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Mars Exploration Rover

STRENGTH OF SOIL DEPOSITS ALONG MER TRAVERSES

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1: DLR Institute of Space Simulation

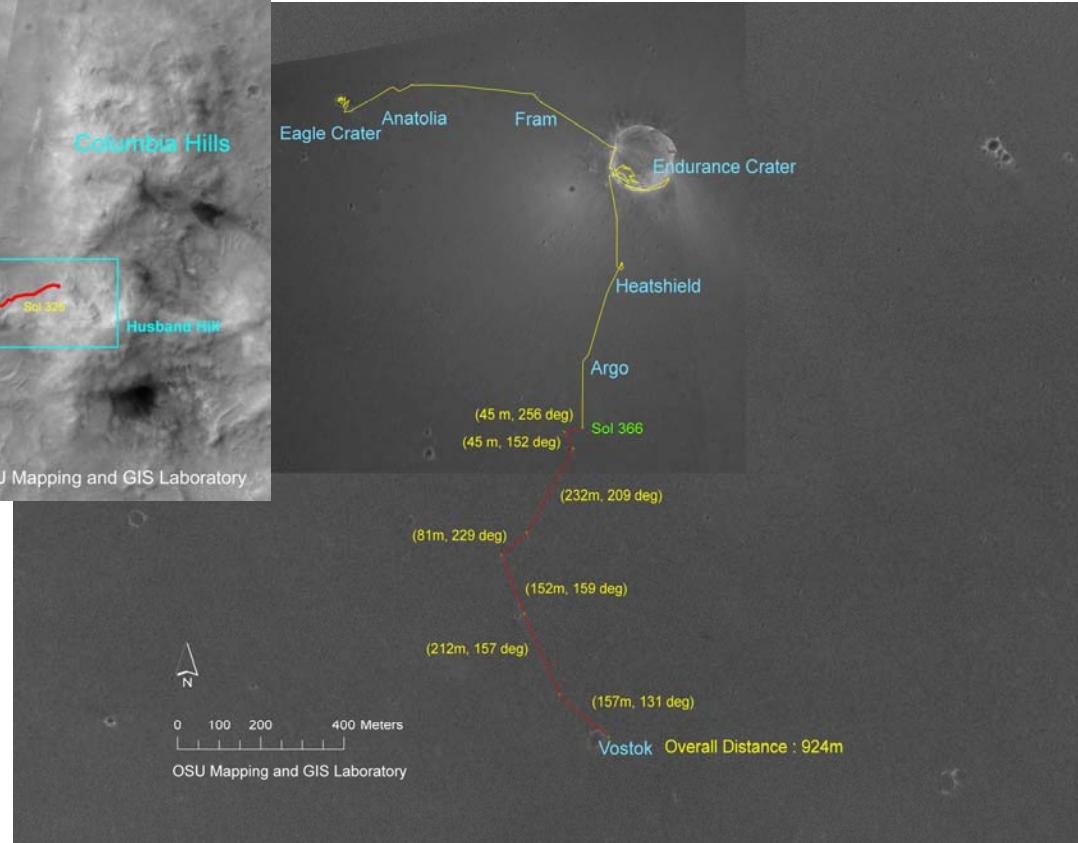
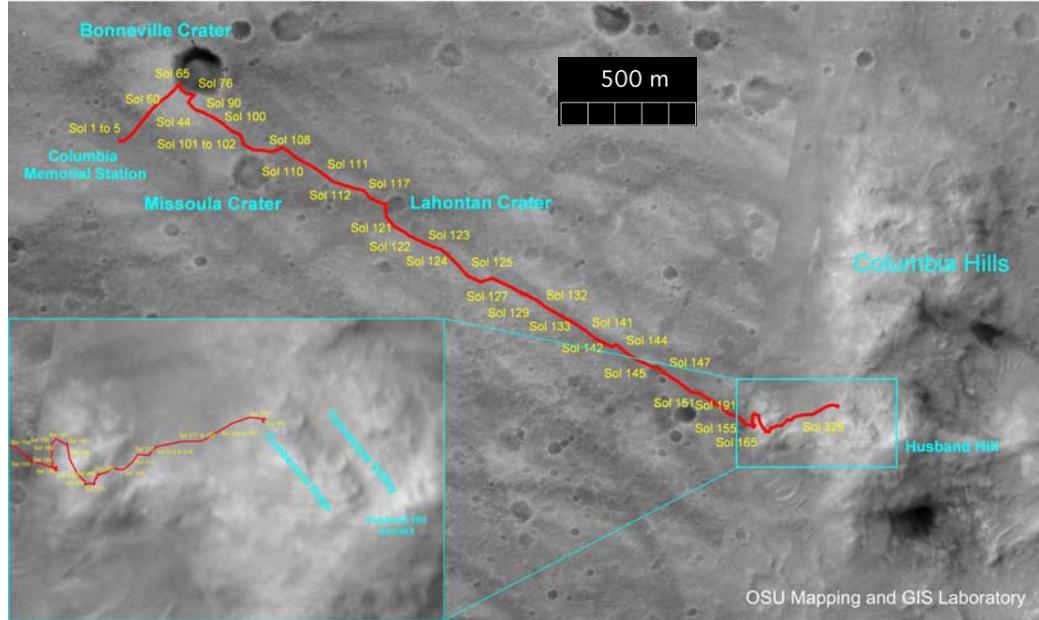


Two traverses on Mars

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Spirit Rover Traverse Map (Sol 343)

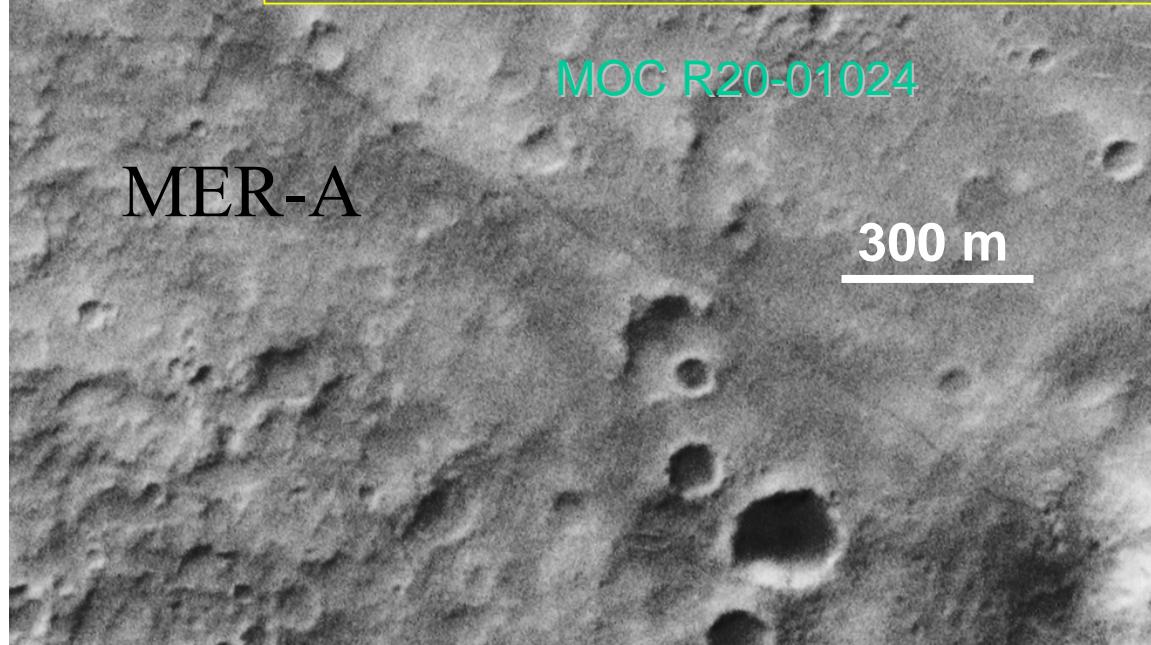
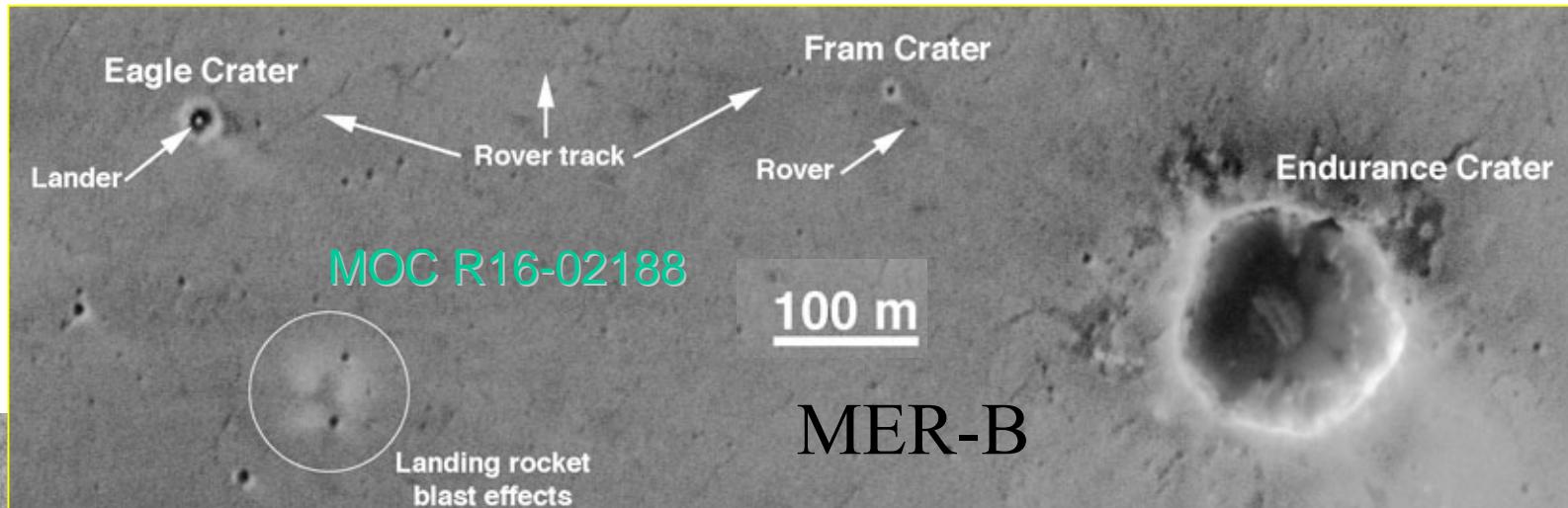




Wheel tracks from orbit

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MER physical properties studies **JPL**

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- **Relevance of physical properties: knowledge of strengths of soils and rocks, and observation of physical processes, to:**
 - identify different materials from their strengths
 - compare physical and compositional properties
 - correlate materials with geologic units
 - identify stratigraphic relationships
 - identify modification and transport processes
- **No dedicated instruments -> works across multiple disciplines and payload elements**
- **Reported here: soil strength and other properties from wheel sinkage**



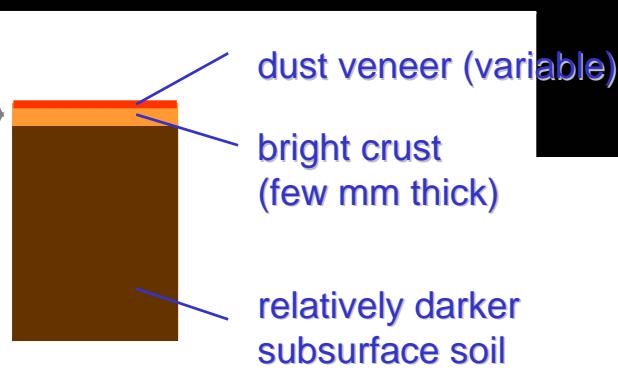
'Spirit' wheel tracks

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MER-A Sol 61 Legacy Pan
(Middle Ground Hollow)

~few mm

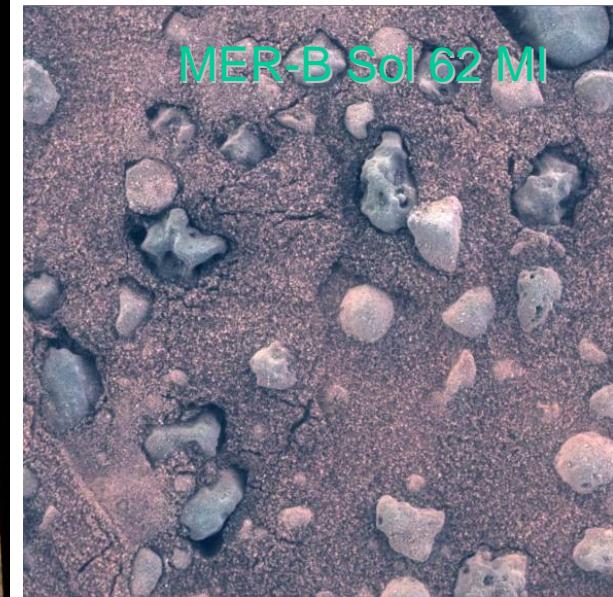
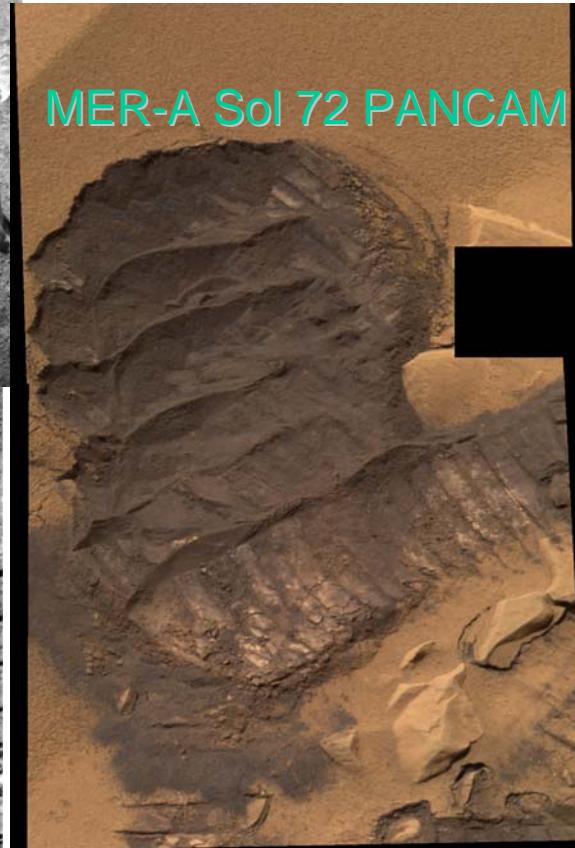
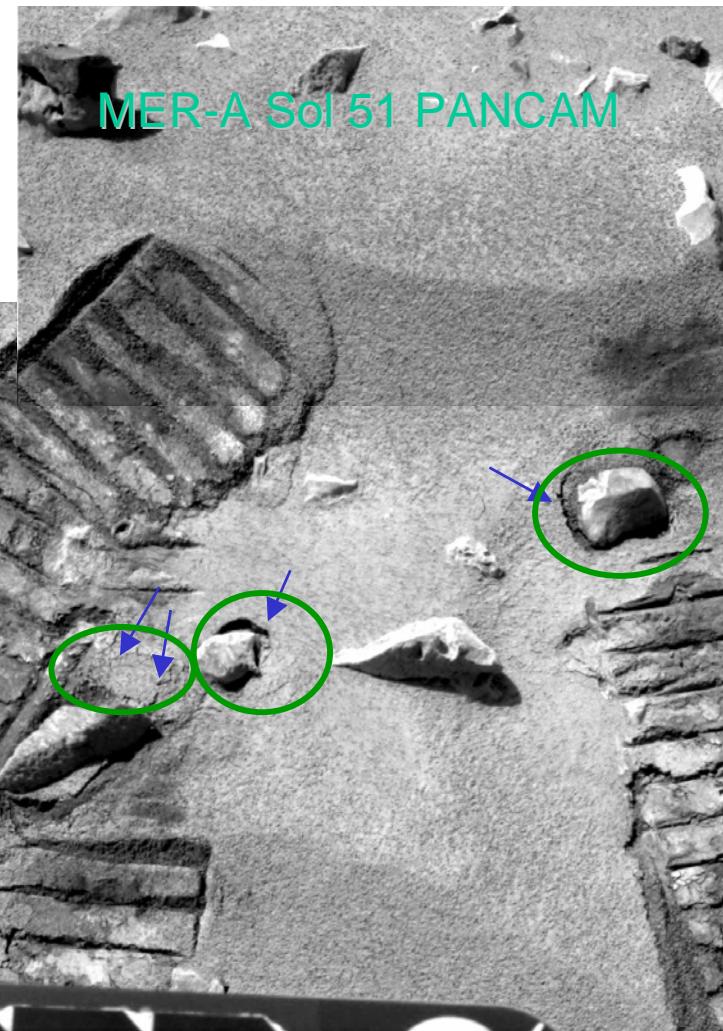




Soil crusts evidence

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30 mm

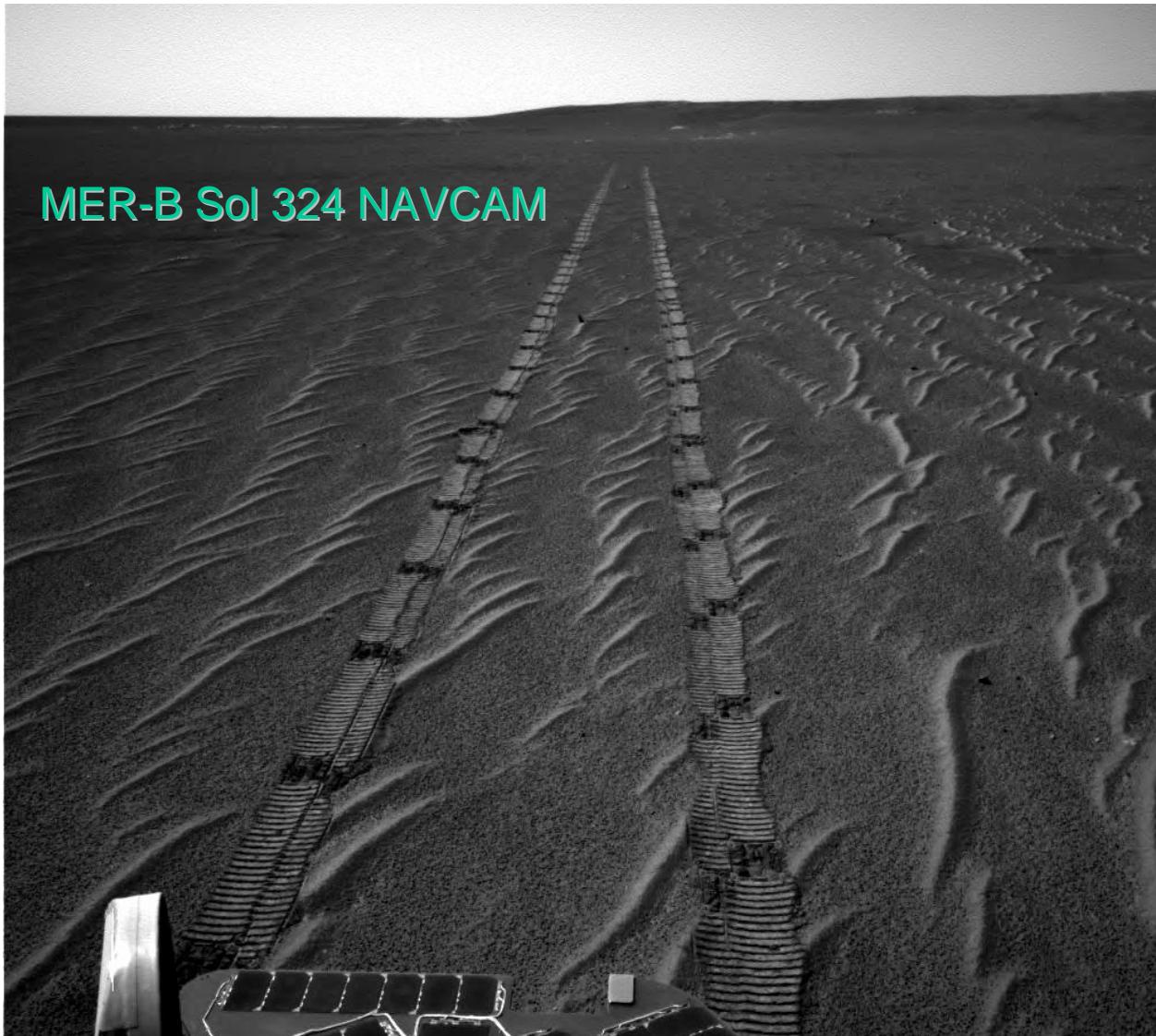
DLR



'Opportunity' wheel tracks

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MER-B Sol 324 NAVCAM





Tracks PANCAM observations **JPL**

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MER-A Sol 42 ,True Color'

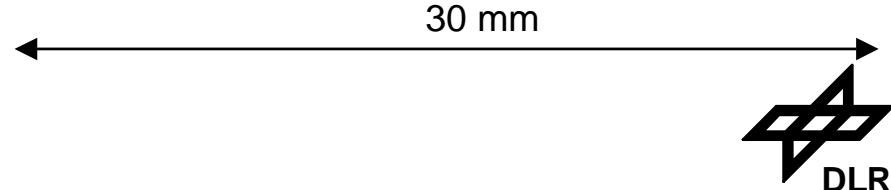




Tracks MI observations

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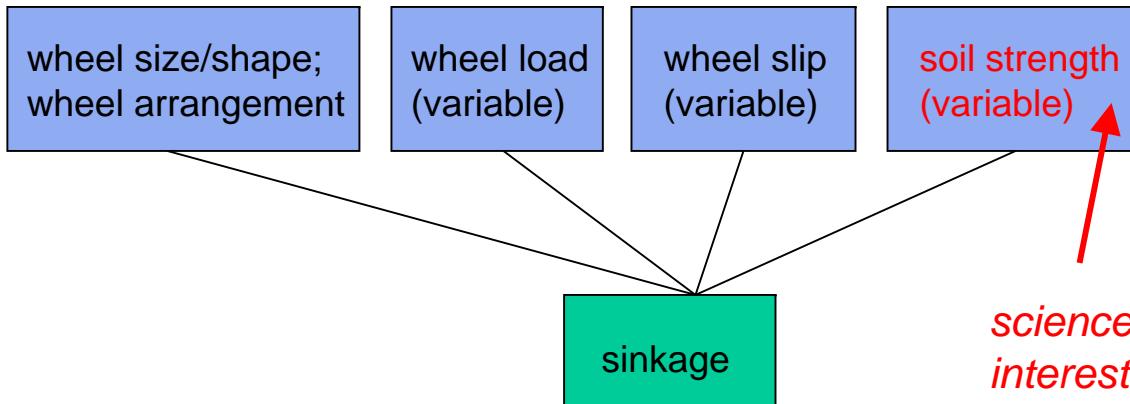
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Soil strength from wheel sinkage **JPL**

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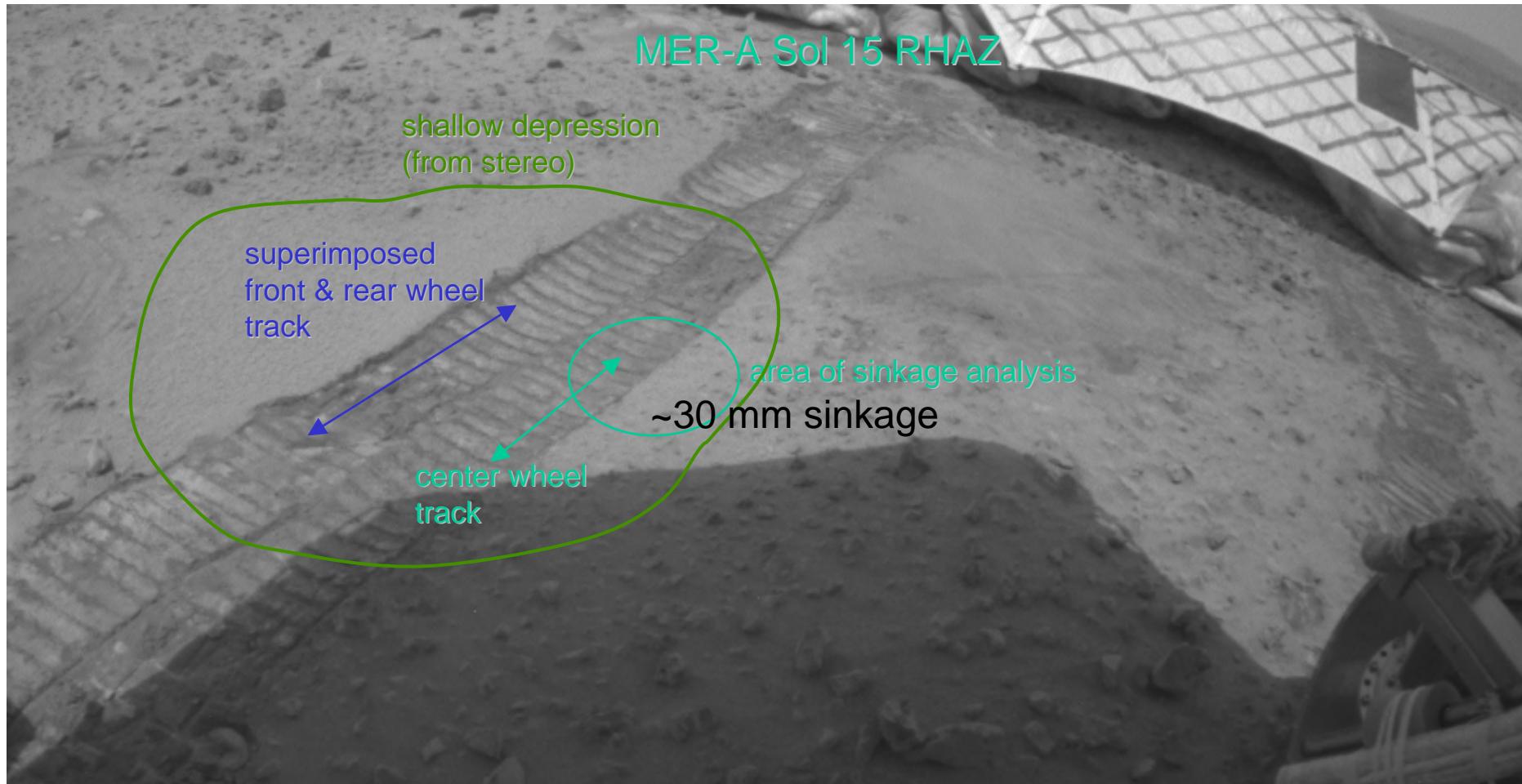


- Measure wheel rut depth in stereo images (< 1 mm error in near-field)
- Wheel-soil theory calibrated for MER wheel: obtain soil strength parameters for vertical loading
- Compare with known soils: estimation of soil strength parameters for shear loading
- Results applicable to upper 20-30 cm of soil (stress dissipation for wheel width)
- Soil strength -> correlated with bulk density -> correlated with thermal inertia and dielectric properties -> comparison with datasets from orbital missions/remote sensing



Soil strength from wheel sinkage **JPL**

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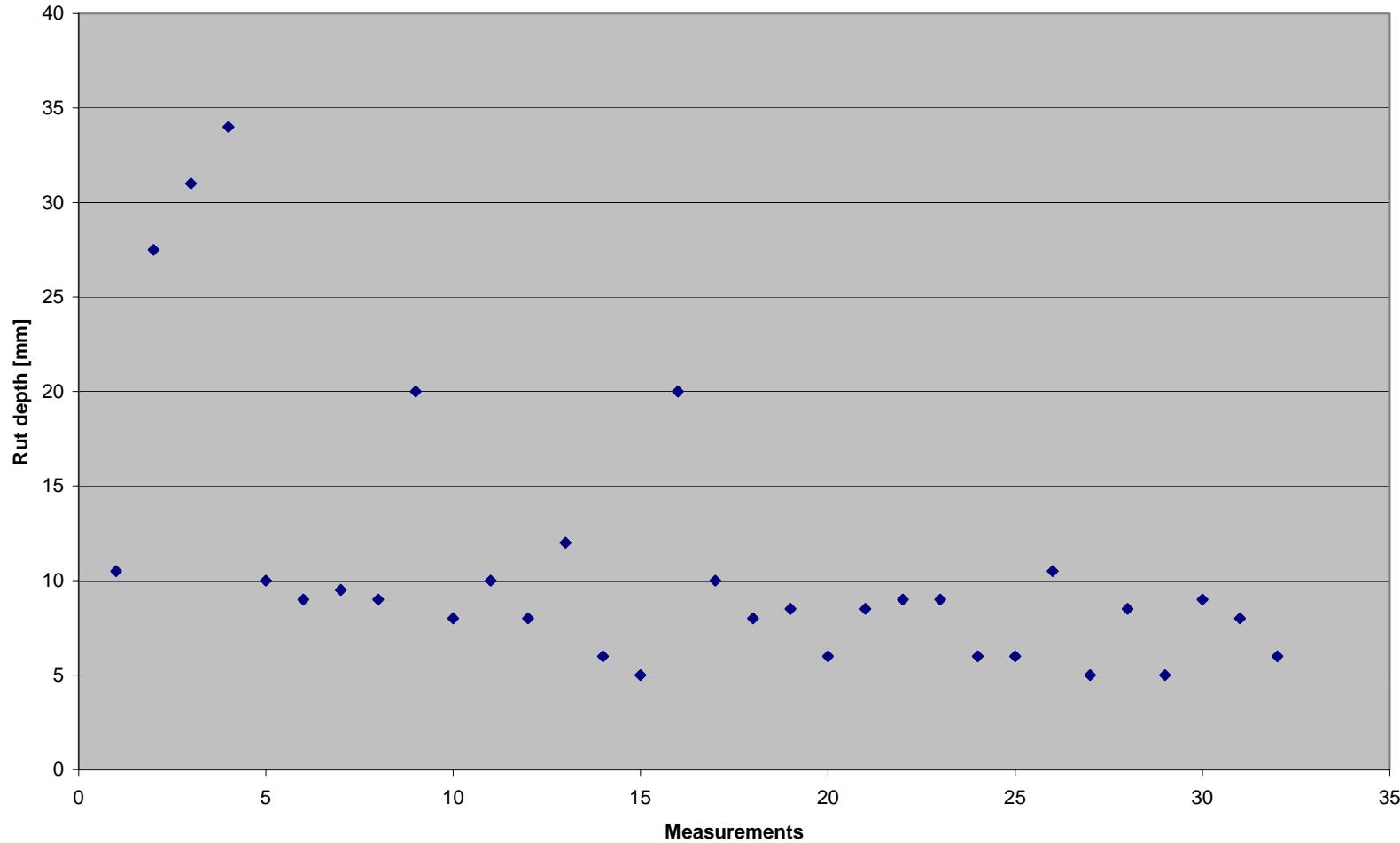


Systematic rut depth results

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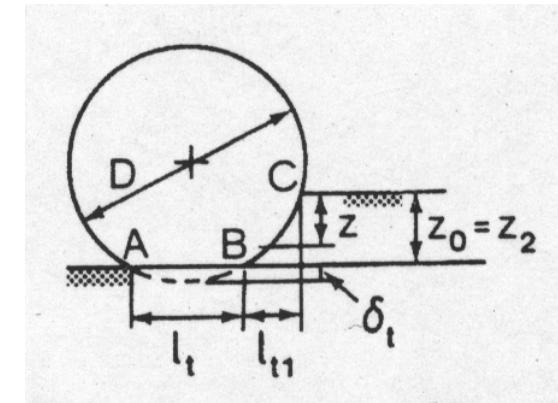
MER-A (through Sol 103)



- Wheel slip estimate from traverse reconstructions (absolute localization vs. wheel turns) -> account for wheel slip-sinkage
- Relation between wheel dimensions (D, b), wheel load (W), soil parameters for vertical loading (k_c, k_ϕ, n) and load-sinkage z_0 :

$$W = -b \cdot k^* \cdot \int_0^{z_0} z^n \frac{2(z_0 - z) - D}{2\sqrt{D(z_0 - z) - (z_0 - z)^2}} dz$$

$$k^* = \frac{k_c}{b} + k_\phi$$



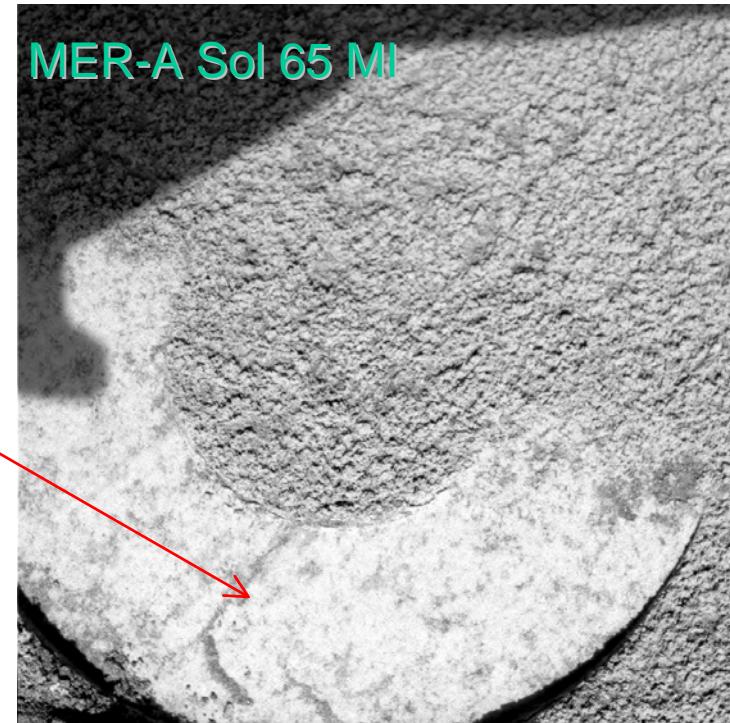
- Caveat: no unique solution for k^*
- Ways out:
 - Additional loading surface in same soil unit (Mössbauer contact plate)
OR
 - Comparison with known soils -> also gives values for shear strength parameters by analogy



MB ,noseprint‘

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30 mm

 DLR



Example: MER-A egress soil

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- Solving for k^* for different values of n (exponent of soil deformation for vertical loading):

n	1.0	0.8	0.5	0.3
k*	2.375E+06	8.786E+05	1.947E+05	7.031E+04

Comparative
database:

best match

Soil	n	k_c [N/m $^{n+1}$]	k_ϕ [N/m $^{n+2}$]	k^* [N/m $^{n+2}$] (for b=160 mm)	ϕ [°]	c [kPa]
Dry sand (Land Locomotion Laboratory)	1.1	9.90E+02	1.53E+06	1.53E+06	28	1.04
Sandy loam (Land Locomotion Laboratory)	0.7	5.27E+03	1.52E+06	1.55E+06	29	1.72
Sandy loam Michigan (Strong, Buchele)	0.9	5.25E+04	1.13E+06	1.46E+06	20	4.83
LETE sand (Wong)	0.79	1.02E+05	5.30E+06	5.94E+06	31.1	1.30
DLR mechanical Mars soil simulant	0.8	5.79E+03	1.80E+05	2.16E+05	17.8	0.30
Lunar nominal soil	1	1.40E+03	8.20E+05	8.29E+05	35	0.17
Sandy loam (Land Locomotion Laboratory)	0.2	2.56E+03	4.31E+04	5.91E+04	38	1.38
Sandy loam (Hanamoto)	0.3	2.79E+03	1.41E+05	1.59E+05	22	13.79
Clayey soil (Thailand)	0.5	1.32E+04	6.92E+05	7.75E+05	13	4.14
Heavy clay (WES)	0.11	1.84E+03	1.03E+05	1.15E+05	6	20.69



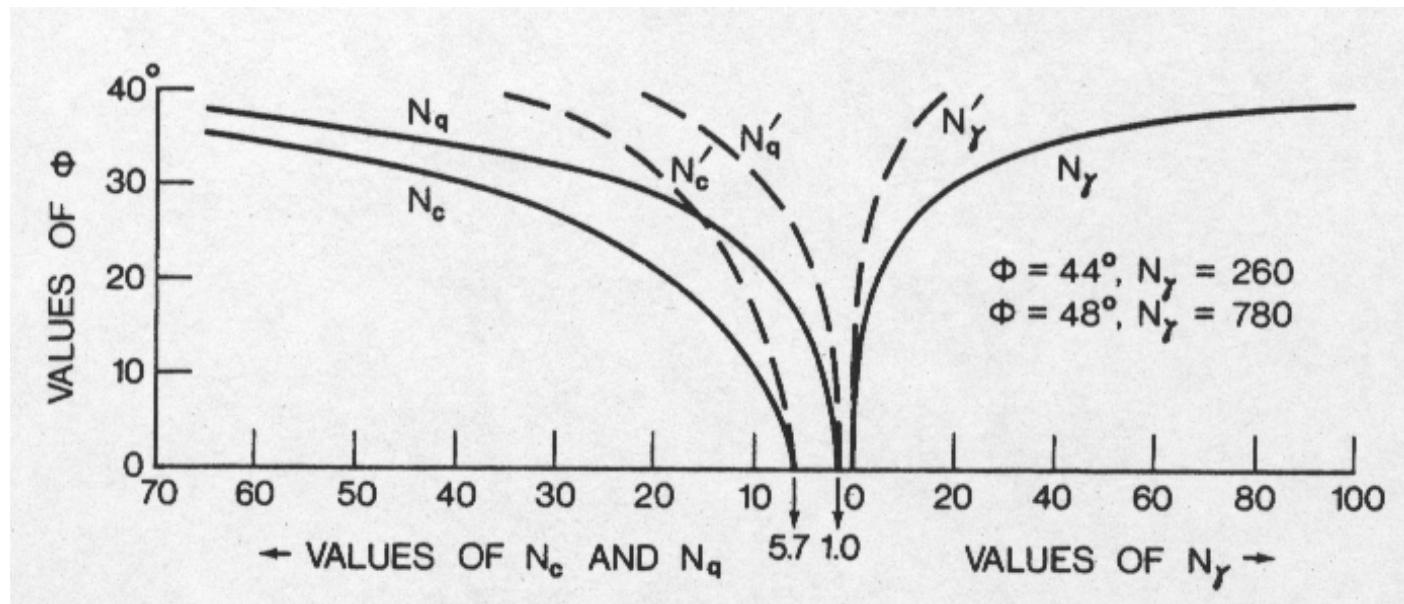
Example: MER-A egress soil

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- Apply bearing capacity theory to derive bearing strength

$$W'_c = 2\gamma_s l \cdot b^2 N'_\gamma + 2l \cdot b \cdot q N'_q + \frac{4}{3} l \cdot b \cdot c N'_c$$





Some results and comparisons

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	Bearing strength	Cohesion	Internal friction angle	Bulk density (soil)	Thermal inertia (soil)
<i>Gusev loose soils</i>	5 kPa	1 kPa	20°	1230 kg/m ³ *)	134 J m ⁻² s ^{-1/2} K ⁻¹ **)
<i>Gusev dense soils</i>	200 kPa	15 kPa	25°	1483 kg/m ³ *)	153 J m ⁻² s ^{-1/2} K ⁻¹ **)
<i>Meridiani Eagle crater floor & M. plains</i>	80 kPa	5 kPa	20°	1333 kg/m ³ *)	142 J m ⁻² s ^{-1/2} K ⁻¹ **)
<i>Meridiani Eagle crater wall (loose soil)</i>	8 kPa	0.5 kPa	20°	1186 kg/m ³ *)	130 J m ⁻² s ^{-1/2} K ⁻¹ **)
VL-1 ,drift'		1.6±1.2 kPa	18±2.4°	1150±150 kg/m ³	
VL-1 ,blocky'		5.5±2.7 kPa	30.8±2.4°	1600±400 kg/m ³	
MPF ,cloddy'		0.17±0.18 kPa	37.0±2.6°	1530±110 kg/m ³	
Lunar soil (inter-crater areas) (Kemurdjian et al., 1978)	36-55 kPa	2.7-3.4 kPa	22-27°		
Coarse sand (Earth)	150-300 kPa	0.1-2.0 kPa	30-39°		

*) assuming behavior like lunar regolith: $\tan\phi=0.725*1/E$, $c=0.720*E^{**}(-5.664)$

**) assuming $\kappa=0.01 \text{ W m}^{-1} \text{ K}^{-1} + [6.4*10^{-6} \text{ m}^4 \text{ s}^{-3} \text{ K}^{-1}] \rho$, and $c_p=820 \text{ J kg}^{-1} \text{ K}^{-1}$

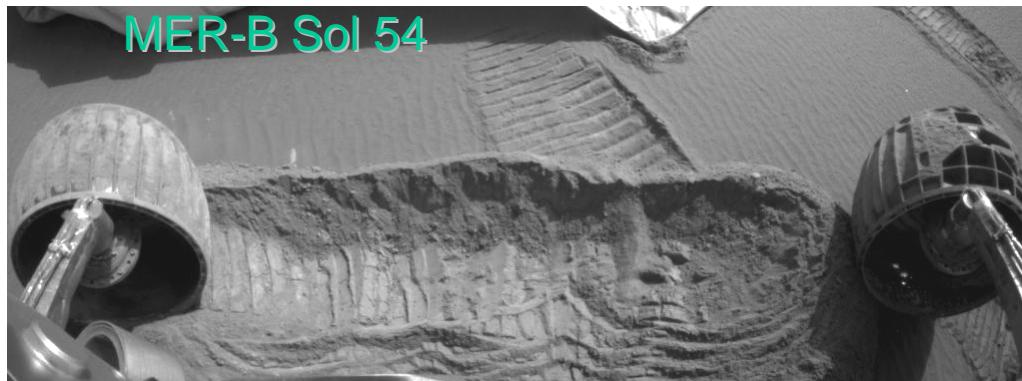


Wheel trenching

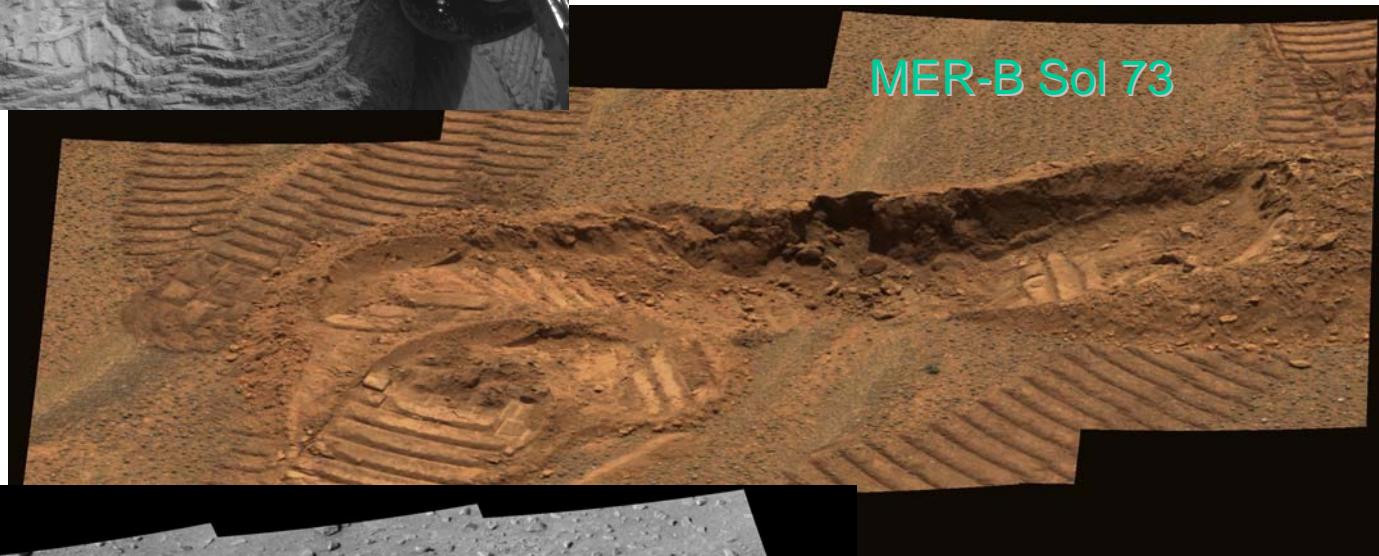
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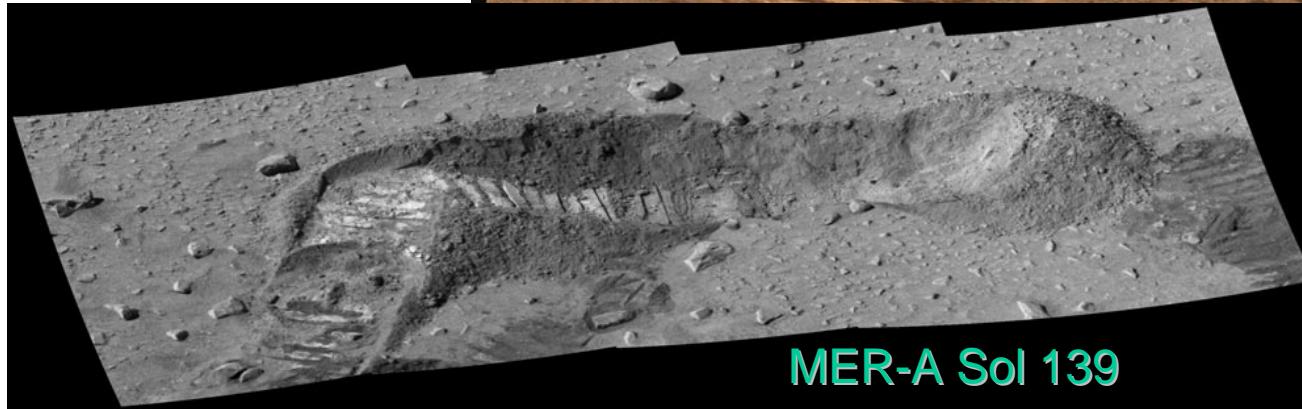
MER-B Sol 54



MER-B Sol 73



MER-A Sol 139



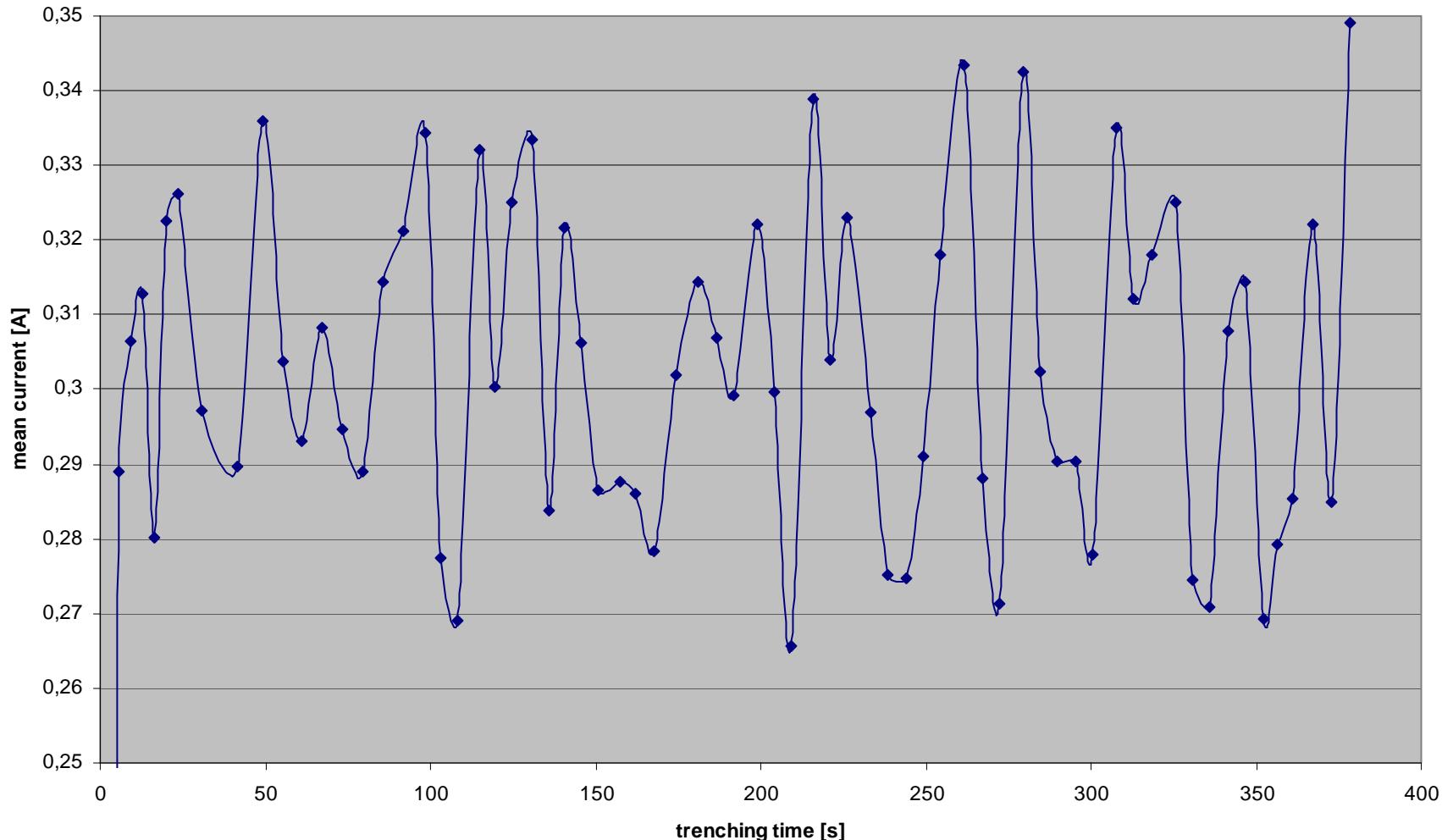


MER-B Sol 73 trench

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mean current vs. time (10% winsorized mean)

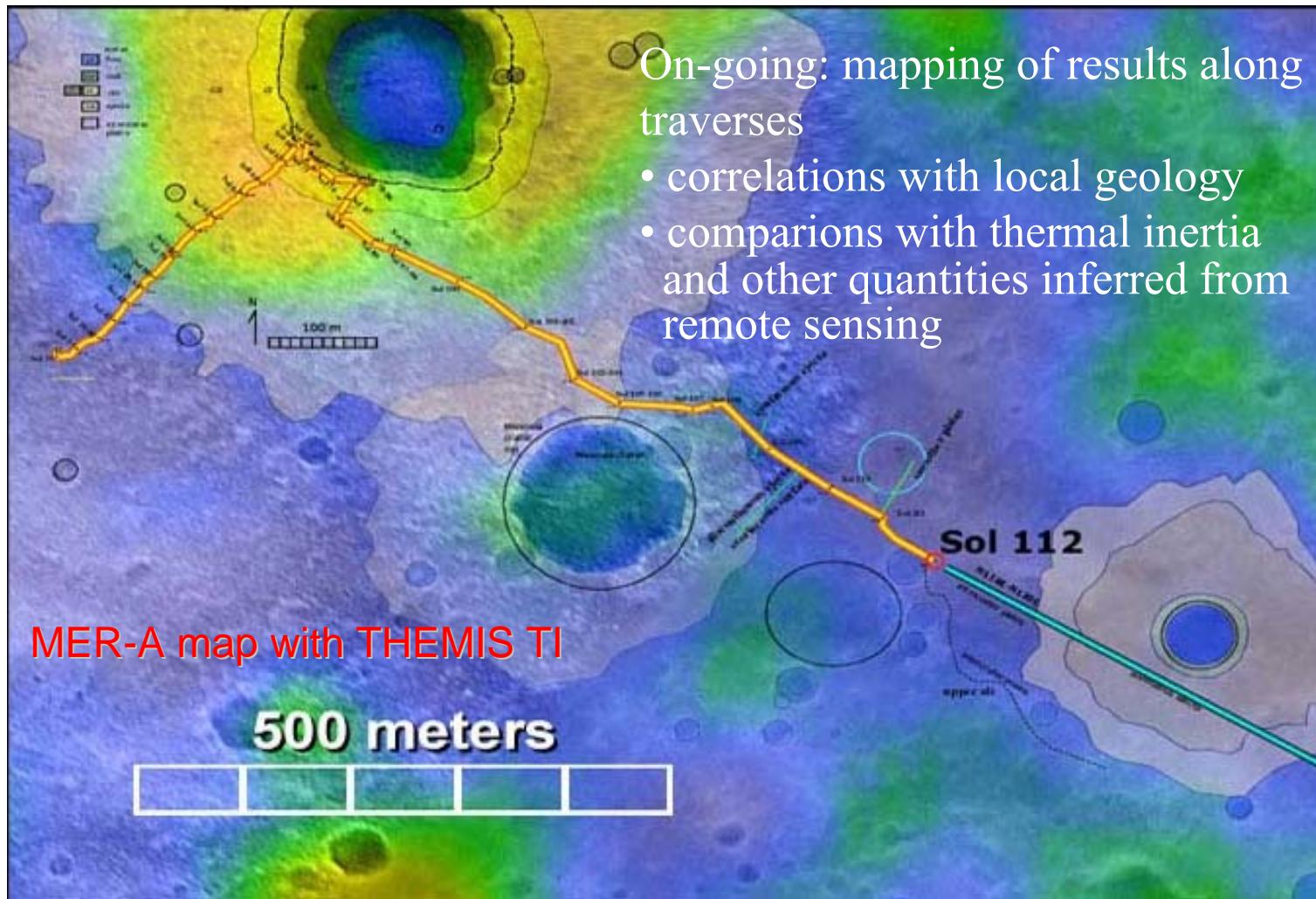


-> being used to infer soil shear strength vs. depth



Geologic context

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Implications

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- Soil deposits with distinct mechanical properties
- Cohesion inferred for all soil types observed with the described methods
- Soil thermal inertia retrievals: slightly lower than fine-component TI from orbital measurements? (comparisons in progress, also with Mini-TES derived TI)
- Soil crusts formation as recent or modern process (action of moisture & salts?)