

# Non Thermal Continuum radiation observed from the Cluster fleet



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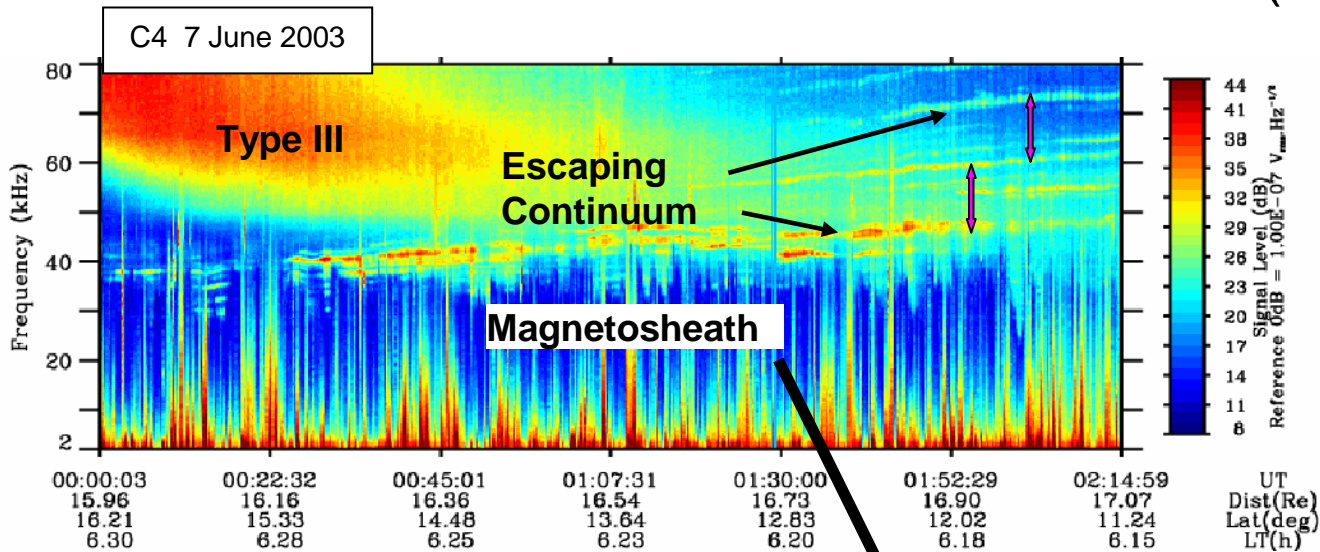
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5th Anniversary of Cluster in Space  
19-23 September, ESTEC

# NTC observed from the Cluster fleet

1. Generalities
2. Advances with Cluster observations
  - Ray path from multipoint observatory
  - Sources
3. Summary and perspectives

# 1.1 What is Non Thermal Continuum (NTC)?



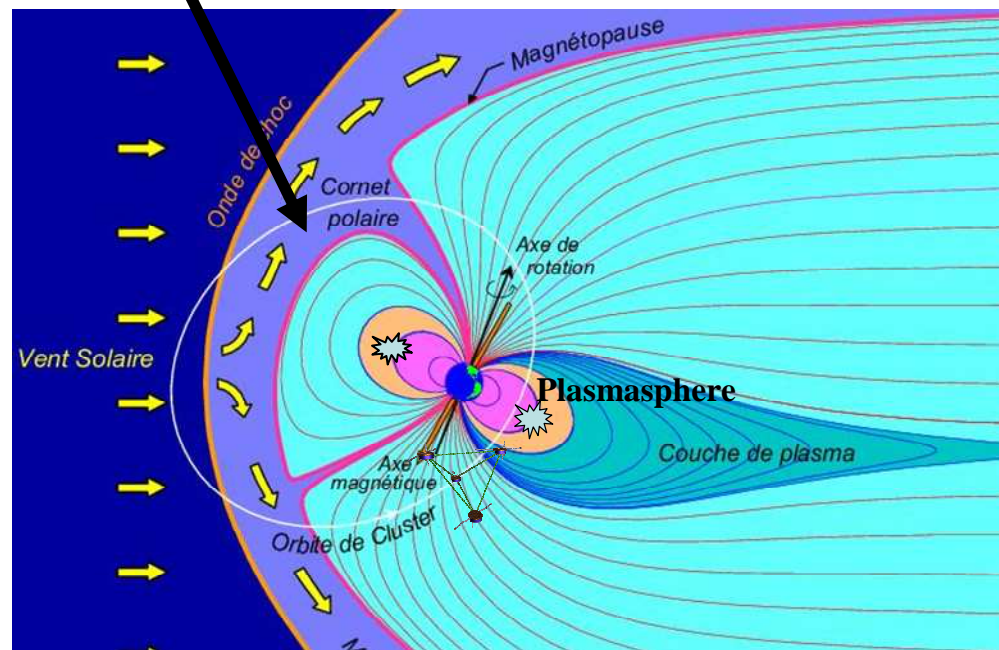
NTC and Type III:

$$\sim 10^{-16} \text{ V}^2 \text{ M}^{-2} \text{ Hz}^{-1}$$

$$\Delta F \sim F_{ce} \text{ at source}$$

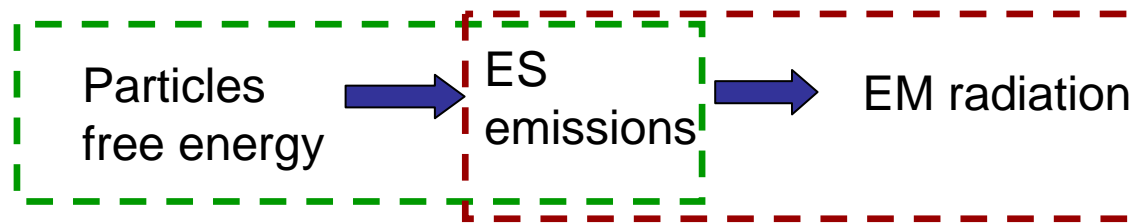
NTC radiations are:

- weak, long lasting, narrow band, EM emissions
- observed (since 1970s), in almost all regions of magnetosphere, they can escape in the interplanetary medium



## 1.2 What is the source of NTC?

*General scheme of NTC generation:*



*This happens at the plasmopause boundary layer*

NTC formation is largely not understood. How to explain:

Duration of emissions?

Spectral characteristics (narrow bands, harmonic structures)?

Wave intensity?

Beaming properties?

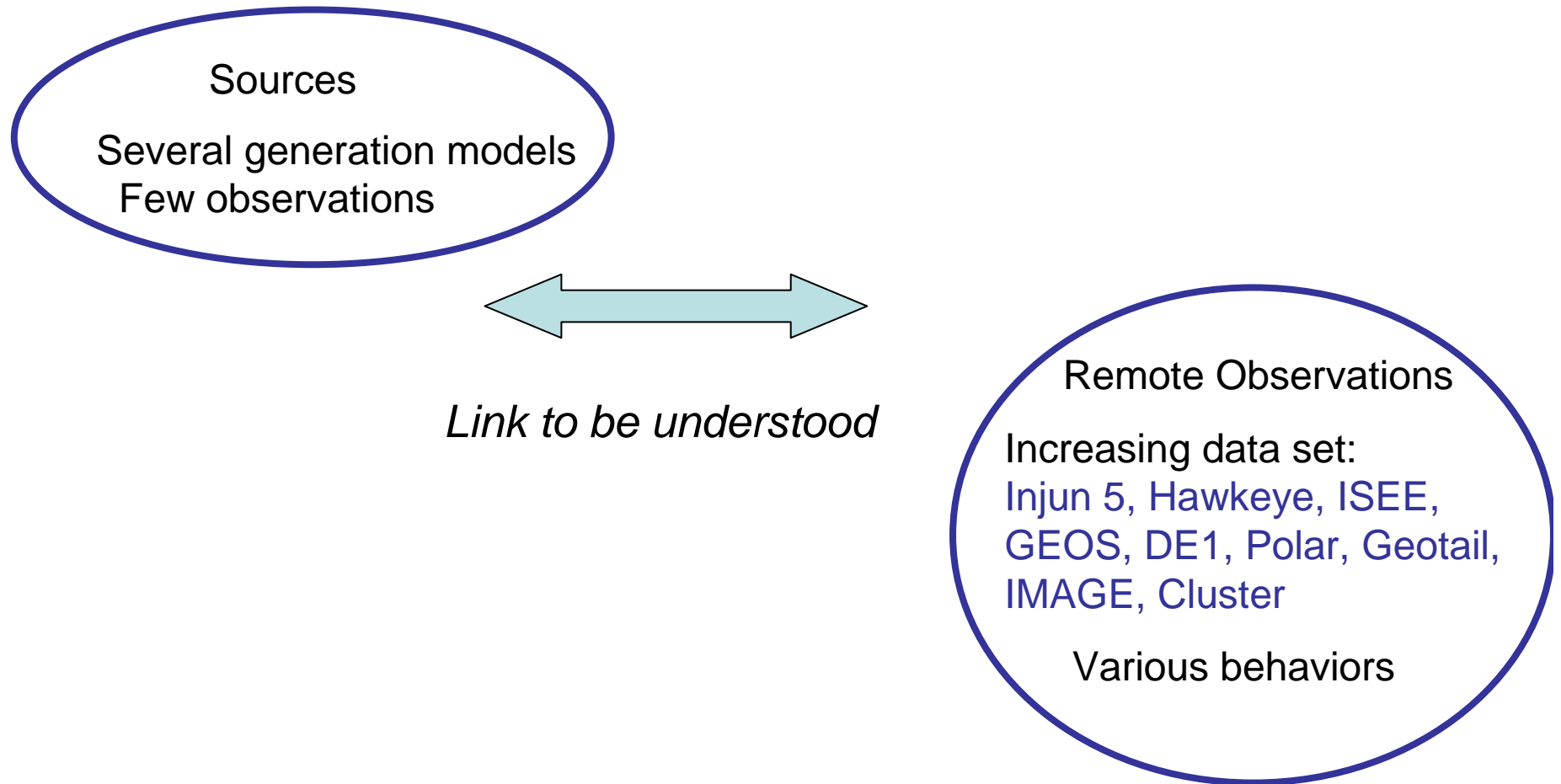
### *Models*

Radio window  
at grad (Ne)  
*Jones, 1982*

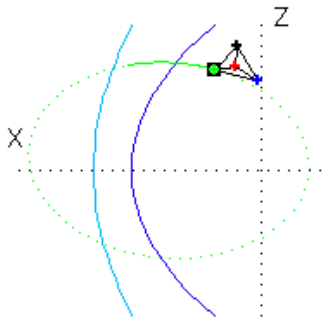
N.L. processes

- coalescence with LF waves  
*Melrose, 1981*
- decay of Bernstein waves  
*Rönmark, 1985*

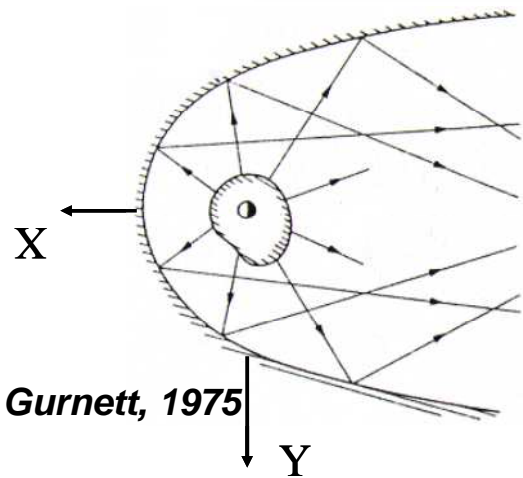
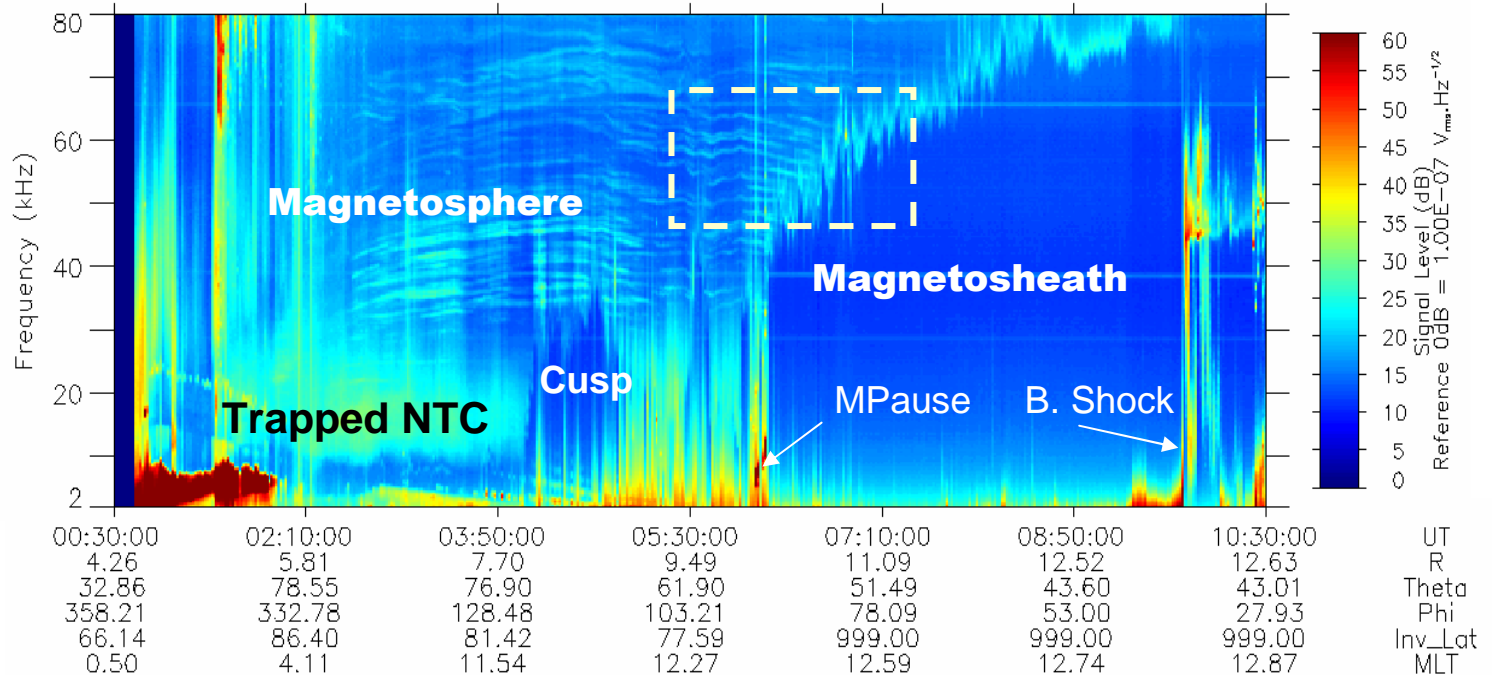
## 1.3 Current studies



# 1.4 How does NTC propagate?



26/02/2001



Gurnett, 1975

5th

- NTC radiation emerges from sources at PPause boundary layer. It is beaming predominantly near equatorial plane.
- The lower frequency band is trapped inside the magnetospheric cavity, bouncing between MPause(/cusp) and Plasmapause

## 2. Advances with Cluster observations

- Ray path from multipoint observatory  
2 examples
- Sources



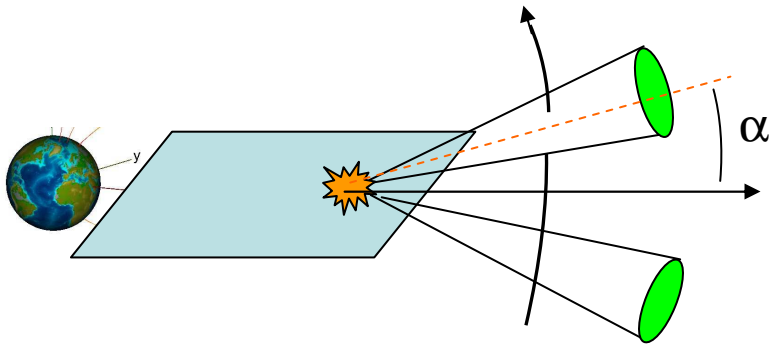
# Example 1: Beaming properties

Test of Jones theory

( ES to EM conversion )

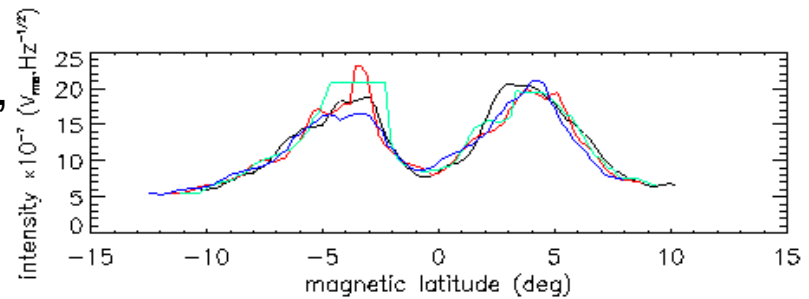
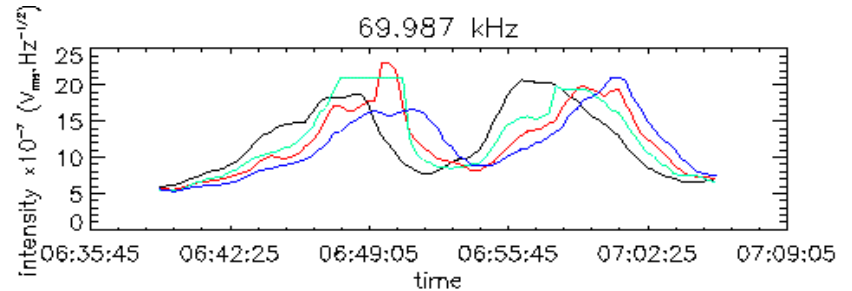
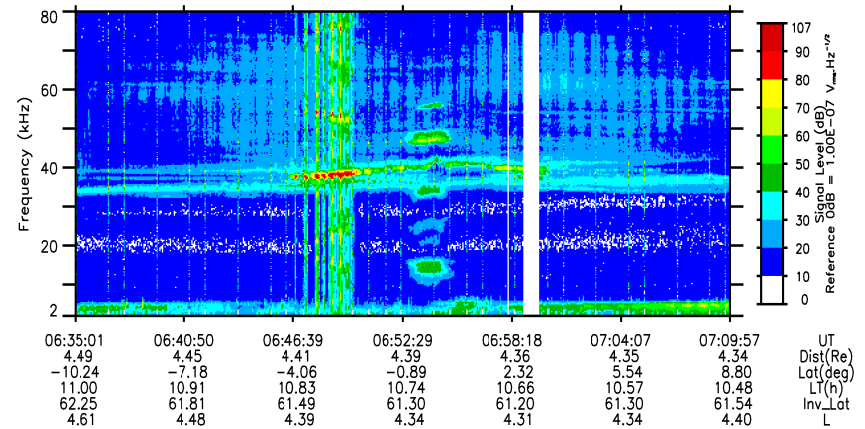
2 narrow beams, symmetric / equator

$$\alpha = \arctan(f_{ce}/f_{pe})^{1/2}$$



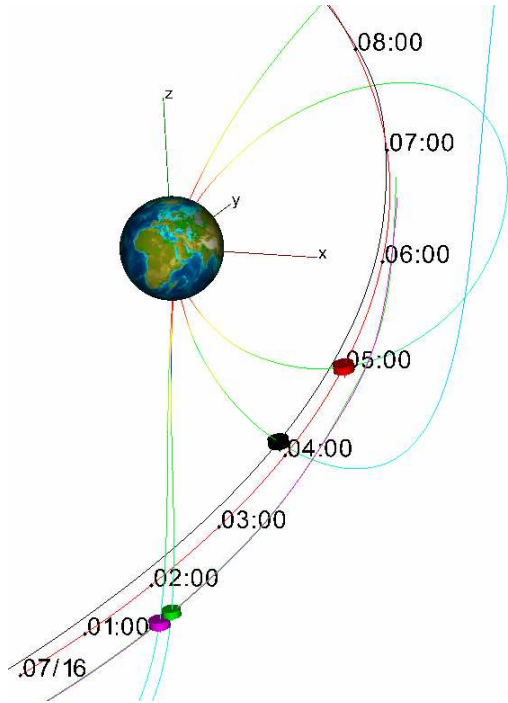
- No definitive conclusion has been reached, concerning the validity of Jones theory
- Cluster orbit does not cross a beam pair issued from the same source

26/09/03 Separation ~ 250 km

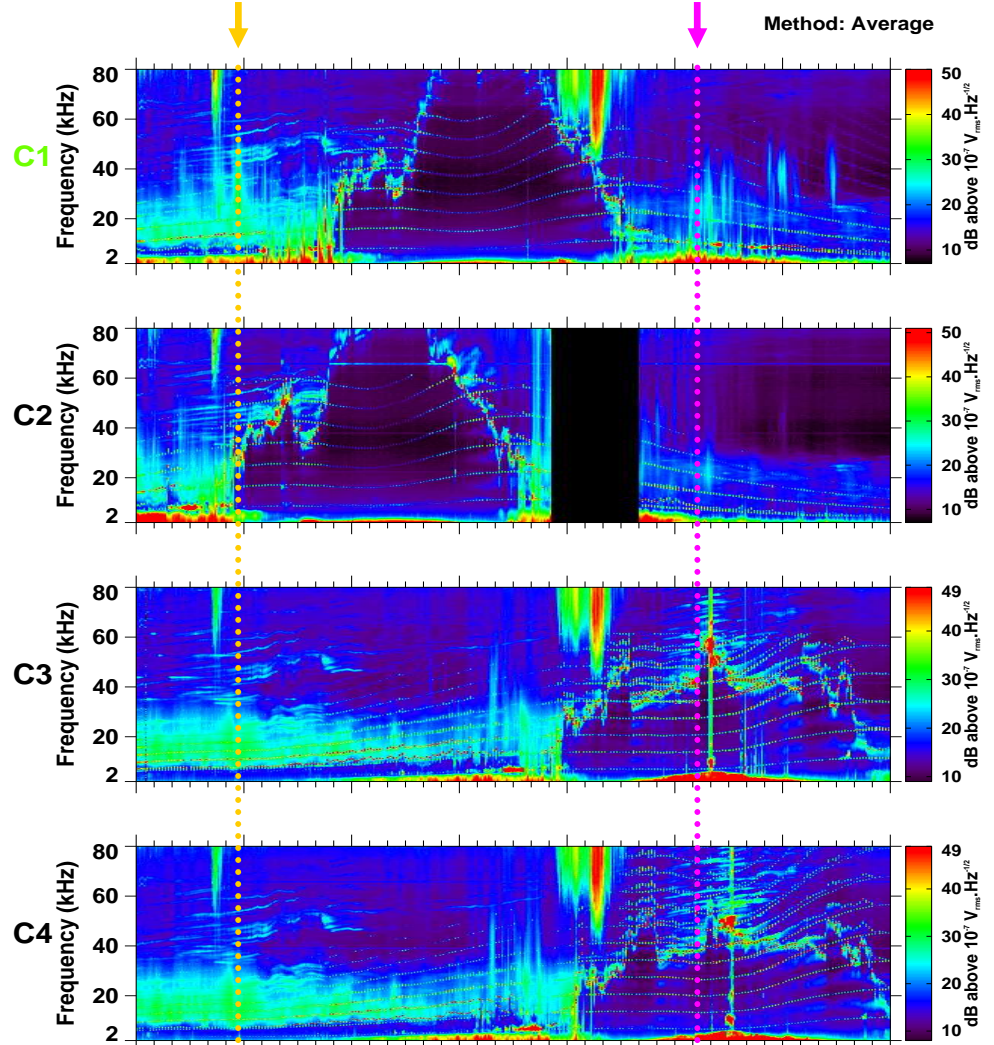




## Example 2: Observation of narrow band emissions



CLUSTER-WHISPER Spectrogram / JUL 16, 2005 (Day 197)



Far from equator, the four SC see exactly the same emission

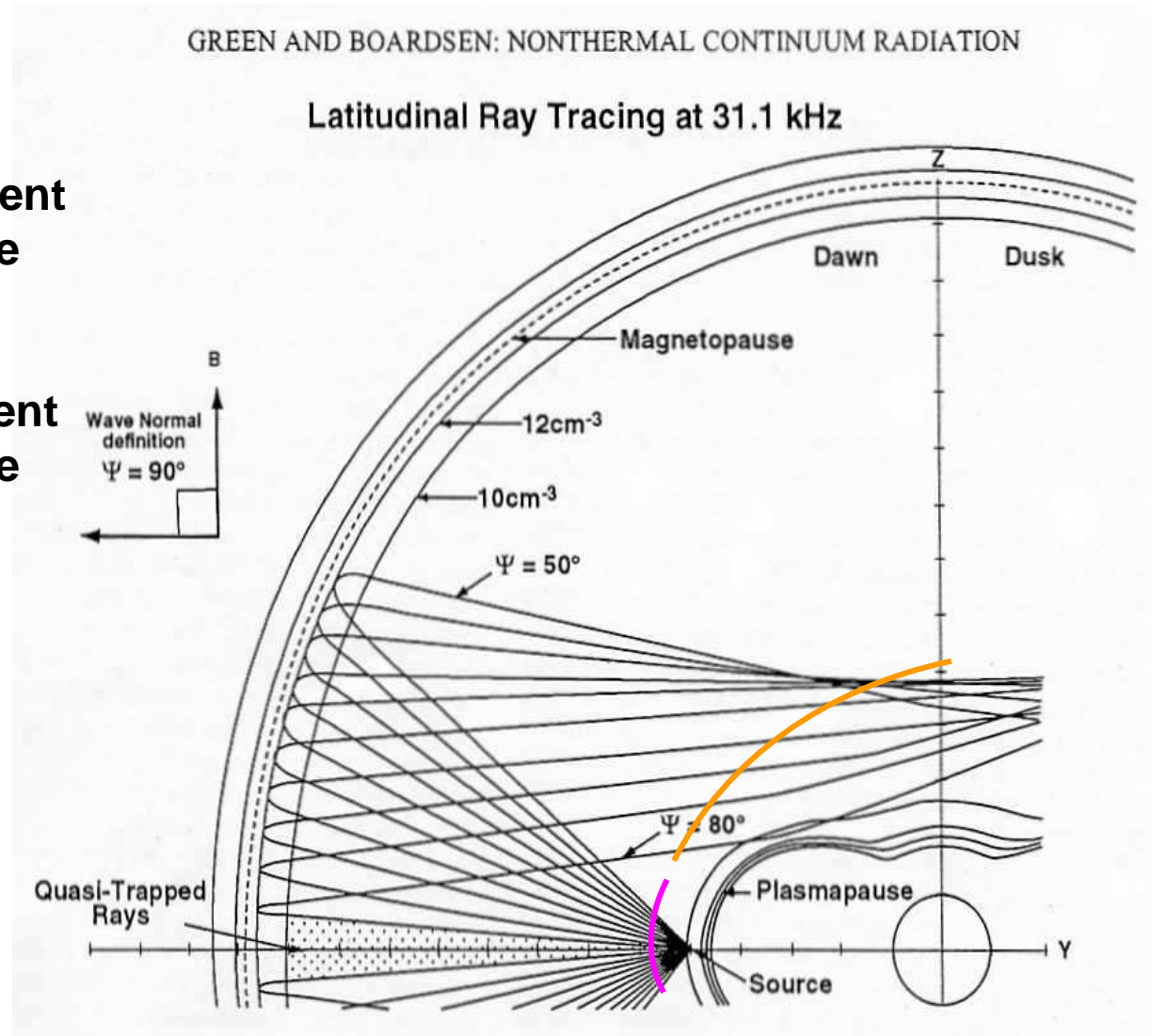
Only part of the constellation is illuminated near the equator



UT	04:00:00	05:00:00	06:00:00	07:00:00	08:00:00	09:00:00	10:00:00	11:00:00
X_gse(Re)	-0.82	0.42	1.62	2.60	3.12	3.09	2.61	1.88
Y_gse(Re)	4.04	4.50	4.57	4.04	2.81	1.08	-0.82	-2.66
Z_gse(Re)	-5.78	-4.18	-2.23	-0.00	2.21	4.04	5.31	6.09
Lat_sm(deg)	-62.41	-50.30	-31.75	-5.89	24.54	53.39	76.84	84.15
Cluster ID	15.39	15.30	15.59	15.39	15.30	15.59	15.39	15.30

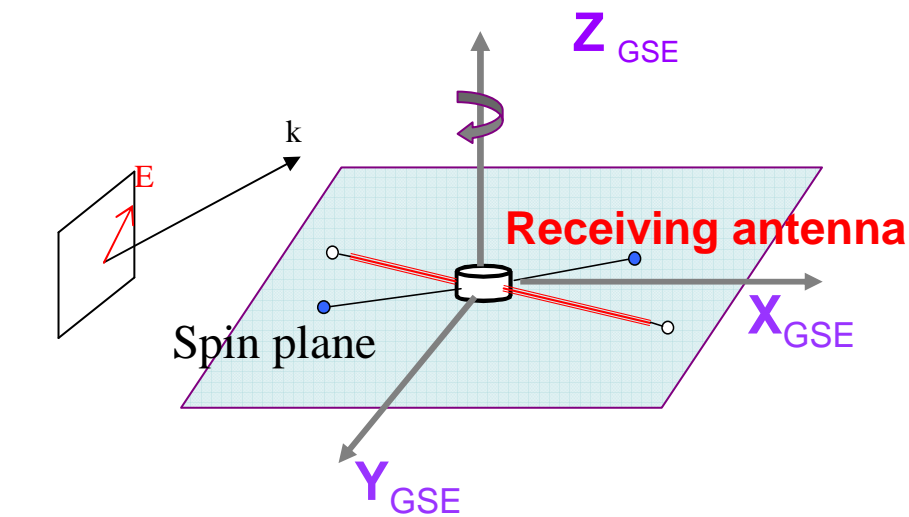
# Interpretation

- A small orbit element is illuminated by the direct path
- A large orbit element is illuminated by the reflected radiation

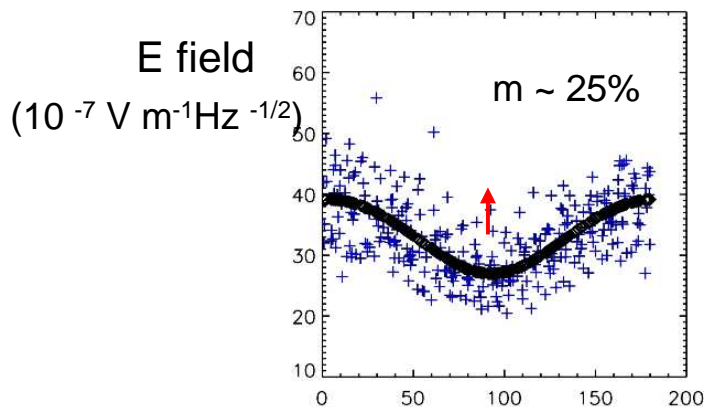
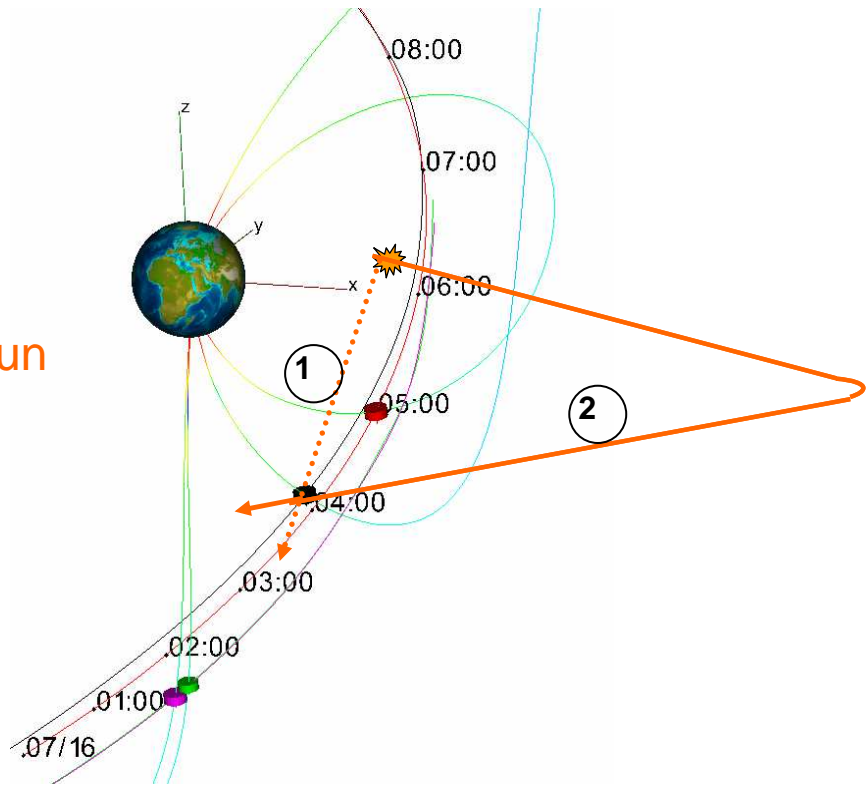


Meridian plane

# Spin modulation as a clue to ray path orientation



Sun

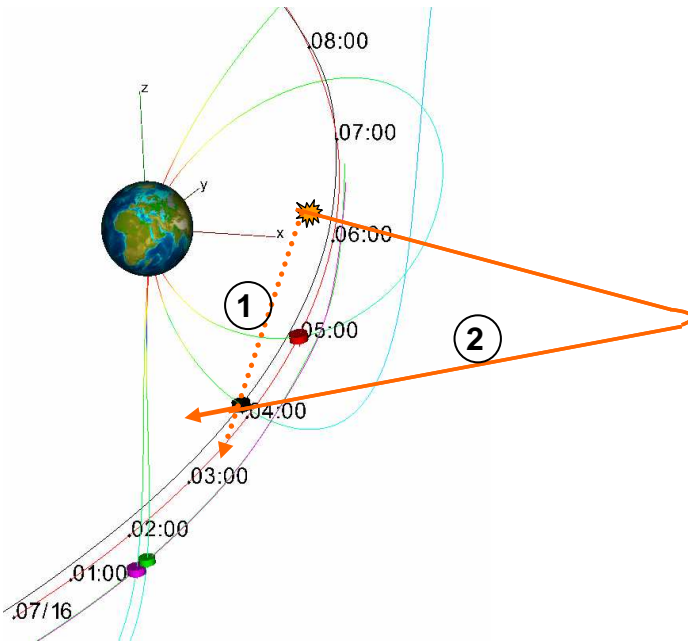


Phase angle (0-180°)

The spin modulation gives an estimation of the inclination of the ray wrt the spin plane

It allows to find out if the ray reaches the observatory directly (1) or after a reflection (2)

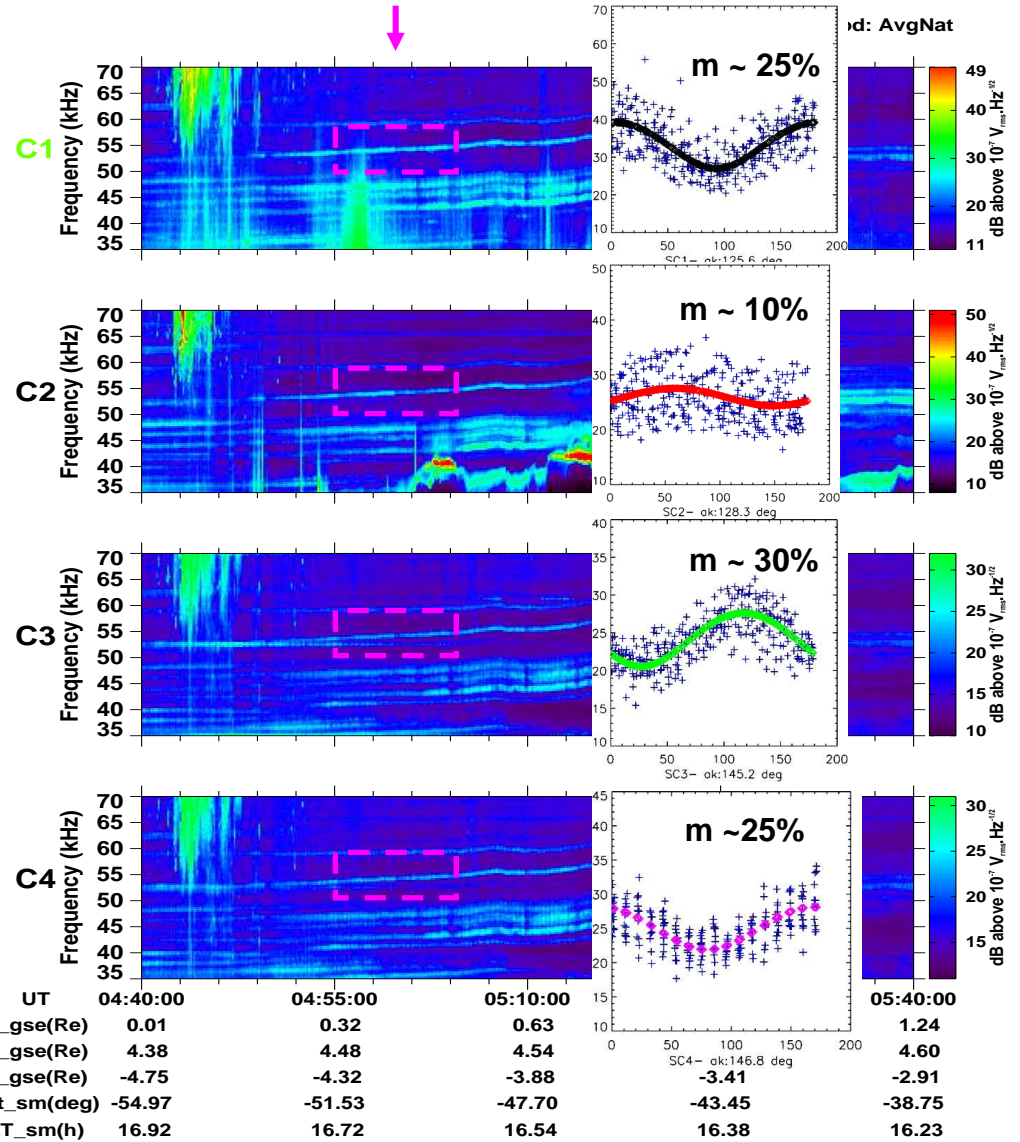
# Compared spin modulations



Modulation values indicate that beams have been reflected

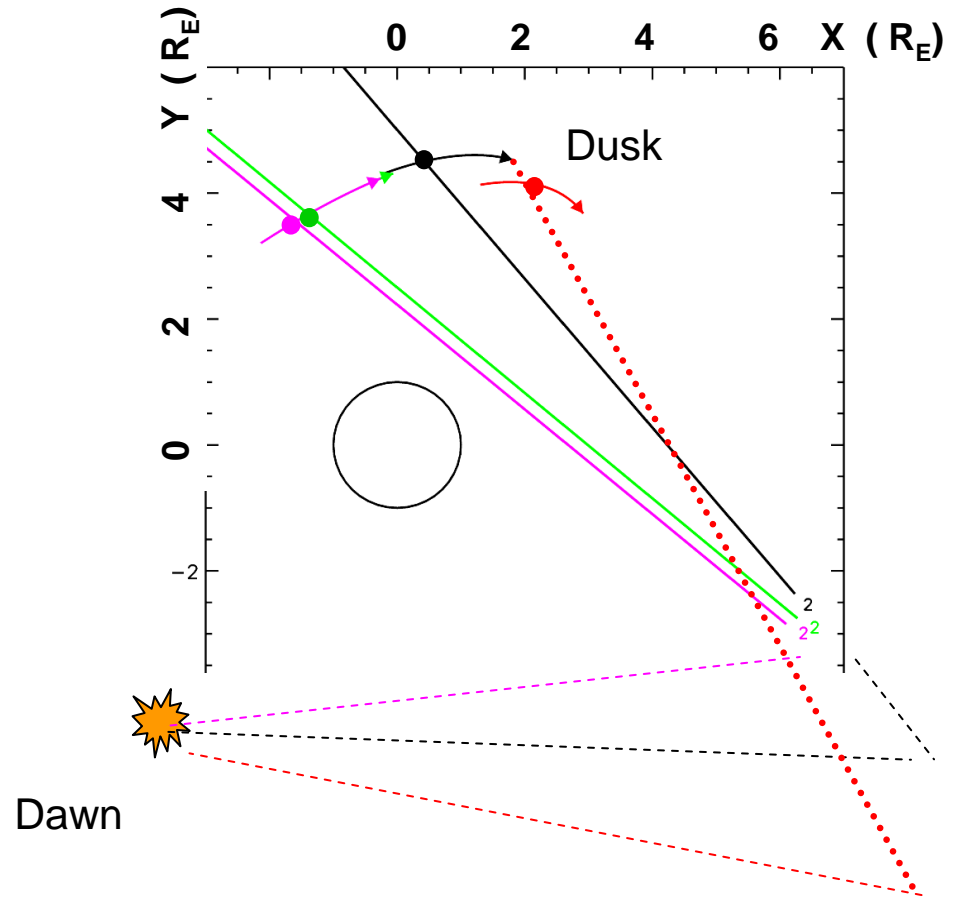
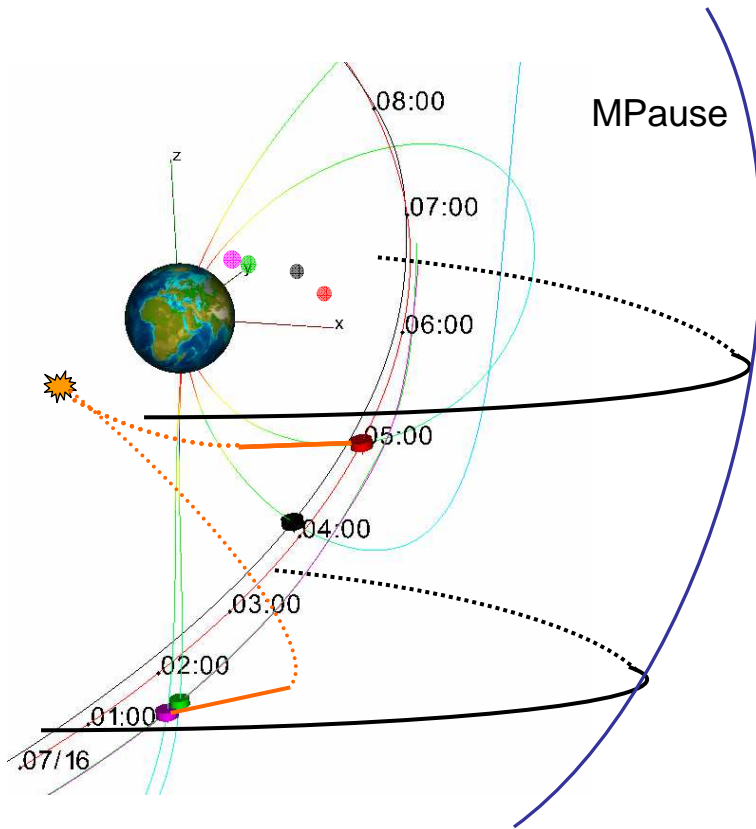
Directivity angles complete the information

## CLUSTER-WHISPER Spectrogram / JUL 16, 2005 (Day 197)





Triangulation points toward high X value, it is compatible with the proposed scenario

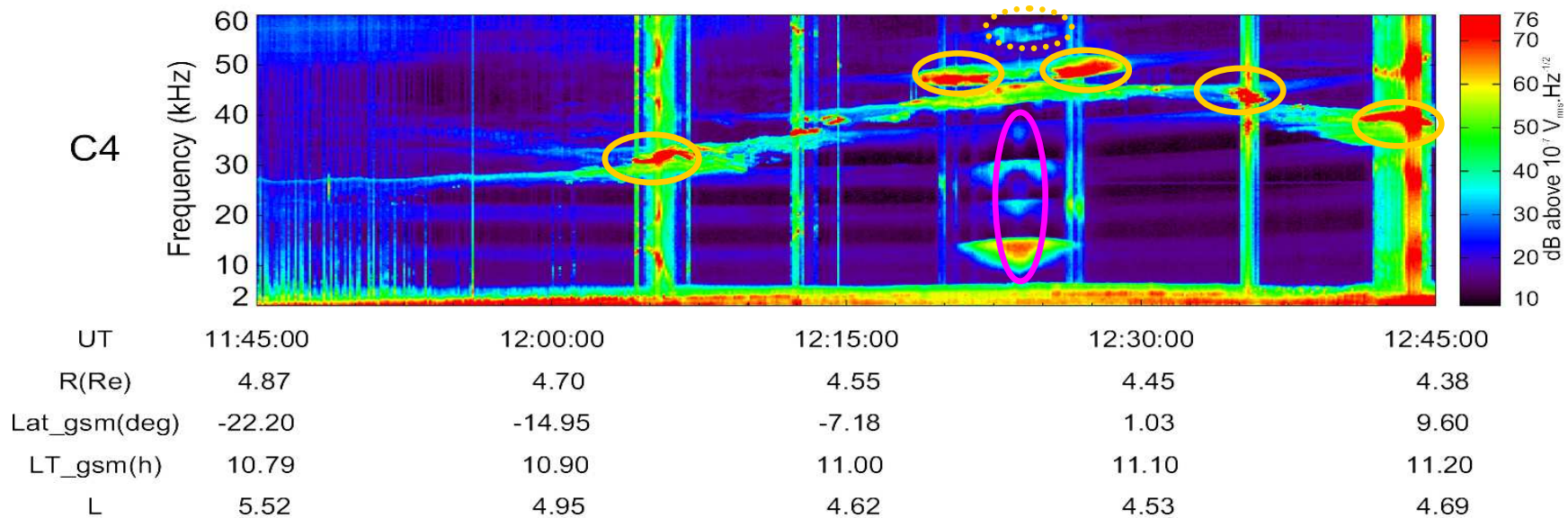


- Conclusion: the origin of beams observed in the dusk sector is probably a source placed in the dawn sector

## 2. Advances with Cluster observations

- Ray path from multipoint observatory
- Sources

## NTC Sources: Which candidate Electrostatic Emissions ?



- ***Equatorial  $n+1/2$ ?***

Too far from density gradient (inside plasmasphere)

- ***Intense (saturating) emissions at PPause density gradient ?***

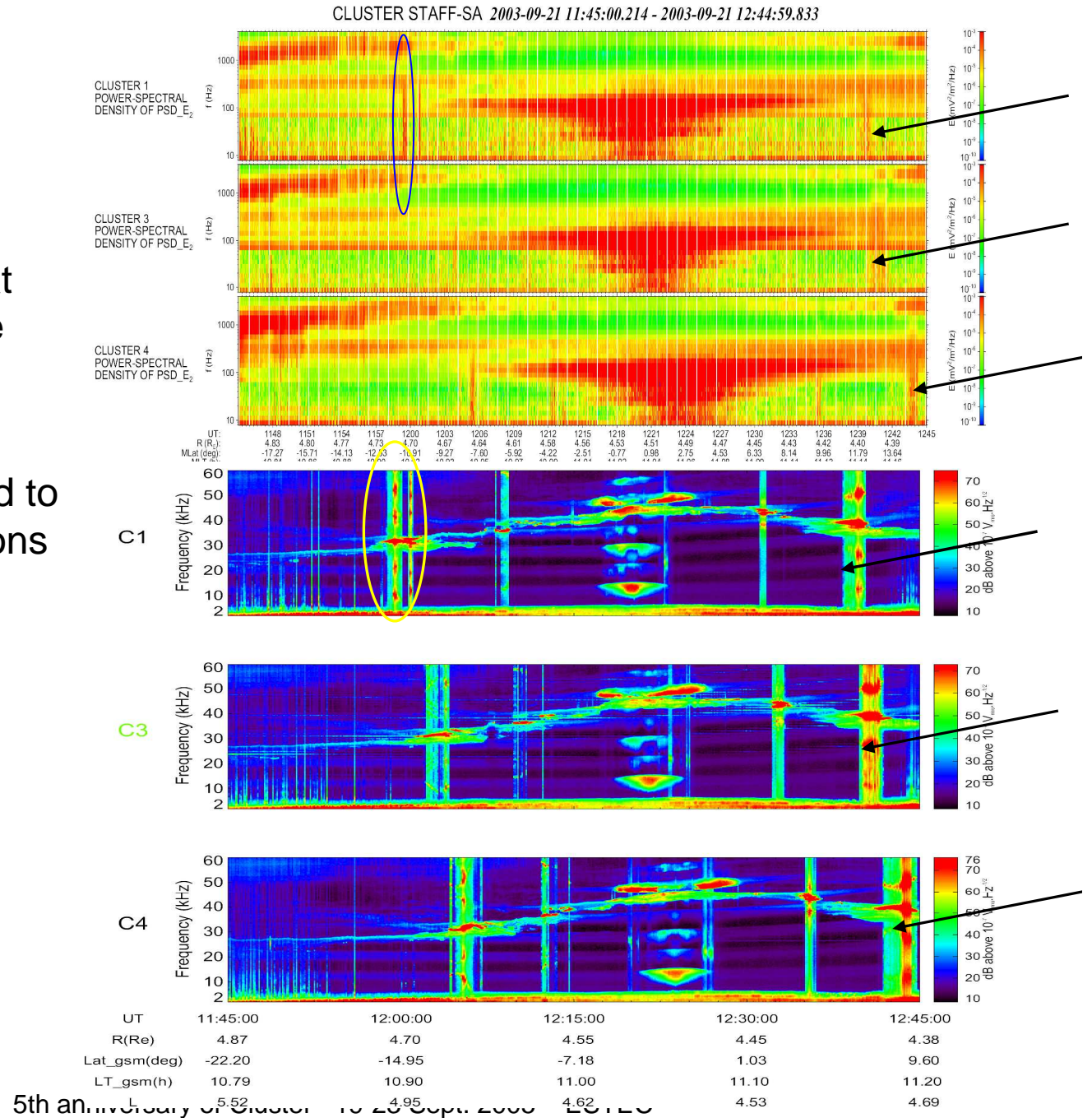
Observed just above  $F_p$  (in the band  $F_p - F_q$ ), close to equator

Good candidates for the main forms of NTC

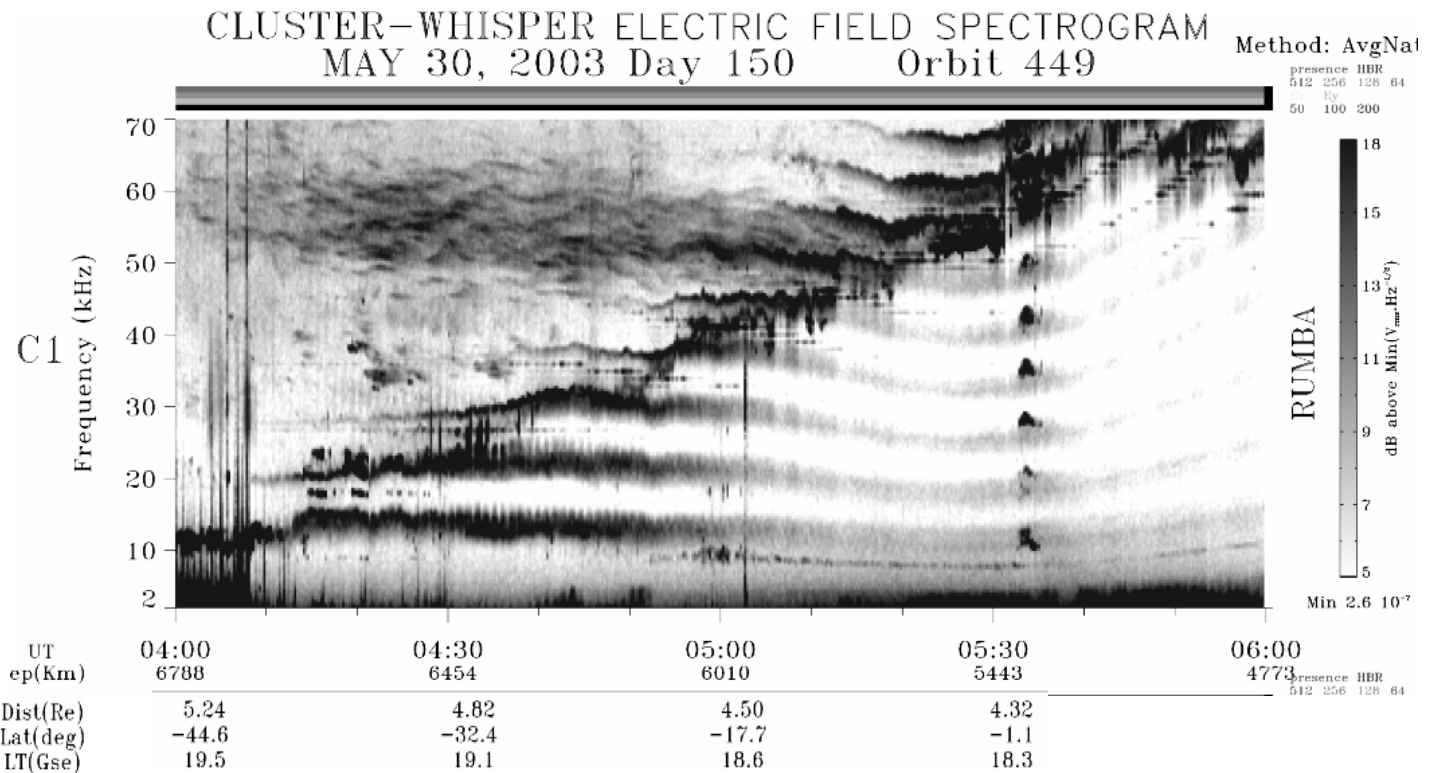


# NTC Sources

- Intense ES emissions at plasmopause surface are often associated with LF broadband emissions
- More analysis is needed to interpret these observations

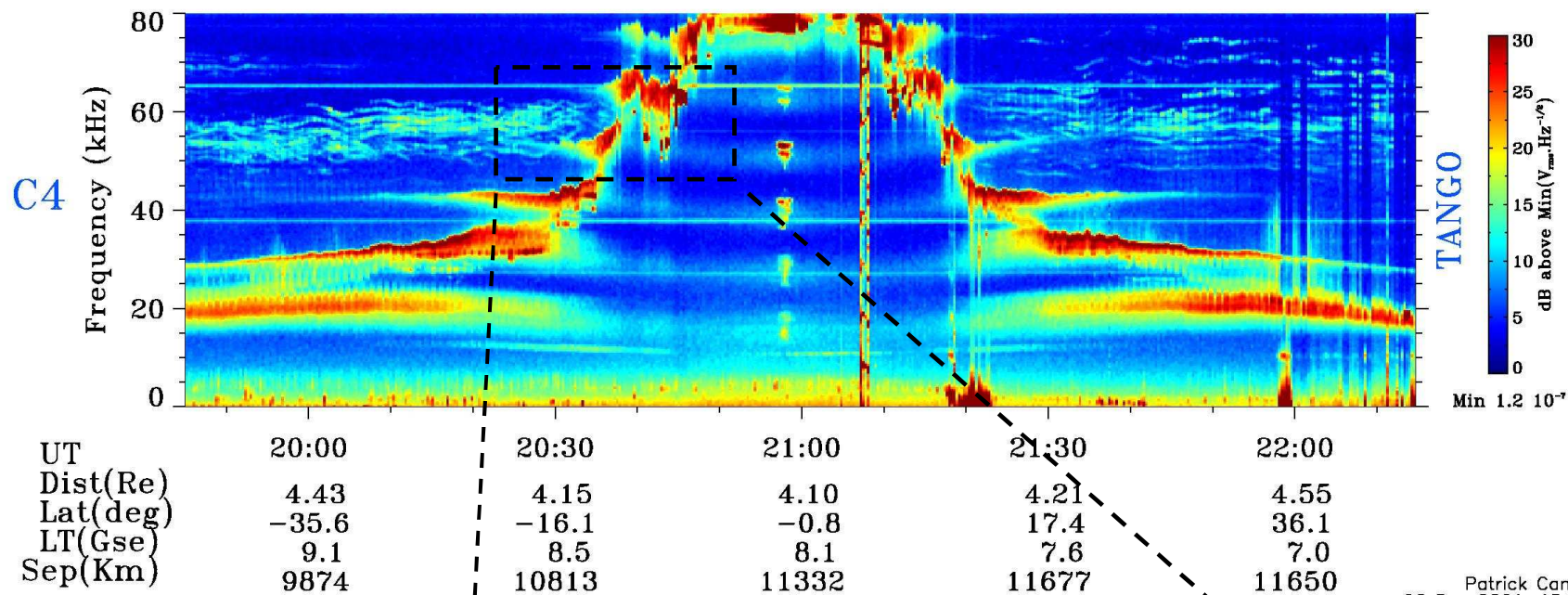


# Role of Bernstein waves



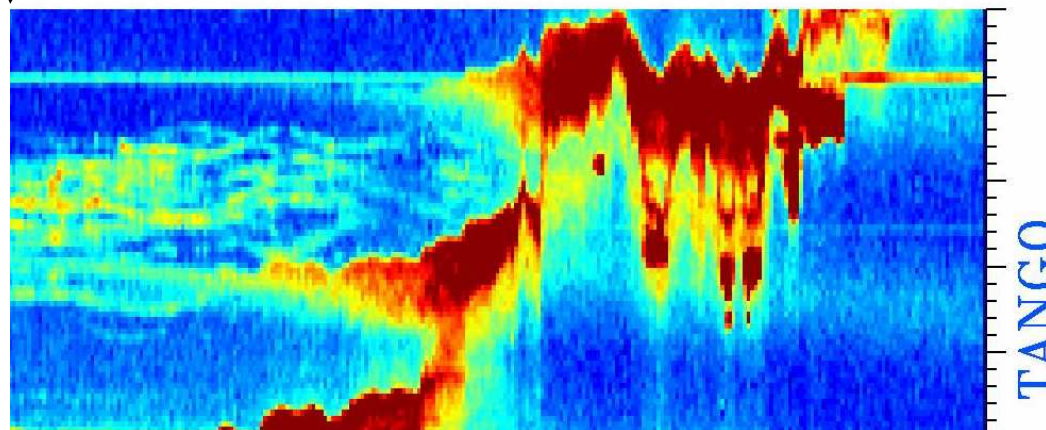
Several examples show convincing evidence of a link between ES emission at local Fqs (Bernstein mode) and simultaneously observed NTC

Canu et al., 2005



Patrick Canu  
06 Dec 2001, 15:51

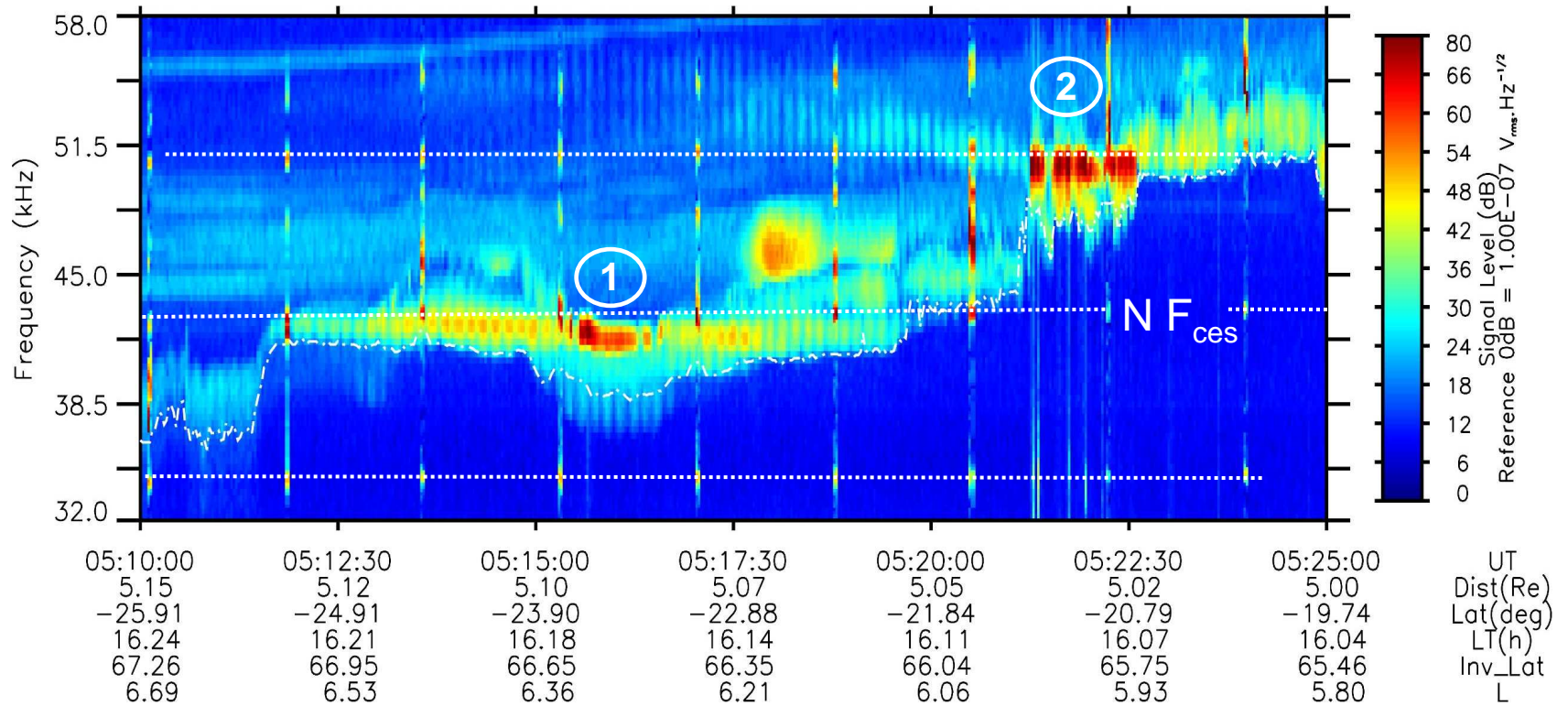
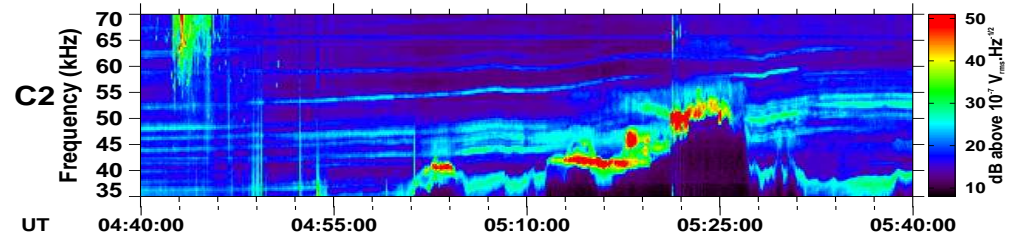
Possibly linked with  
small density cavities



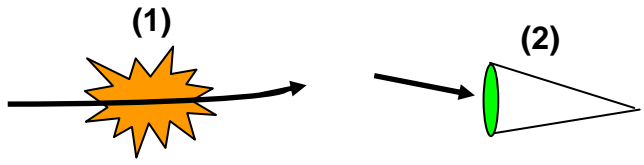
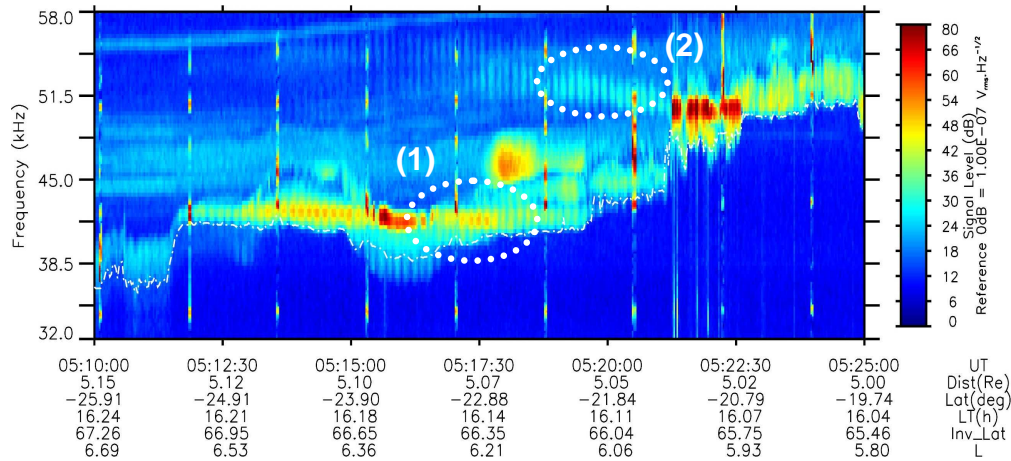


# NTC observed at sources

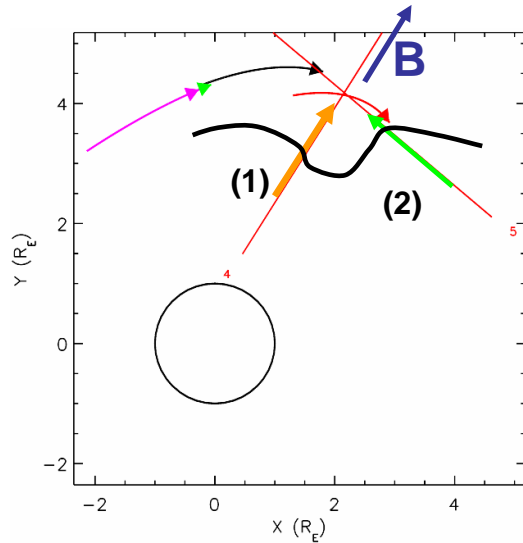
More from the 16 July 2005 event  
EM emissions near density cavities



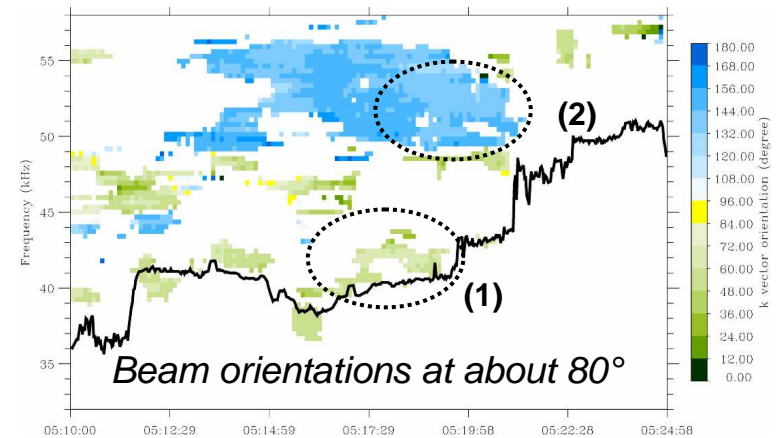
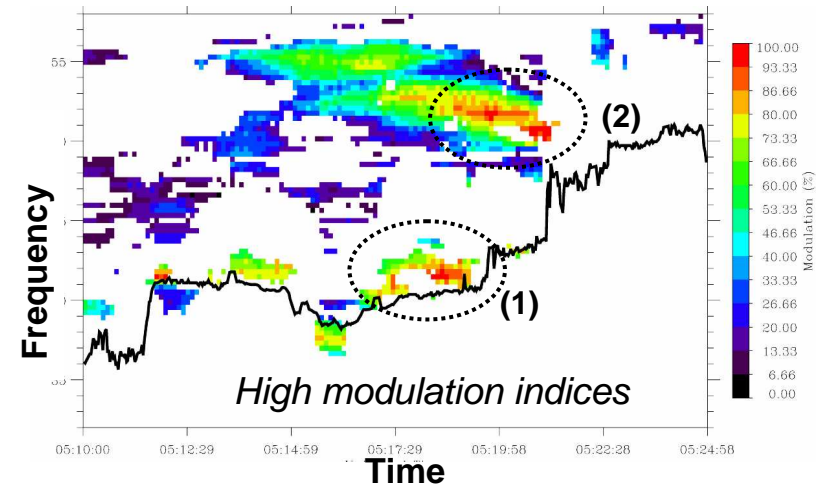
# Analysis of directivity properties



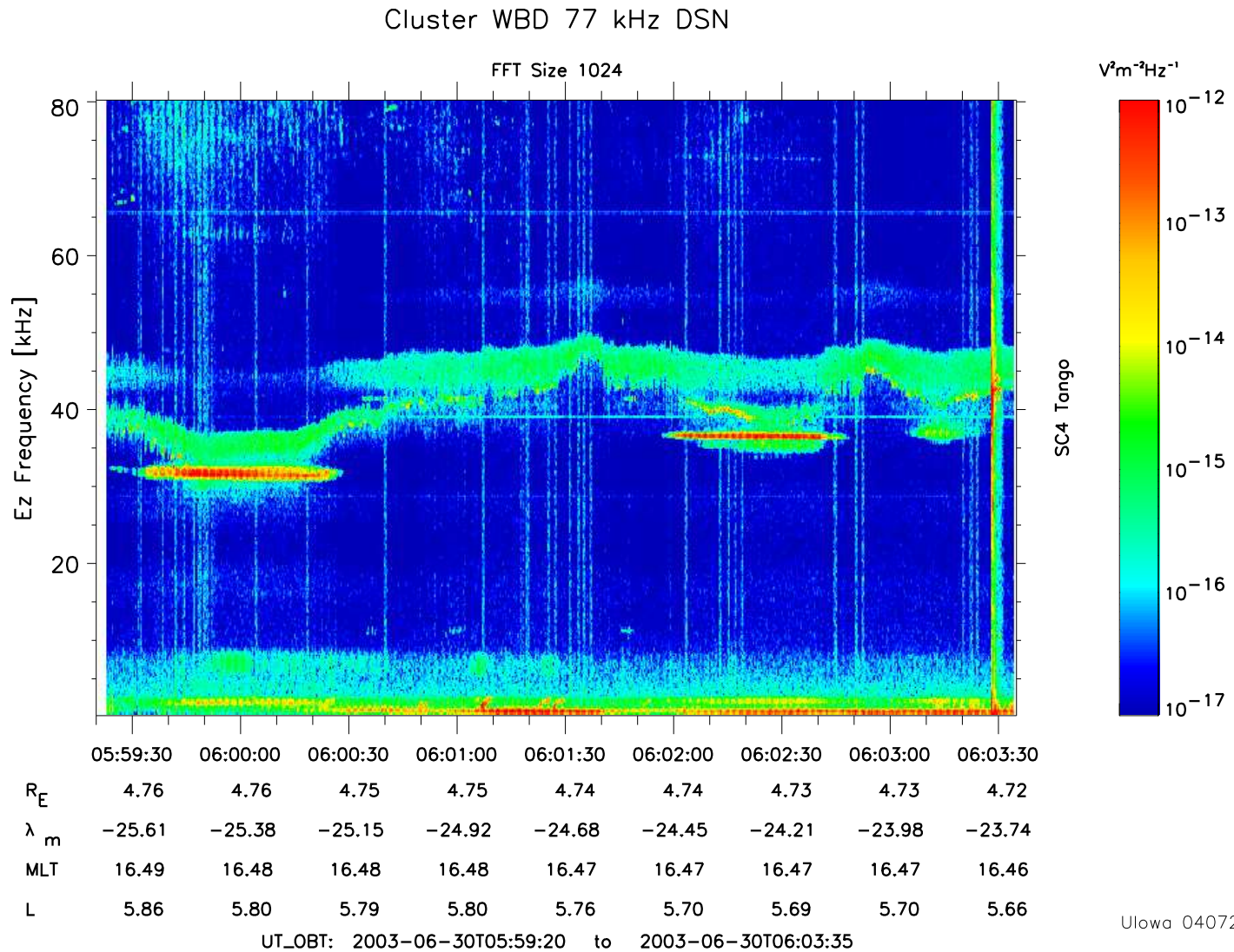
Ripples in the PPause boundary (rays aligned with density gradient) could explain different beam orientations



*Study in the spin plane (parallel to XY GSE)*

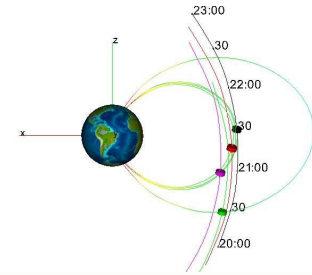


# WBD instrument offers a good time/frequency resolution of 'micro' sources

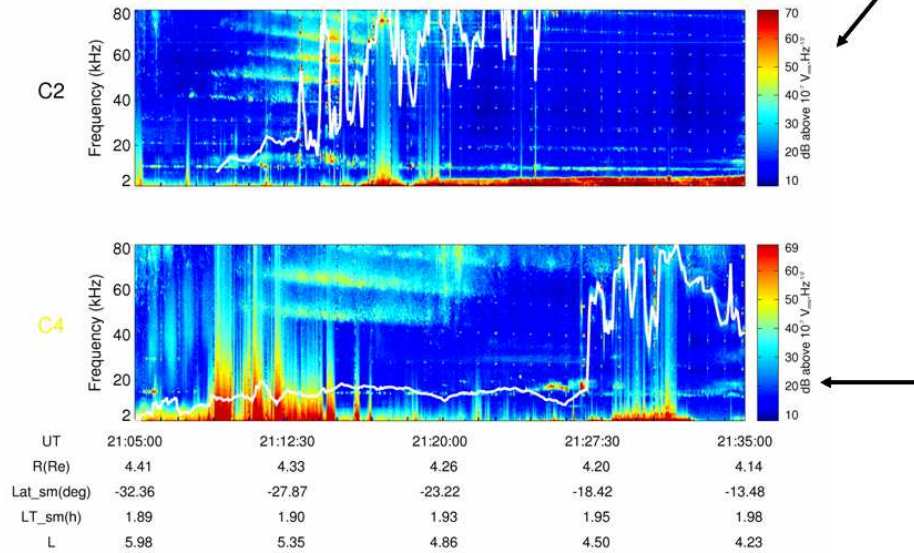
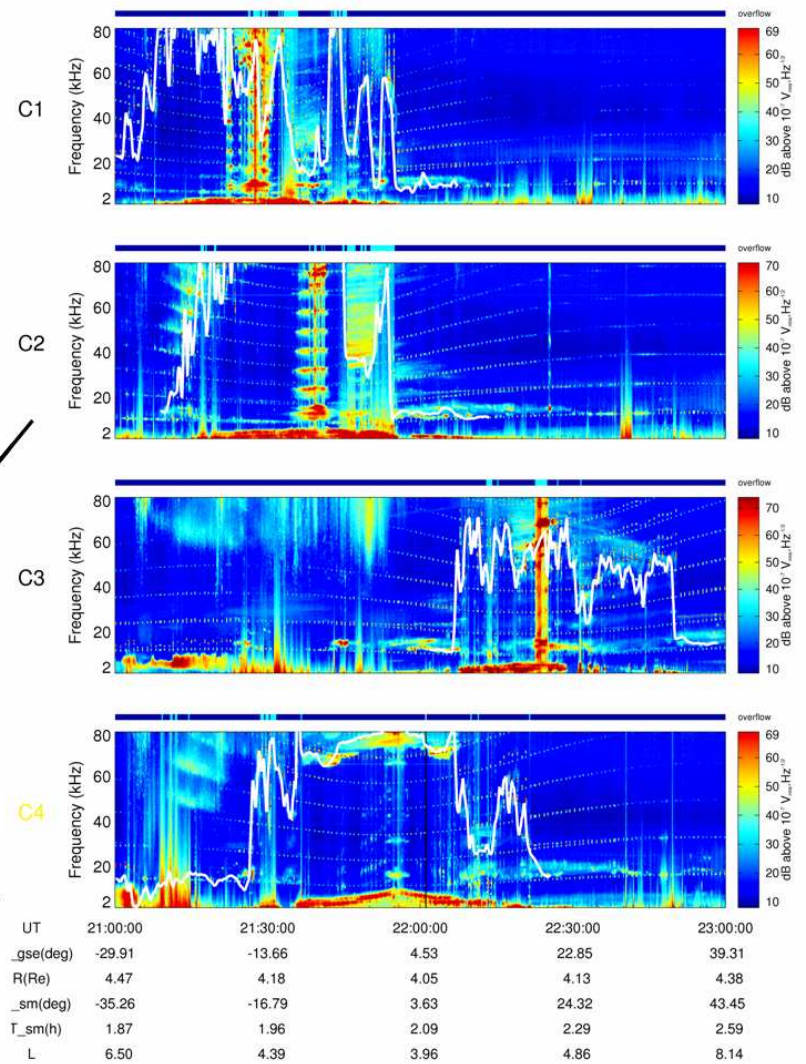




# Other large scale findings



- Simultaneous observations of harmonic 'large band' NTC structures
- Different spectral characteristics when viewed from different SC
- Local sources (off equator) are illuminating short orbit elements





## 5. Summary and conclusion

- Various types of ES emissions observed in the Ppause boundary layer
  - (a) Intense emissions in  $F_{pe} - F_q$  band
  - (b) 'Micro' sources, in density holes, some at  $n F_{ce}$
- Two main types of structured EM radiations observed in magnetosphere:
  - (a) multiple narrow spectral lines, radiating over a long range, having undergone one reflection at Msheath in many cases
  - (b) 'harmonic wide bands', structured according to local magnetic field conditions.

Both forms could result from many small size sources. Respective source spatial distribution might be different.

- Challenging question: how to conciliate numerous, micro sources, with well established beams of large aperture ?

## Perspectives

Close analysis of ES sources (PEACE, WBD, WHISPER) plus simulation, toward a better understanding of source behavior (stability, power)