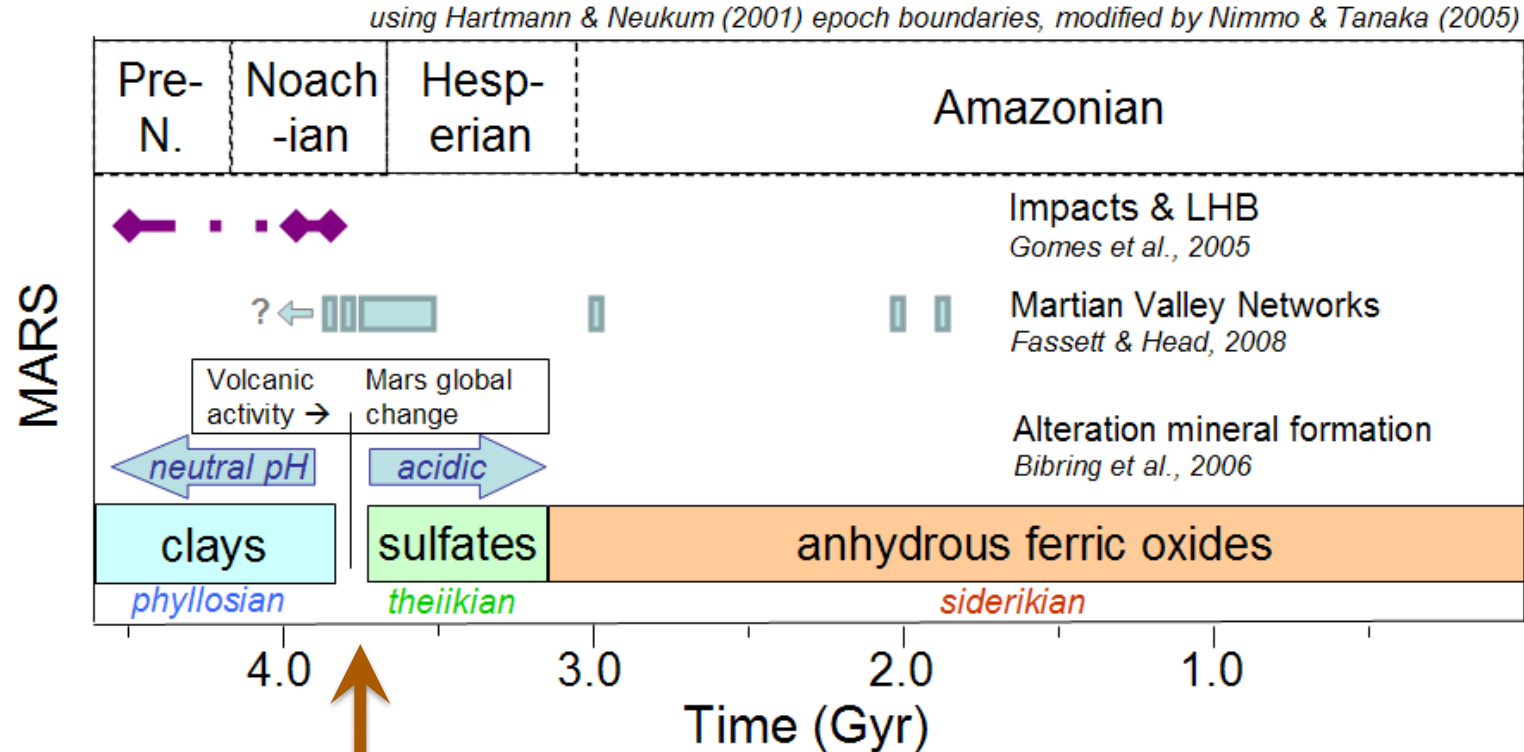


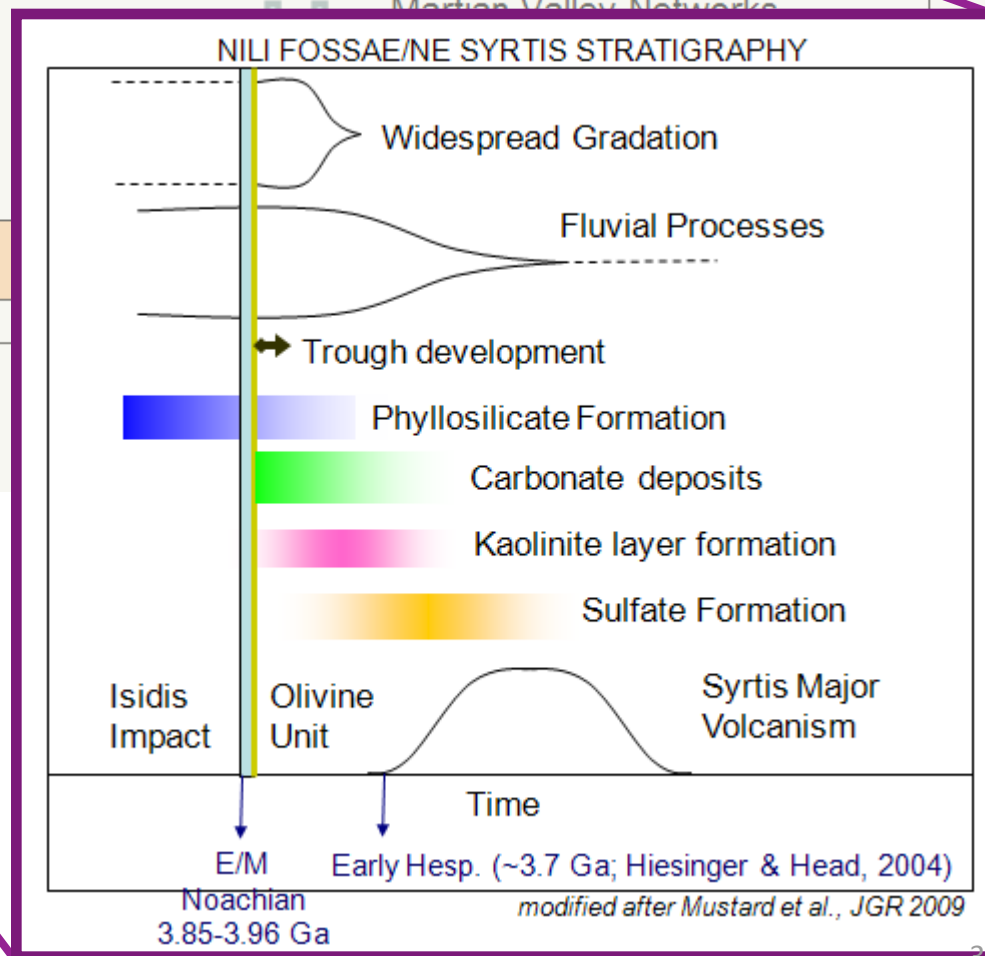
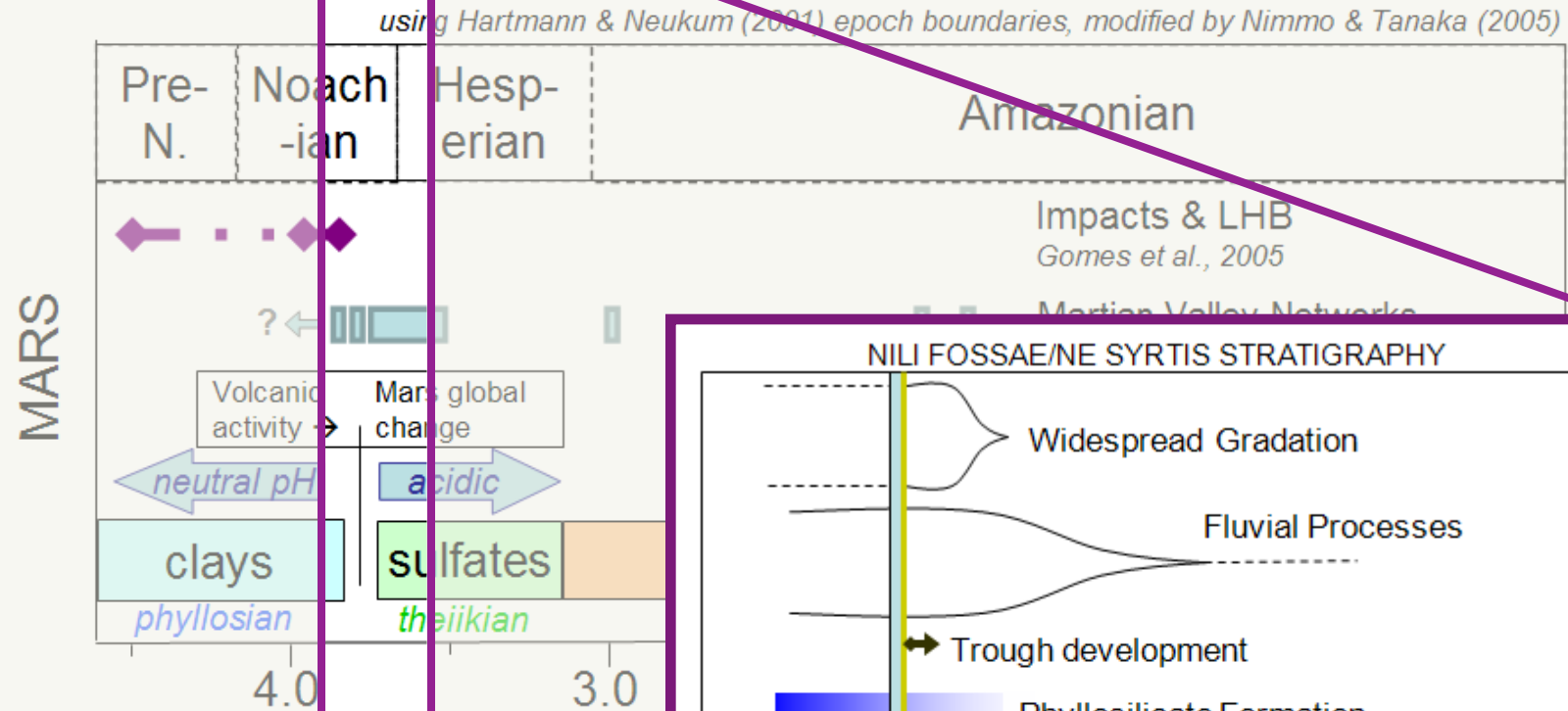
DIVERSE HABITABLE ENVIRONMENTS DURING THE NOACHIAN-HESPERIAN TRANSITION

Jack Mustard, Bethany Ehlmann, J.R. Skok
Brown University and IAS-Paris

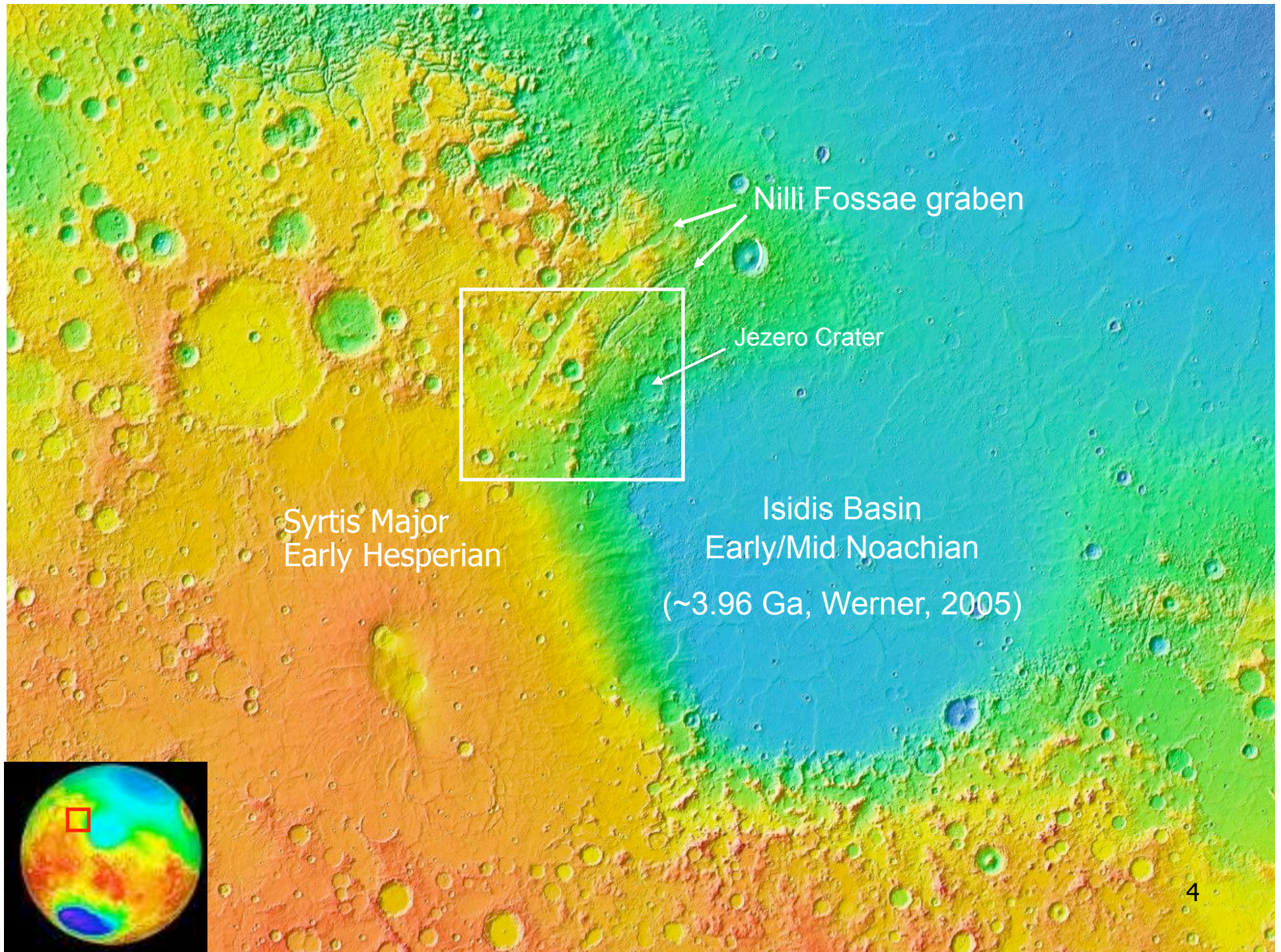
Era of Major Reorganization of Planetary-scale processes



Distinct Geologic Units in Stratigraphic Relationship and Relative Ages
Do they Mineral Assemblages and Geologic Contexts Define Habitable Boundary Conditions?



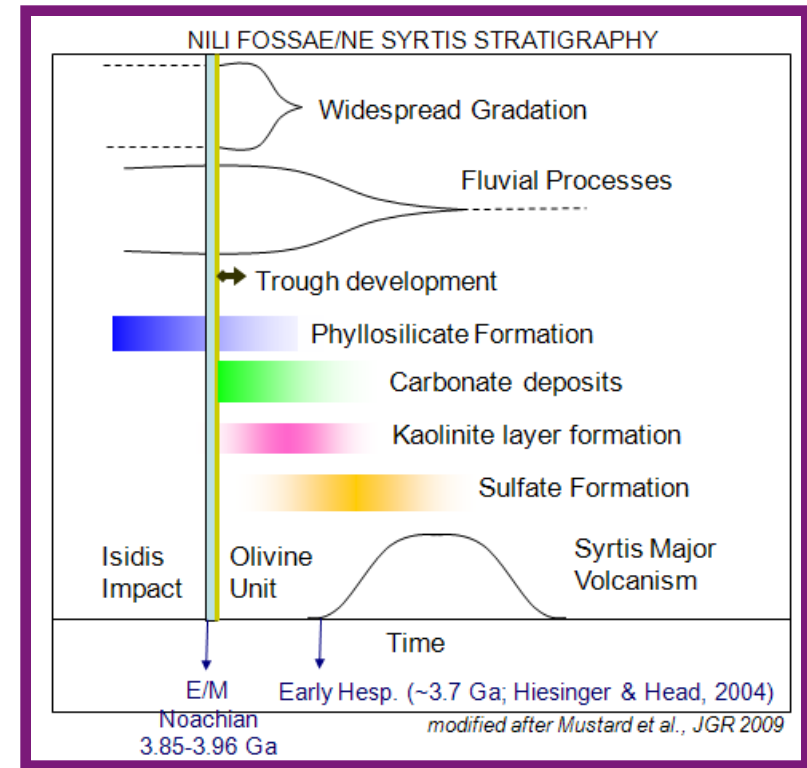
Stratigraphy of Nili Fossae/NE Syrtis records multiple aqueous environments from the Middle Noachian to Early Hesperian



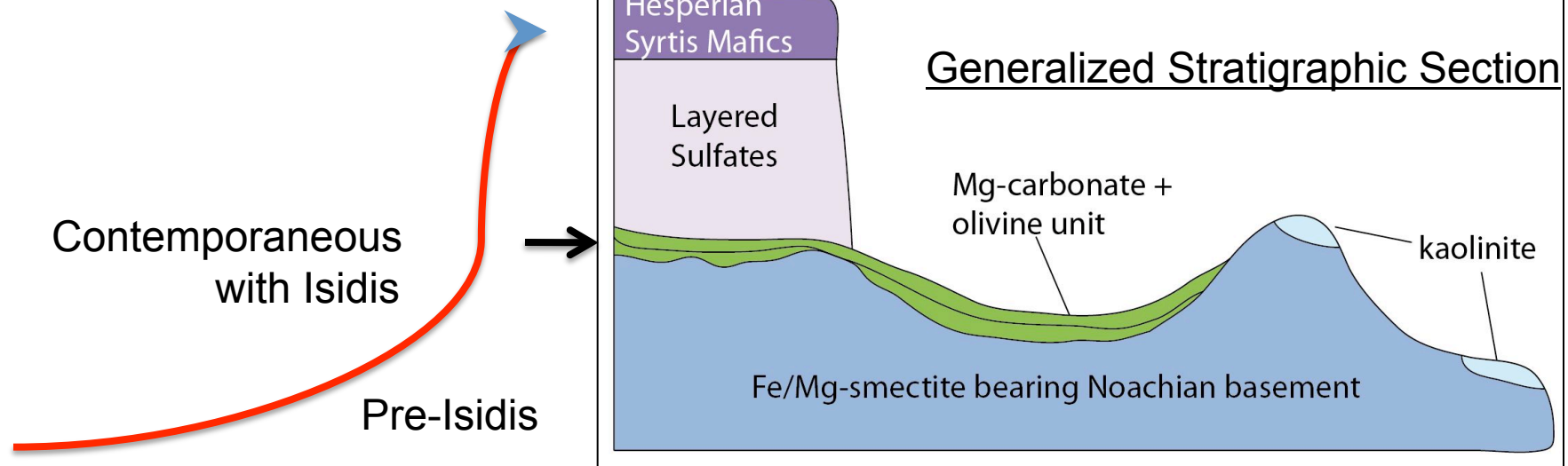
What are the processes and environments that formed early phyllosilicates?

How did conditions change with carbonate formation?

How rapid was the transition to a more acidic, sulfate (jarosite) forming era?



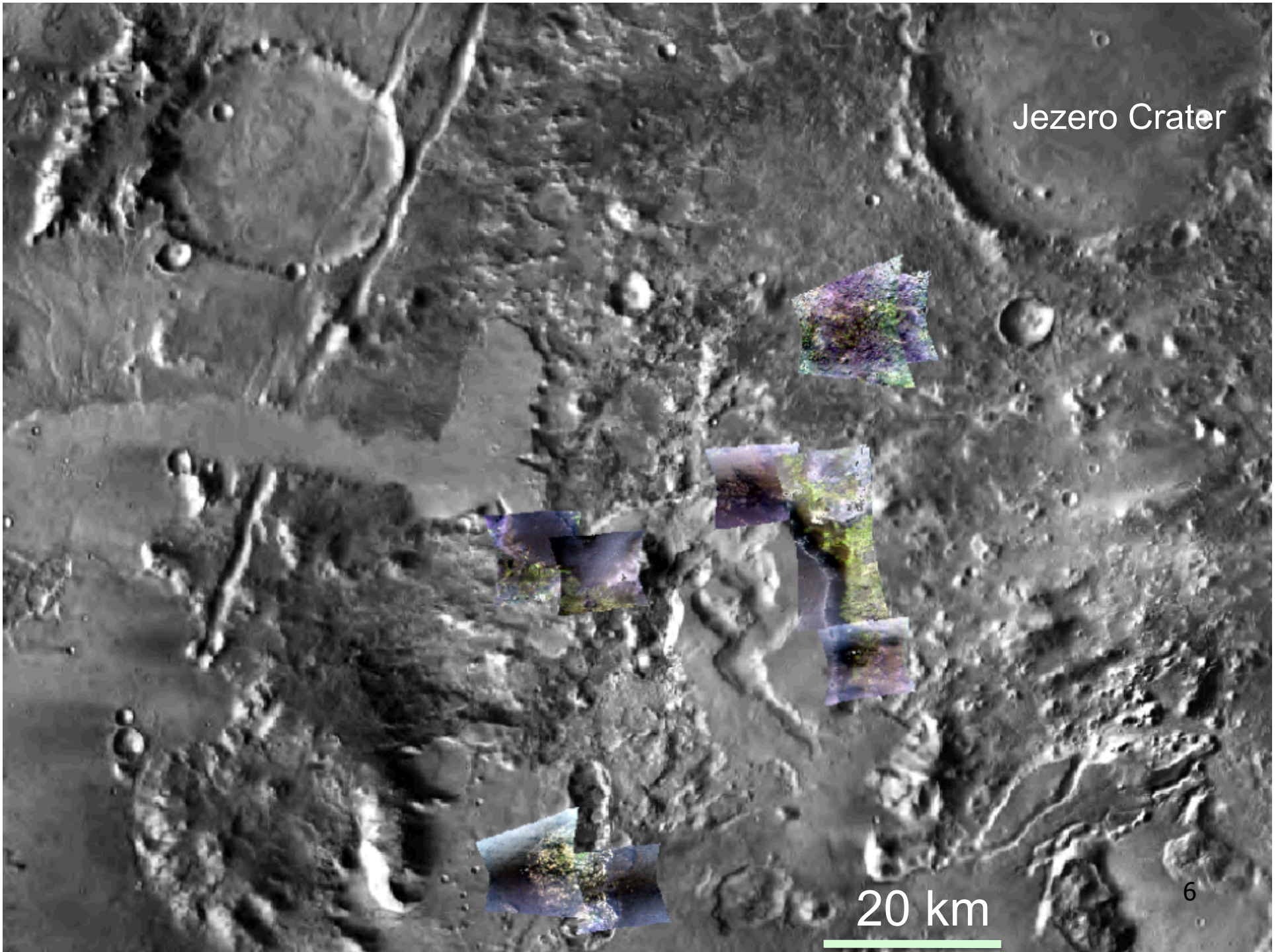
A Traverse Up the Section



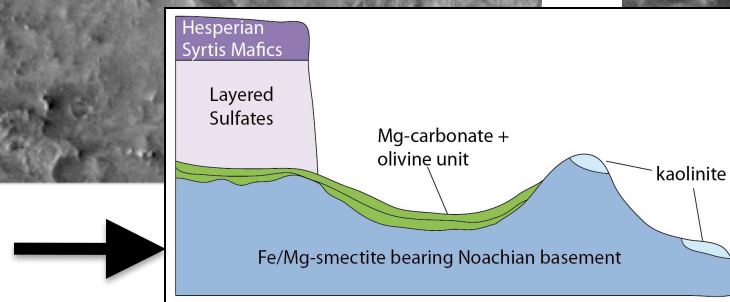
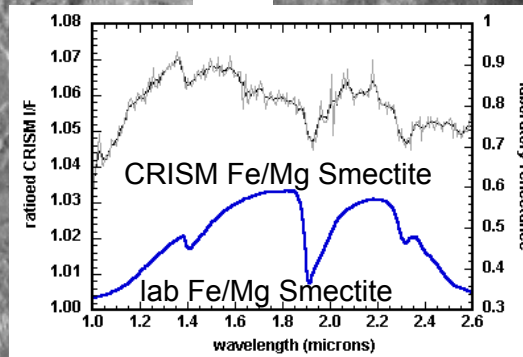
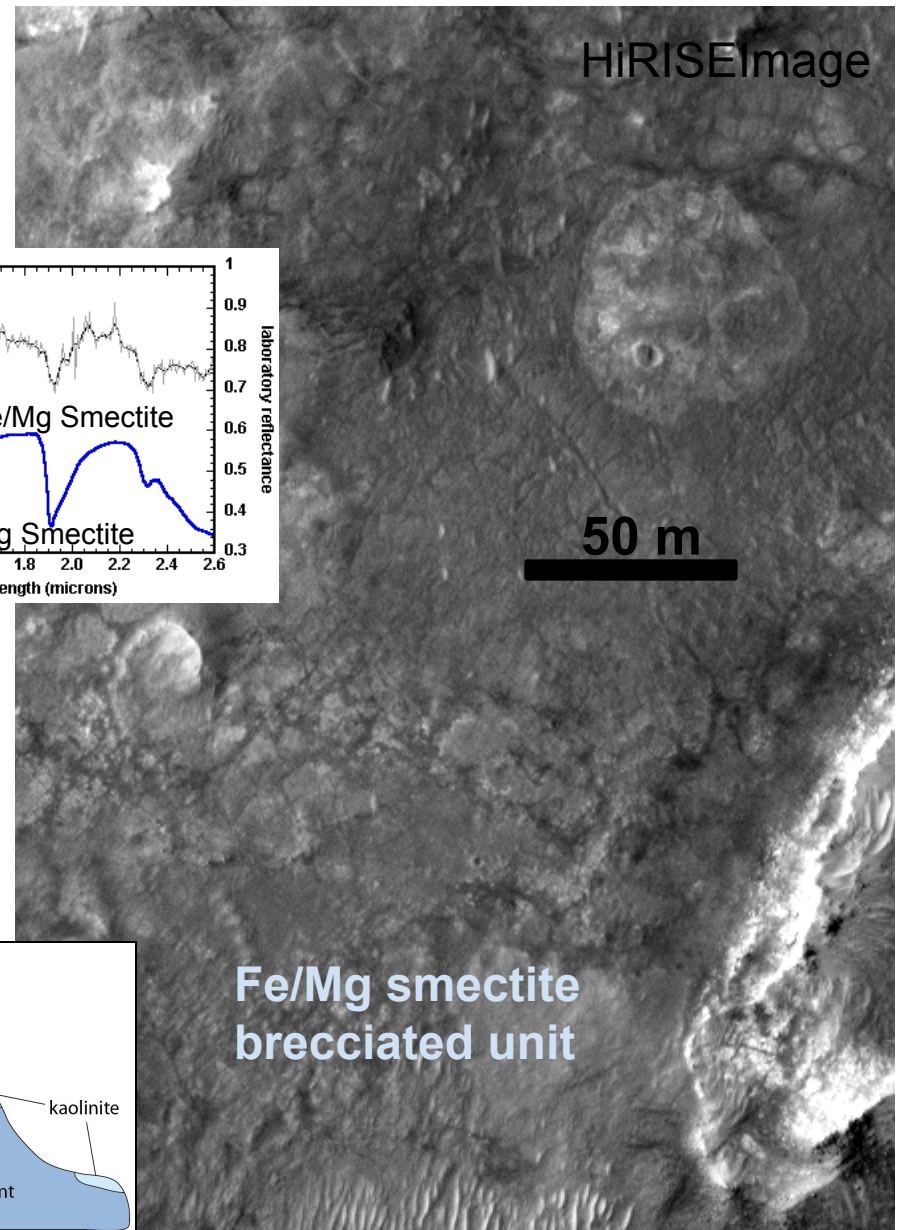
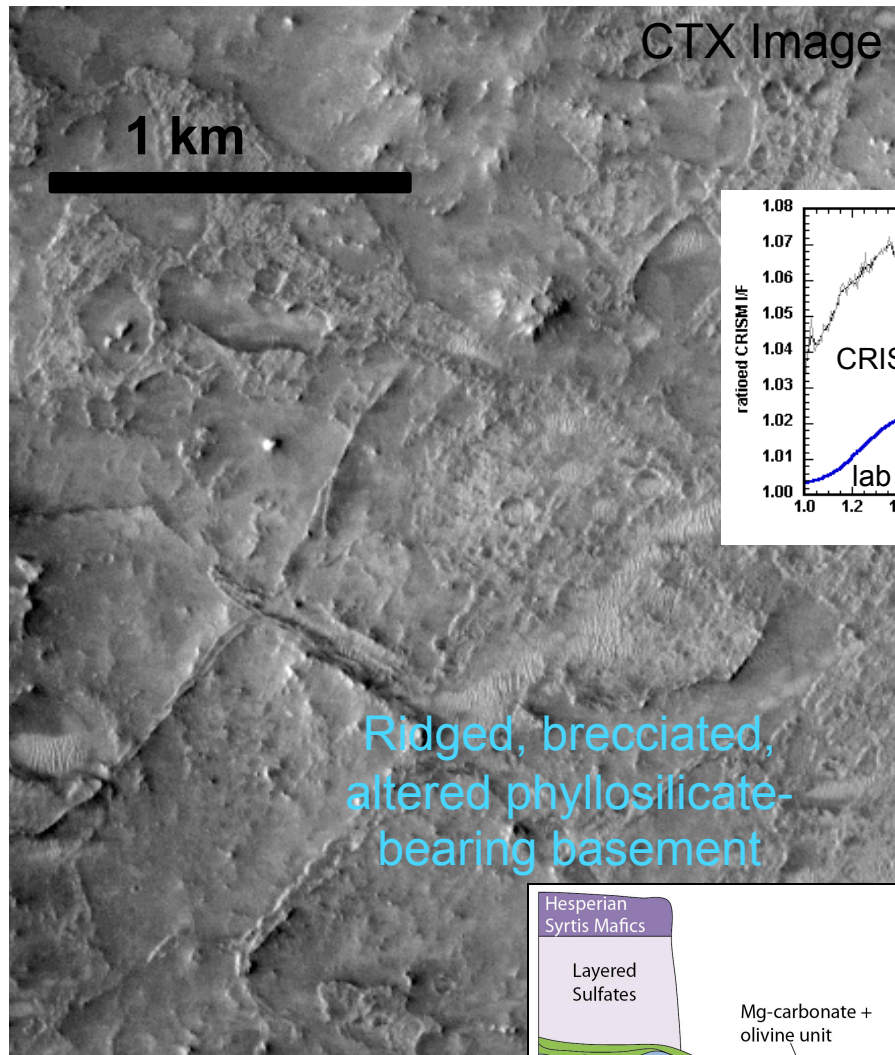
Jezero Crater

20 km

6

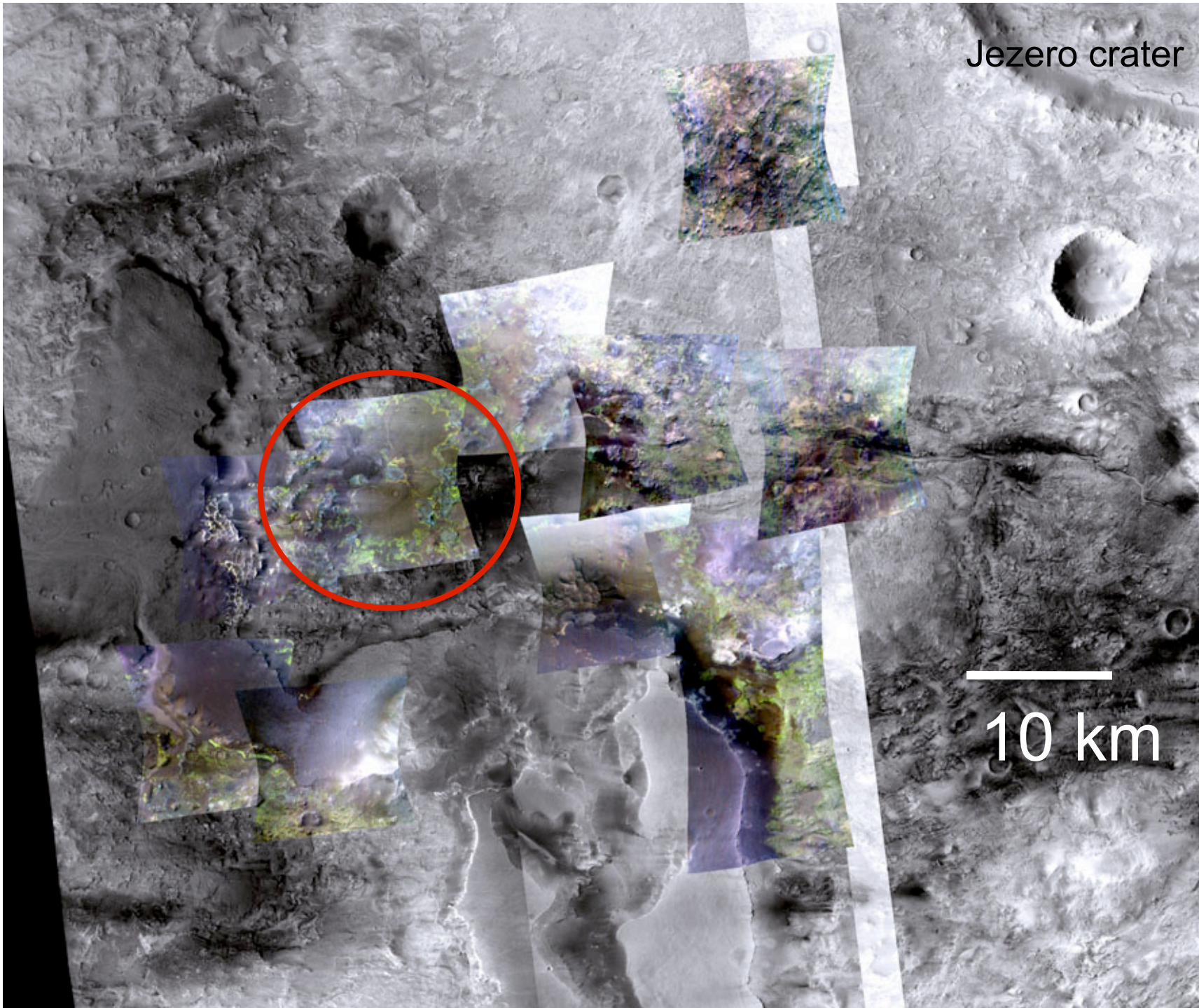


Noachian Basement exposed on the plains shows breccia blocks of unaltered pyroxene-bearing rocks set in a matrix of Fe/Mg phyllosilicate and laced with raised ridges

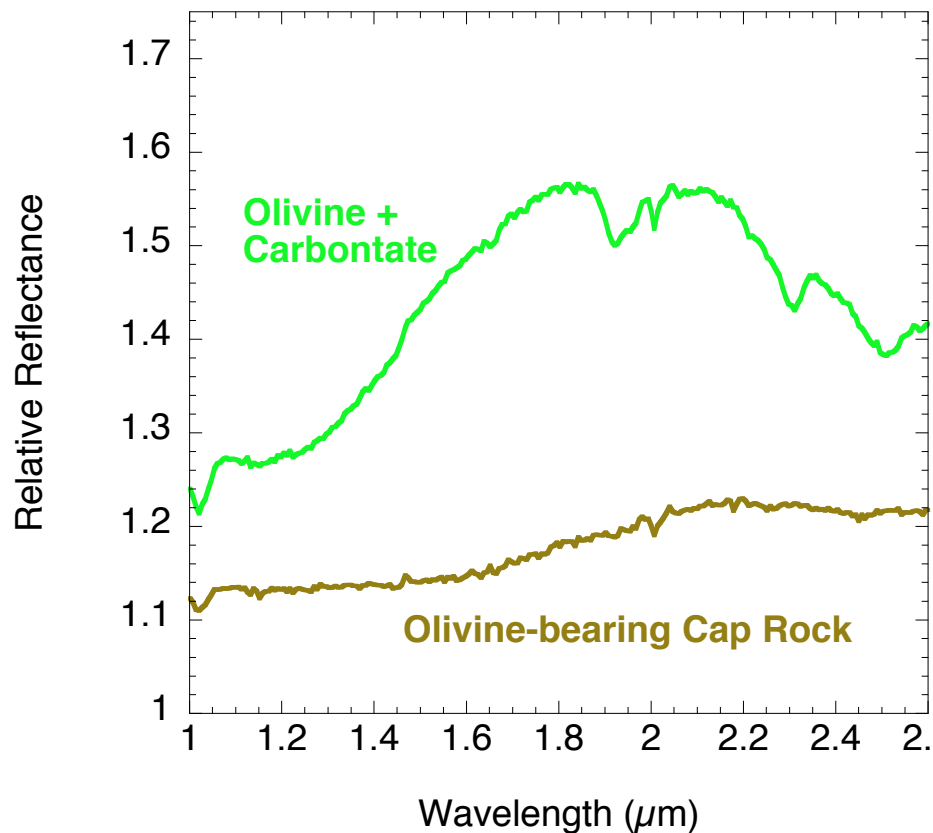
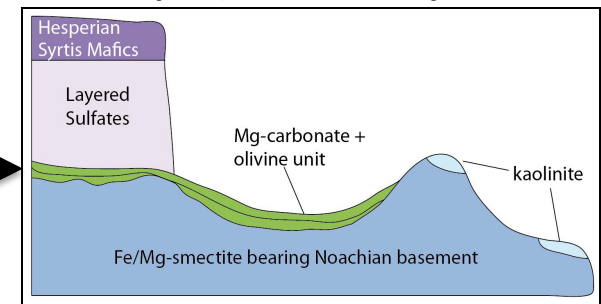


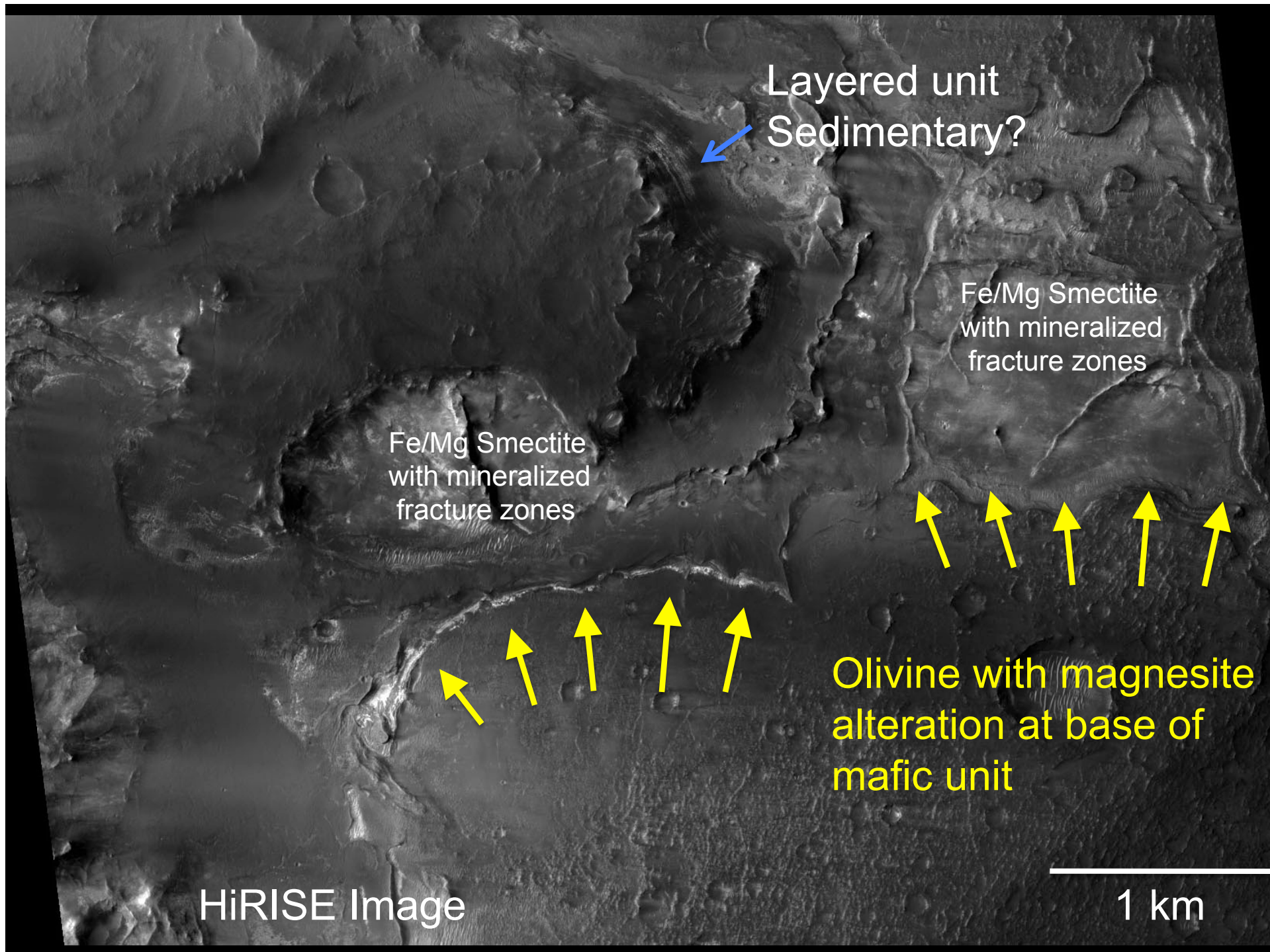
Jezero crater

10 km

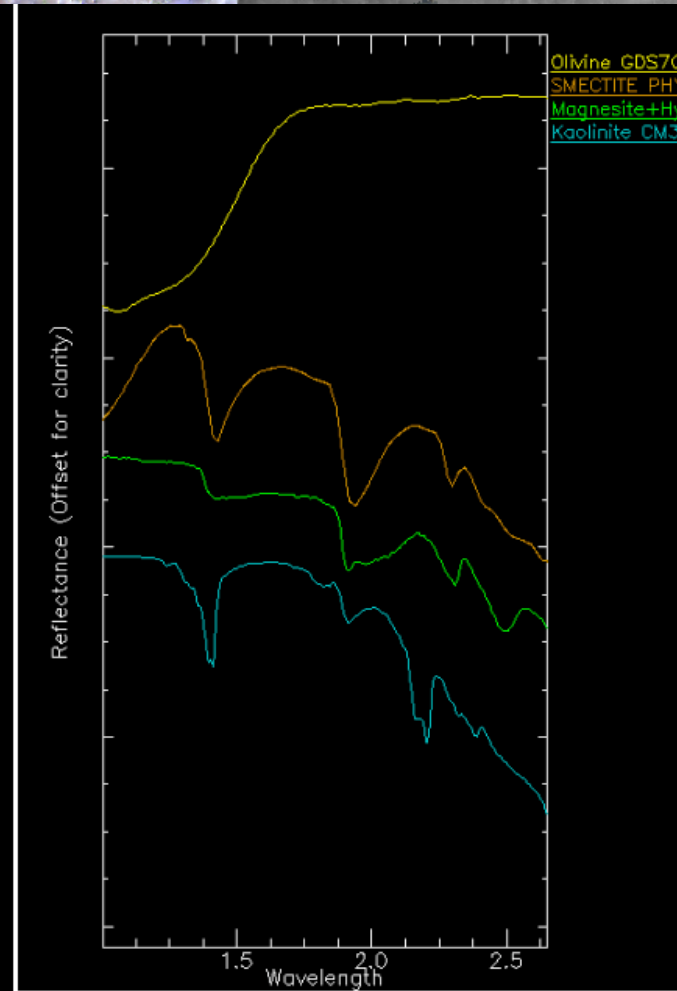
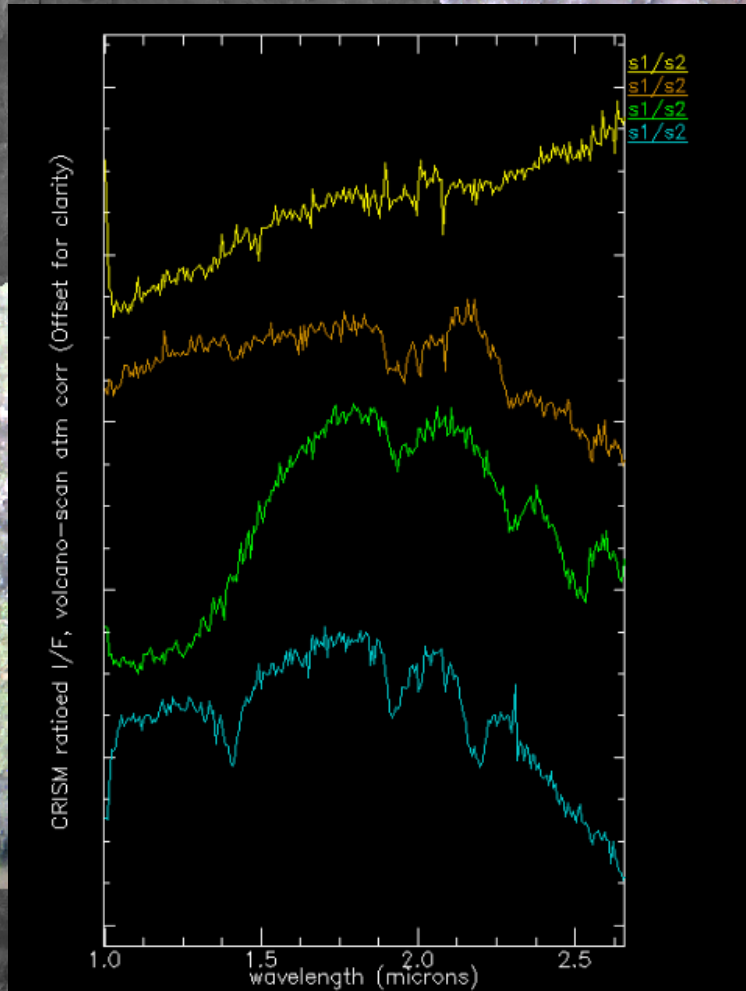
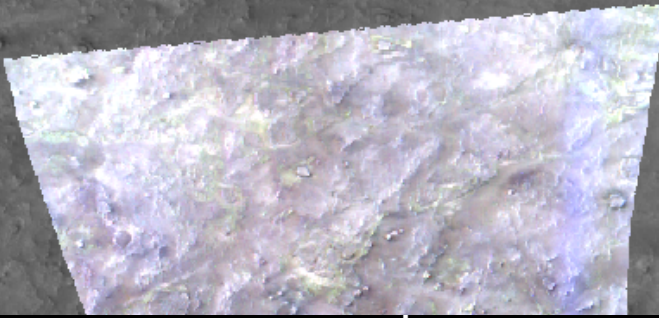


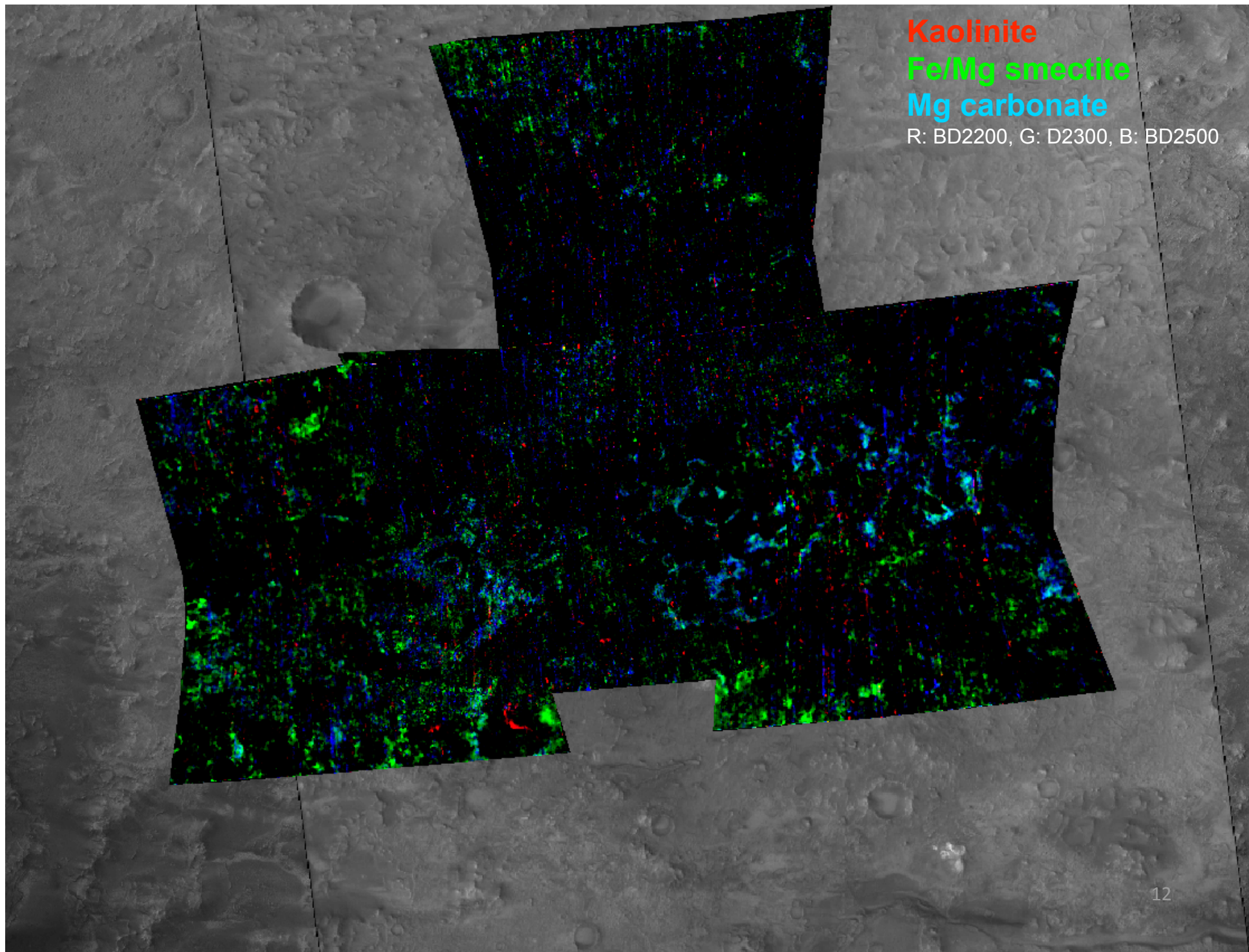
Distinct Mafic to Ultramafic Deposit (regional in scope) that is post
Isidis in age but pre-fossae:
Impact melt or ultramafic lava flows
Rests directly on Fe/Mg altered basement
The base of the deposit hosts the
Mg-carbonate deposits





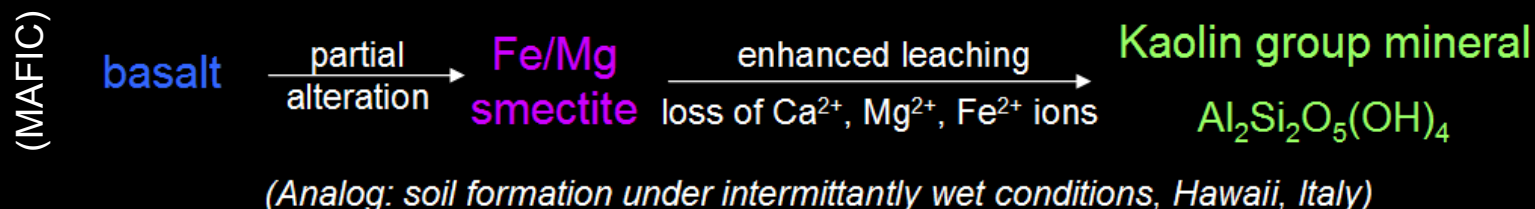
Plains Deposits Show
Kaolinite Mineral
Deposits Indicative of
Surface Water



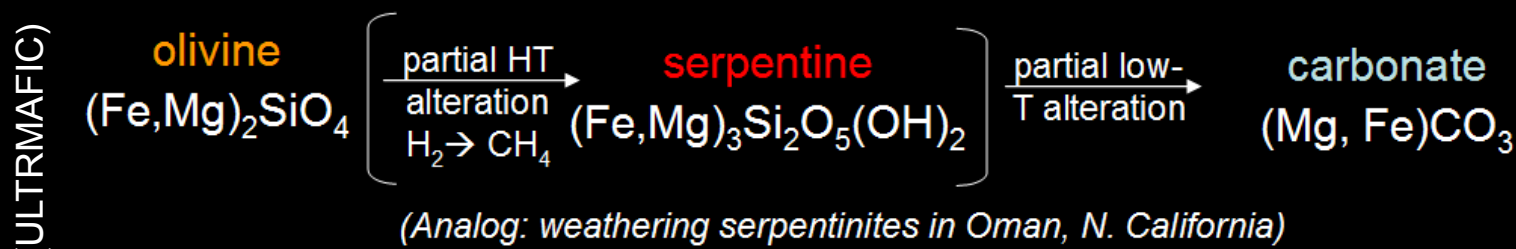


Kaolinite, carbonate, and serpentine mineralization

- Chemistry controlled by precursor mineralogy



--or--



- Carbonate(/Serpentine): Carbonate alteration mostly related to near-surface hydrology? Or mostly generated by hydrothermal systems?
 - H_2 an energy source for organisms, potential for methane production during serpentine formation
- Kaolinite: Leaching from an active hydrologic system?

Layered Sequence beneath Syrtis Lavas
How do mineral assemblages change with location?

Jezero Crater
floor age
3.03 Gy

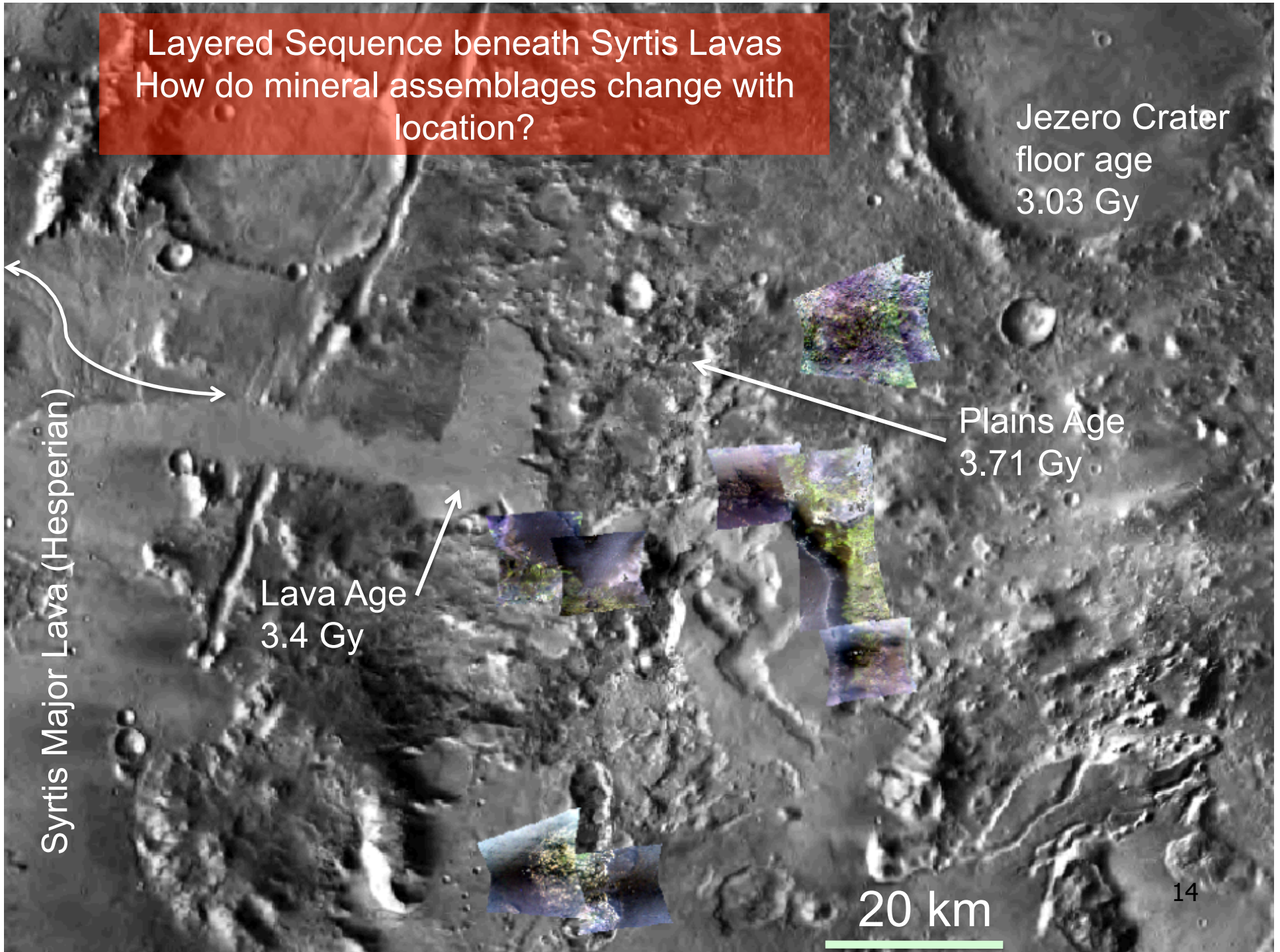
Plains Age
3.71 Gy

Lava Age
3.4 Gy

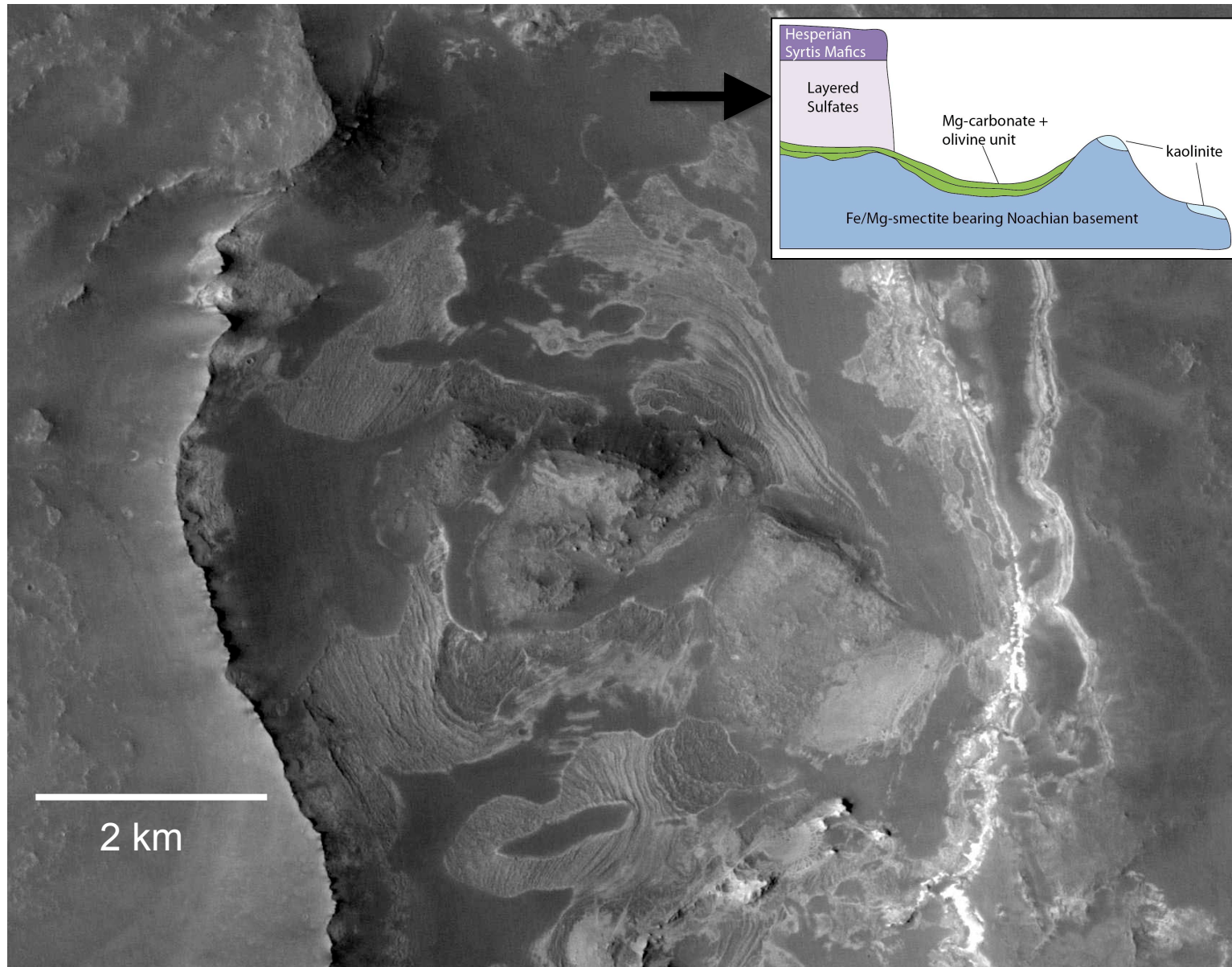
20 km

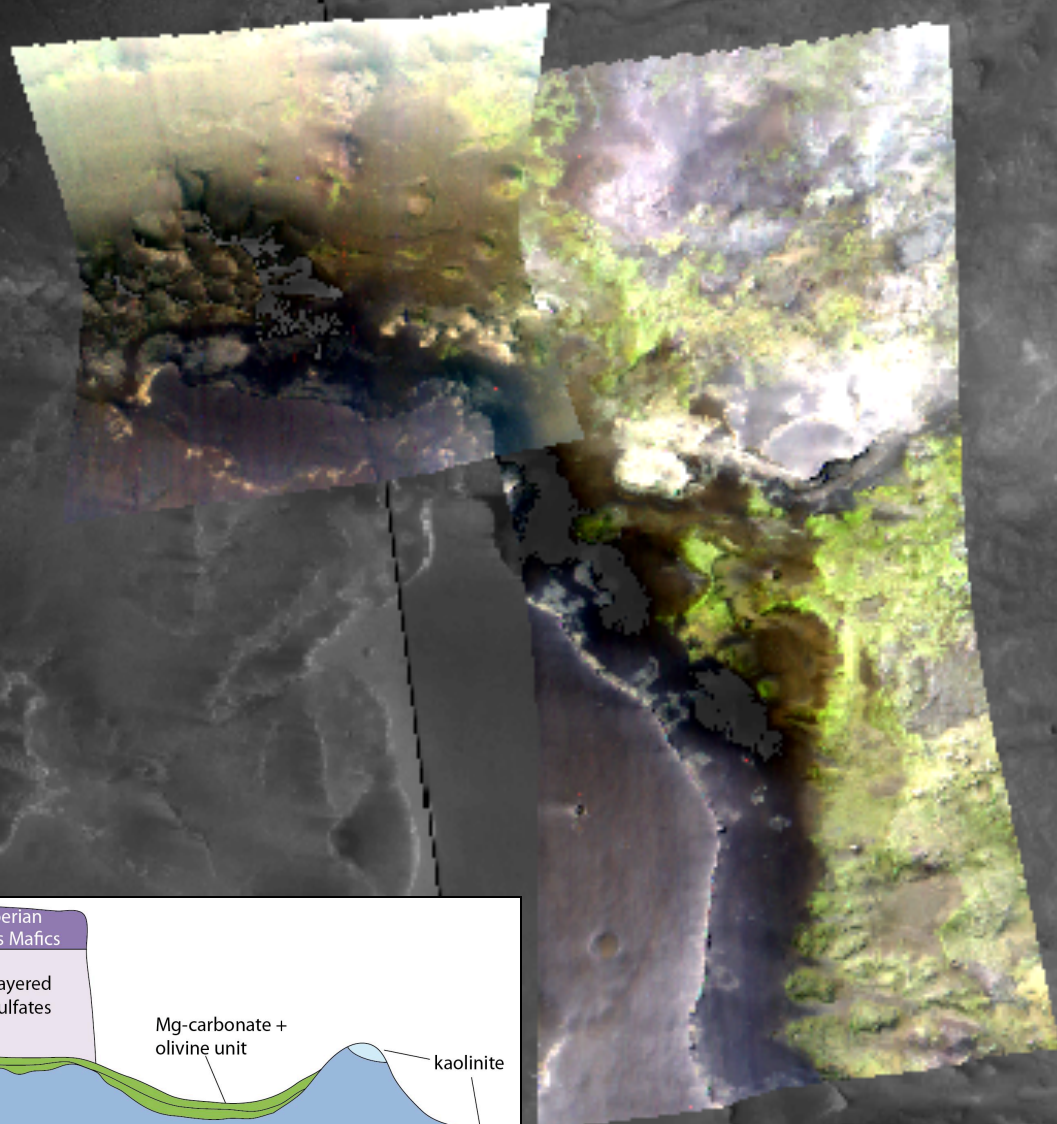
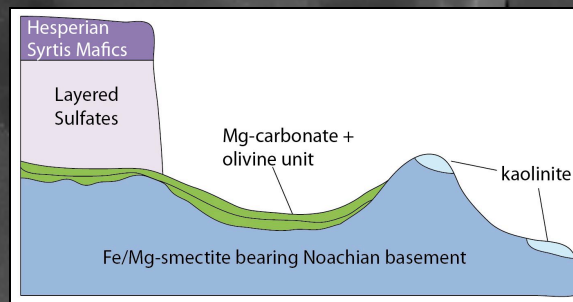
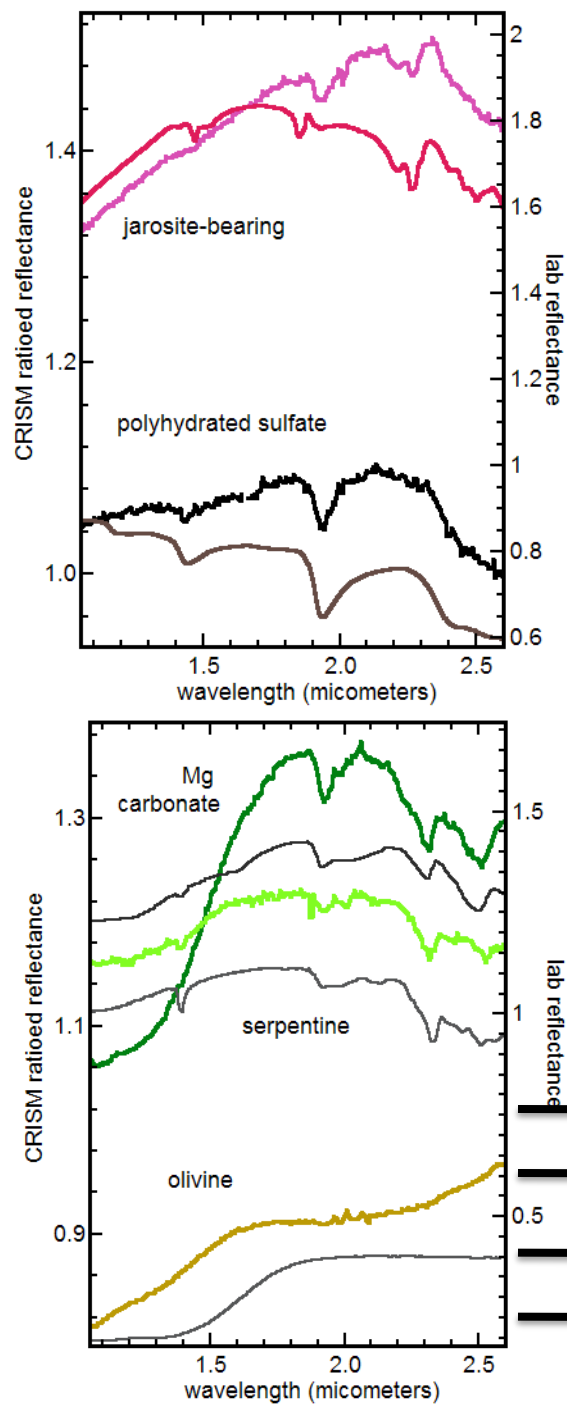
14

Syrtis Major Lava (Hesperian)

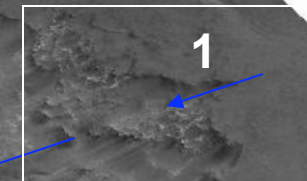


Beneath the capping lava is a 500 m stack of layered rock that appears sedimentary

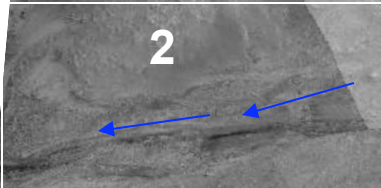




Syrtis Major Materials, capping



1



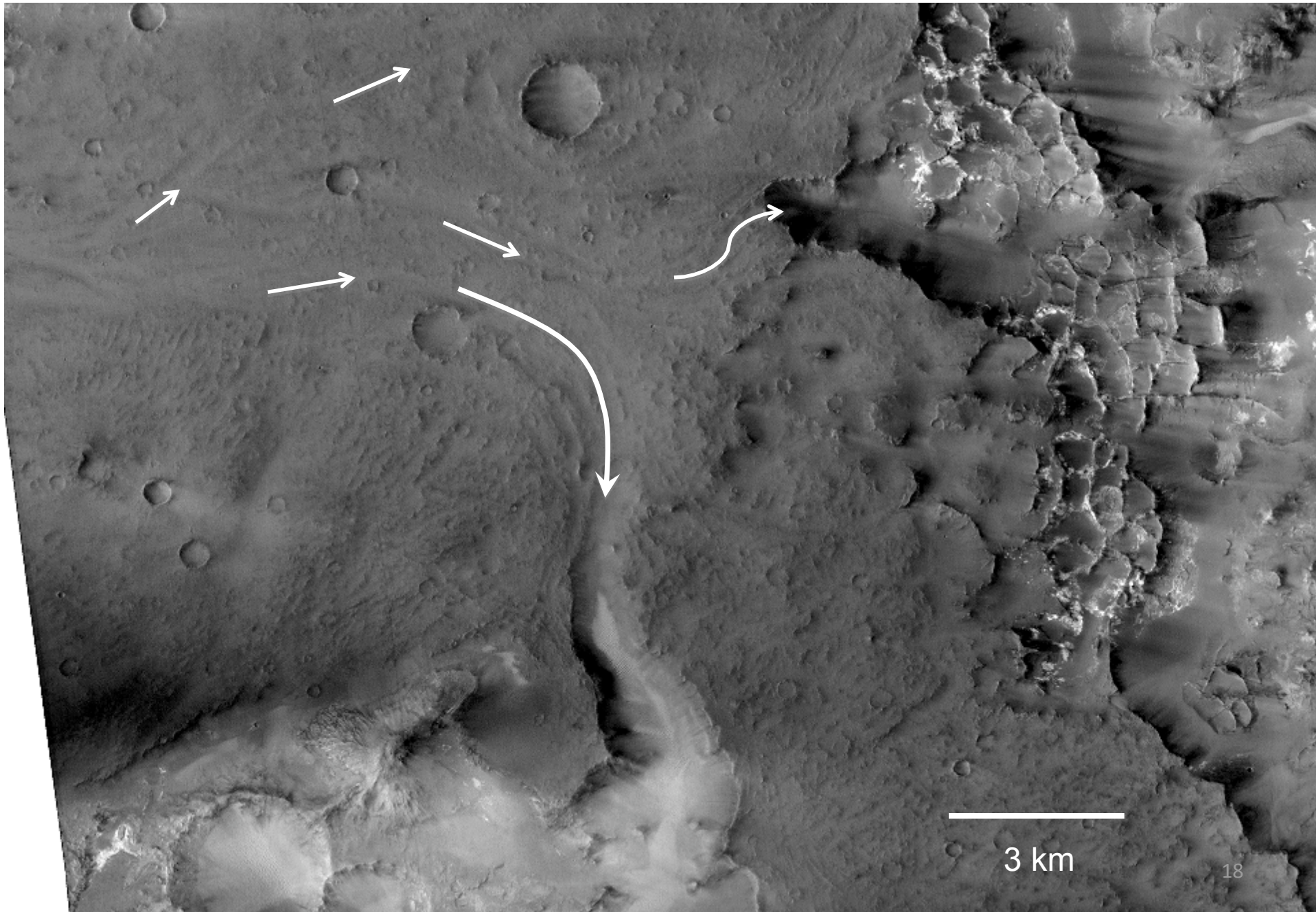
2

topographic low



CTX+HiRISE over MOLA
topography perspective view, 7x¹⁷

1. NE Syrtis inflow



Noachian to Hesperian Transition: An Acid-Alkaline Processes

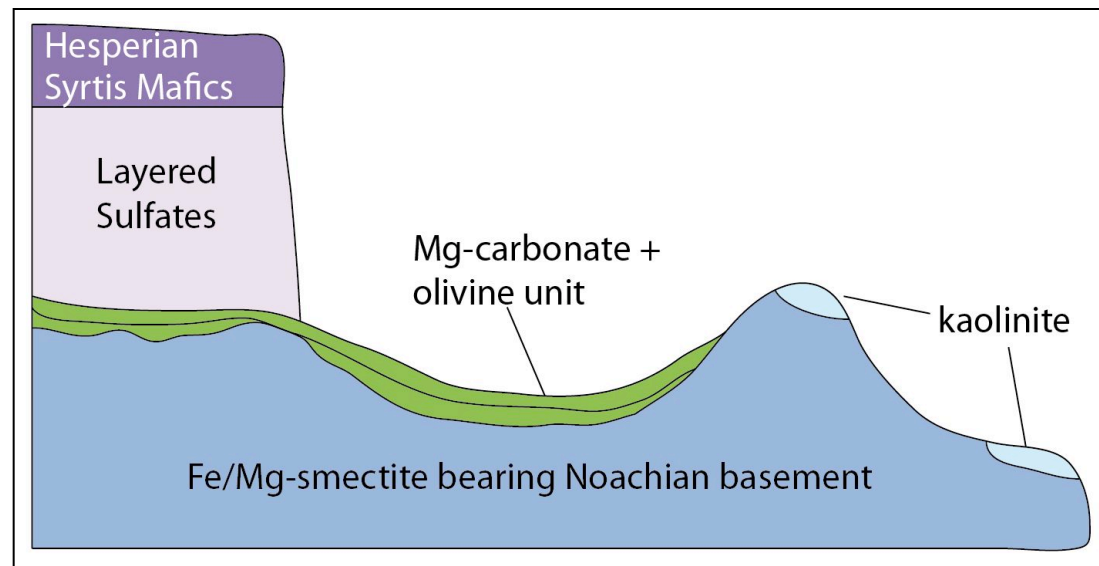
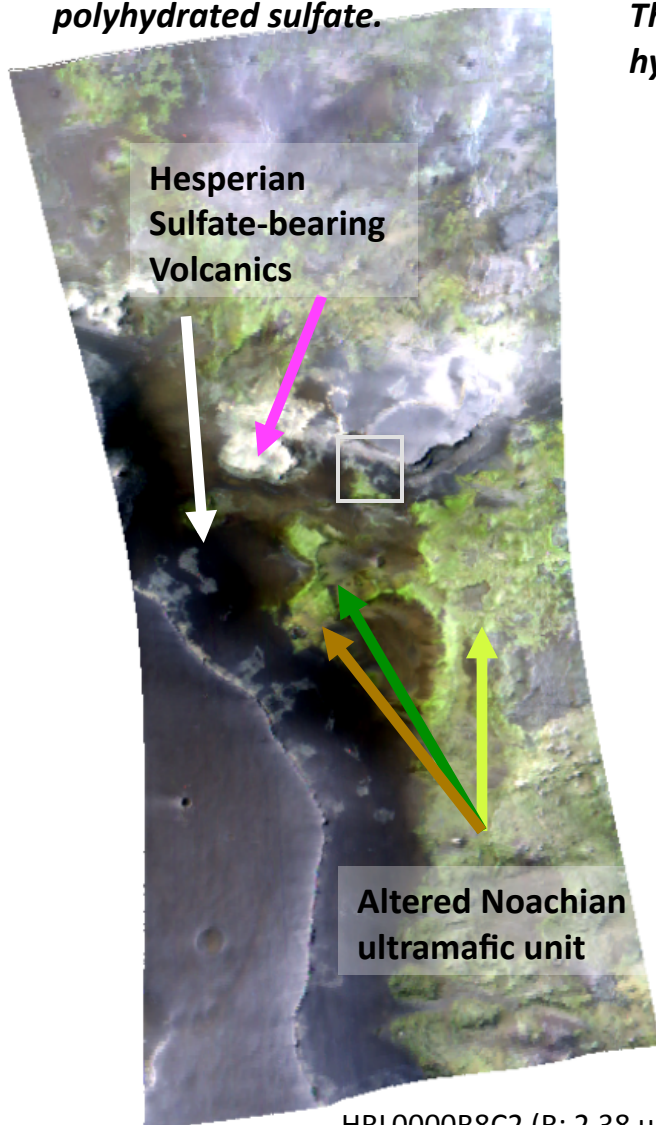
This Noachian-Hesperian stratigraphy transitions from ancient phyllosilicate to carbonate (alkaline) and kaolinite to sulfate (acidic)

Lavas emplaced in a volatile-rich environment overlie (sedimentary?) sulfate-bearing deposits that include jarosite and polyhydrated sulfate.

There is evidence for circulation of fluids by heat of lava source and thus hydrothermal systems in a volcano-ice environment

Clear stratigraphy of unaltered volcanic units, sulfates, carbonate-serpentine-olivine unit and Fe/Mg smectite basement

Long record of aqueous geochemistry recorded by alteration minerals



HRL0000B8C2 (R: 2.38 μm , G: 1.80 μm , B: 1.15 μm)

Diversity of Habitable Environments (1)

- Shallow Subsurface Hydrothermal Systems
 - Fe/Mg Phyllosilicate with evidence of fluid flow
- Surface Fluvial Systems
 - Network of fluvial channels and low regions showing open basin lake system (e.g. Jezero)
 - Kaolinite in capping rocks indicative of water flow in regolith-pedogenesis
- Reactant-product Mineral associations
 - Olivine => Serpentine: production of H_2 - CH_4
 - Olivine => Magnesite (Mg-carbonate)

Diversity of Habitable Environments (2)

- Volcano-ice Interaction at the Isidis Basin Edge
 - Volcanic morphology suggests lava-ice interaction
- Layered Material Beneath Hesperian Lava
 - Sulfate rich with polyhydrated and jarosite minerals
 - Pre-Syrtis sedimentary deposits (in an area predicted to have upwelling groundwater (e.g. Andrews-Hanna)
 - Alteration of Syrtis lava flows
- Transition in preserved habitable environments and major Mars Environmental Change