

# Studying Mars with OMEGA

(Observatoire pour la Minéralogie, l'Eau, les Glaces et  
l'Activité)



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## Agenda

- OMEGA
- Reflectance spectroscopy
- Tyrrhena Terra, Mars

## OMEGA details

- Imaging spectrometer
- Spectral range:  $0.38\mu\text{m} - 5.1\mu\text{m}$ 
  - Two channels: VNIR and SWIR
  - VNIR pushbroom
  - SWIR whiskbroom
- Spatial resolution ~350m
- Spectral resolution 7-20nm
  - 352 ‘spectels’

## What can we do with OMEGA?

- Determine the composition of
  - Dust
  - Ices
  - Atmospheric gases
  - Surface rock
- Through spectra showing
  - Absorption features
  - Thermal emission

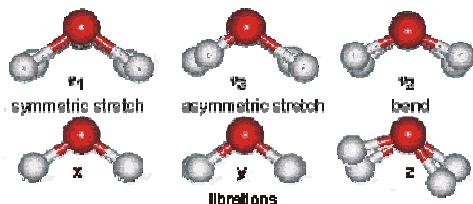
## Reflectance spectroscopy

- “*The study of light as a function of wavelength that has been reflected or scattered from a solid, liquid, or gas*”
- Diagnostic absorption processes
  - Crystal field absorptions (VIS-NIR)
  - Charge transfer (UV-VIS)
  - Vibrational modes of atomic bonds (NIR)

## So, what are we looking for?

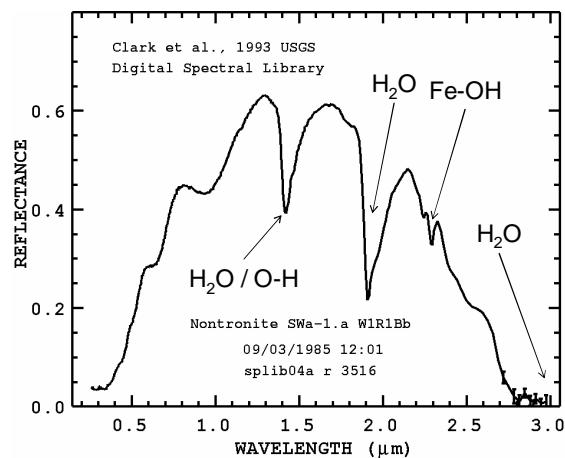
- Evidence for liquid water
  - Phyllosilicates (clay)
  - Sulphates (gypsum, monohydrates)
- Volcanics
  - Olivine  $(\text{Mg}, \text{Fe})_2\text{SiO}_4$
  - Pyroxenes  $\text{XY}(\text{Si}, \text{Al})_2\text{O}_6$
- Organic activity
  - Carbonates

## $\text{H}_2\text{O}$ , vibrational modes



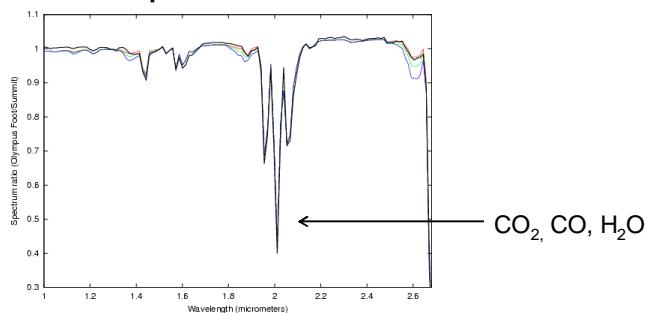
	Feature	Gas	Liquid	Ice
Fundamentals	$v_1$	2.74 $\mu\text{m}$	3.05 $\mu\text{m}$	3.17 $\mu\text{m}$
	$v_3$	2.66 $\mu\text{m}$	2.87 $\mu\text{m}$	2.96 $\mu\text{m}$
Combinations	$v_2+v_3$	1.875 $\mu\text{m}$	1.93 $\mu\text{m}$	2.02 $\mu\text{m}$

## Lab spectrum - Nontronite



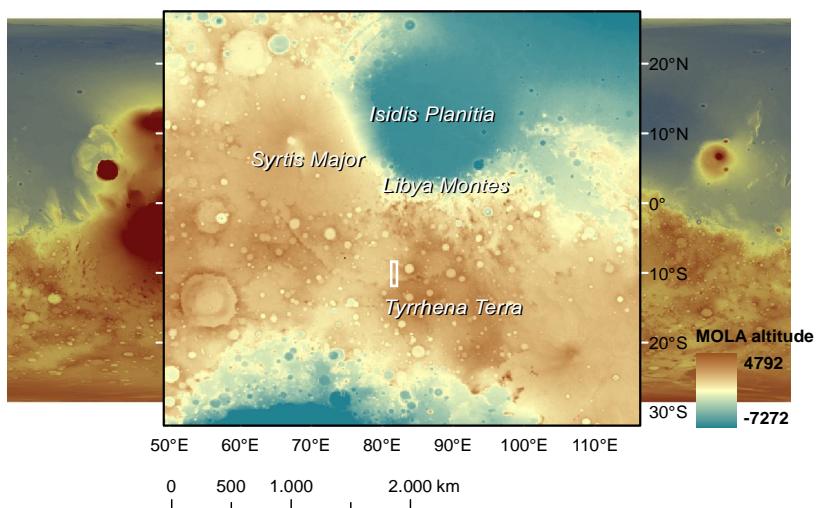
## Spectral detection

- Atmospheric correction

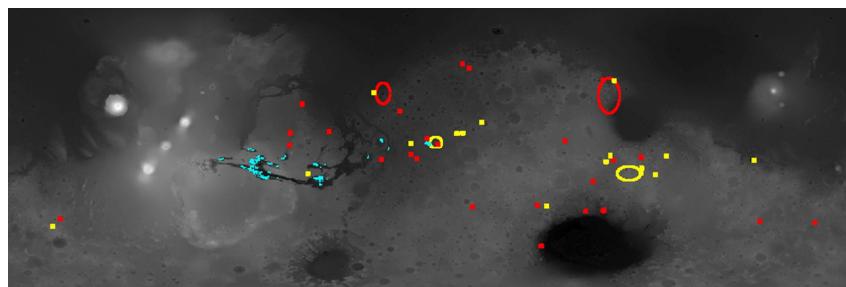


- Summary (ratio) products
  - Ratio band center with shoulders
  - Pelkey et al. 2007, JGR 112

## Tyrrhenia Terra, Mars

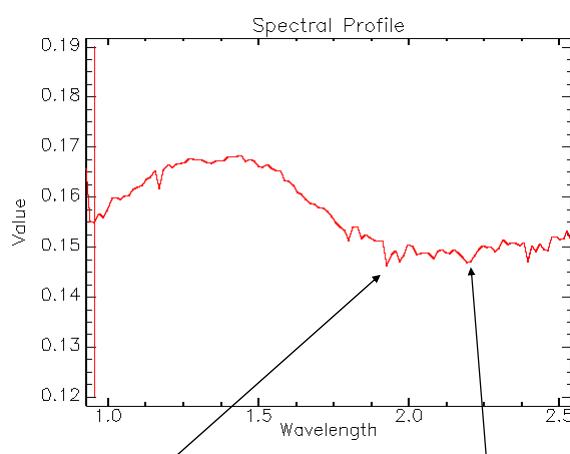


## Hydrated mineralogy



- Which process?
  - Sedimentary?
  - Hydrothermal?

## Sample OMEGA spectrum



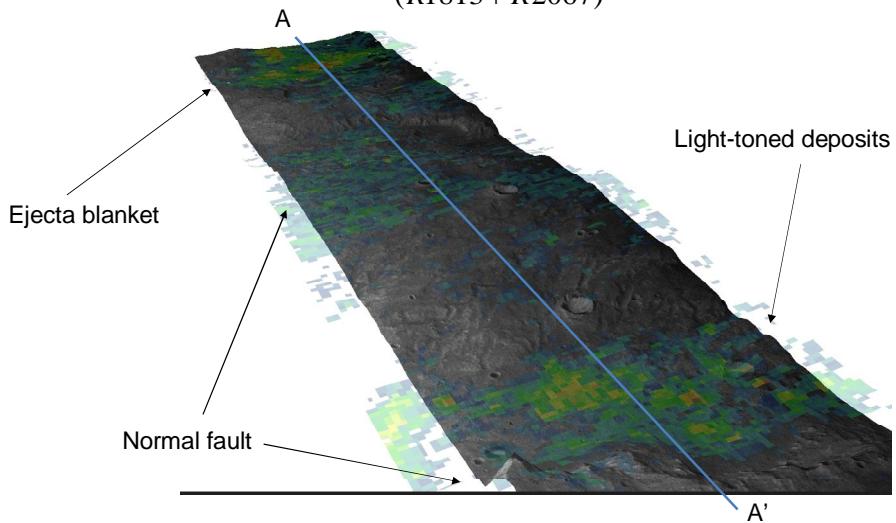
- Feature: 1.93  $\mu\text{m}$ , Feature: 2.2  $\mu\text{m}$
- Overall shape: Pyroxene  $\rightarrow$  Intimate mixture

## OMEGA issues

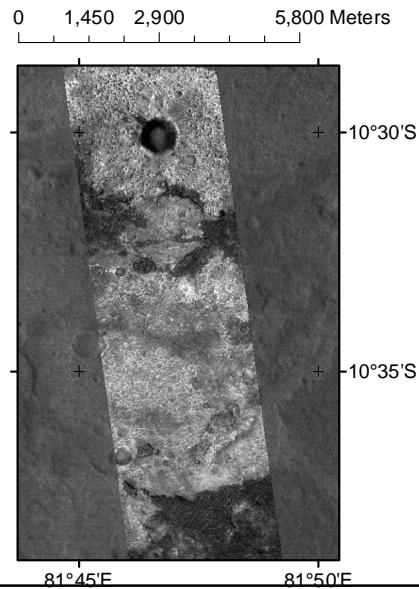
- Relatively low S/N ratio
- Hot / cold spectels
- Atmospheric variability
- “Mixels” due to spatial resolution
- Dust wipes out features

## Geological context

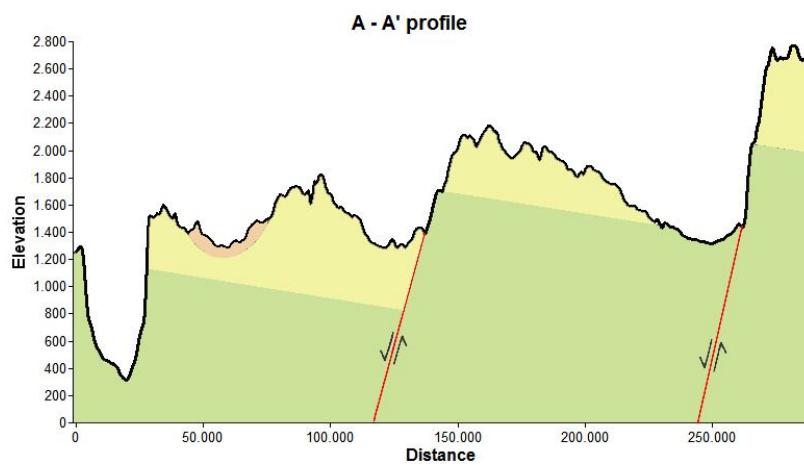
- Spectral ratio  $1 - \frac{(R1930 + R1940)}{(R1813 + R2067)}$



## Light-toned deposits



## Cross-section



## Conclusions

- OMEGA S/N troublesome
  - Visible geology verification needed
- Hydrated layer  $10^2$ m thick
- Two episodes of hydration
  - Channels and hydrated unit
- Likely to be of hydrothermal origin
  - No layering visible
  - Thickness
  - Proximity to major fault region