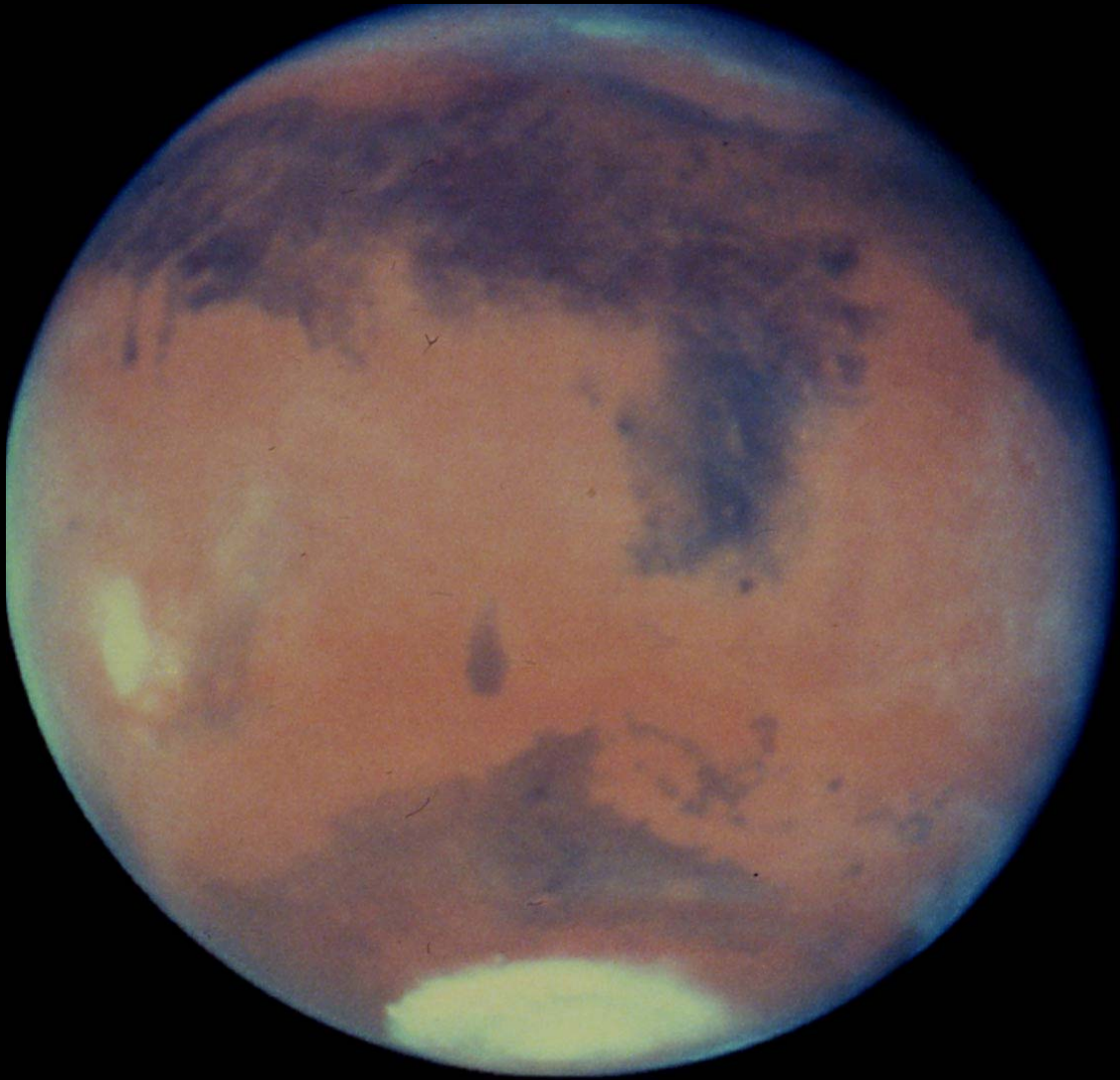
➤ **Loss of surface volatiles ~3.8 Ga**

→ **Freezing and drying of planet**

→ **Cryosphere formation**



HABITABILITY OF MARS TODAY?



HABITABILITY OF MARS TODAY?

ENVIRONMENTAL CONDITIONS

- **Some water (mostly frozen)**
- **Carbon**
- **Energy sources**
(geochemical, hydrothermal, solar)
- **Radiation $\sim 54\text{W/m}^2$**
- **Limited geological renewal of nutrients**

HABITABILITY OF MARS PAST AND PRESENT?

MARS IS HABITABLE

- **In melt pockets in the subsurface cryosphere**
- **Organisms using a very primitive metabolism**

ASSESSING THE HABITABILITY OF MARS PAST AND PRESENT?

- **Understand the workings of the planet
(life cannot last long on a dead planet)**
- **Much more information about the geological/
environmental history of the planet**
 - Impact history
 - Volcanic/hydrothermal activity
 - Sedimentological environment
 - Volatile inventory

ASSESSING THE HABITABILITY OF MARS PAST AND PRESENT?

- **Did life appear on Mars? What was its fate?**
 - ➔ **search for biosignatures
(structural, biogeochemical)**
- **What was/is the nature of martian life?
(similar to terrestrial life or different
e.g. different chirality, biomolecules?)**
- **Distribution of life, past and present?**

ASSESSING THE HABITABILITY OF MARS PAST AND PRESENT?

Major requirements:

- | | |
|--------------------|---|
| Orbital
mapping | <ul style="list-style-type: none">- morphology- mineralogy- water (H, liquid/ice H₂O)- location of hot spots
(present volcanic activity)- CH₄ |
|--------------------|---|

ASSESSING THE HABITABILITY OF MARS PAST AND PRESENT?

Major requirements:

In situ
studies

- global seismic structure
- geomorphology

RETURN
SAMPLES

- mineralogy
- petrography
- chemistry
- biochemistry
- micropalaeontology
- dating

surface/
subsurface

- local weather
- local atmosphere
- radiation levels

Human
exploration

ASSESSING THE HABITABILITY OF MARS PAST AND PRESENT?

- *All require dated rocks
(in situ methods or sample return)*
- *Instruments have to be able to travel*
- *Context is of paramount importance
- difficult in subsurface*

ASSESSING THE HABITABILITY OF MARS PAST AND PRESENT?

- **Mars can give us information about the geological and environmental conditions during the earliest evolution of the terrestrial planets**
- **Mars has rocks old enough to contain evidence of the origin and earliest evolution of life**
 - **distinction between inert/living**
 - **testing panspermia**



VENUS

VENUS



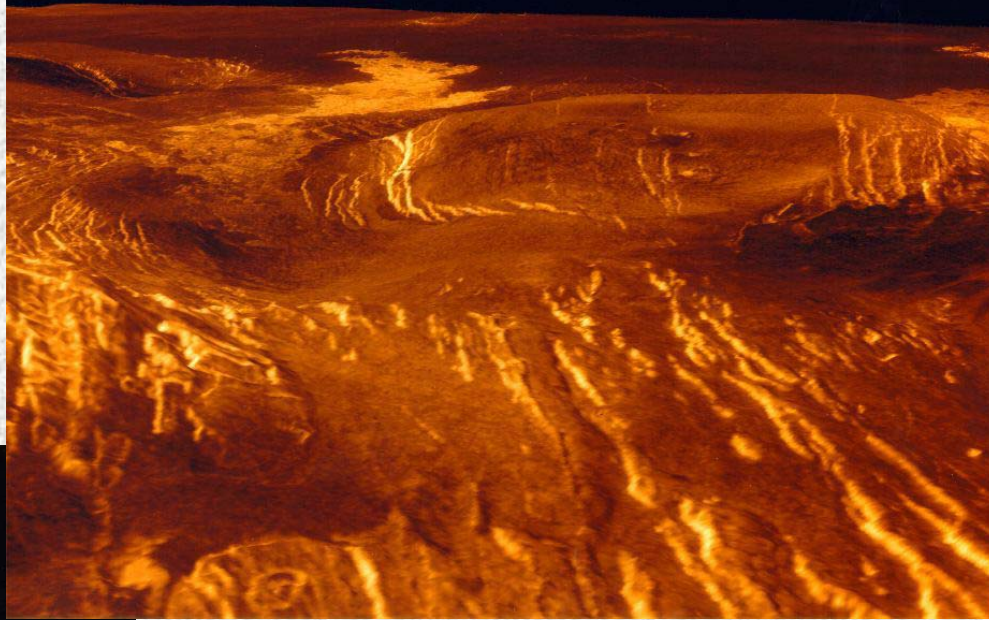
- **Terrestrial planet**
- **Had an early geological/environmental evolution similar to the Earth**
- **Had water, organics, energy**
- **Was an HABITABLE planet**
- **May have been able to support life**

VENUS



- Runaway greenhouse effect → surface temperatures $> 480^{\circ}\text{C}$
- No liquid water nor usable organics (at surface)
- Is **not** an HABITABLE planet now (surface)
- Volcanically active planet
- Coronae structures (c.f. early Earth?)

**Maat Mons
volcano**



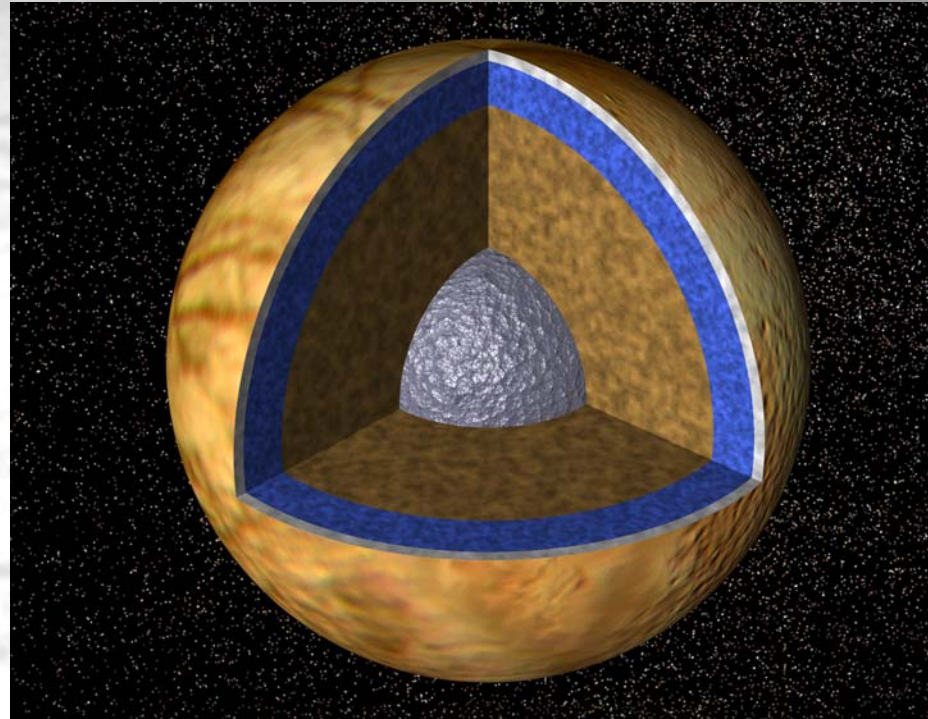
Corona structure

EUROPA



EUROPA

- Core (Fe rich)
- Silicate mantle
- Subsurface ocean
- 100-150 km ice crust
- Gravity field



EUROPA – PAST AND PRESENT HABITABILITY ?

2. Icy surface

- Structurally complex – cryovolcanism
- Dark patches = sulphate salt deposits on surface

3. Geological evolution

- Cooling down of planet
- Internal heat source
(radioactive decay or heat provided by flexing caused by Jupiter?)

EUROPA – PAST AND PRESENT HABITABILITY ?

- **Liquid water**
- **Organics**
- **Energy (thermal, geochemical)**
- **Nutrients**

Europa was/is potentially habitable

OTHER BODIES

GASPRA



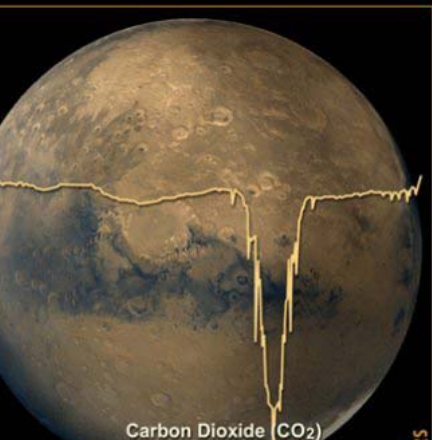
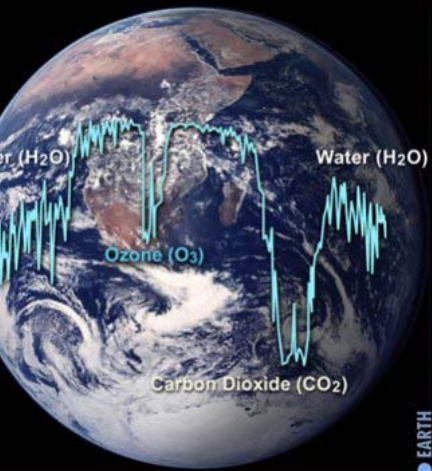
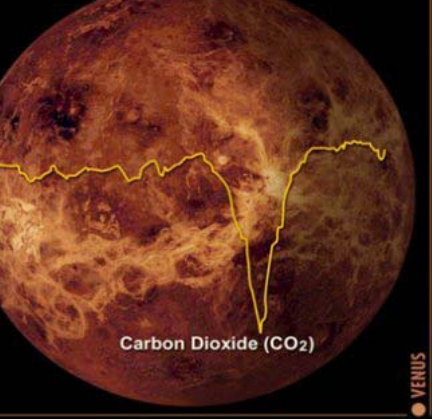
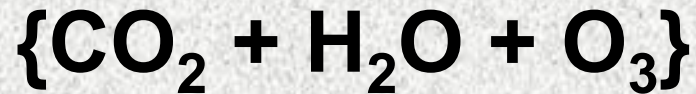
**Prebiotic molecules
- building bricks of life**

Panspermia ?

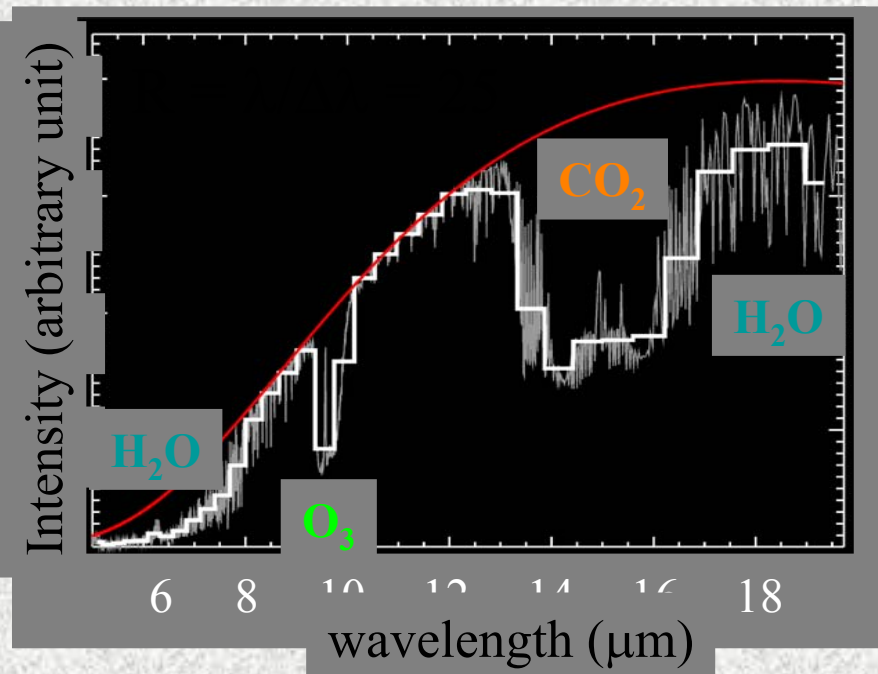
ALH84001,0



Detection and characterisation of exoplanets



Analysing the light from the planet with a spectrograph allows identification of key gases



The background of the slide is a dark grey-blue space filled with various celestial bodies. On the left is a large, detailed Earth showing continents and clouds. To its right is a large, smooth, light blue sphere. Further right is a smaller, cratered Moon. Below these are several other celestial bodies of varying sizes and colors, including a yellowish-orange sphere, a greyish sphere, and a bright yellow-orange sphere. The overall composition is a collection of diverse planets and moons.

THANK YOU TO:

Alain Léger

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Helga Stan-Lotta