

Weather Sentinel

Piero Spillantini¹, Marco Casolino², Guido Castellini³

¹Institute of Nuclear Physics (INFN), branch of Firenze, Firenze, Italy

²Institute of Nuclear Physics (INFN), branch of Roma2, Roma, Italy

³Consiglio Nazionale delle Ricerche, area ricerca di Firenze, Italy

Joint Scientific Space Mission Chinese Academy of Science (CAS) - European Space Agency (ESA)

Science objectives

Main science objective:

understanding emission and transmission through heliosphere of energetic particles in correspondence of different kinds of SEEs, by measuring energy spectra and the angular distributions, of the energetic tail of SEEs (electrons, protons, ions and neutrons), and their time evolution, continuously and on a long period of time, possibly a full solar cycle

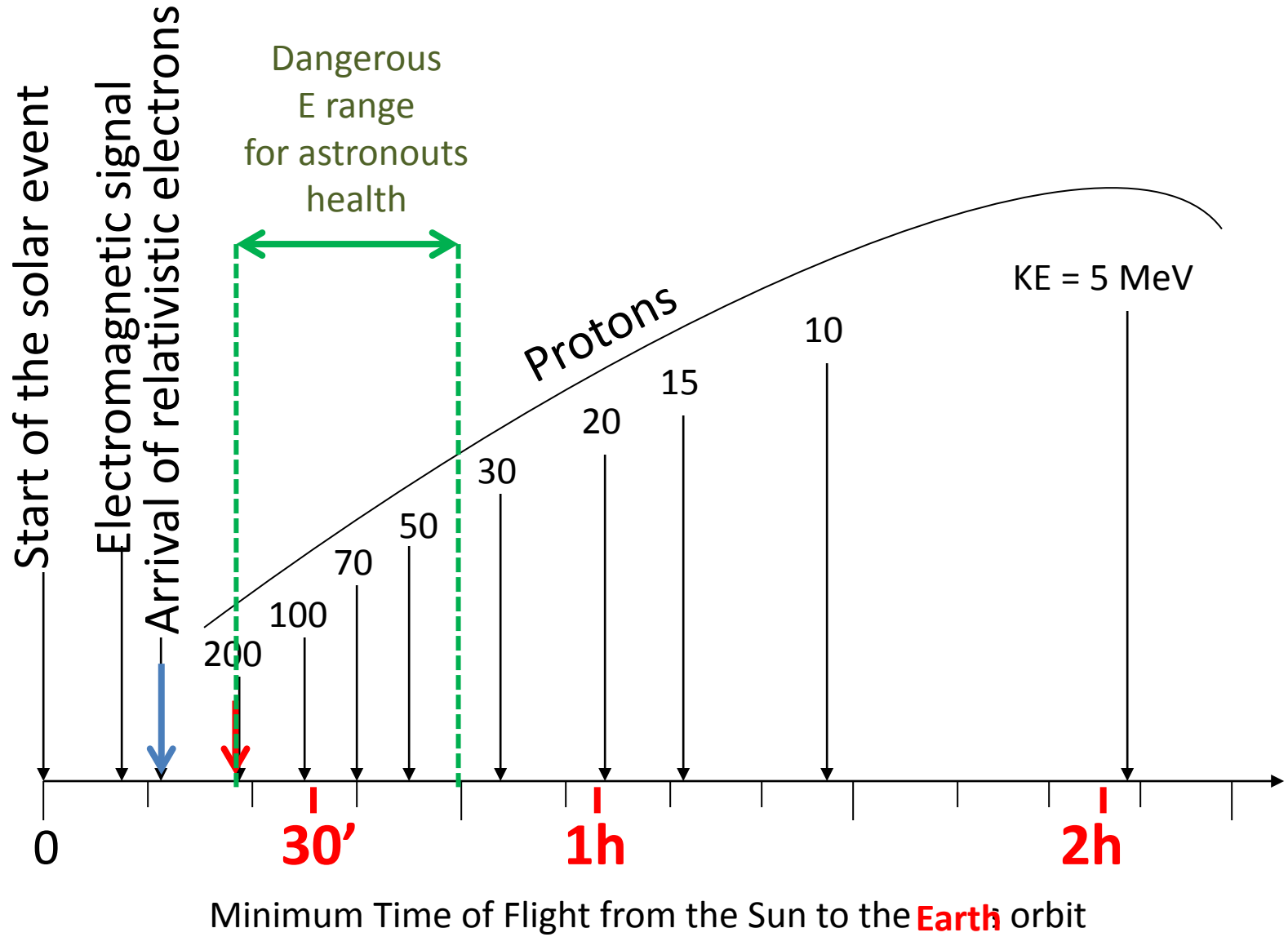
In particular:

knowledge of evolution in time of the angular distribution as a function of energy of electrons and protons in the CME type of events, the most energetic and high fluence SEEs.



maximum importance for the future manned exploration of the solar system:

“knowledge of the arrival direction of the energetic electrons allows the spacecraft to take countermeasures against the shortly arriving, by the same direction, protons and ions”.

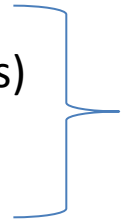


The Instruments

Main instrument:

SPhERA

- **Arrival angle** of the charged particle
(by coincidence of pixels of two concentric spheres)
- **Energy** of the charged particle
(by ToF between the two pixels)



PAMELA and AMS
heritage

Possible complementary instruments:

INTERACTOR

- **neutron energy**
(by position of the neutron interaction +
angle, amplitude and time of the recoiling proton
in the SPhERA instrument)

SUNSPOT MAP

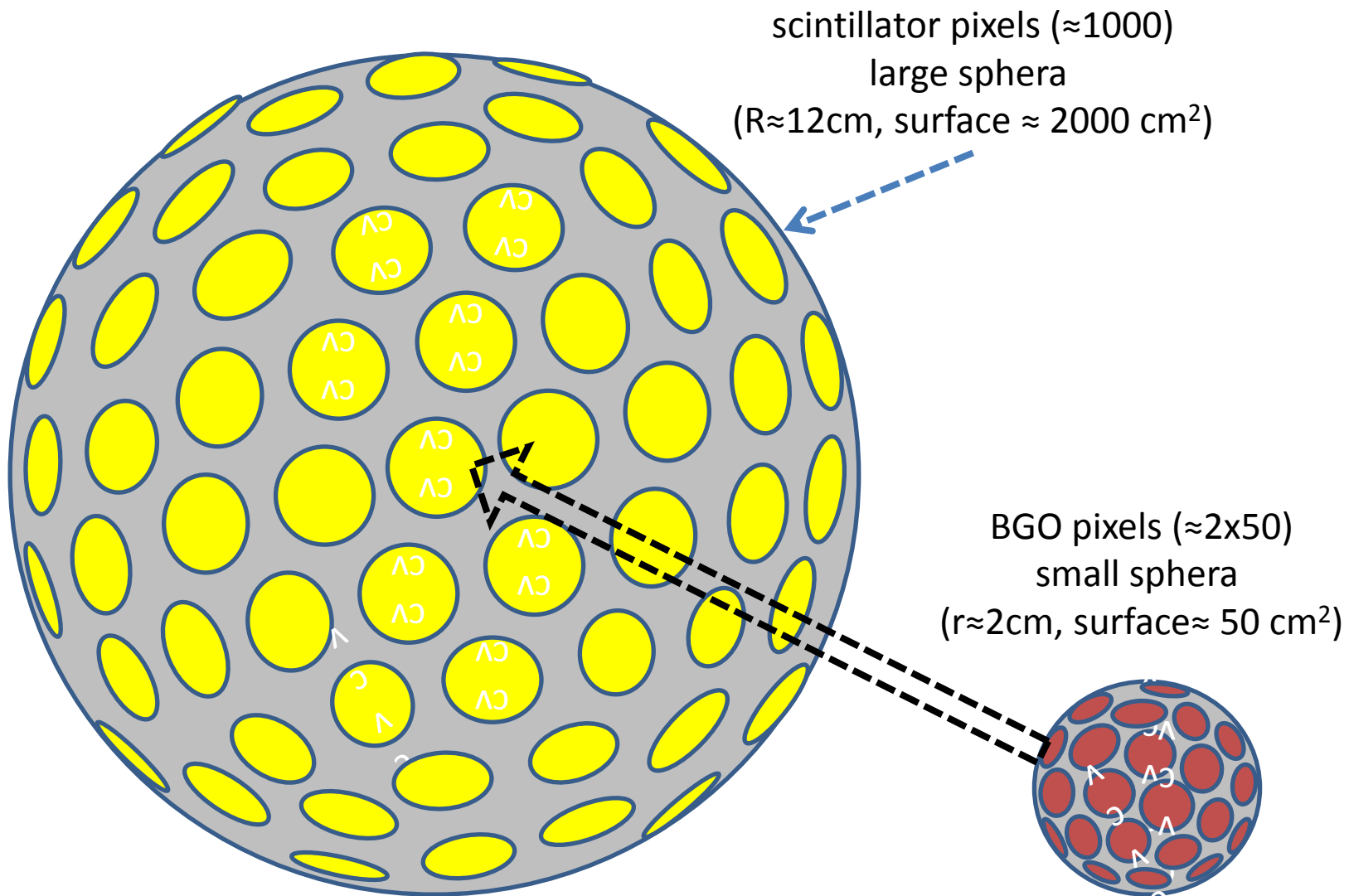
- time evolution in several bands (UV+optics)
(by filters + CCD image)

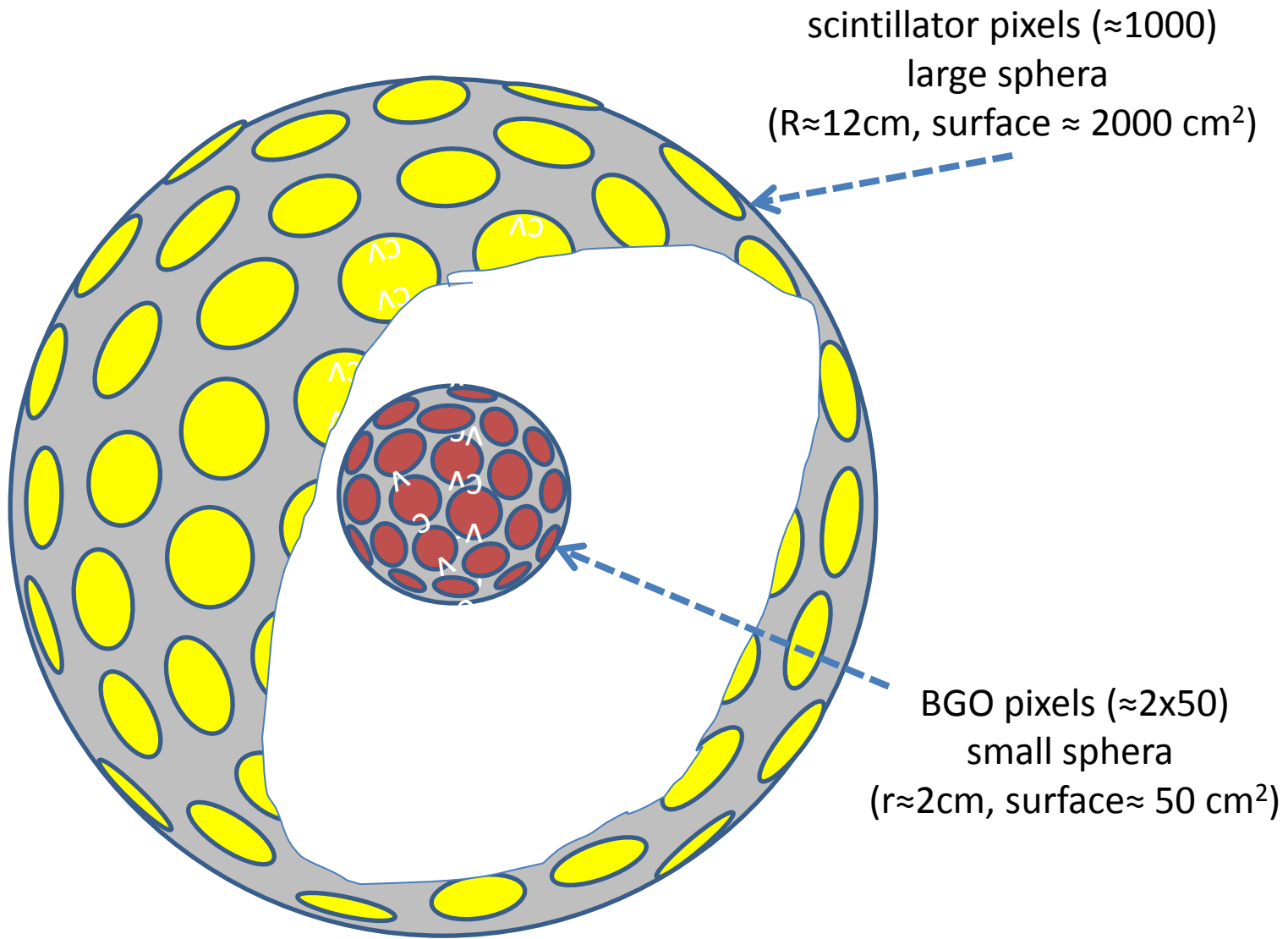
RADIATION DOSIMETRY

- time evolution in diamond detectors
(by two dosimeters)

.....

SPhERA sensors

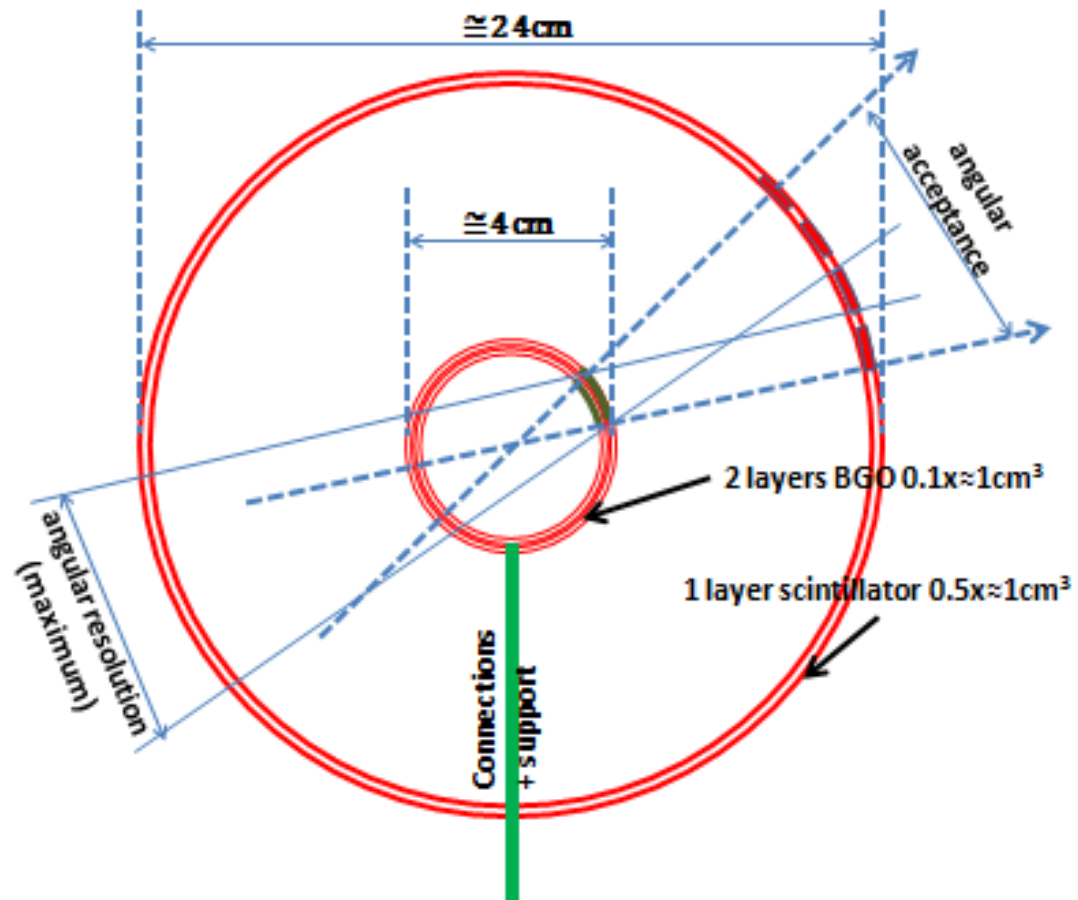


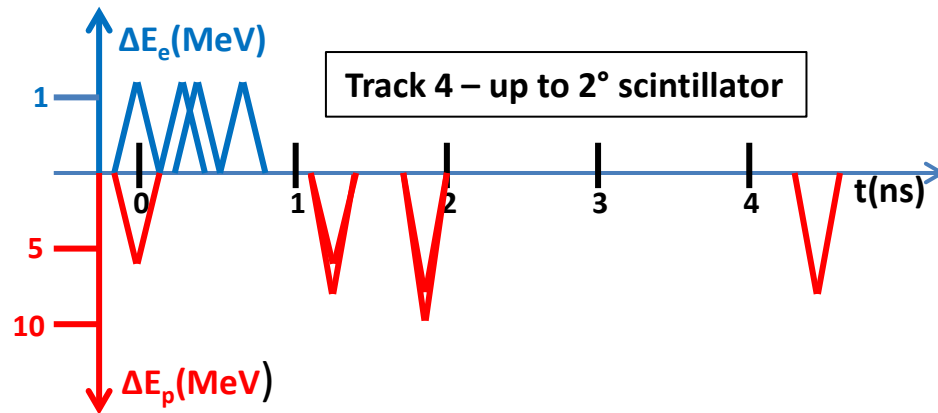
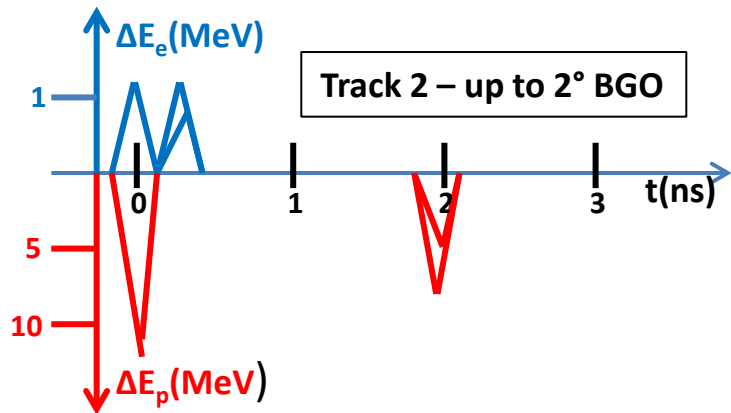
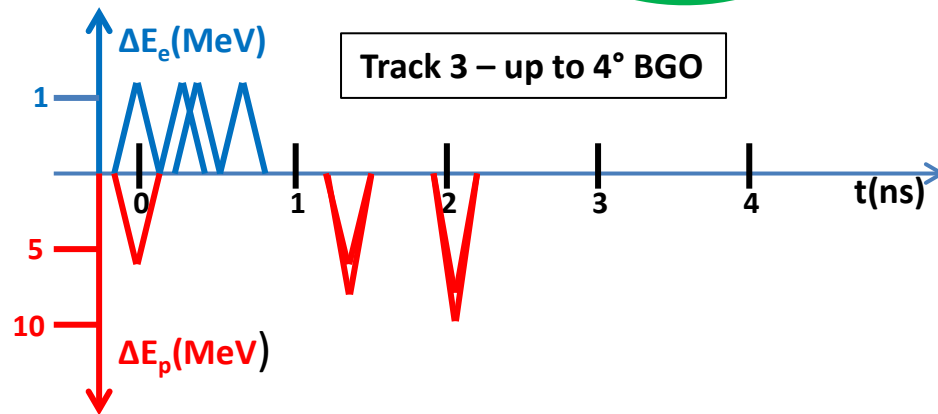
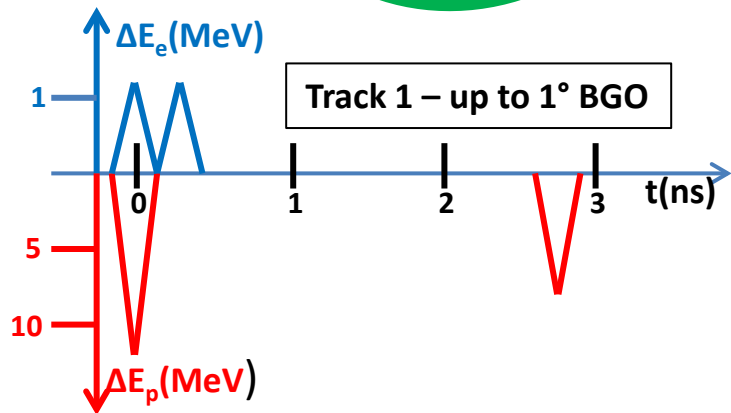
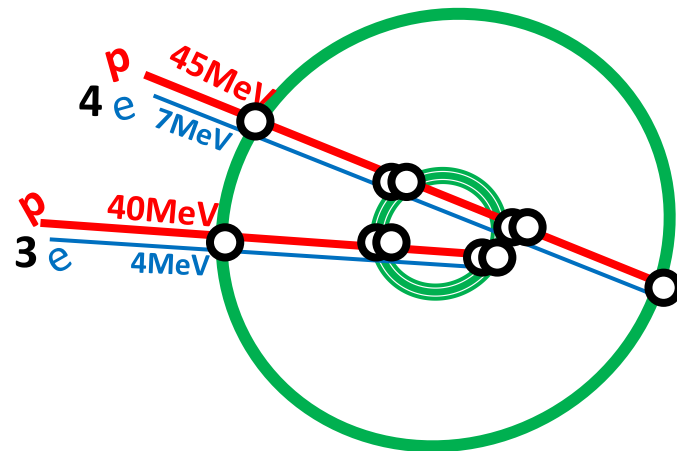
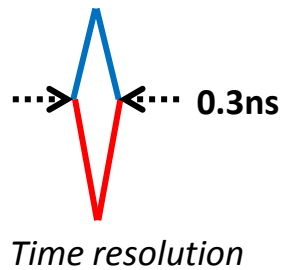
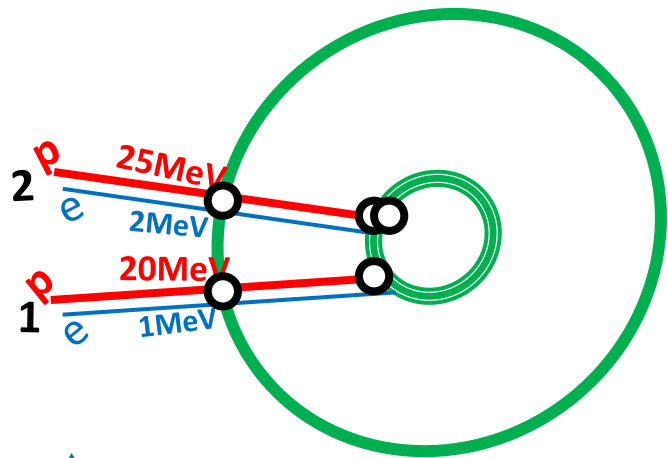


scintillator pixels (≈ 1000)
 large sphaera
 ($R \approx 12\text{cm}$, surface $\approx 2000\text{ cm}^2$)

BGO pixels ($\approx 2 \times 50$)
 small sphaera
 ($r \approx 2\text{cm}$, surface $\approx 50\text{ cm}^2$)

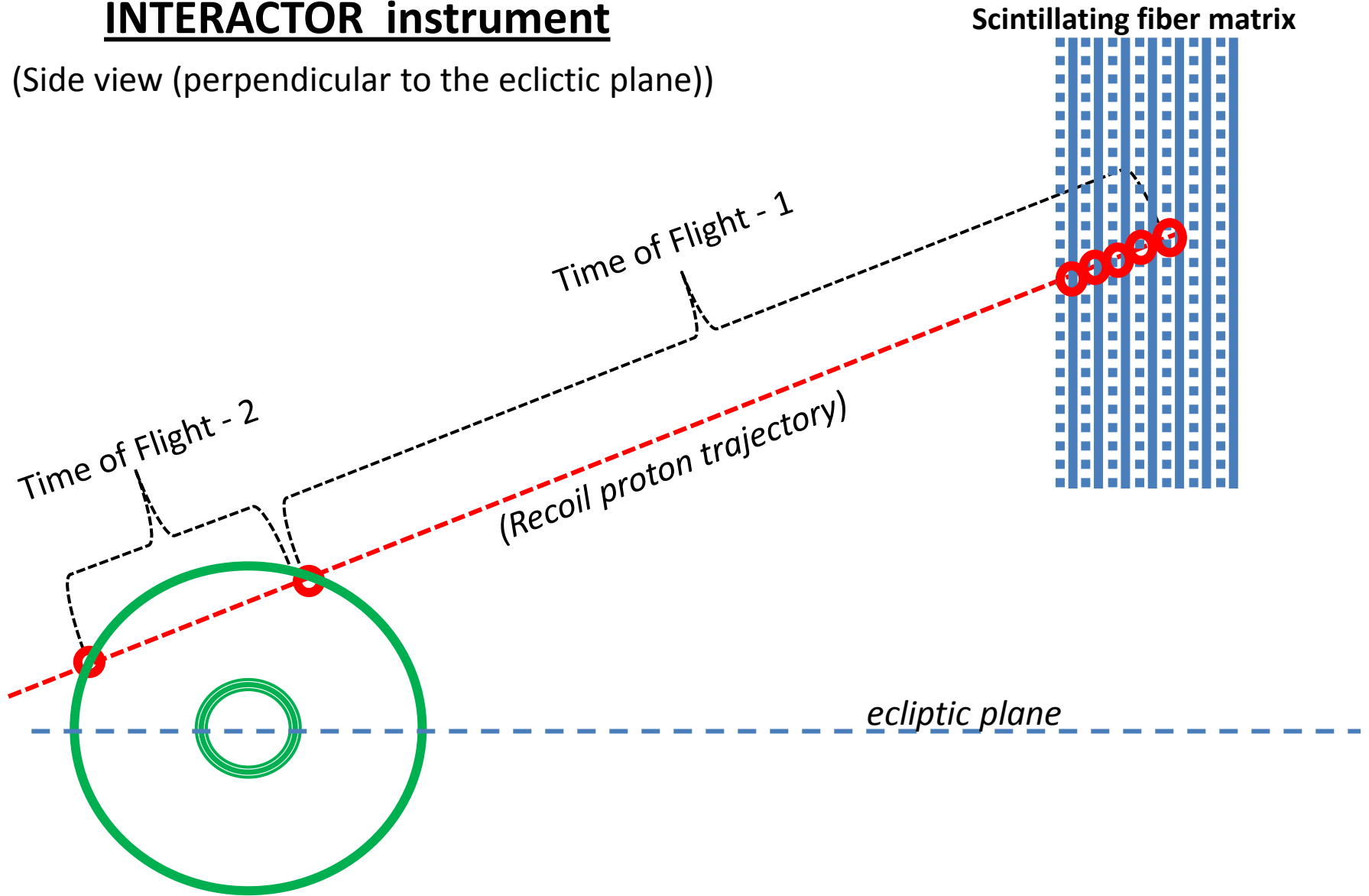
SPhERA sensors





INTERACTOR instrument

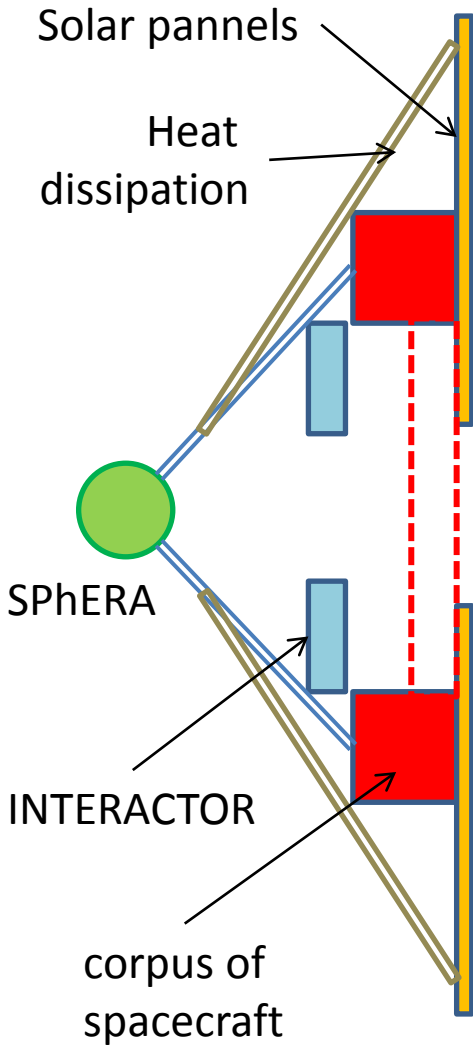
(Side view (perpendicular to the ecliptic plane))



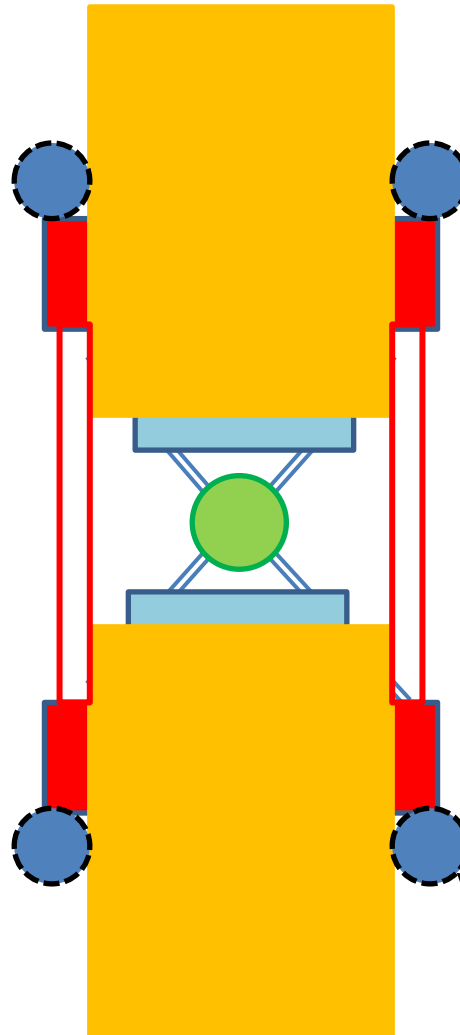
Main parameters of the instruments:

SPhERA	Mass of the sensors	3kg	}	Total: ≈30kg 20W
	Number of channels	4k		
	Power consumption	10W		
	Mass of the electronics	5kg		
INTERACTOR	Mass of the sensors:	20kg	}	
	Number of channels	50k		
	Power consumption	10W		
	Mass of the electronics	5kg		
SUNSPOT MAP	few kg + few W		}	≤ 20kg ≤20W
RADIATION DOSIMETRY	few kg + few W			
.....				

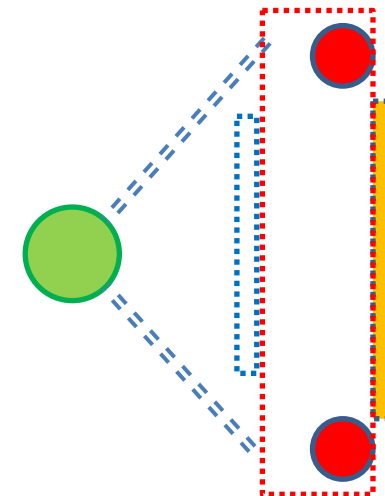
SIDE VIEW



FRONT VIEW

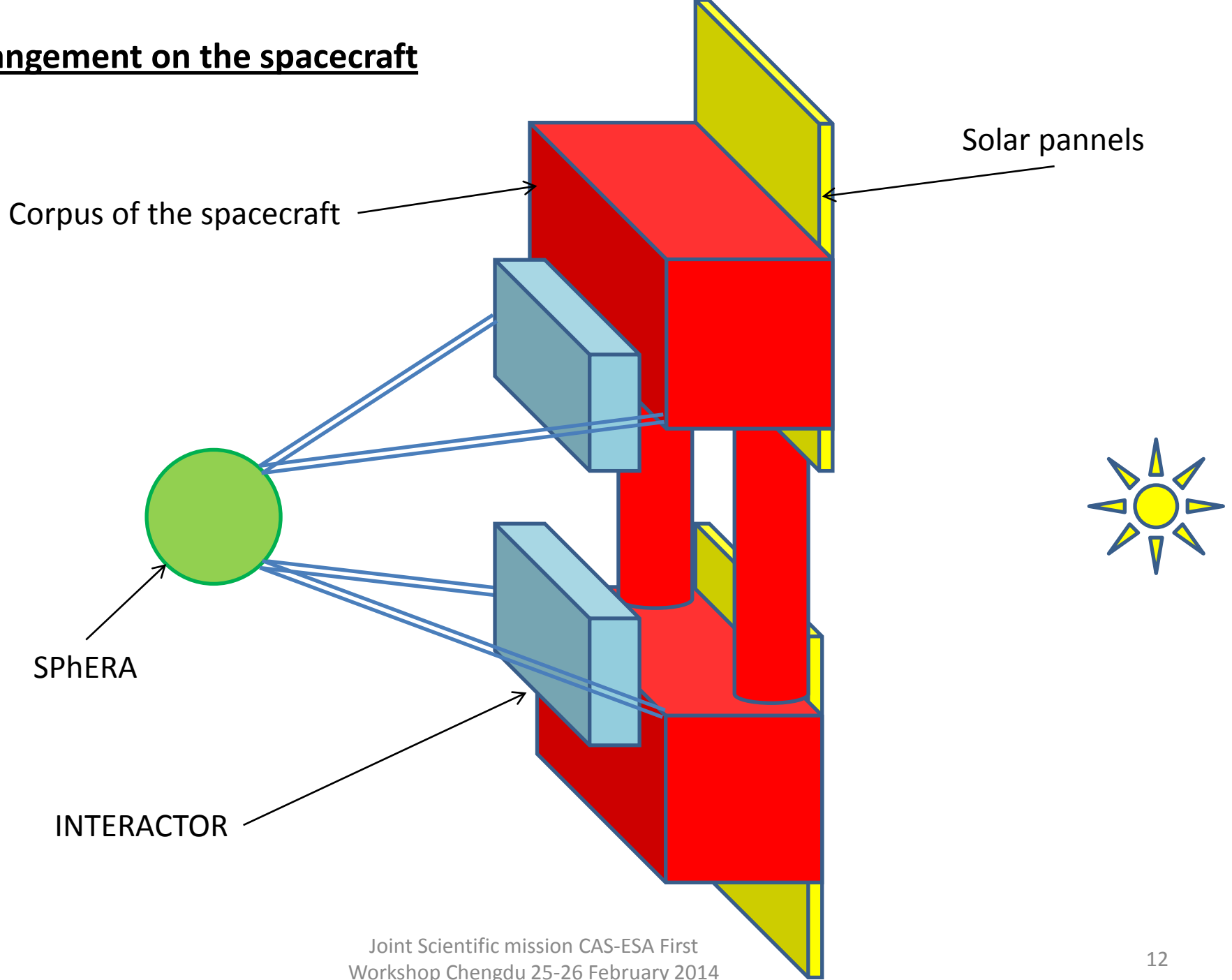


TOP VIEW
(on the ecliptic plane)

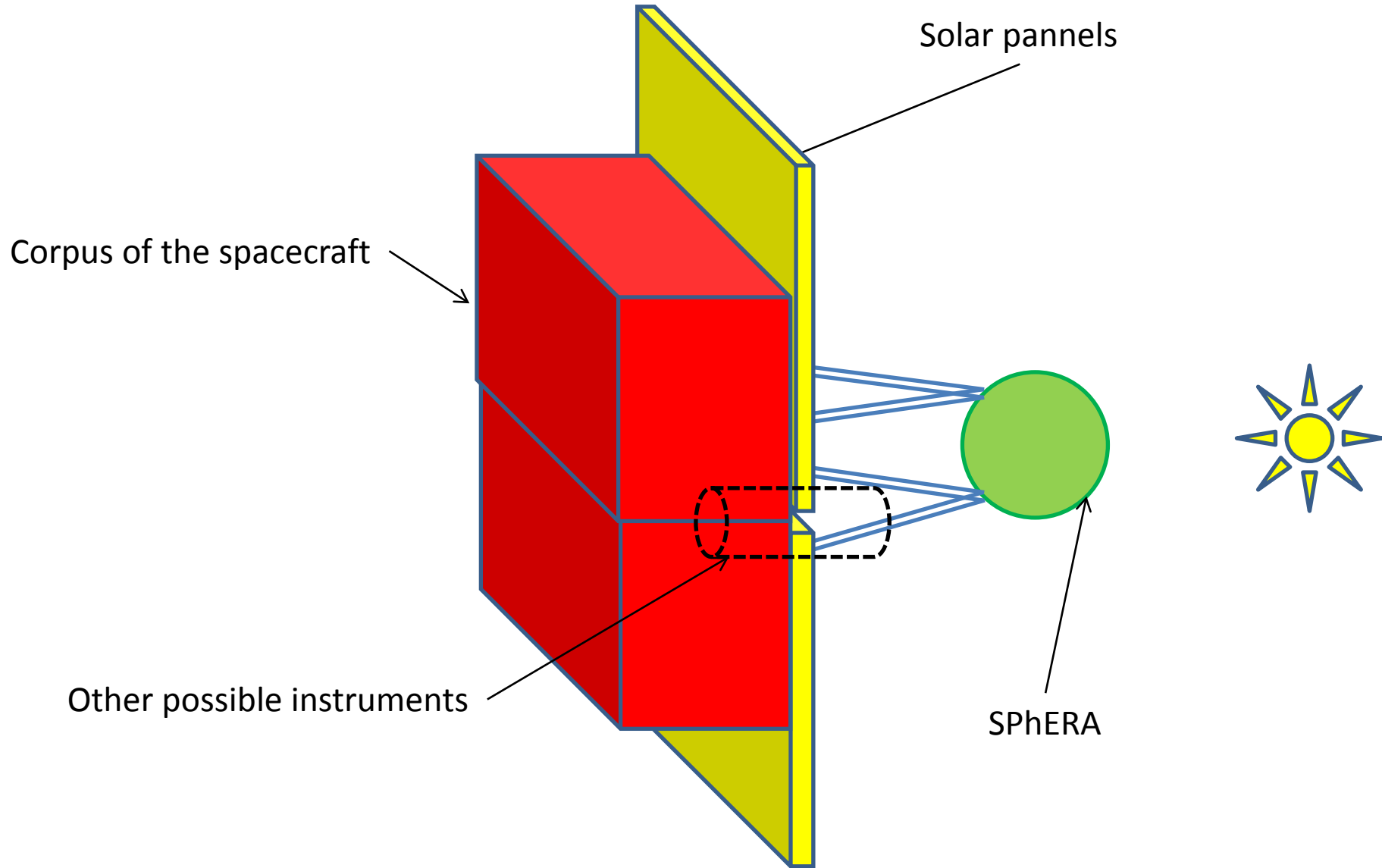


Position of possible
Other instruments

Arrangement on the spacecraft



Arrangement on the spacecraft (without detection of neutrons)

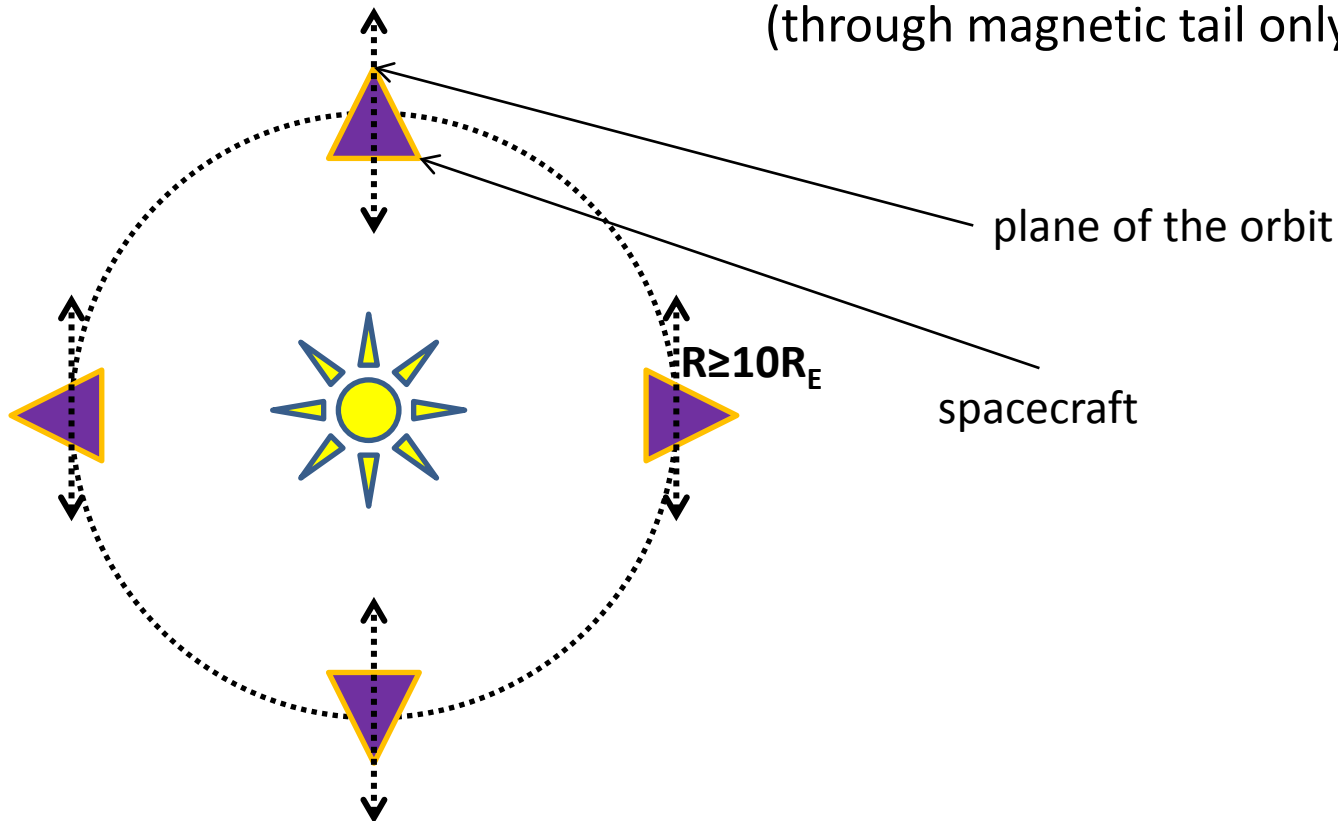


Other possible instruments

Suggested locations in space

(a) – ‘Sun soriented’ spacecraft in very high Earth orbit

Very High Earth Orbit, plane fixed in space
‘Looking to Sun’ attitude of the spacecraft
Orbit fully outside the terrestrial magnetic field
(through magnetic tail only for few % of time)



preferred location in space

(b) -- Earth-Sun L1

Orbit around L1

Spacecraft attitude: looking to Sun

Time interval to alarm Earth: several minutes



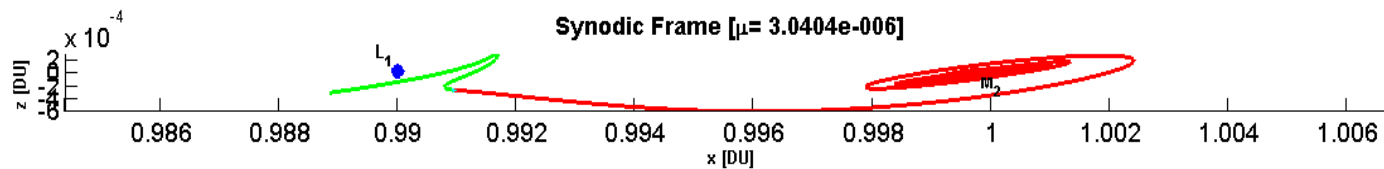
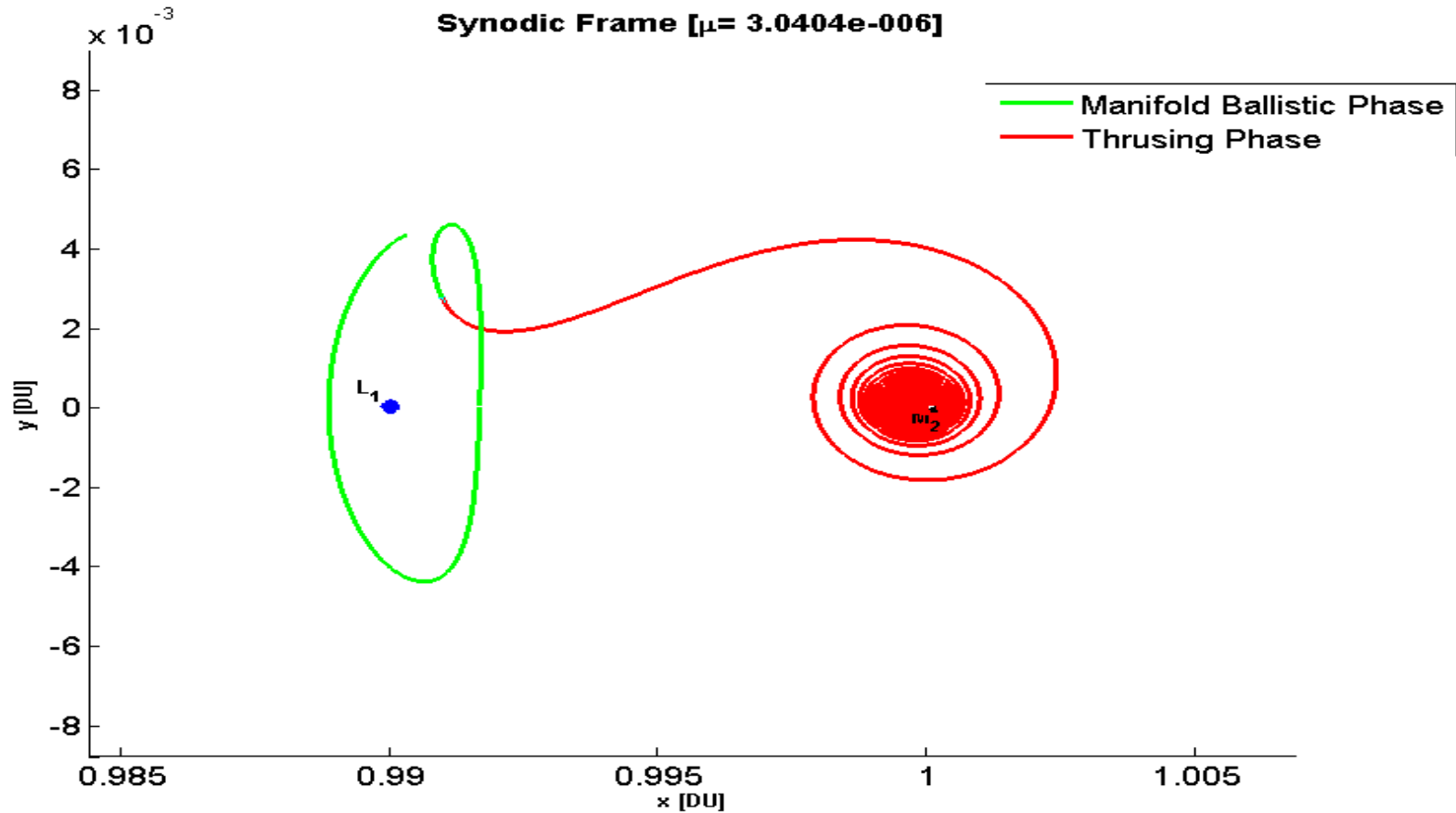
-- Chemical fuel: about 60% of the total mass!

-- Ion thrusters:

fuel (Xe) 25% of the total mass

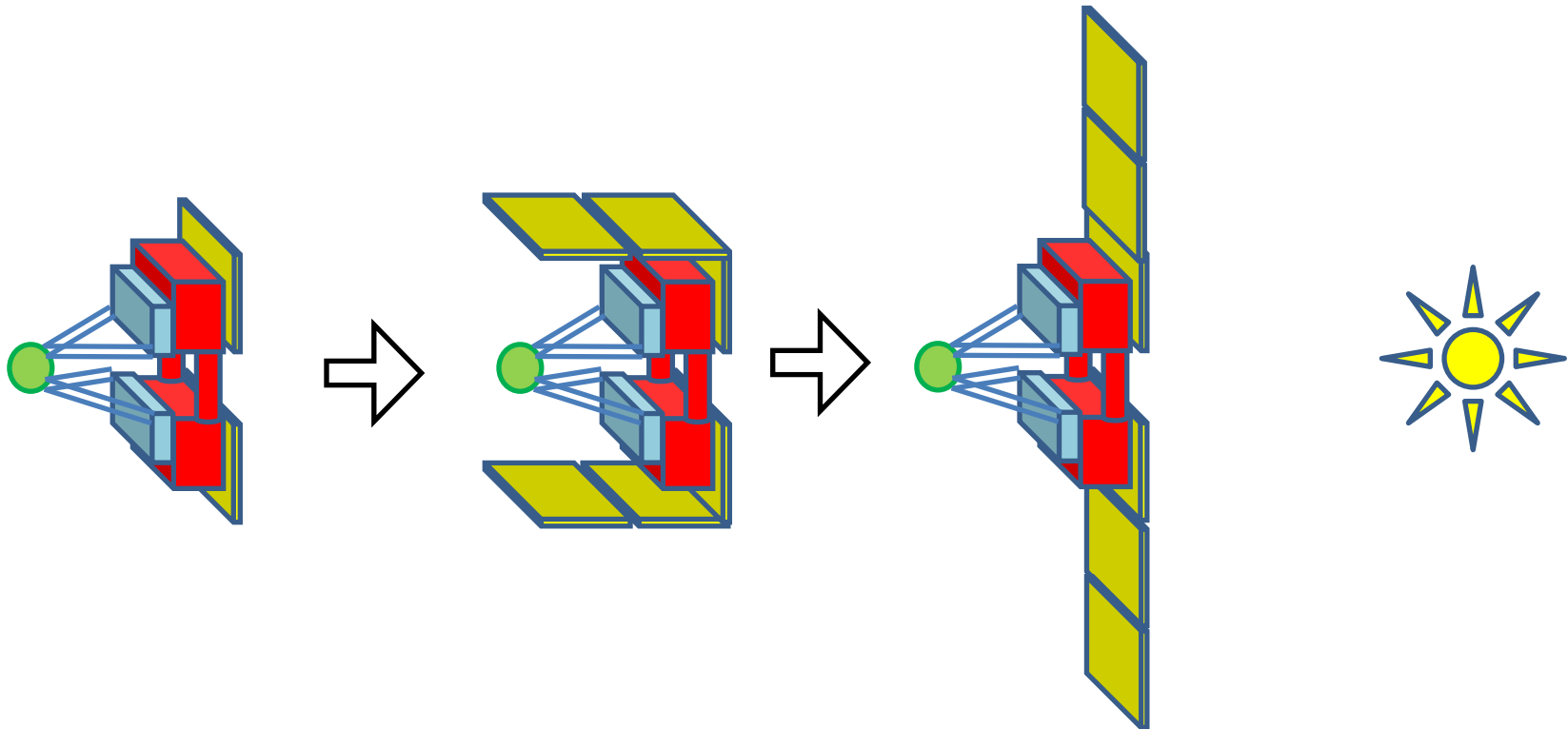
with one Hall effect Truster HT400 (ALTA-Space)

→ spiraling + ballistic insertion (10+7 months)



for each HT400 thruster needed 500W

→ several m² of solar pannels



Thanks for the attention